

(A Universe of) Massive
Black Holes:
Birth, Growth and Impact

Organizers:

T. Bogdanovic, E. Gallo, L. Mayer, M. Volonteri J. Baker, M. Begelman, C. Miller I like black's authority: its gravity, its self-evidence, its radicalism. Its powerful strength of contrast gives an intense presence to all other colour, and when it illuminates the darkest of them, it confers to them a sombre grandeur. Black has unsuspected possibilities, and remaining alert to what I do not know, I go to meet them.



P. Soulages 2005

Pierre Soulages (born 1919) French abstract painter, engraver, lithographer and designer

Rationale

 Physics of massive black holes and their link to the evolution of cosmic structures

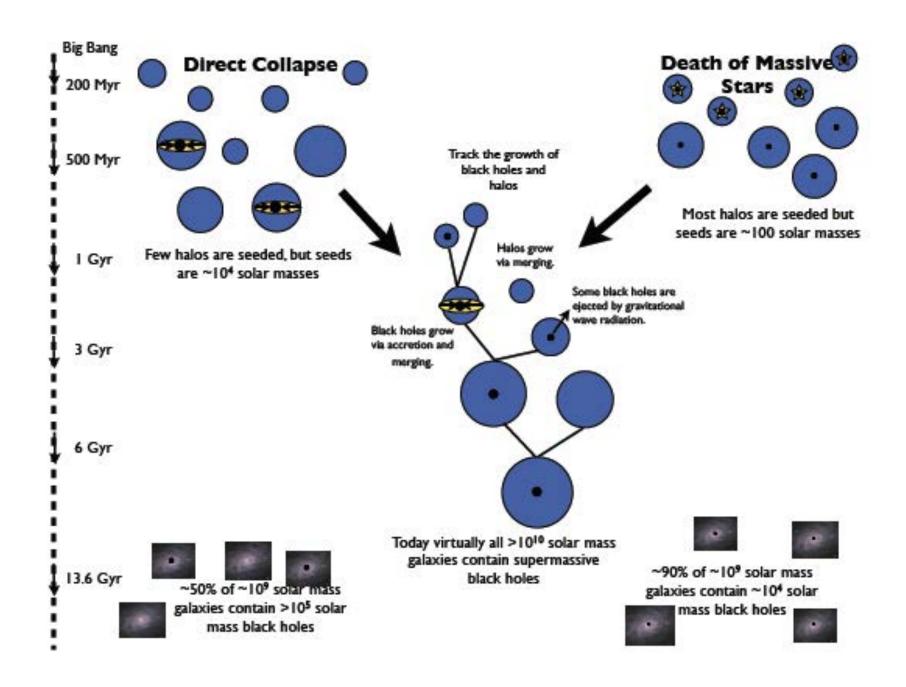
 Explore the relevant processes operating on different scales and in different physical regimes relevant to black holes

I. BH formation

 Identify the main physical ingredients of each formation model

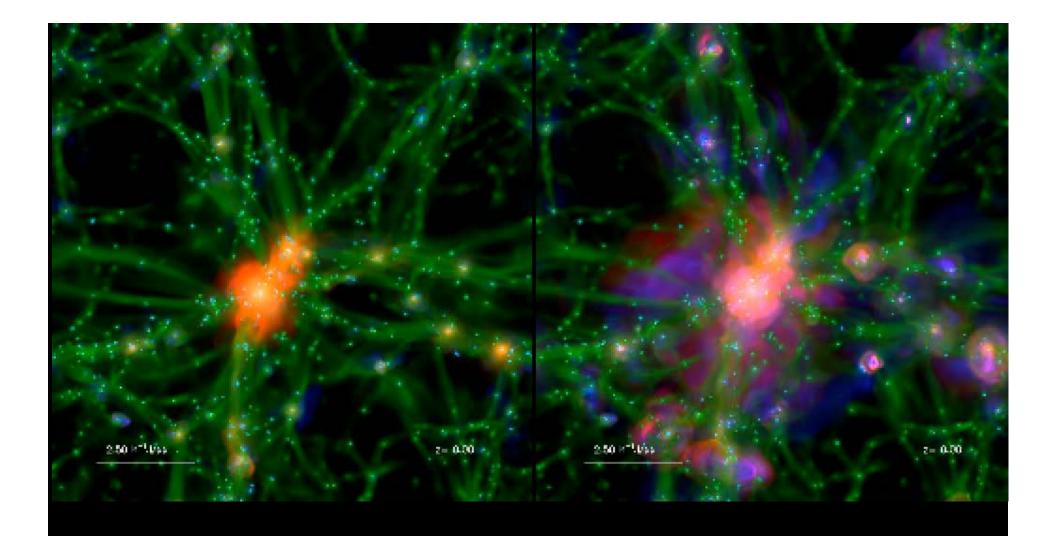
Bridge the gap between Newtonian and general relativity models

 Investigate observational consequences that can distinguish different models



2. Accretion and feedback

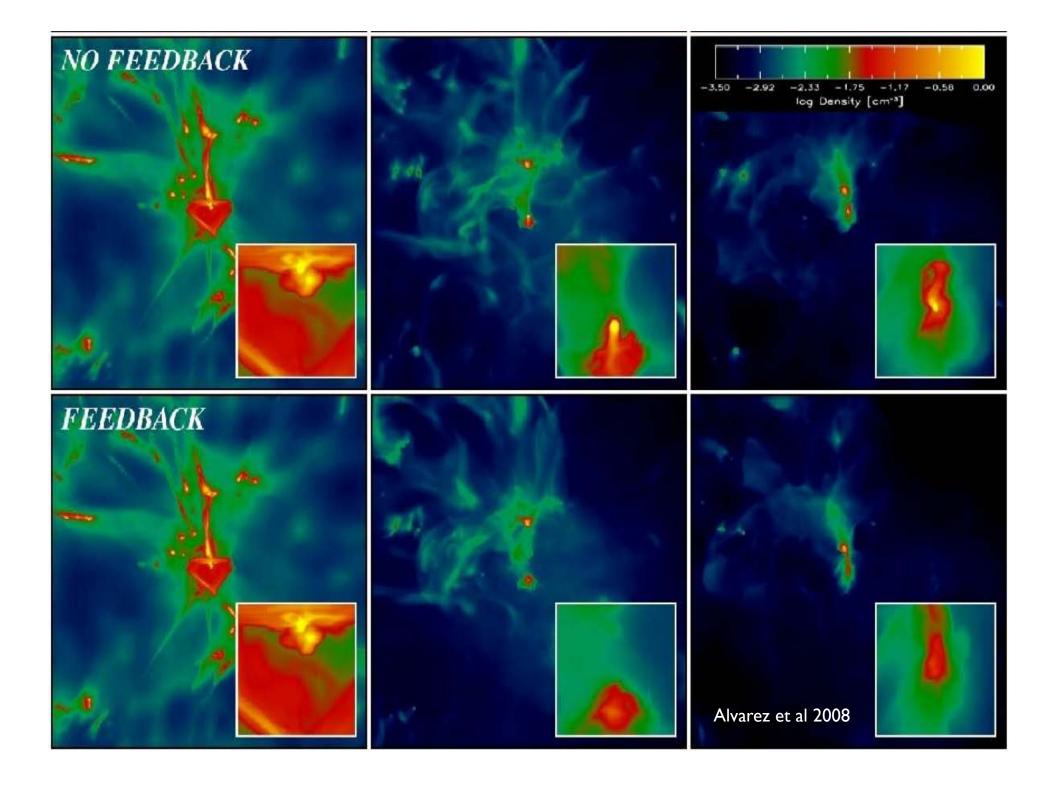
- How models of accretion and feedback on different scales relate to observations
- How observations and physical models on small scales can inform large scale simulations in order to improve 'sub-grid' physics
- Assess the role of AGN in energizing the intracluster/interstellar medium and estimate the implications for accretion/feedback themselves



Courtesy of Y. Dubois

NO AGN

WITH AGN

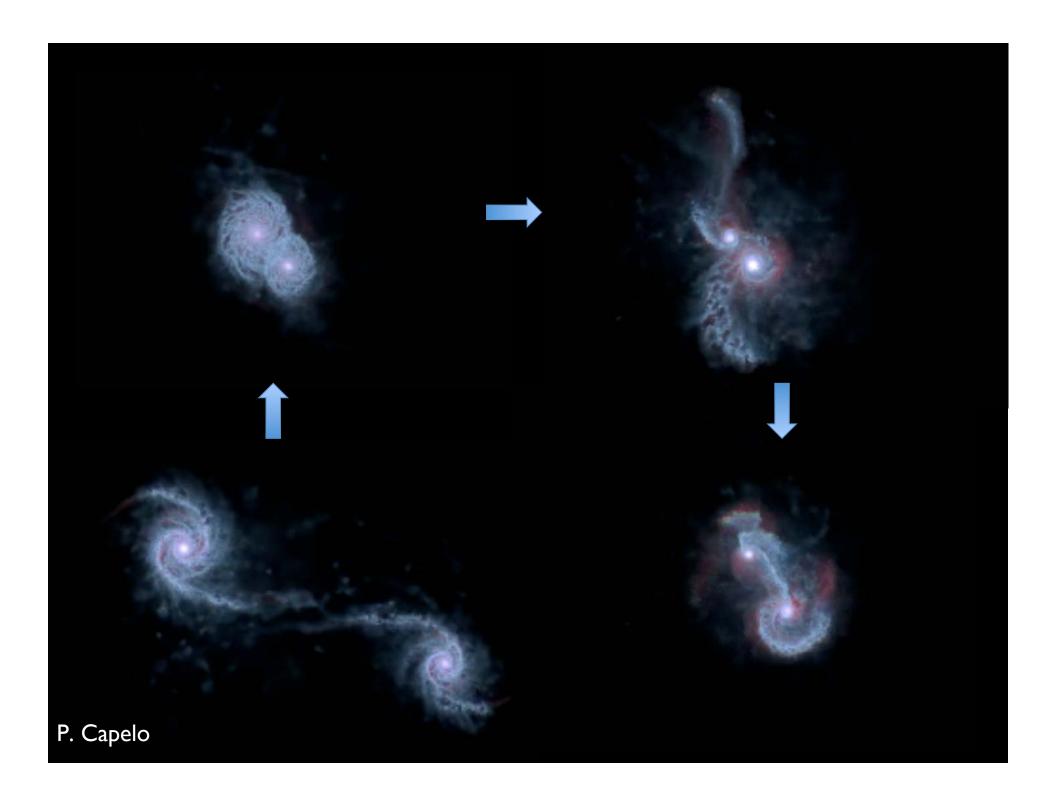


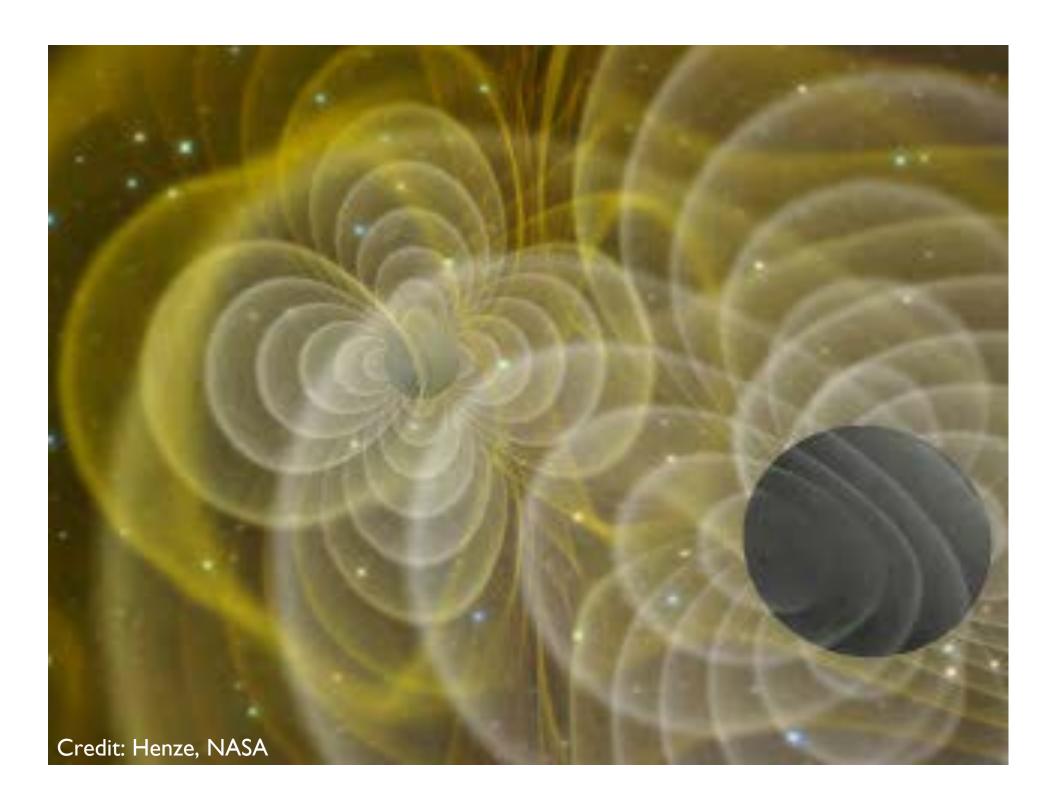
3. BH mergers

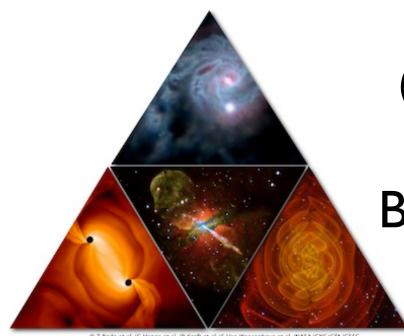
 The role of gas vs stars in shrinking massive BH binaries

Gravitational waves and electromagnetic counterparts

 Timescales associated to sinking from the galactic scales of many kiloparsecs to the milliparsecs relevant gravitational radiation







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