Horizon entropy, higher curvature, and spacetime equations of state

Ted Jacobson University of Maryland

Outline

- 1. Einstein equation of state
- 2. Higher curvatures and black hole entropy
- 3. Equation of state with higher curvatures
- 4. Lessons?

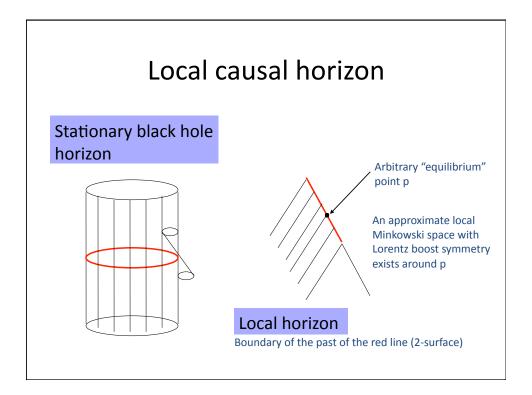
QFT & Thermo

From relativistic QFT:

- In a Rindler wedge the vacuum is thermal at $T = \hbar/2\pi$ wrt the boost Hamiltonian, and has (entanglement) entropy $S = \alpha$ A, with $\alpha = \infty$.
- The entropy scales with the area because the entanglement is dominated by vacuum correlations which diverge at short distances.

From thermodynamics:

• The Clausius relation dS = dQ/T gives the entropy increase when heat dQ enters a thermal bath at temperature T.



Spacetime Thermodynamics

- Assume a horizon entropy $S = \alpha A$ with universal α .
- <u>Assume</u> the Clausius relation dS = dQ/T holds for all LCH's, with dQ the (local) boost energy flux across the horizon.
- Raychaudhuri focusing eqn then implies causal structure of spacetime must respond via

$$\alpha R_{ab} = \frac{2\pi}{\hbar} T_{ab} + \Phi g_{ab}$$

Einstein equation of state

Local energy conservation $\nabla^a T_{ab} = 0$ and Bianchi identity $\nabla^a R_{ab} = \tfrac{1}{2} \nabla_b R$ imply $\Phi/\alpha = \tfrac{1}{2} R + \Lambda$

hence we obtain Einstein's equation

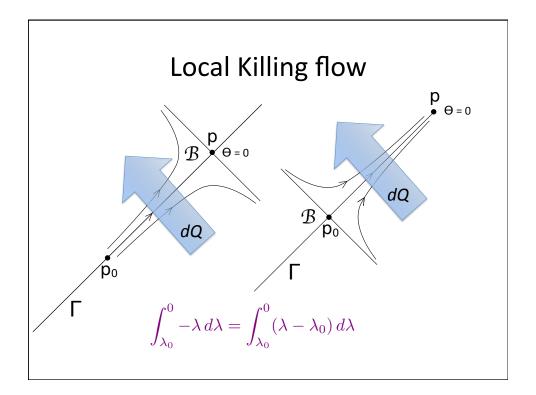
$$R_{ab} - \frac{1}{2}Rg_{ab} - \Lambda g_{ab} = \frac{2\pi}{\hbar\alpha}T_{ab} \qquad G_N = \frac{1}{4\hbar\alpha}$$

- 1. More entanglement \rightarrow weaker gravity (spacetime more rigid).
- 2. $1/G_N$ always tracks the net thermodynamic entropy (worries about multiple species, nonminimal coupling, gauge fields and gravitons notwithstanding)

(Also, cf. poster of work by Donnelly and Wall showing that 1-loop renormalization of $1/G_N$ by gauge fields is not negative in 2d, and perhaps higher d, after all.)

Remarks on Einstein equation of state

- If $\alpha = \infty$ then dS = dQ/T cannot be satisfied for general stress tensors. Clausius relation not apply to infinite entropy?
- Gravity itself provides a Lorentz invariant cutoff: entanglement at a scale L is "cloaked" by horizon fluctuations if $L^2 < \hbar G(L)$. (TJ, 1204.6349)
- Shear included via dS = dQ/T + dS_i, \rightarrow shear viscosity/ $\alpha = \hbar/4\pi$ (Eling, Guedens, TJ, gr-qc/0602001)
- Move bifurcation surface to the past?



Higher curvatures & horizon entropy

For stationary horizons, Wald entropy for $L = R + aR^2 + ...$ is

$$S_{BH} = \frac{A}{4\hbar G_N} + \text{curvature terms} = \frac{2\pi}{\hbar} \oint_{\Sigma} Q^{ab}[\hat{\chi}] N_{ab} dA$$

$$Q^{ab}[\xi] = W^{abc}\xi_c + X^{abcd}\nabla_c\xi_d$$

Noether potential for norizon-generating Killing vector

For
$$L=L[g_{ab},R_{abcd}]$$
, can choose $X^{abcd}=\frac{\partial L}{\partial R_{abcd}}$ and $W^{abc}=2\nabla_d X^{abcd}$ (Lopes Cardoso, de Wit, Mohaupt: hep-th/9904005)

Dependence on Killing vector field disappears on stationary Killing horizon.

Higher curvatures & horizon entropy

Remarks:

- 1. 2^{nd} law *not* established, except for (a) perturbations, where it follows from the 1st law and the null energy condition, and (b) f(R) theories.
- 2. Causal structure of such theories is generally *not* the metric light cone ... so what is a dynamical "horizon"?
- 3. A curious example: pure curvature² theories have black holes with zero entropy (Solodukhin, 1203.2961). Is it because such theories have no consistent quantization with a stable vacuum?

Higher curvature equation of state - I

Idea: adopt new horizon entropy density $s(g_{ab}, R_{abcd}, N_{ab})$, impose dS = dQ/T on all LCH's, infer field equation.

Problem 1:
$$\delta S \sim \nabla R_{abcd} \cdot \Delta \lambda$$
, whereas $\frac{\delta Q}{T} \sim (\Delta \lambda)^2$

Problem 2: Can't use Raychaudhuri eqn...

Non-equilibrium solution for $s(g_{ab},R)$ case: choose nonzero expansion $\Theta \neq 0$ at p to cancel the offending term, and allow for bulk viscosity. Works *only* for this special case. (Eling, Guedens, TJ, gr-qc/0602001)

Other approaches:

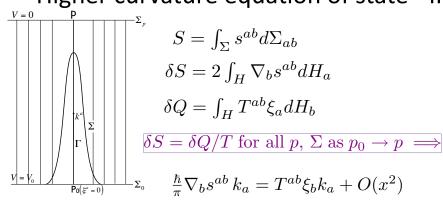
Elizalde and Silva (0804.3721) – based on lyer-Wald "boost invariant" proposal

Brustein and Hadad (0903.0823) Parikh and Sarkar (0903.1176) Padmanabhan (0903.1254)

Guedens, TJ, Sarkar (1112.6215)

- -- based on "Noether potential"

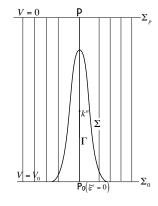
Higher curvature equation of state - II



Assume Noetheresque form for entropy wrt approximate local Killing vector:

$$s^{ab} = \frac{2\pi}{\hbar} (W^{abc} \xi_c + X^{abcd} \nabla_{[c} \xi_{d]})$$

Local Killing vector



There exists an approximate Killing vector satisfying:

$$\begin{aligned}
\xi^{\alpha}|_{\Gamma} &= (V - V_0)k^{\alpha} \\
\nabla_{\alpha}\xi_{\beta}|_{p_0} &= (k_{\alpha}l_{\beta} - l_{\alpha}k_{\beta})|_{p_0} \\
\nabla_{(\alpha}\xi_{\beta)} &= O(x^2) \\
\nabla_{\alpha}\nabla_{\beta}\xi_{\gamma}|_{\Gamma} &= (R^{\delta}{}_{\alpha\beta\gamma}\xi_{\delta})|_{\Gamma}
\end{aligned}$$

Higher curvature equation of state - III

Using the properties of the local Killing vector and the narrowness of the slice, the Clausius relation implies

$$W^{arb} = \nabla_s \left(X^{sarb} + X^{sbra} + X^{srba} \right)$$
$$R^{(a)}_{rst} X^{b)rst} + 2\nabla_r \nabla_s X^{(a|s|b)r} + \Phi g^{ab} = \frac{1}{2} T^{ab}$$

Conservation of the stress tensor then implies

$$\nabla^a \Phi = -\nabla_b \left(R^{(a}_{rst} X^{b)rst} + 2\nabla_r \nabla_s X^{(a|s|b)r} \right)$$

This imposes an integrability condition on X, which can be satisfied for

$$X^{abcd}=rac{\partial L}{\partial R_{abcd}}, \ \ {
m with} \ L=L(g_{ab},R_{abcd})$$
 Then the Clausius relation implies the field equation corresponding to this L.

We have not proved this is the *only* way to satisfy the integrability condition. It doesn't seem to work if derivatives of Riemann are included in L...

What's not to like...

- Entropy, and entropy changes between two local slices, depends on the arbitrary choice of bifurcation point for the local Killing vector
- In the GR case, the entropy change is not the area change for generic horizon slices.

Should it make sense?

No! It seems the *local* thermodynamic analysis can only capture the leading order, area term in the entropy:

The local Killing vector, and therefore the heat flux, is ambiguous at order $(L_{patch}/L_{curv})^2$, where L_{patch} = scale of the horizon patch.

If the entropy is $\propto (A + L_1^2 \int R)$, the curvature correction is of order $(L_1/L_{curv})^2$.

To capture the correction unambiguously, need $L_{patch} < L_1$.

If $L_1 = L_{planck}$, this patch is too small for the analysis to be justified.

If $L_1 = L_{\text{string}} > L_{\text{planck}}$, it is still too small. For any L_1 , it is probably too small!

What about the "virial expansion"?

$$\frac{p}{T} = \frac{\partial S}{\partial V} = \frac{N}{V} + B_2(T) \left(\frac{N}{V}\right)^2 + \dots \qquad \begin{array}{l} \text{small expansion parameter:} \\ \text{particle size or interaction range} \\ \text{intermolecular distance} \end{array}$$

intermolecular distance

leading order simple because of Independent particles; corrections allow for small correlations.

By contrast, the leading order horizon entropy is simple (proportional to area) because of underlying complexity, summarized by G_N. Curvature corrections represent a small deformation of that complexity, not the addition of a small physical effect.