# **Rindler Quantum Gravity**

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## **Motivations:**

How does AdS/CFT work? -How/why does spacetime emerge? -How/why does gravity emerge?

Can we extend to more general spacetimes, e.g. cosmological?







### vacuum state of CFT on S<sup>d</sup>

 $\left|0\right\rangle_{S^{d}}$ 

## An alternate description:



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"Rindler" description of pure global AdS



 $\left|0\right\rangle_{S^{d}} \leftrightarrow \left|0\right\rangle_{R^{d}} \leftrightarrow \sum_{i} e^{-\beta E_{i}/2} \left|E_{i}\right\rangle_{Rind} \otimes \left|E_{i}\right\rangle_{Rind} \leftrightarrow \sum_{i} e^{-\pi R_{H}E_{i}} \left|E_{i}\right\rangle_{H^{d}} \otimes \left|E_{i}\right\rangle_{H^{d}}$ 

## Single CFT picture:



Each H<sup>d</sup> CFT in thermal state  $1/T = 2\pi R_H$  (Casini, Huerta, Myers)

$$\rho = \sum_{i} e^{-2\pi R_{H}E_{i}} |E_{i}\rangle \otimes \langle E_{i}|$$

Dual to ``hyperbolic black hole" = wedge of pure AdS = region accessible to accelerated observer

No information about geometry behind the horizon

## "Microstates" of pure AdS

How do we interpret typical state  $|E_i\rangle$  in ensemble  $\rho$ ?



Density matrix for smaller region almost identical for  $|E_i\rangle$ or  $\rho$ : dual to smaller wedge of pure AdS But: pure state implies divergent  $T^{\mu\nu}$  on Rindler horizon  $\rightarrow$  bulk horizon replaced by something singular

## Different descriptions of a hyperbolic black hole



 $\left|E_{i}\right\rangle$ 

Microstate: wedge of pure AdS with "singular" edge (~ Mathur fuzzball)

Thermal state: wedge of pure AdS with horizon.

Different purifications = different spacetimes with this wedge

$$\sum_{i} e^{-\pi R_{H}E_{i}} |E_{i}\rangle \otimes |E_{i}\rangle$$

Entangled state of 2 CFTs: canonical extended spacetime: information behind horizons = entanglement information

## Decomposing spacetime (cf MVR09, Mathur)



 Pure AdS = quantum superposition of disconnected Rindler wedge microstates (explicit realization of Mathur proposal)
These are the microstates counted by Rindler horizon area
Choice of wedges arbitrary: -observer-dependent horizon area
-gives observer-dependent entropy

-counts observer-dependent microstates

## Lessons for cosmological spacetimes?



Wild extrapolation to dS suggests:

- dS static patch described by density matrix for some degrees of freedom (different for different observers)
- dS entropy best viewed as entanglement entropy (microstates like static patch with horizon replaced by end of spacetime, but physical spacetime NOT a microstate)

#### Differences: need finite number of d.o.f., different dynamics

## **Emergence of spacetime via entanglement**

Have seen: connected spacetimes emerge when subsets of degrees of freedom (maximally) entangled

Approximations to ground state incorporating entanglement at successively longer scales correspond to (t=0) spacetimes that extend further and further to IR

## **Emergence of spacetime (dynamical)**

Field theory evolution: subsystems maximize entanglement for given energy

Starting with atypical high-energy state, density matrix for successively larger subsystems approach thermal

Classical spacetime grows from UV  $\rightarrow$  IR