

Temporal Power-Law Characteristics of the Atmospheric General Circulation

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Papers:

V & K, submitted to J. Climate. (Methods, obs analysis.)

V & K, in prep. (Model/obs comparison.)

V, K & M, in prep. (IPCC AR4 analysis.)

Outline

- Introduction

Characterizing Dec-Cen Climate Variability

- Methods & Data

Estimating the Hurst exponent

- Observational analysis

Analyzing ERA40 temperature data.

- Model-observation intercomparison

Insights and a surprise.

- Conclusion

Introduction: Dec-Cen Variability

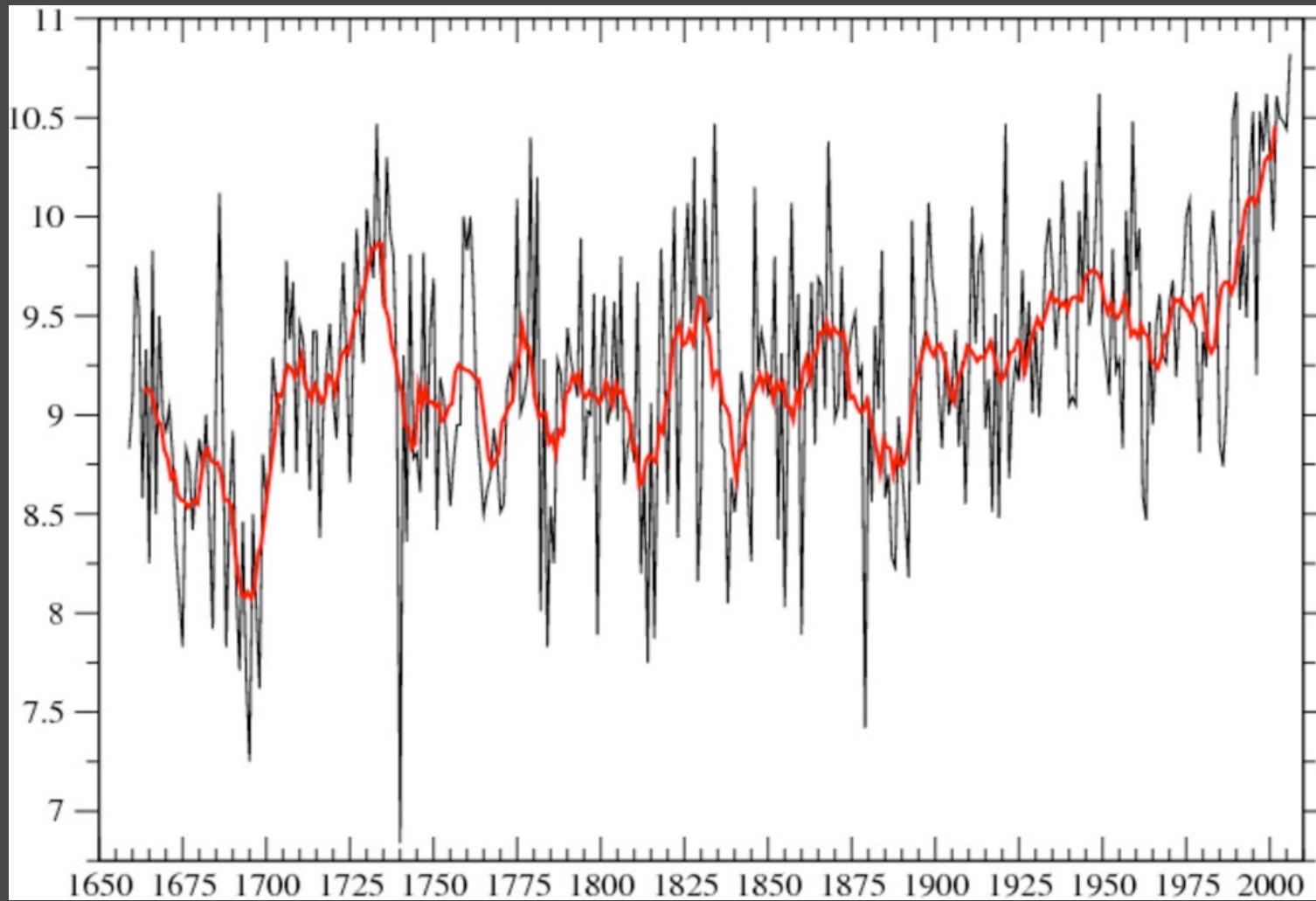
We explore “internal” 10-100 y climate variability.

- Distinct from externally forced variability such as greenhouse warming forced trends.
- Arises from ocean-atmosphere(-biosphere?) variability on timescales longer than ENSO.
- “Virtually nothing is known” about it (Wigley and Raper 1990).

Applications:

- Distinguishing forced from spontaneous trends.
- Identifying periodic modes.

Annual Mean Central England Temperature



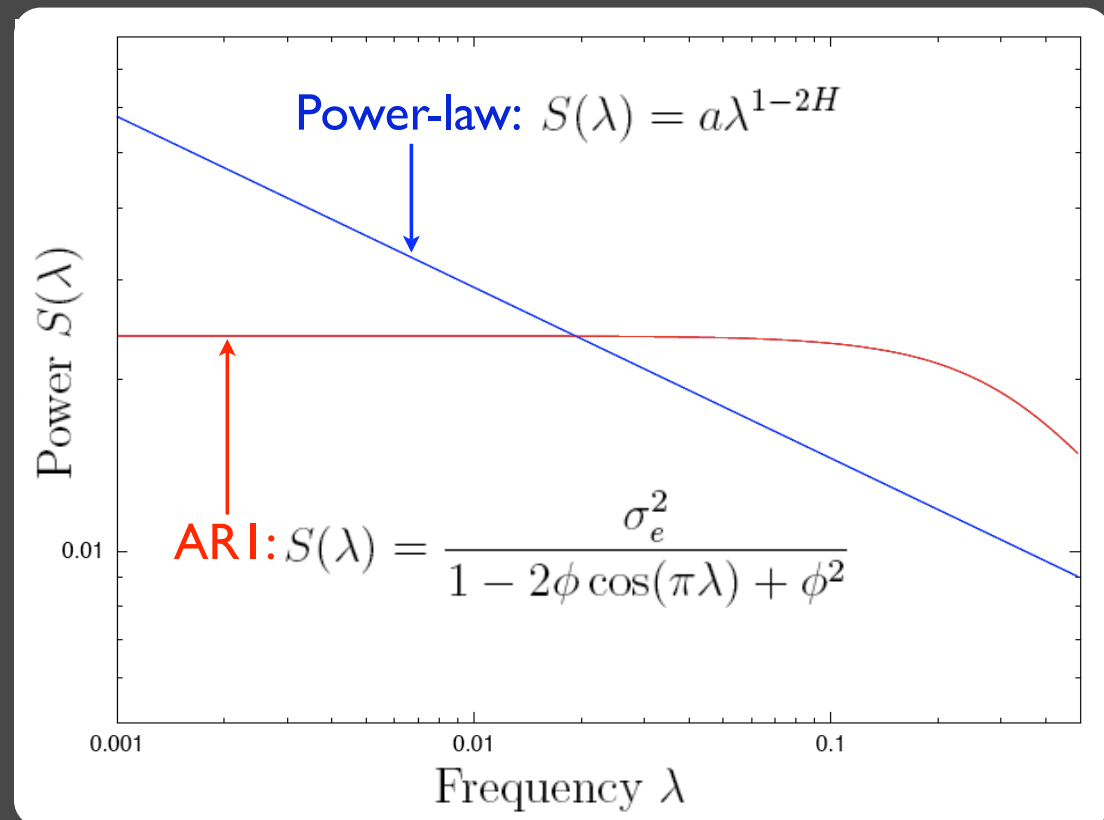
Two-parameter Climate Noise Models

AR1/red noise:

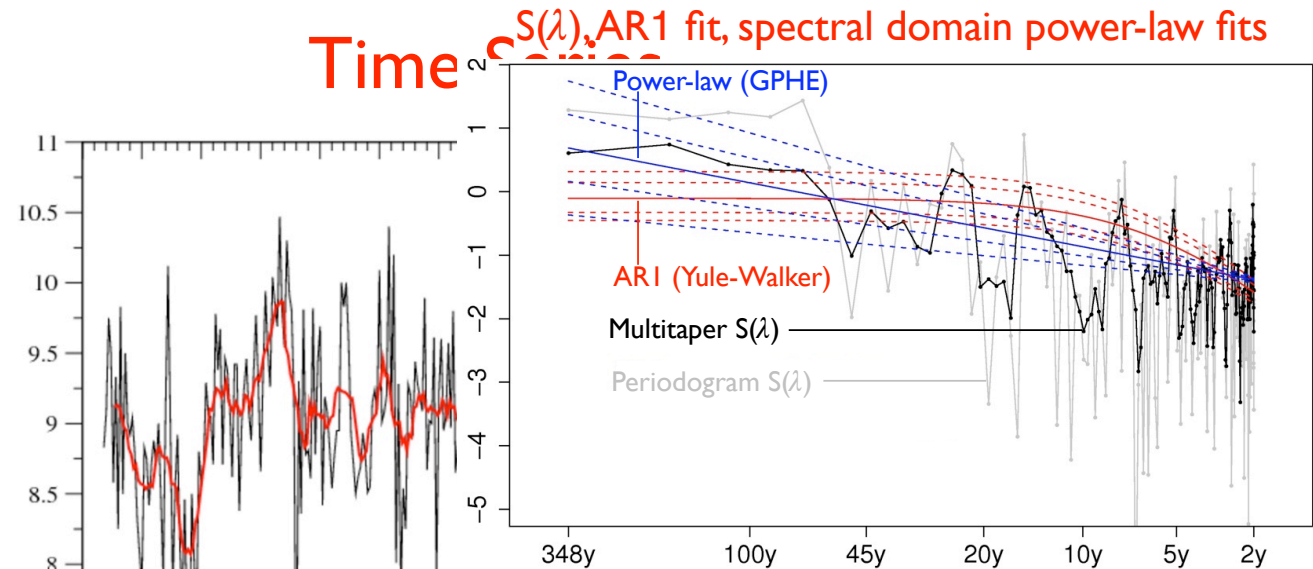
- Parameters: variance & lag-1 autocorrelation ϕ , $0 \leq \phi \leq 1$.
- Classical Hasselman model, fast noisy atmosphere + slow damped ocean.

Power law/scaling:

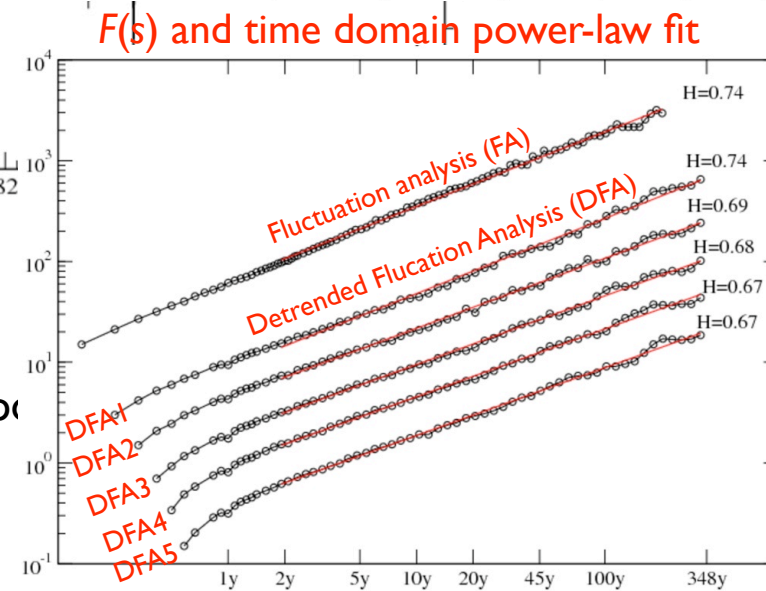
- Parameters: variance and Hurst exponent H , $1/2 \leq H \leq 1$.
- No simple unambiguous physical model.



Central England Temperature: A Best Case

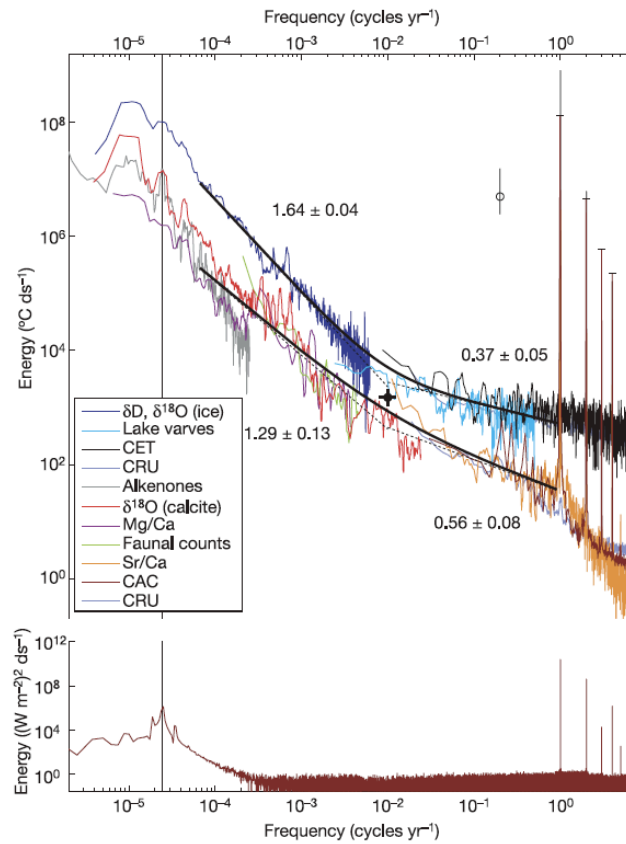


CET unambiguously fits power-law
 ("Portmanteau test").



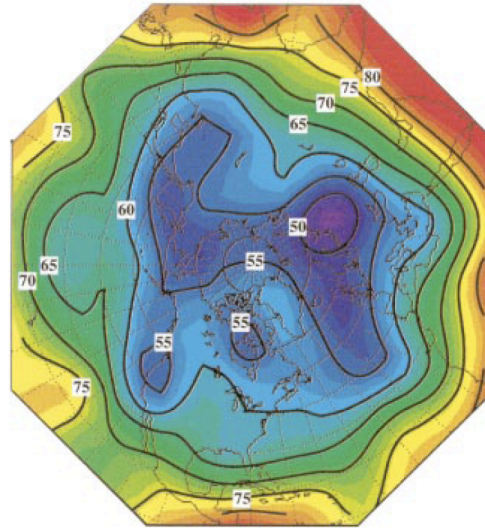
Previously Documented Powerlaw Behavior

$S(\lambda)$ in paleo temperature proxies



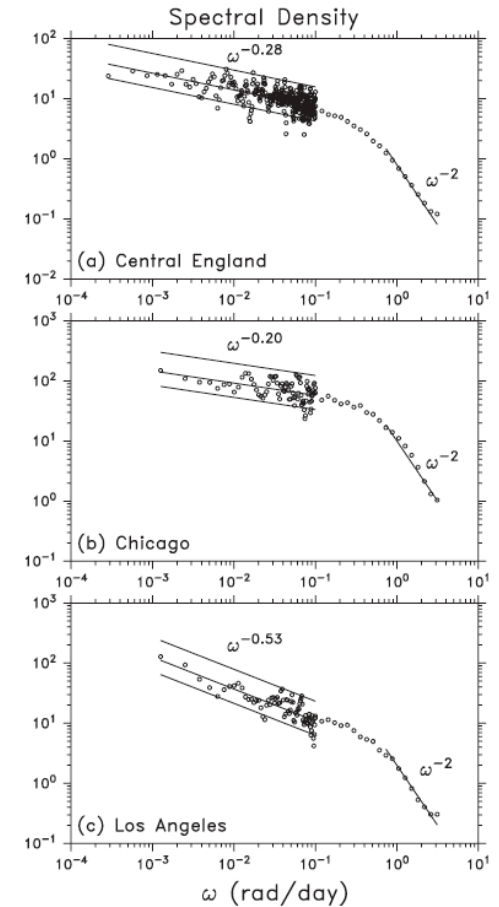
Huybers and Curry 2006

$H \times 100$ in midtroposphere



Tsonis et al. 1999

$S(\lambda)$ in surface temperature records



Caballero et al. 2002

Model-Observational Intercomparisons

Global Climate Models Violate Scaling of the Observed Atmospheric Variability

R. B. Govindan,^{1,2} Dmitry Vyushin,^{1,2} Armin Bunde,^{2,*} Stephen Brenner,³
Shlomo Havlin,¹ and Hans-Joachim Schellnhuber⁴

PRL 2002

Volcanic forcing improves Atmosphere-Ocean Coupled General Circulation Model scaling performance

D. Vyushin,¹ I. Zhidkov,¹ S. Havlin,¹ A. Bunde,² and S. Brenner²

GRL 2004

Past results have proven controversial and sensitive in surprising ways.

Introduction

Power-law models

- Fit dec-cen $S(\lambda)$ at least as well as AR1.
- Capture buildup of power at low frequencies.
- Provide alternative “null hypothesis” for detecting trends and periodicities.

Working assumption: if you're careful and aware of the ambiguities, power law characterization can be useful.

Objective: to characterize and interpret power-law behavior for the atmospheric general circulation for the recent (50 y) record, from the surface to the stratosphere.

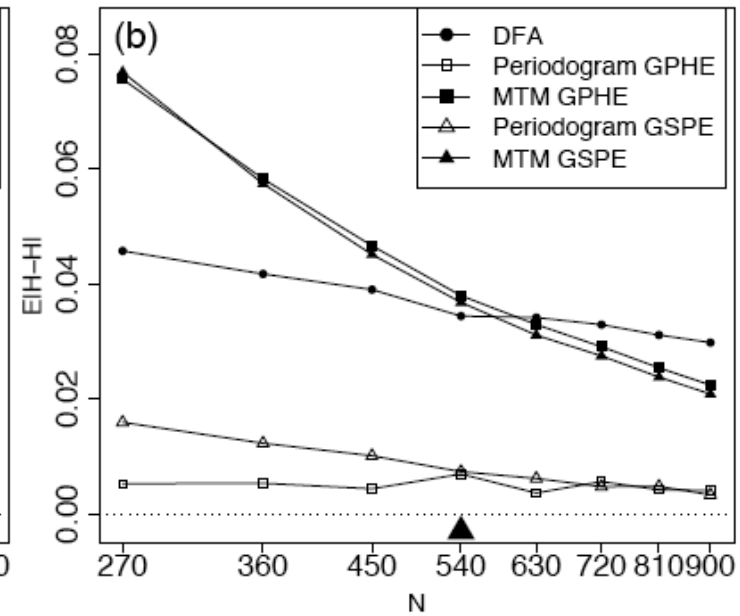
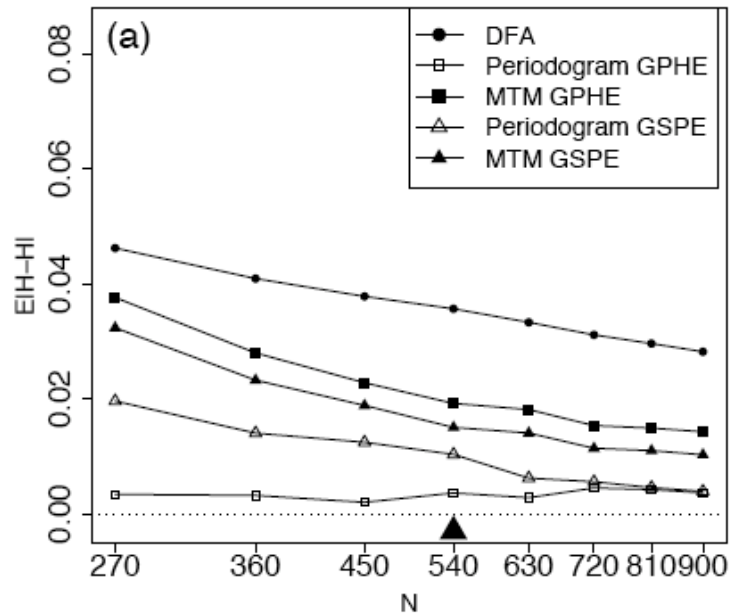
Methods

- We test and apply several methods for calculating Hurst exponent H from $S(\lambda) \sim \lambda^{1-2H}$.

		Dependence on timescale/ λ range		Physics & stats literature
		HF cutoff	LF cutoff	Remark
Time Domain	DFA(t)	$s_{\text{short}}=18\text{m}$	$s_{\text{long}}=11\text{y}$	Kantelhardt et al. (2001)
	DFA(a)	$s_{\text{short}}=18\text{m}$	$s_{\text{long}}=45\text{y}$	this study
Spectral domain: • Least squares & max. likelihood • Periodogram & multitaper spectra	GPHE(t)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/15\text{y}$ ($l=2$)	Robinson (1995b)
	GPHE(a)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/45\text{y}$ ($l=0$)	Hurvich et al. (1998)
	MTM GPHE(t)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/15\text{y}$ ($l=2$)	McCoy et al. (1998)
	MTM GPHE(a)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/45\text{y}$ ($l=0$)	this study
	GSPE(t)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/15\text{y}$ ($l=2$)	this study
	GSPE(a)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/45\text{y}$ ($l=0$)	Robinson (1995a)
	MTM GSPE(t)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/15\text{y}$ ($l=2$)	this study
	MTM GSPE(a)	$\lambda_{\text{high}}=1/18\text{m}$	$\lambda_{\text{low}}=1/45\text{y}$ ($l=0$)	this study

- All methods work well for pure power law stochastic processes (ARFIMA).

E.g. convergence of H for ARFIMA



Methods

Data used

- Focus on temperature as primary meteorological field.
- “Reanalysis products” that merge data + models for last 50 years: NCEP and ECMWF
- Climate models from GFDL and from PCMDI/CMIP3/IPCC archive.

Seasonal cycle removed.

Test impact of filtering for well known climate signals using simple linear filtering.

Observational Analysis

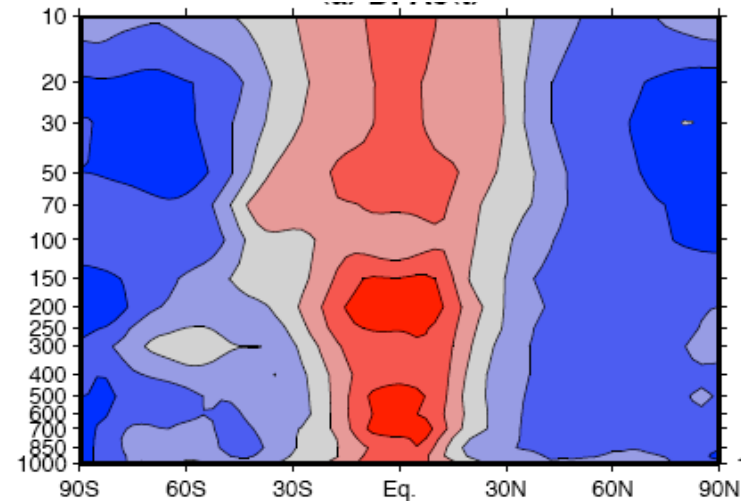
Estimating H in the ERA40 Reanalysis

First Results

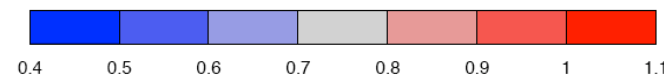
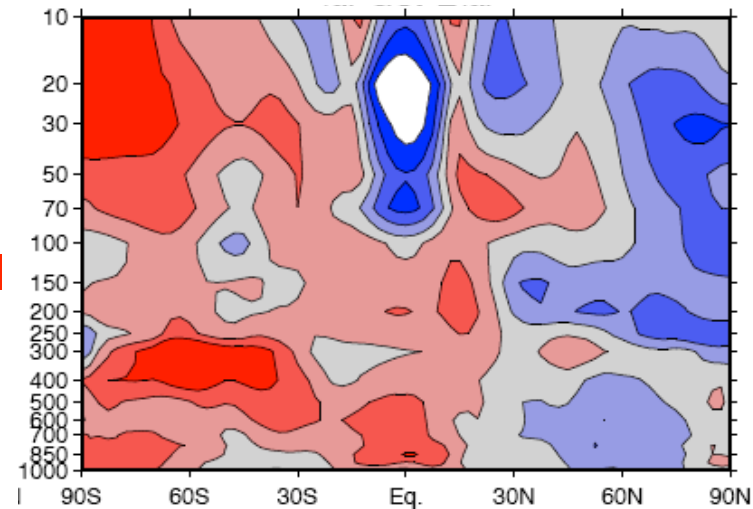
- We here compare $H(T)$ for ERA40 data using DFA (time domain) and GSPE (spectral domain) methods.
- The methods produce very different pictures!
- We account for these differences by:
 1. Making time scale ranges consistent.
 2. Filtering well known climate signals.

Zonal mean of H — $[H(T)]$

DFA
(Time
domain)



GSPE
(Spectral
domain)



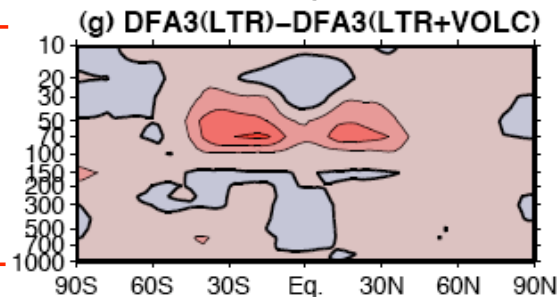
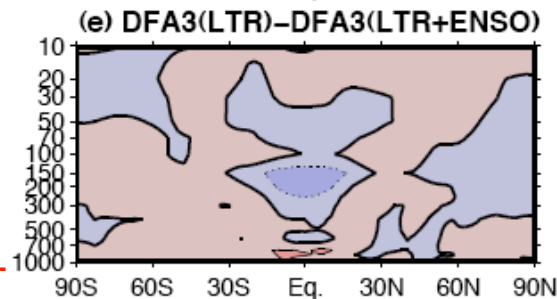
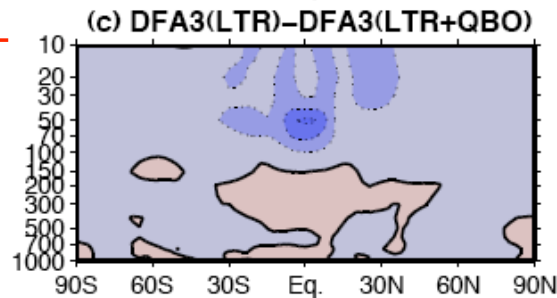
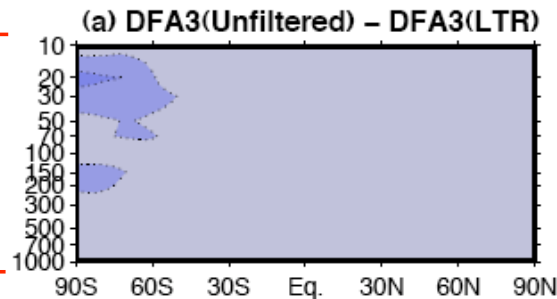
Impact of Filtering on H

Linear trends increase H (steeper slope) for GSPE but not for DFA.

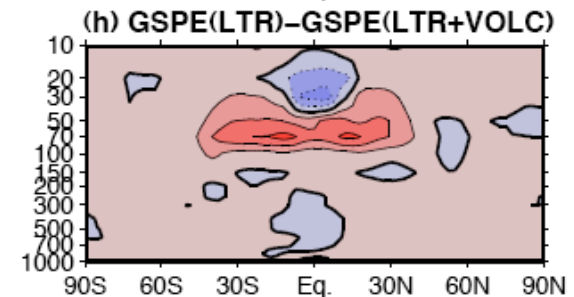
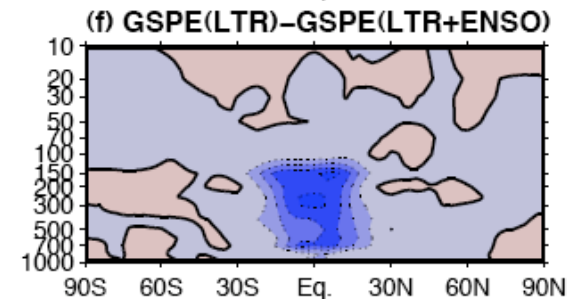
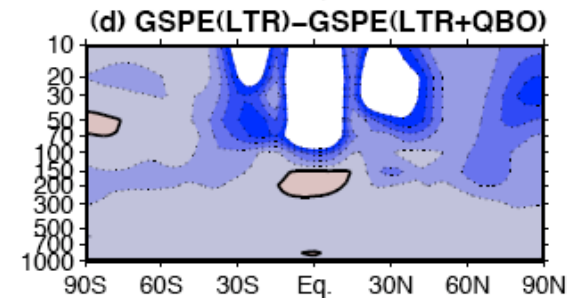
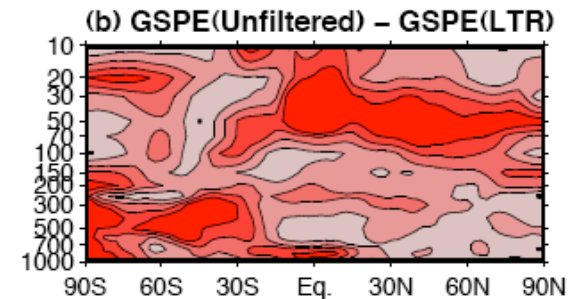
High-frequency periodicities like QBO and ENSO decrease H (shallow slope), more strongly for GSPE.

Volcanic forcing increases H similarly for DFA & GSPE.

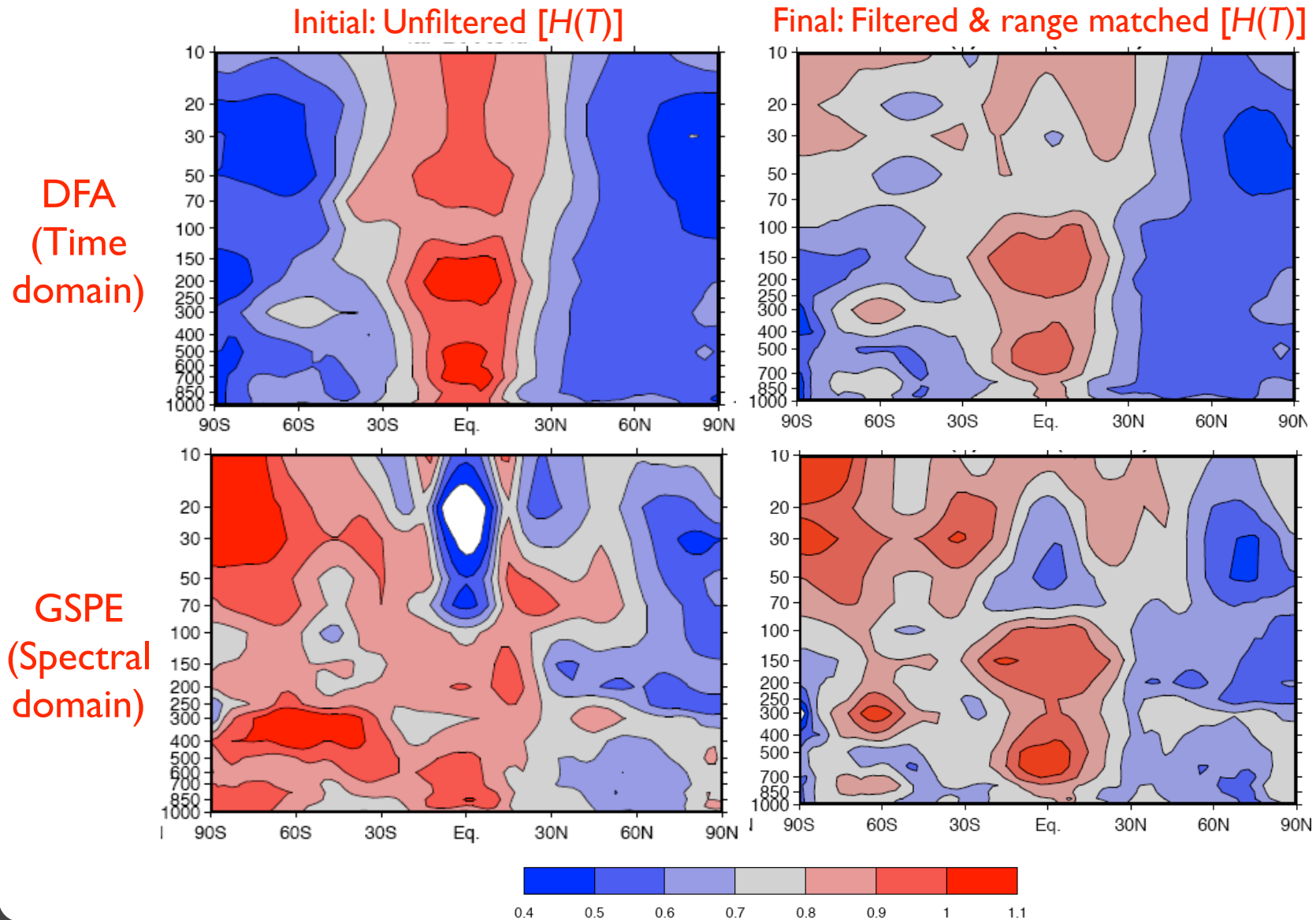
DFA



GSPE



Convergence of H Estimation Methods

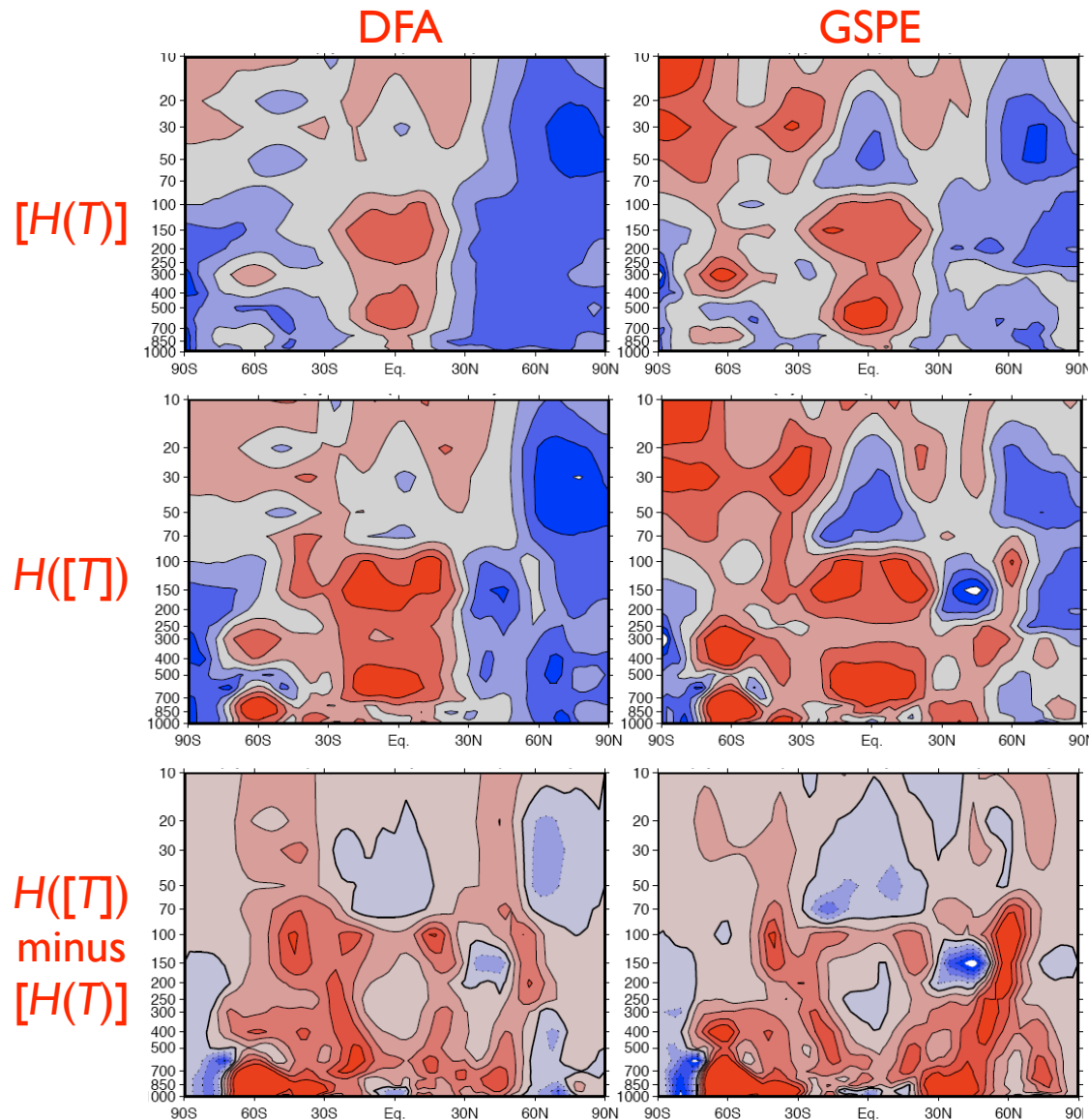


Power law behavior in the zonal-mean circulation

The previous analysis applied to H at each point.

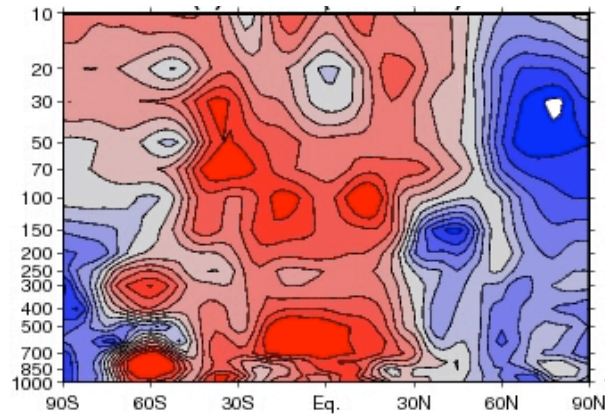
The extratropical zonal mean circulation has a relatively large H compared to point values for H .

This reflects eddy zonal flow interactions and energy/angular momentum constraints.

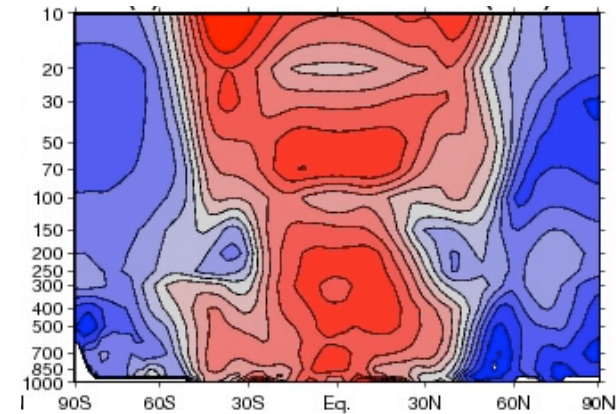


Modeling the Observed H Distribution

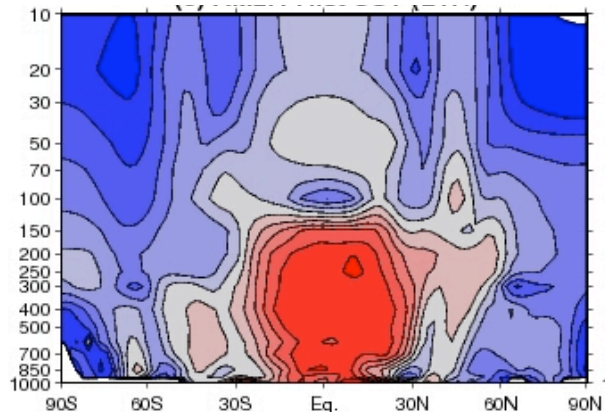
ERA40



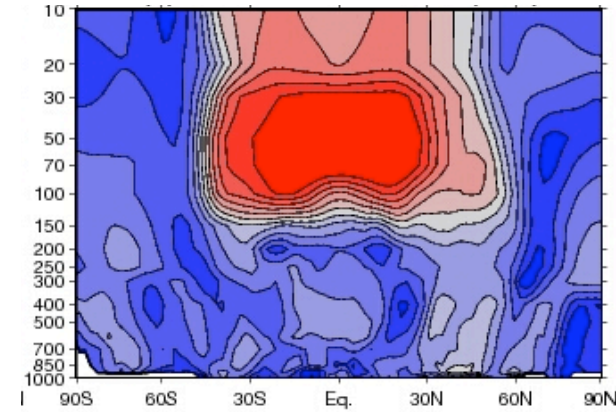
AM2.1 Historical SSTs + All Forcings



AM2.1 Historical SSTs + Fixed Forcings



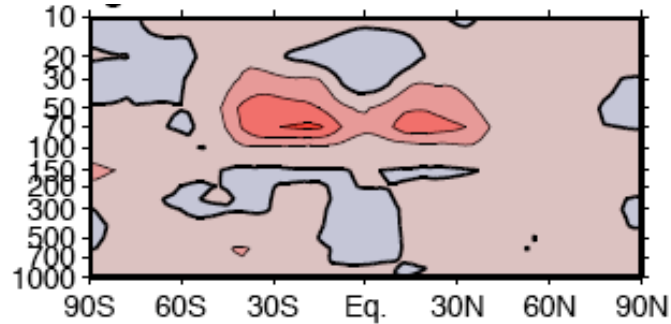
AM2.1 Climo SSTs + Volcanic Forcing



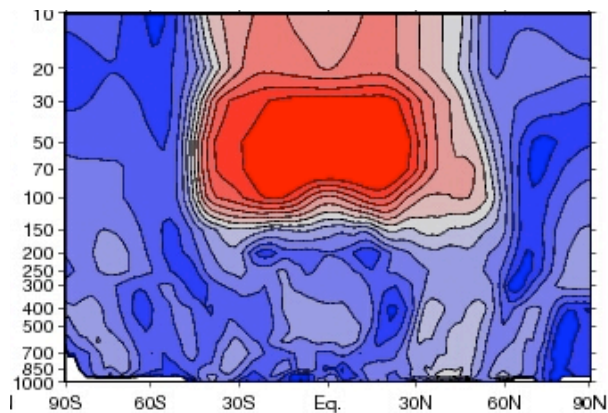
We use climate models to simulate and decompose H .

H signature in the stratosphere: Volcanoes

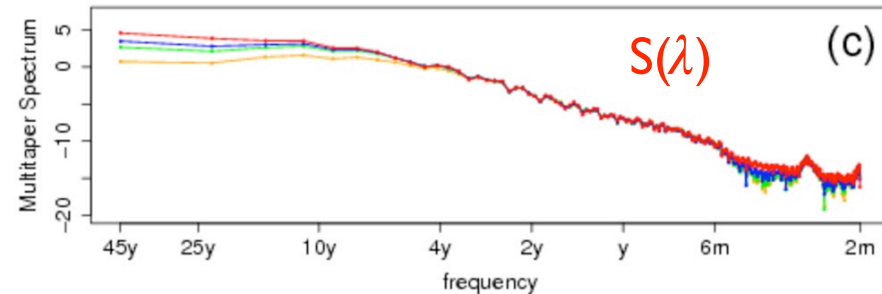
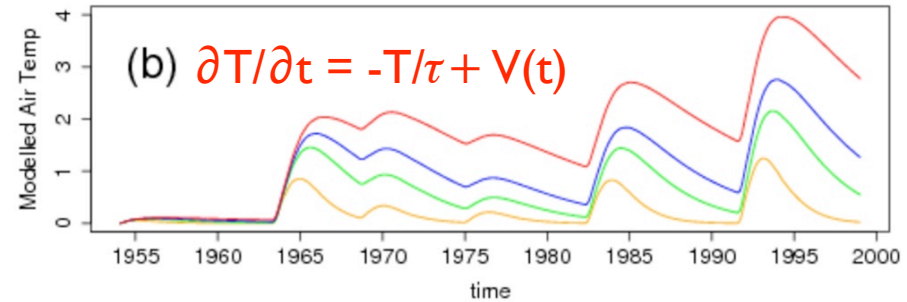
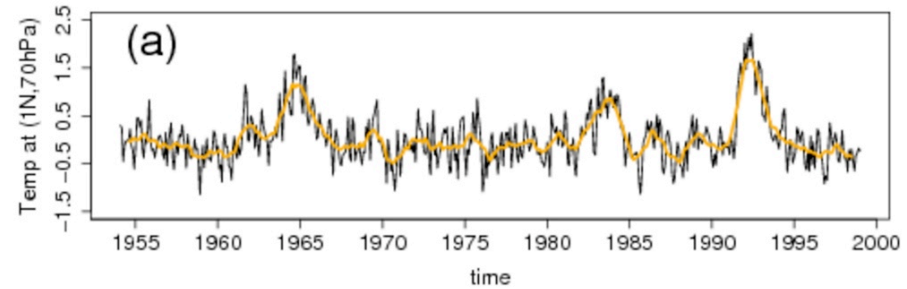
ERA40: Impact of filtering volcanic forcing



AM2.1 with volcanic forcing



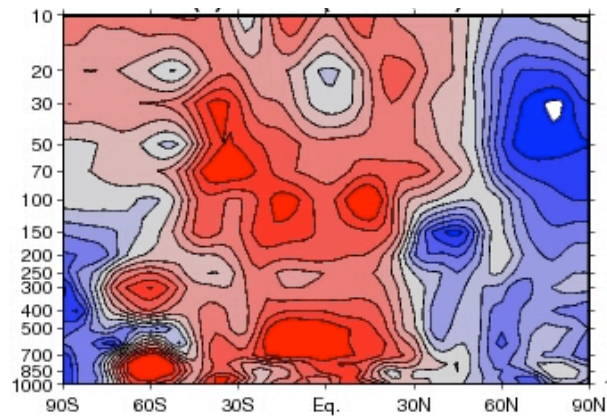
Simple Model for the Volcanic Response



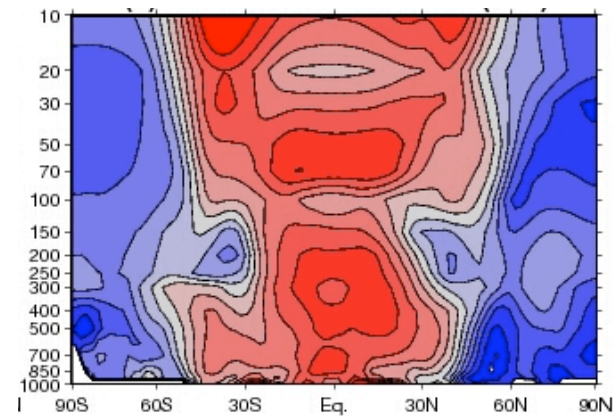
- *H* signature in stratosphere: direct response to volcanic forcing, not internally generated.

H signature in the tropics: SSTs

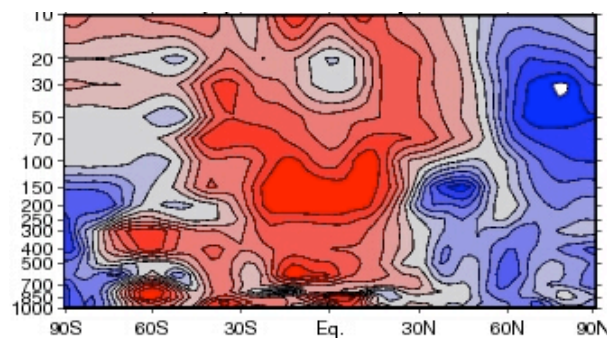
ERA40



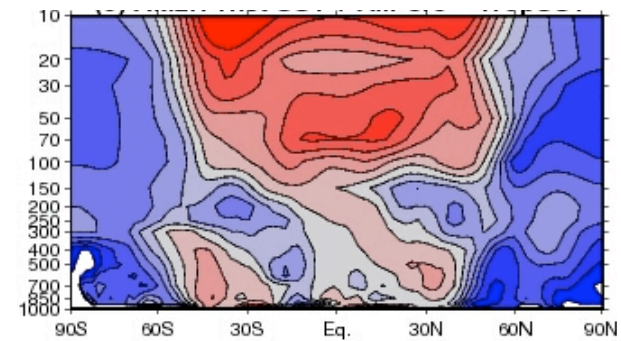
AM2.1



ERA 40 with Tropical SST filtering



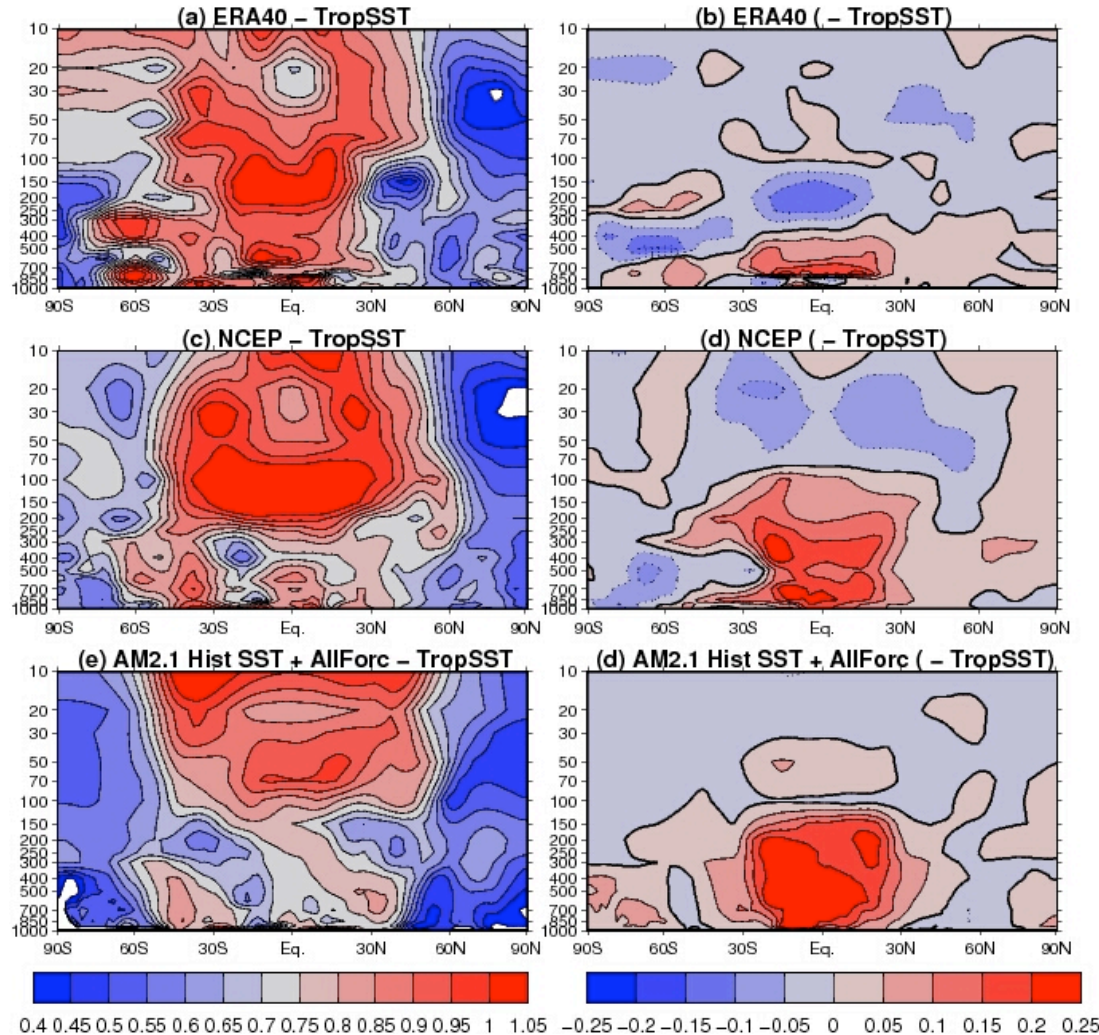
AM2.1 with Tropical SST filtering



- Tropical SSTs account for H signature in the model, but not in ERA40!

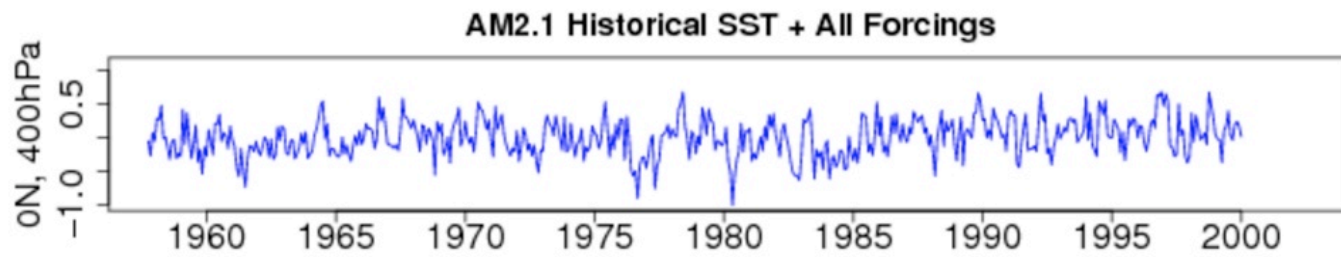
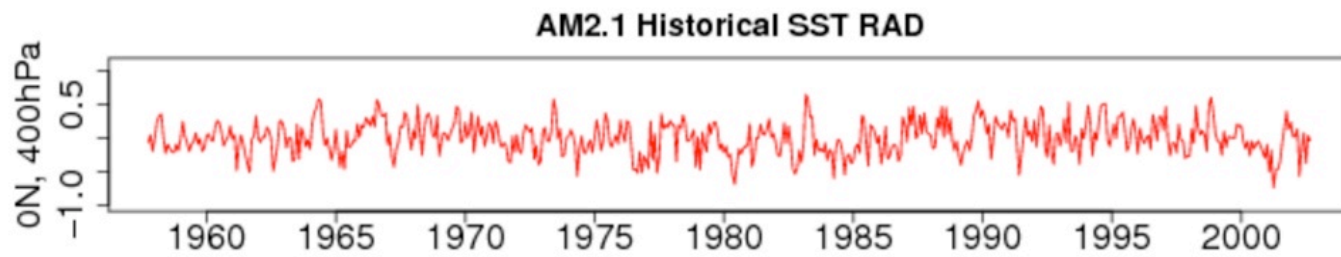
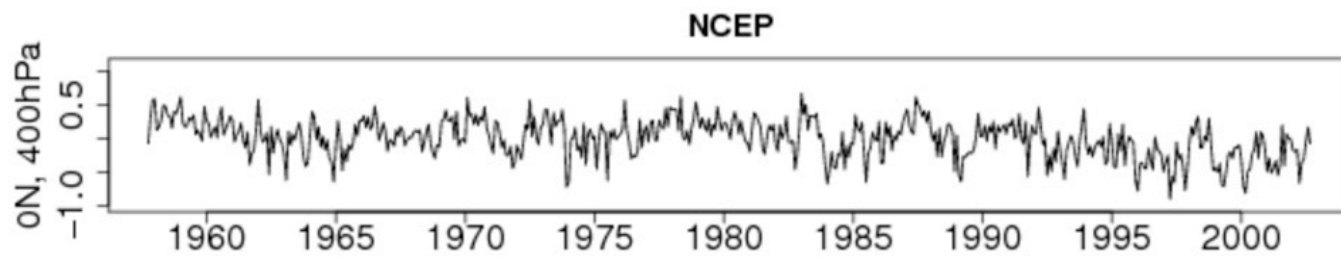
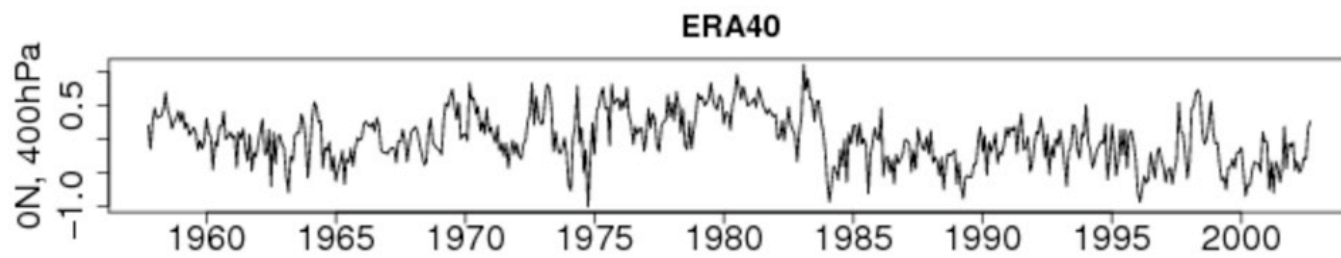
$H([T])$ with tropical SST filtering

Impact of tropical SST filtering



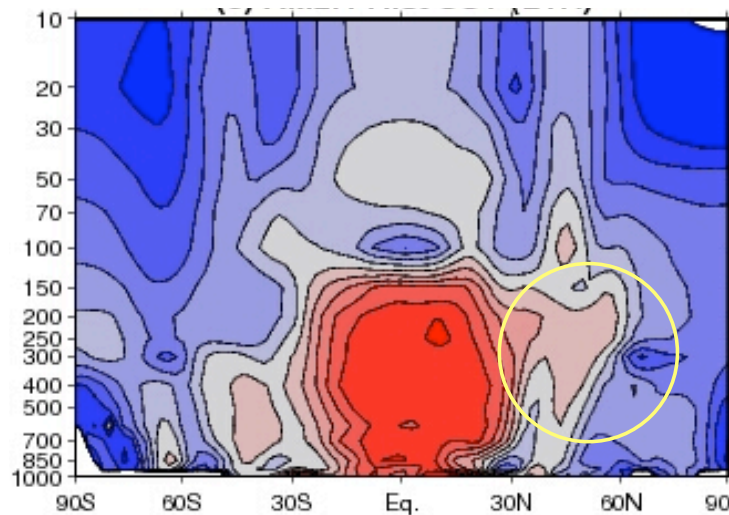
Notice the contrast between the NCEP and ERA40 products!

- NCEP Reanalysis & Model agree, and both disagree with ERA40. Data inhomogeneity issues?

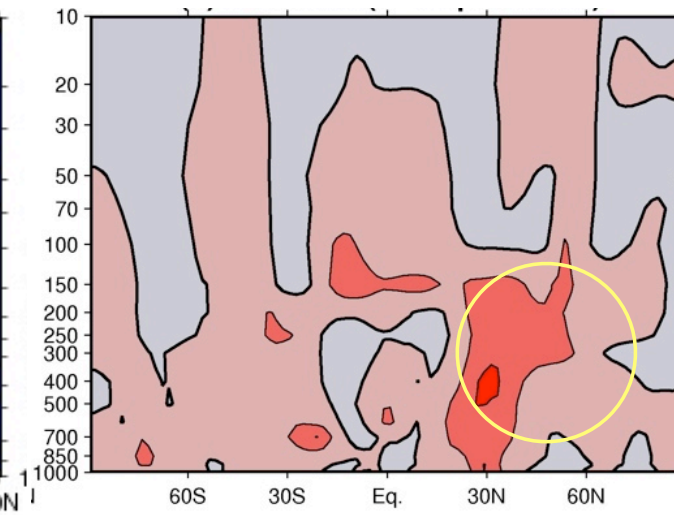


Effect of Eddy Mean-Flow Interactions

AM2.1 Historical SSTs



Impact of filtering index of zonal asymmetry in tropical SSTs.



Part of the extratropical H arises from wave forcing of zonal mean flows (ENSO like).

Conclusions

As an integral measure of low-frequency variability, H is a useful quantity to calculate.

But H signatures require careful interpretation:

- Climate trends,
- Volcanoes,
- Ocean-atmosphere coupling,
- Data problems.

Conclusions

Initial H estimates seemed non-robust, but we could account for the differences.

Three principal regions of large H :

- Tropical-subtropical lower stratosphere,
- Tropical troposphere,
- Low extratropical troposphere (for $[T]$)

Demonstrated that volcanic forcing gives rise to power-law behavior in the lower stratosphere.

Propose that tropical air-sea interaction responsible for tropospheric H signature.

Supplementary Figures

