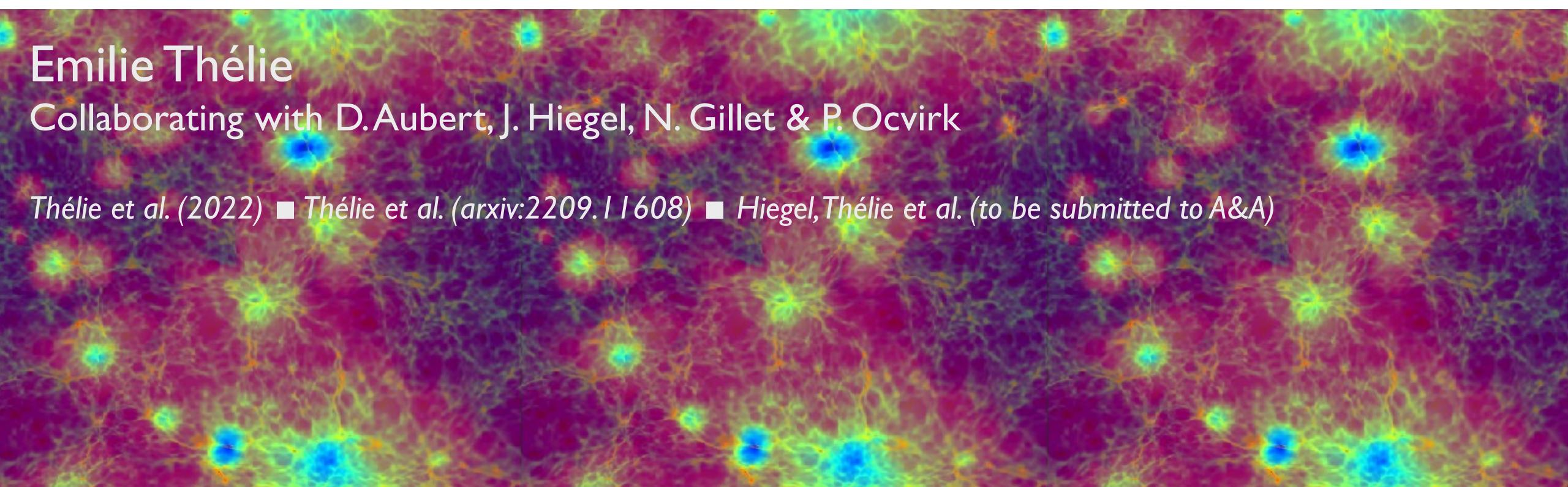


# Exploring the Epoch of Reionisation through its evolving topology

Emilie Thélie

Collaborating with D. Aubert, J. Hiegel, N. Gillet & P. Ocvirk

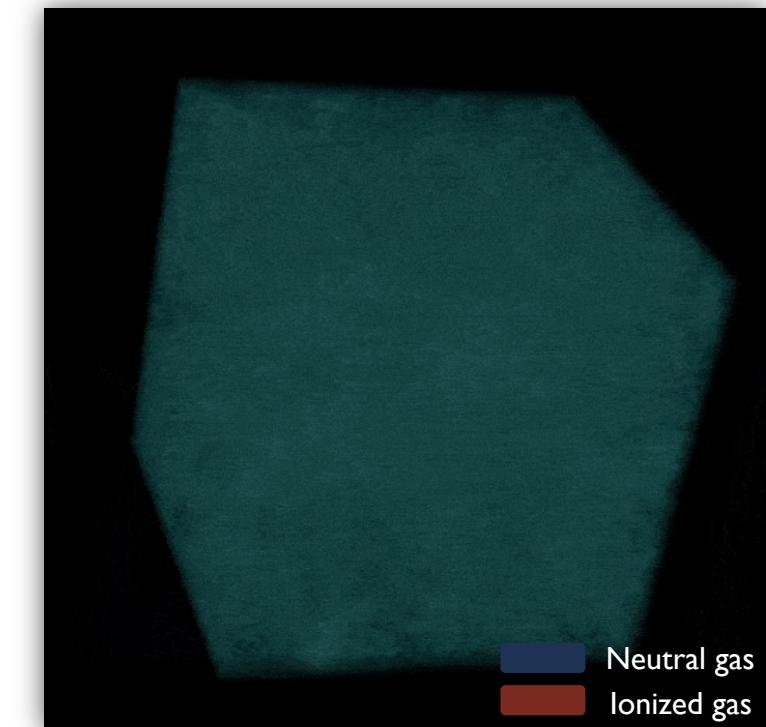
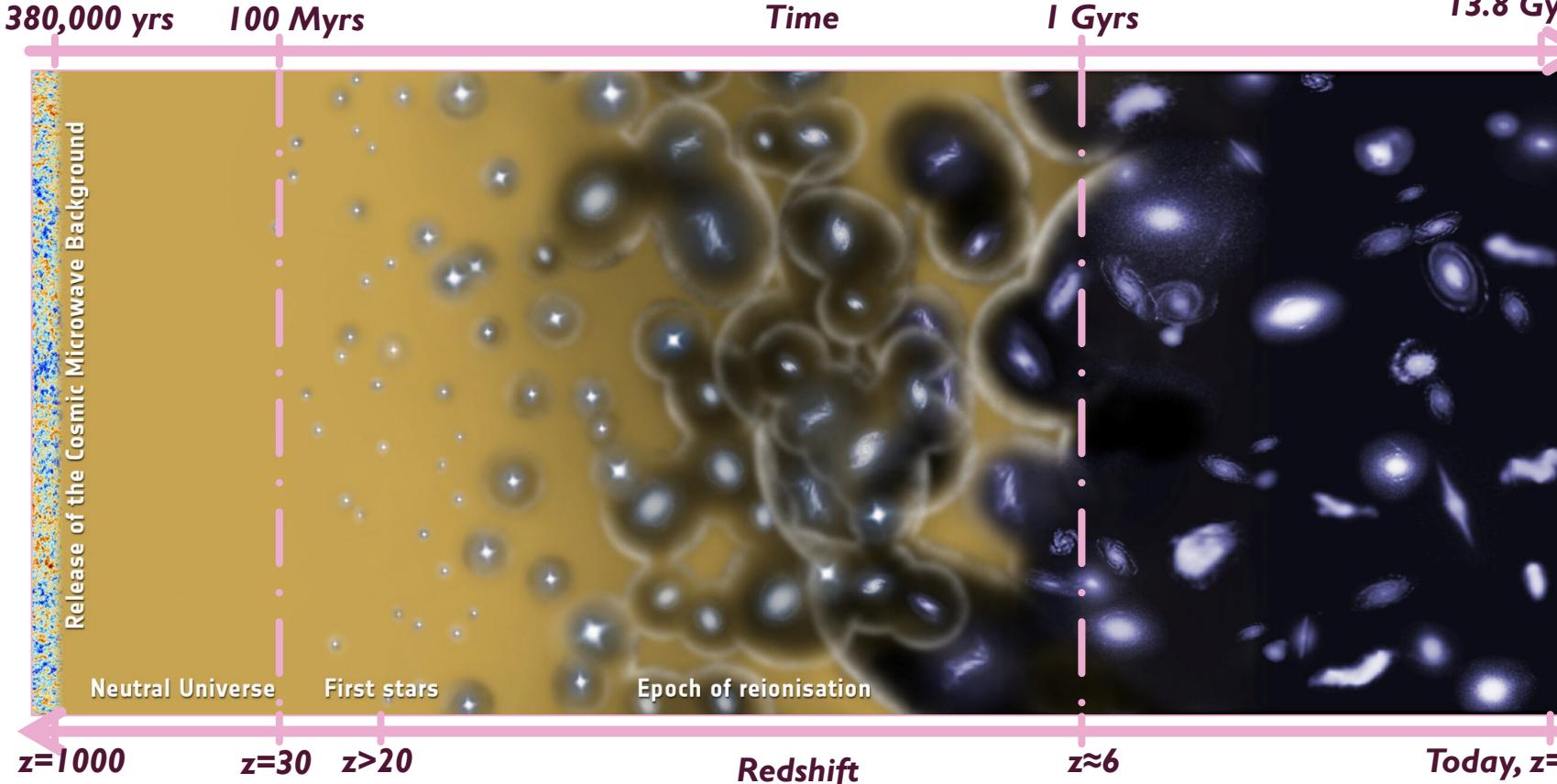
*Thélie et al. (2022)* ■ *Thélie et al. (arxiv:2209.11608)* ■ *Hiegel, Thélie et al. (to be submitted to A&A)*



# INTRODUCTION

## The Epoch of Reionisation (EoR)

Reionisation history – Credit: ESA



Last big transition our Universe has known: the gas goes from totally neutral to totally ionised

# INTRODUCTION

## Observations of the Epoch of Reionisation: *indirect observations*

- Emission from distant quasars → Lyman  $\alpha$  forest  
= distribution of hydrogen clouds along their line of sight

- HST, JWST : deep surveys of distant galaxies ( $z \geq 11$ )  
= galaxy distribution during the EoR

- CMB power spectrum  
= constraints on e.g. the optical depth  $\tau$

SMACS 0723 galaxy cluster  
NIRCam Image, JWST  
Credit: NASA



# INTRODUCTION

## Observations of the Epoch of Reionisation: **direct observations**

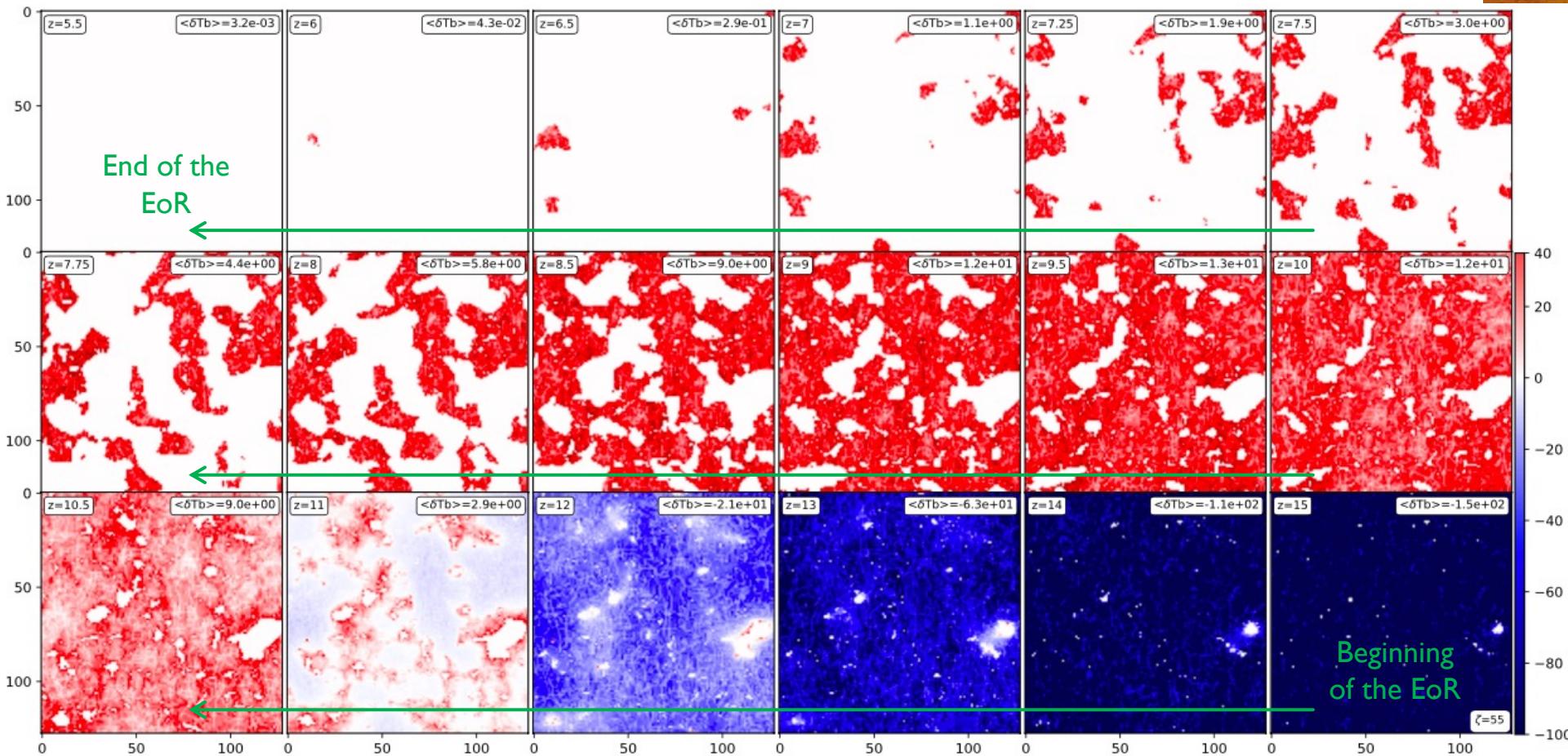
- Upcoming observations of the brightness temperature with the 21 cm signal  
= *distribution of neutral hydrogen gas at many frequencies*
- $\delta T_b(z) \sim x_{HI}(z)(1 + \delta_b(z))F(T)$



# INTRODUCTION

## Observations of the Epoch of Reionisation: **direct observations**

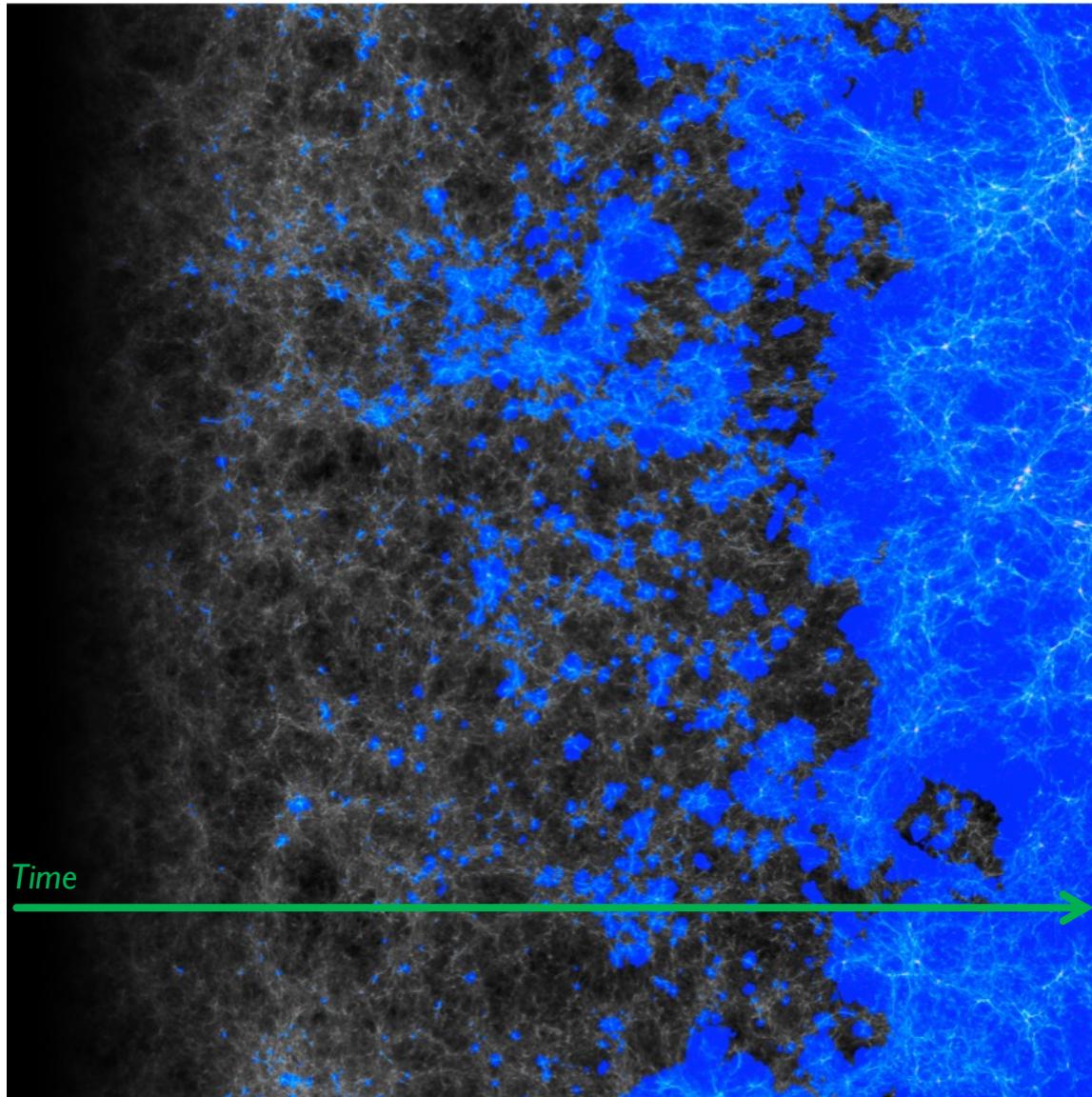
- Upcoming observations of the brightness temperature with the 21 cm signal  
= distribution of neutral hydrogen gas at many frequencies
- $\delta T_b(z) \sim x_{HI}(z)(1 + \delta_b(z))F(T)$



# INTRODUCTION

Why studying the Epoch of Reionisation?

CoDa II (64 cMpc/h)  
Credit: Ocvirk+20



## I. Understanding past and present structures thanks to the EoR

- Galaxy formation, evolution and properties  
= *constraints thanks to the EoR*
- Link cosmic web – EoR  
= *non-uniform distribution of matter at large scale*

## 2. Understanding the evolution of the EoR

- Timing of the EoR: beginning and end?
- Is reionisation uniform everywhere and for all the galaxies?
- Evolution of ionised and neutral bubbles

# INTRODUCTION

## Aims and existing studies

### How does the EoR happen?

Topological studies to analyse:

- *Growth of structures*
- *Ionised/neutral bubbles geometry, distribution, organisation*
- *Percolation, evolution of the process*

#### ***Size of neutral and ionised bubbles...***

... with diverse methods (e.g. FOF in Giri+19,  
spherical averages in Giri+18, ...)

#### ***Analyses of spatial structures...***

... with 21 cm power spectra (e.g.  
Gazagnes+21) and bispectra (e.g. Hutter+20)

A lot of  
existing  
methods  
within this  
field

#### ***Counts of 3D structures like peaks, tunnels and voids...***

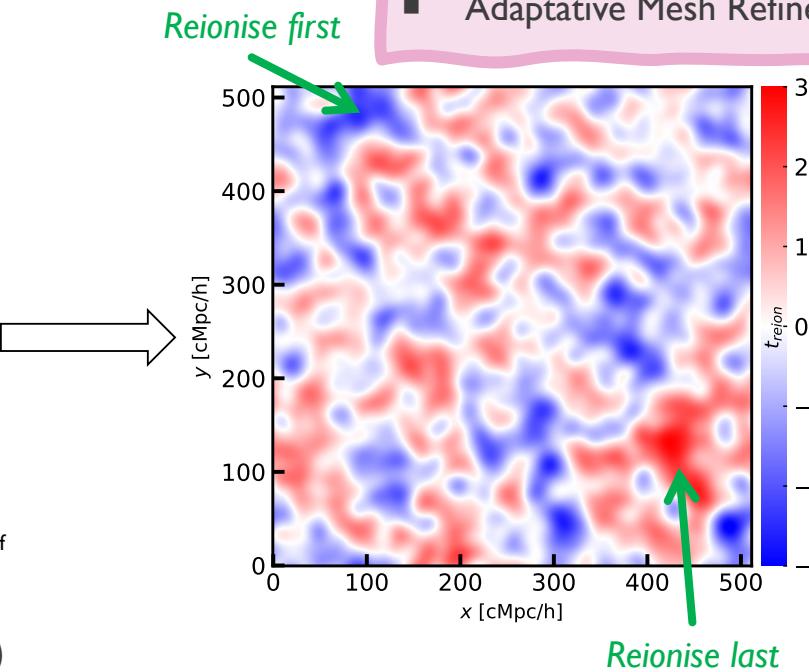
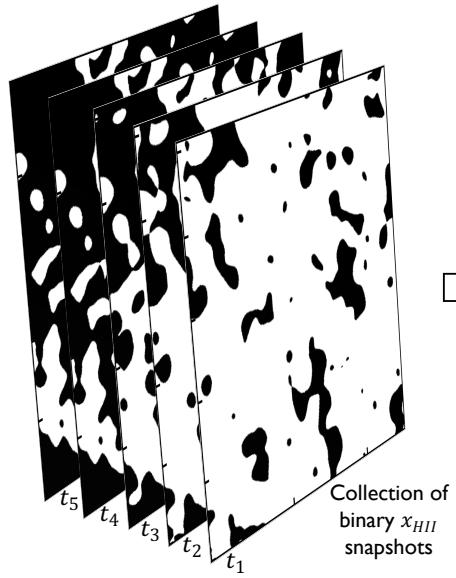
... with Betti numbers (e.g. Giri+20,  
Bianco+21)

#### ***Geometry of bubbles and percolation...***

- ... with Minkowski functionals (e.g. Friedrich+11, Chen+19)
- ... with the triangle correlation function (Gorce+19)

# REIONISATION TIMES

## Definition



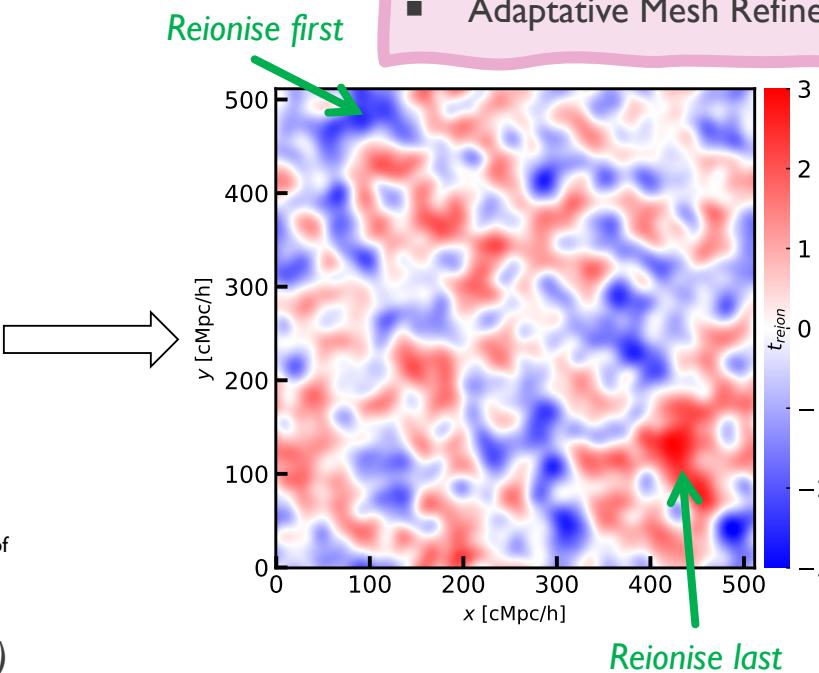
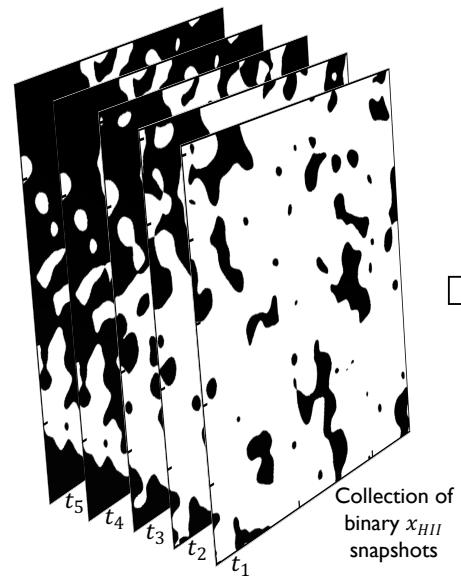
Thélie+ (arxiv:2209.11608)

### Map of reionisation times:

- Cell value = time at which the gas is reionised
- Spatial and temporal information about the reionisation process

# REIONISATION TIMES

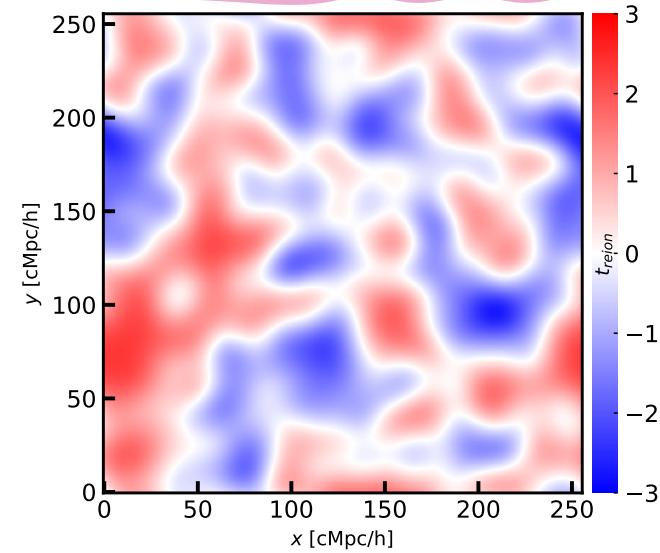
## Definition



**EMMA cosmological simulation (512 cMpc/h):**

- Hydrodynamics + radiative transfer
- Adaptive Mesh Refinement (AMR)

**21cmFAST semi-analytical simulation (256 cMpc/h)**



Thélie+ (arxiv:2209.11608)

## Map of reionisation times:

- Cell value = time at which the gas is reionised
- Spatial and temporal information about the reionisation process

## Reionisation simulations:

- **EMMA cosmological simulations** (Aubert+15) or **21cmFAST semi-analytical models** (Mesinger+11)
- Large scales: box size  $> 128 \text{ cMpc}/h$ ,  $1 \text{ cMpc}/h$  resolution

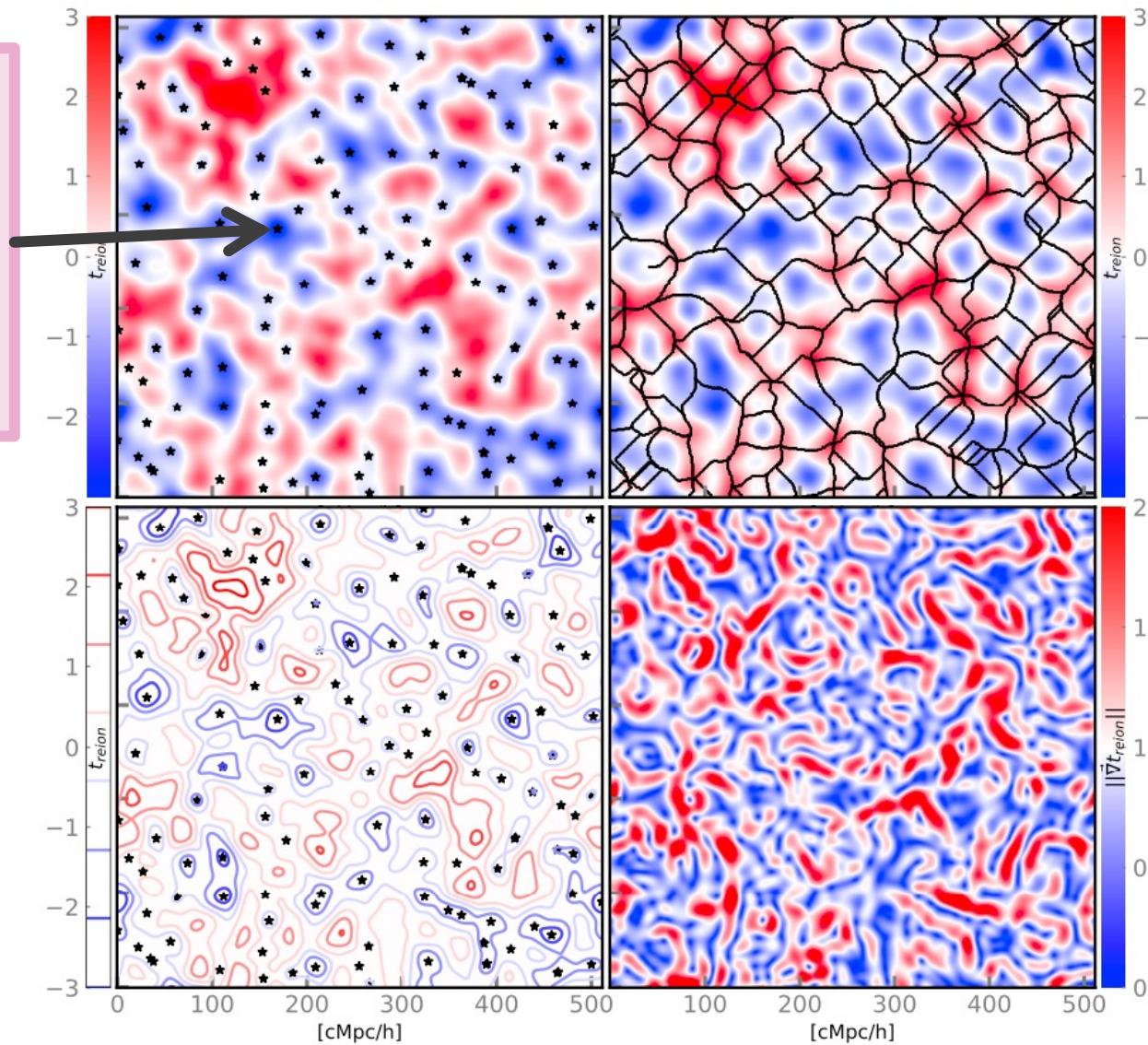
# REIONISATION TIMES

Evolving topology of the EoR

EMMA cosmological simulation maps (512 cMpc/h)  
Thélie+ (arxiv:2209.11608)

## Minima:

- “Reionisation seeds”: sources from which the ionisation fronts propagate
- First places to reionise



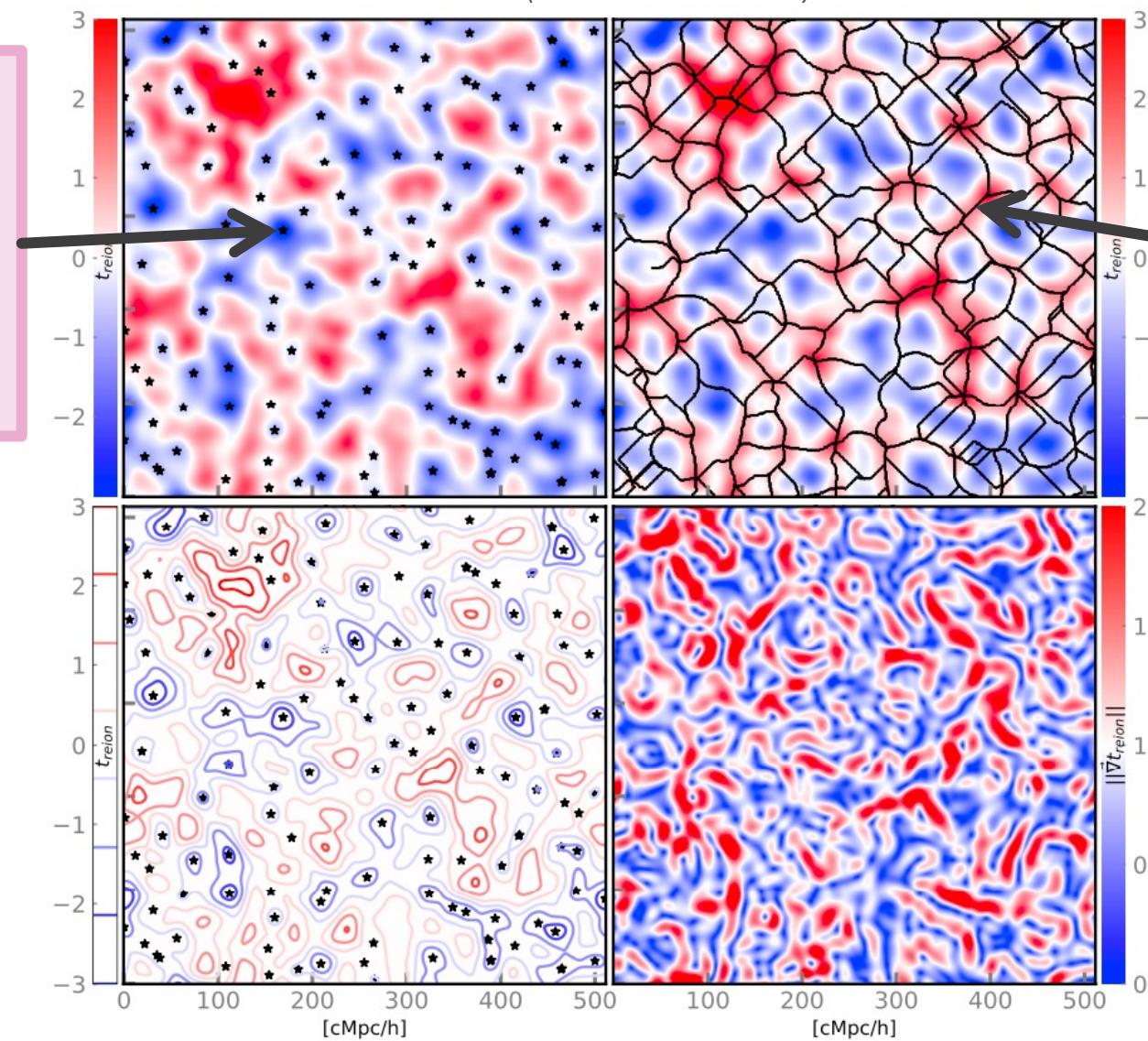
# REIONISATION TIMES

Evolving topology of the EoR

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Thélie+ (arxiv:2209.11608)

## Minima:

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## Reionisation patches:

- Extension of the radiative influence of a source

## Patches edges = skeleton:

- Percolation lines between ionisation fronts

# REIONISATION TIMES

Evolving topology of the EoR

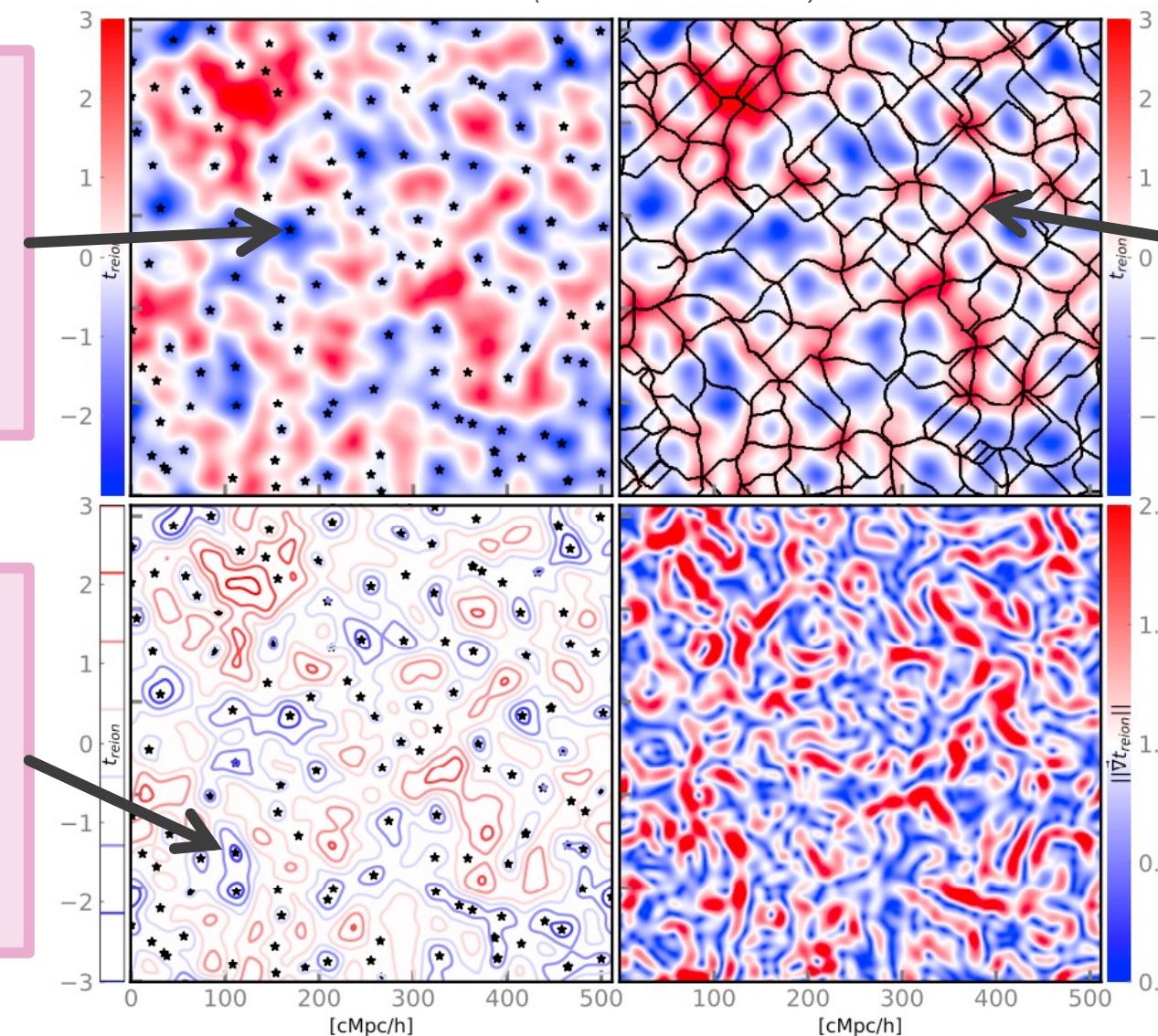
EMMA cosmological simulation maps (512 cMpc/h)  
Thélie+ (arxiv:2209.111608)

## Minima:

- “Reionisation seeds”: sources from which the ionisation fronts propagate
- First places to reionise

## Isocontours:

- Regions reached by ionisation fronts at the same time
- Size evolution of bubbles



## Reionisation patches:

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## Patches edges = skeleton:

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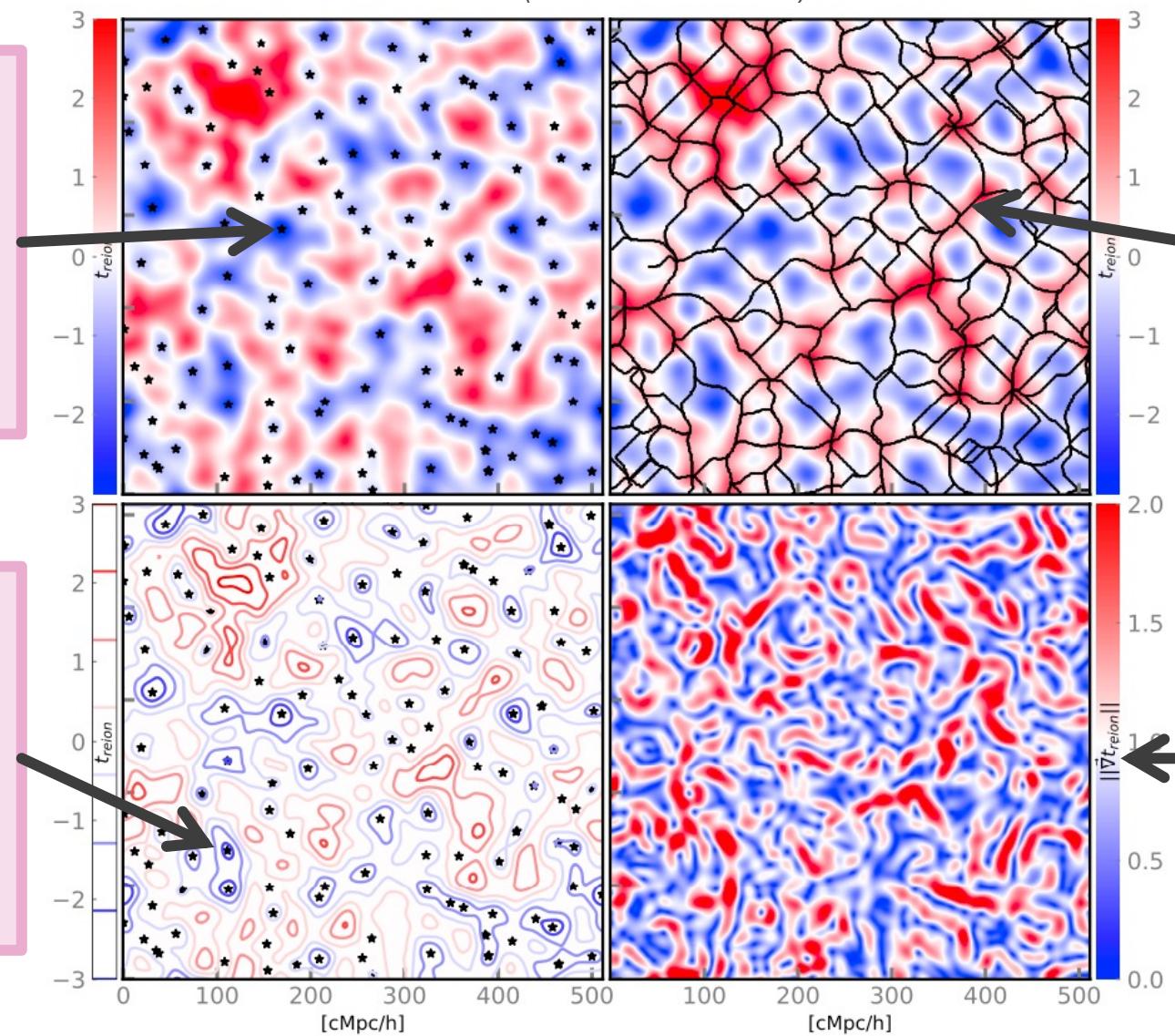
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Thélie+ (arxiv:2209.11608)

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## Isocontours:

- Regions reached by ionisation fronts at the same time
- Size evolution of bubbles

## Gradients:

- $\|\vec{\nabla} t_{reion}\| \sim \frac{\Delta t}{\Delta x} \sim v_{reion}^{-1}$
- Inverse of the ionisation fronts speed

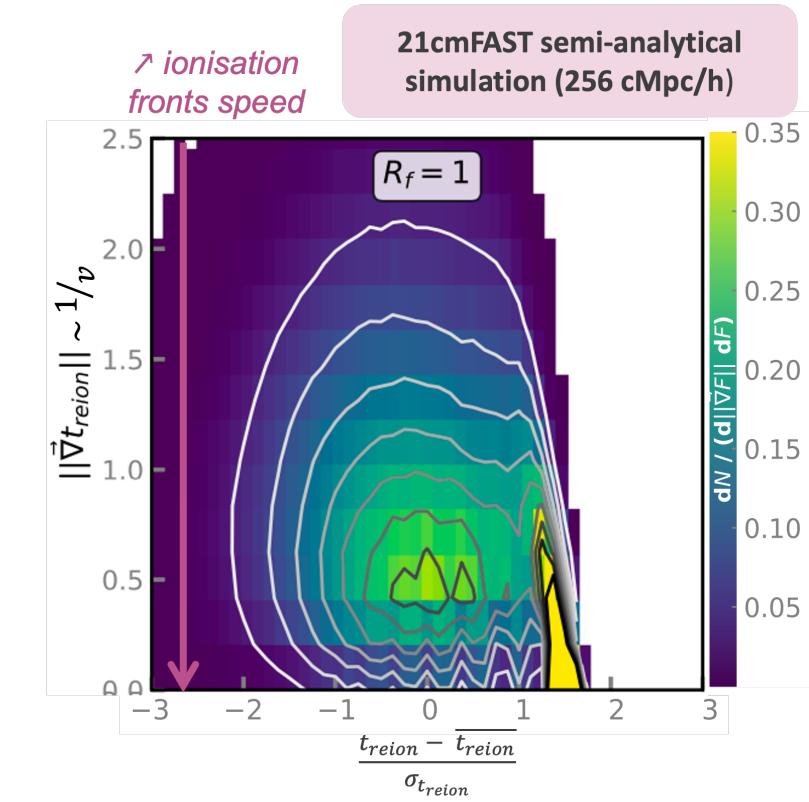
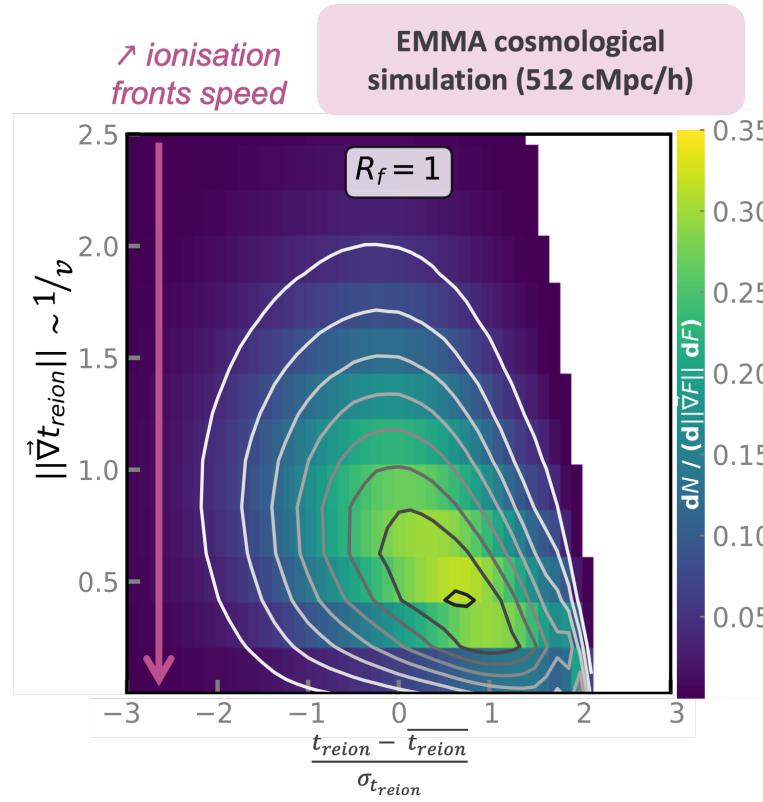
# REIONISATION TIMES

What can we do with it?

## Analysing this map:

- Comparisons between **models of simulations**
- **Geometrical characterisation of models of the EoR**
- Comparison to the **gaussian random fields theory**
- Study of the **evolving topology of the reionisation process**

Thélie+  
(arxiv:2209.11608)



# A. REIONISATION TIMES RECONSTRUCTED FROM THE 21 CM SIGNAL

Hiegel, Thélie+ (to be submitted)

# B. PROPAGATION OF IONISATION FRONTS AROUND MATTER FILAMENTS

Thélie+22 (A&A)

# C. REIONISATION TIMES: TOPOLOGY AND GAUSSIAN RANDOM FIELD (GRF) THEORY

Thélie+ (submitted to A&A, arXiv: 2209.11608)

# A.21 CM TO $t_{reion}$

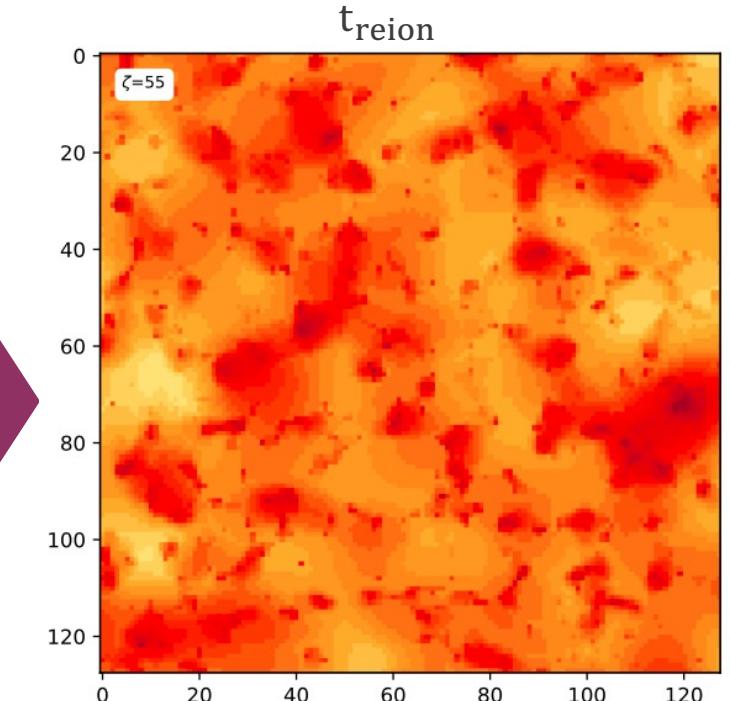
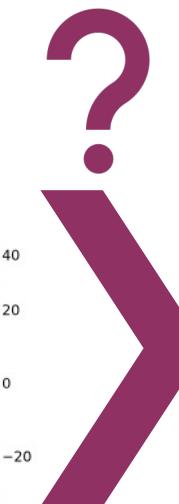
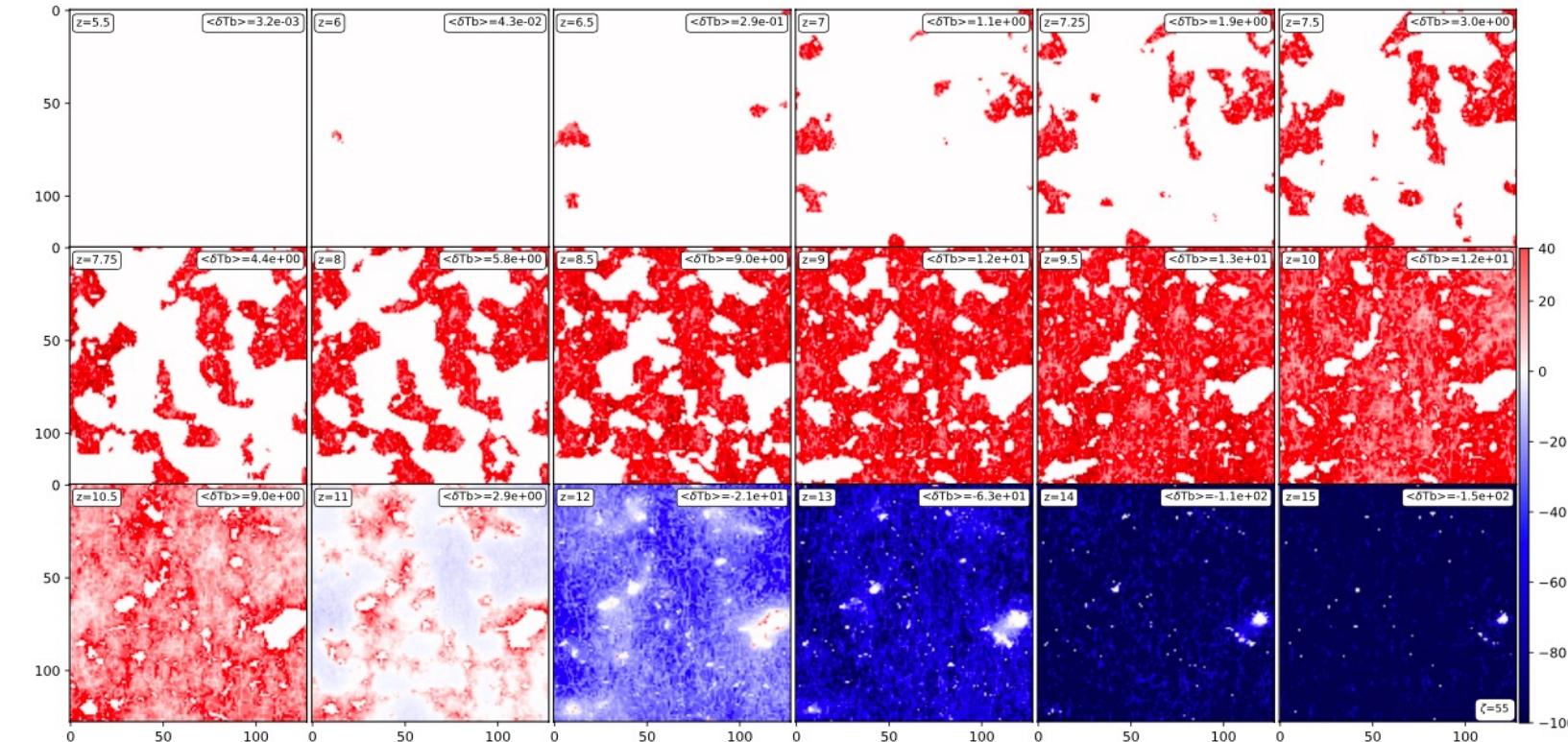
Reconstruction of reionisation times from 21 cm signal maps

Is it possible to reconstruct  
reionisation times maps from  
observations?

Future observations with SKA :

- 2D maps on the plane of the sky...
- ... of neutral gas distribution during the EoR...
- ... at many observational redshifts

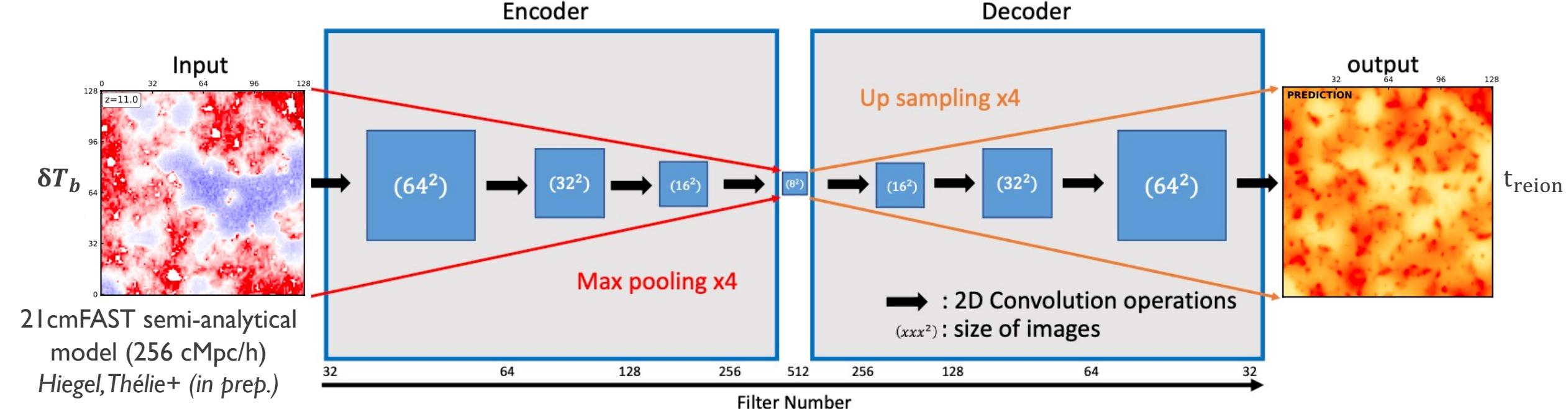
2D images of the 21 cm signal at many frequencies:  $\delta T_b(z)$



21cmFAST semi-analytical model (256 cMpc/h)  
Hiegel, Thélie+ (in prep.)

# A.21 CM TO $t_{reion}$

## Convolutional neural network (CNN)

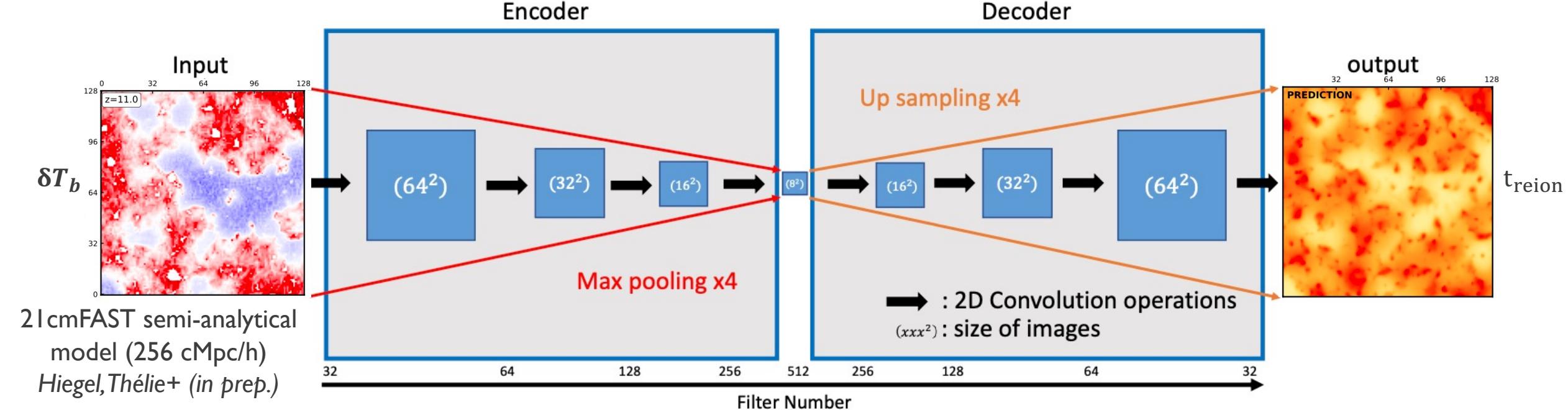


## CNN

- Developed with Tensorflow and Keras
- U-net: 2 parts with the same number of images and filters
  - Encoder = convolve and reduce dimension
  - Decoder = deconvolve and increase back to original dimension

# A.21 CM TO $t_{reion}$

## Convolutional neural network (CNN)



### CNN

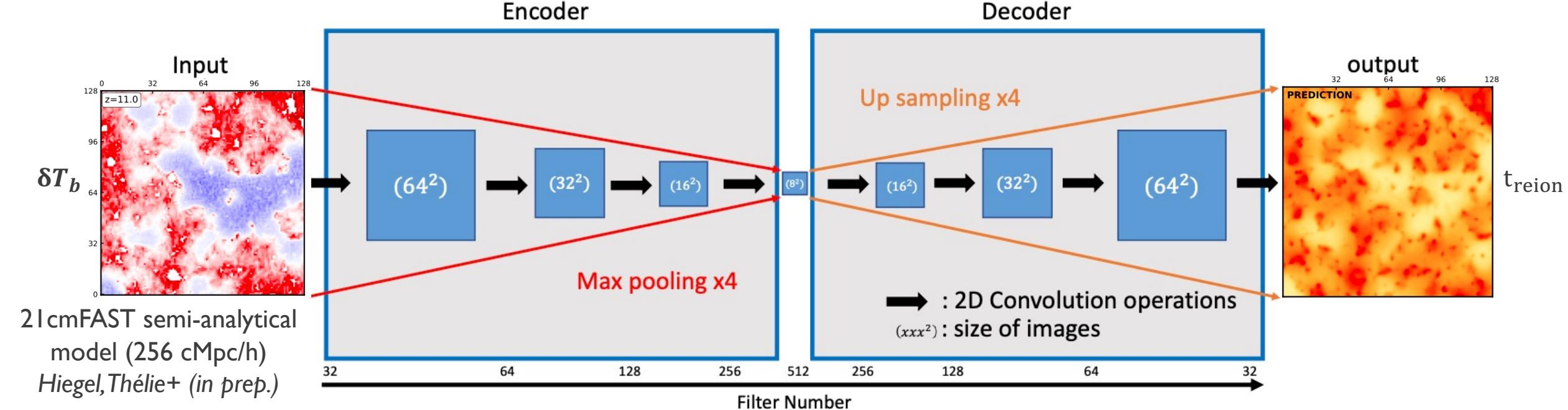
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### SIMULATIONS

- 50 21cmFAST semi-analytical simulations (256 cMpc/h)
- 2 models of reionisation with varying ionisation emissivity of galaxies  $\zeta \in \{30, 55\}$

# A.21 CM TO $t_{reion}$

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### CNN

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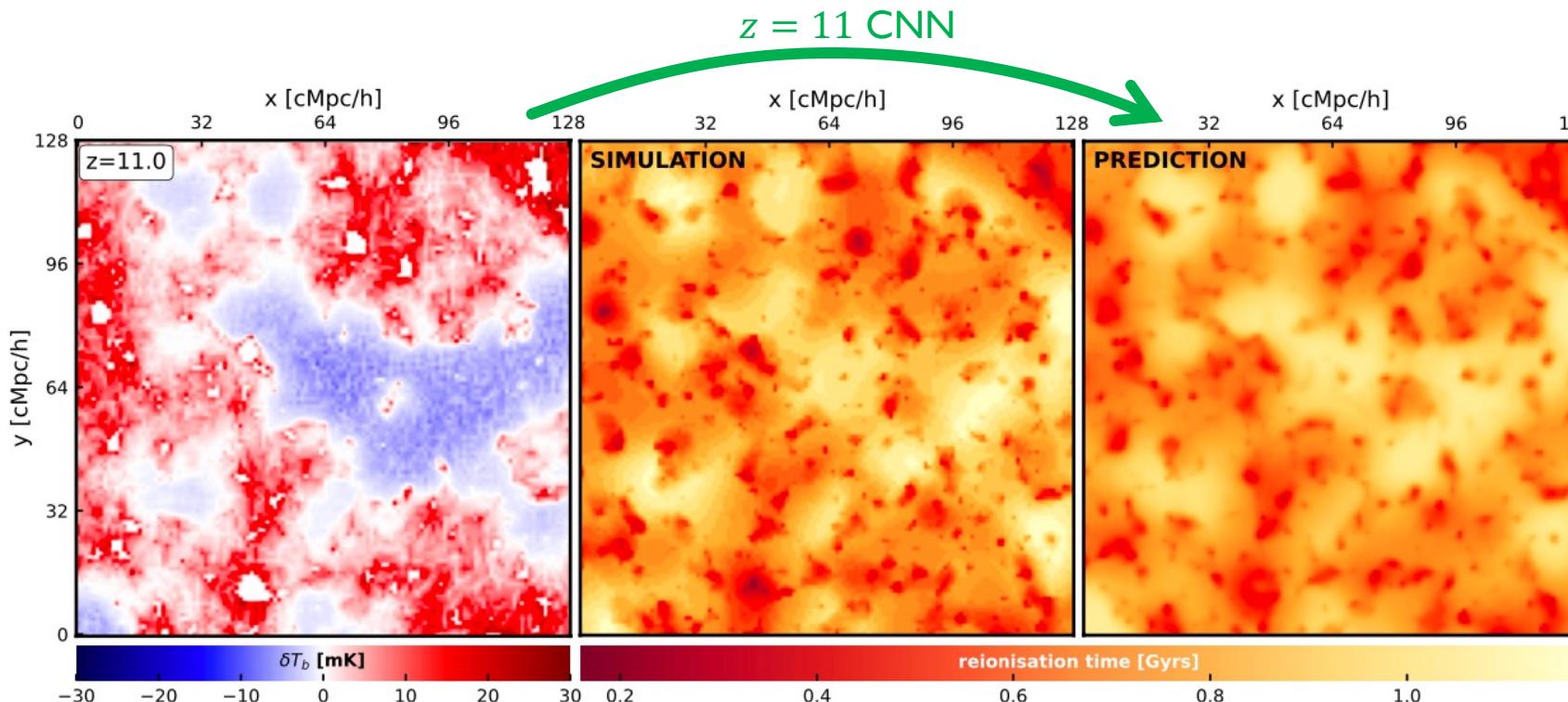
- 50 21cmFAST semi-analytical simulations (256 cMpc/h)
- 2 models of reionisation with varying ionisation emissivity of galaxies  $\zeta \in \{30, 55\}$

### LEARNING

- One prediction per observational redshift
- 35,000 images:
  - 90% for the training set
  - 10% for the test set

## A.21 CM TO $t_{reion}$

Reconstruction of reionisation times from 21 cm signal maps



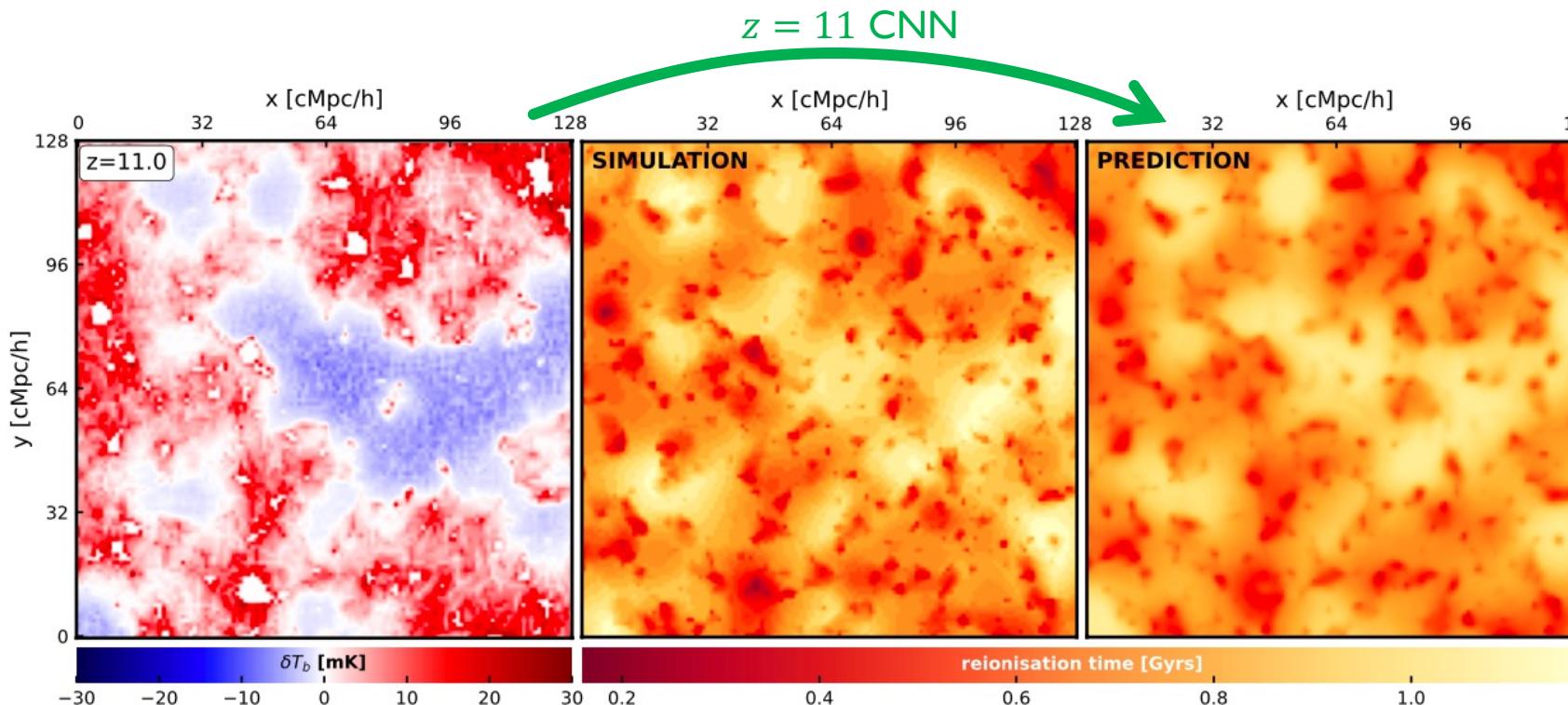
21cmFAST semi-analytical model (256 cMpc/h)

Hiegel, Thélie+ (in prep.)

- Example with  $z_{obs} = 11$ : less small structures but visually really close

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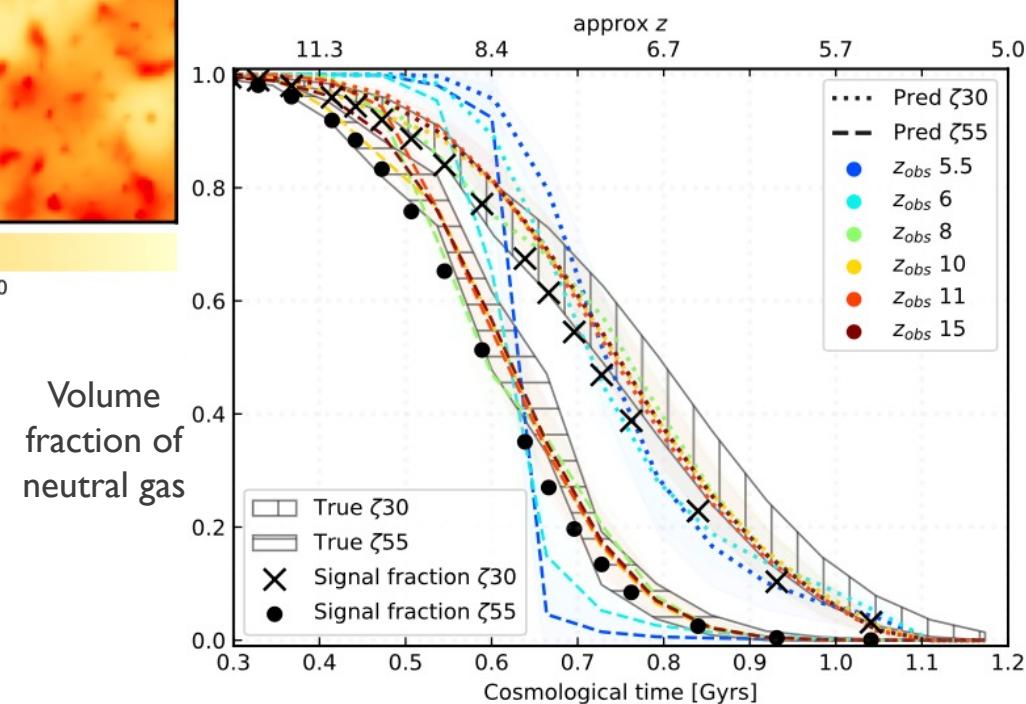
Reconstruction of reionisation times from 21 cm signal maps



- Example with  $z_{obs} = 11$ : less small structures but visually really close
- Access to the **entire** reionisation history with a **unique** observational redshift: *reconstruction of the past and extrapolation of the future*

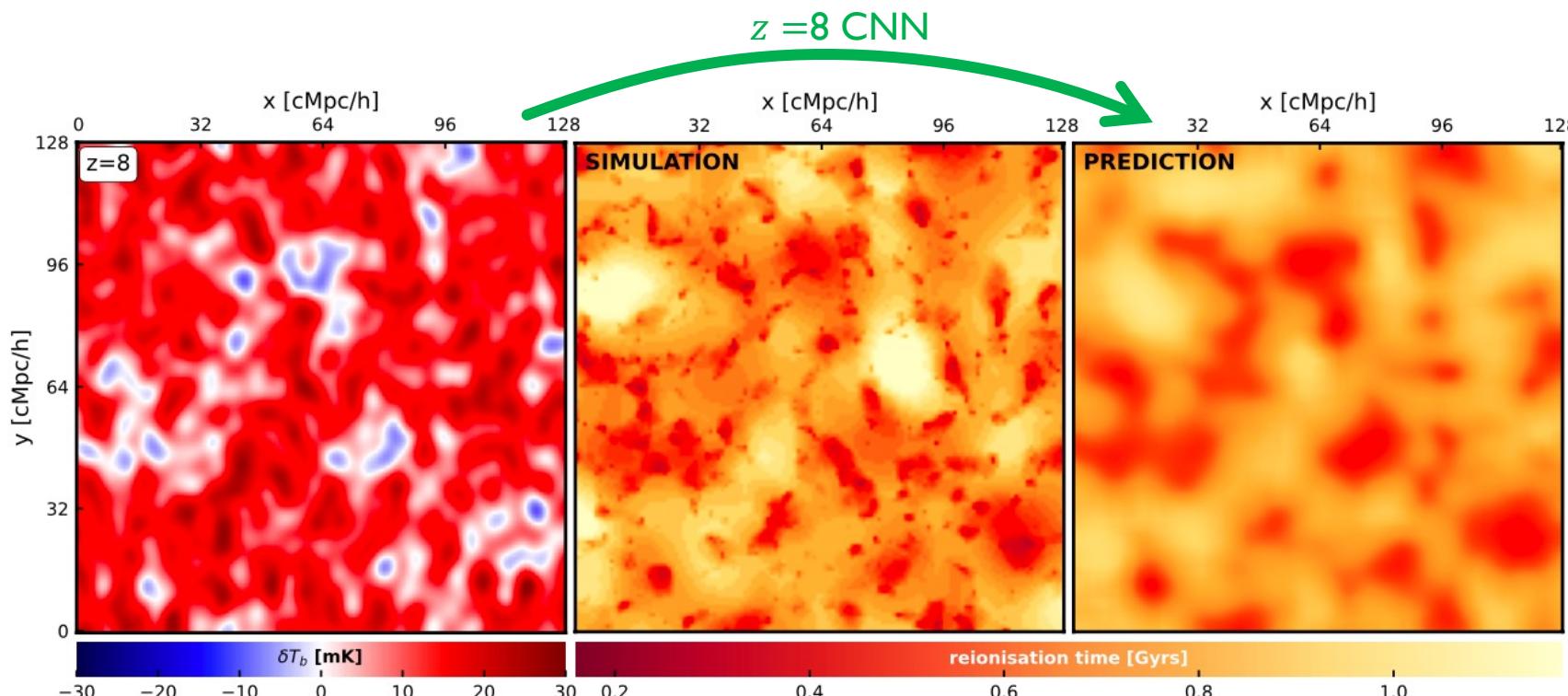
21cmFAST semi-analytical model (256 cMpc/h)

Hiegel, Thélie+ (in prep.)



# A.21 CM TO $t_{reion}$

Reconstruction of reionisation times from 21 cm signal maps including instrumental noise



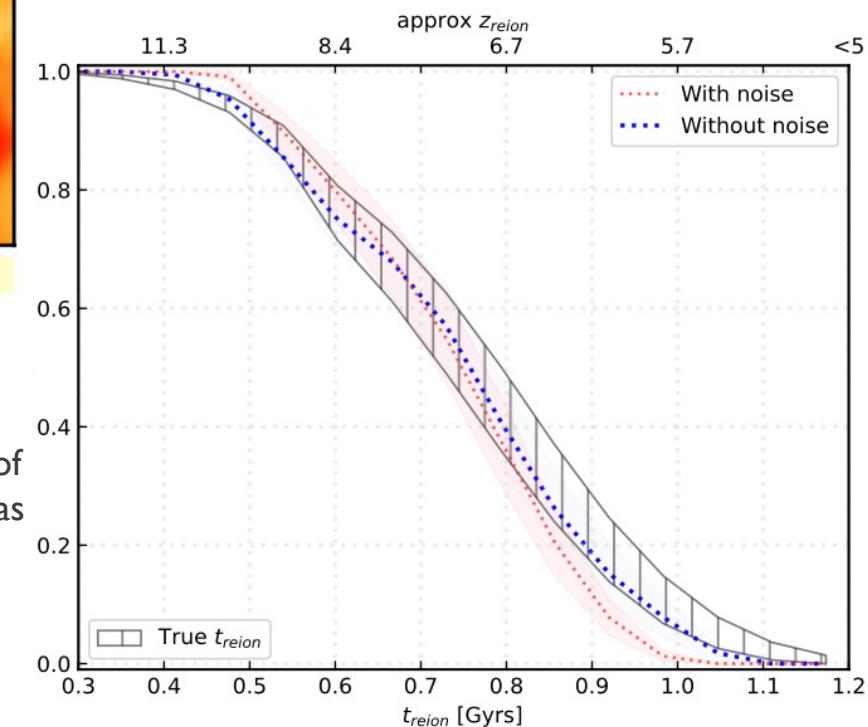
WORK IN PROGRESS

- Example with  $z_{obs} = 8$ : reconstructed maps even more smoothed
- But statistics not that far away from reconstructions without noise

## Noise characteristics (Tool21cm, Giri+20)

- Daily scan: 6h
- Integration time: 10s
- Number of observations hours: 1000h
- Maximum baseline: 2 km

21cmFAST semi-analytical model (256 cMpc/h)  
Hiegel, Thélie+ (in prep.)



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# B. REIONISATION PATCHES

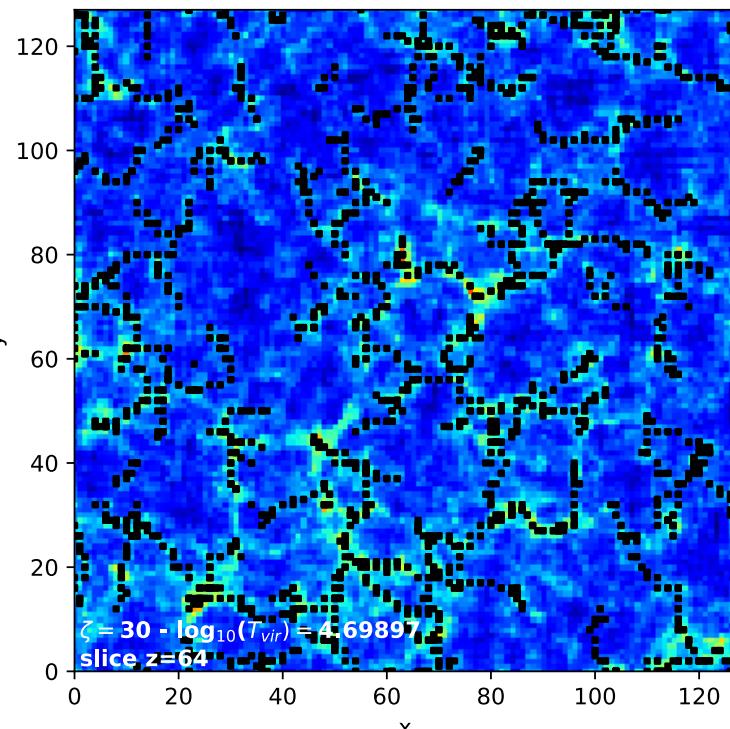
Propagation of ionisation fronts in 21cmFAST maps

How does the radiation from reionisation sources propagate?

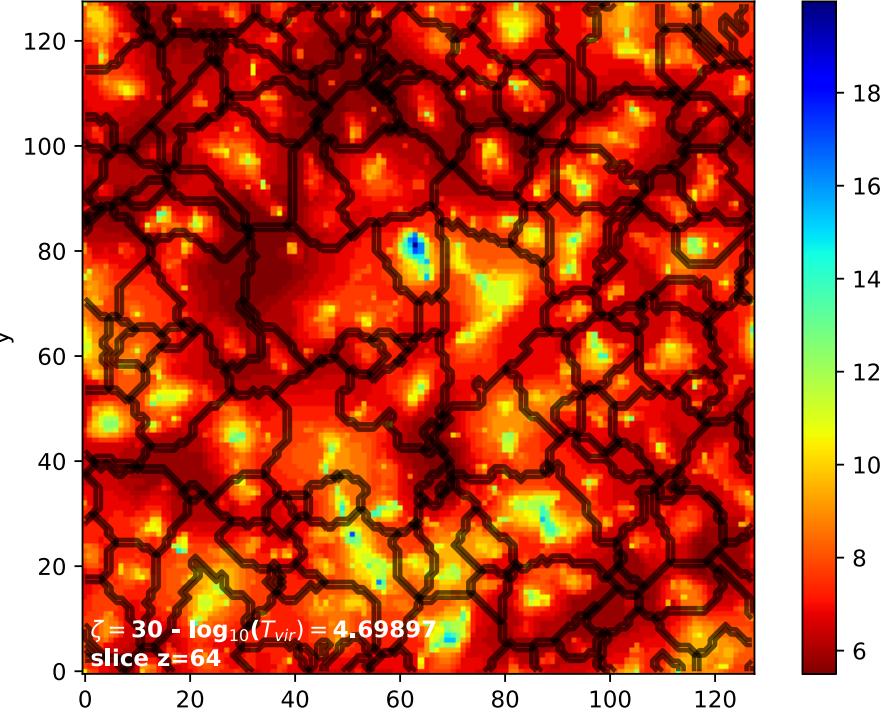
Analysis of the **topology of reionisation redshifts** with **DisPerSE** (Sousbie+11):

- **Radiative influence of reionisation sources on their environment:** orientation of reionisation patches with respect to the matter filaments
- **Geometry of reionisation patches, percolation**

Gas density + filaments



Reionisation redshifts + patches



Maps from a 21cmFAST semi-analytical model (128 cMpc/h)

Thélie+ (2022)

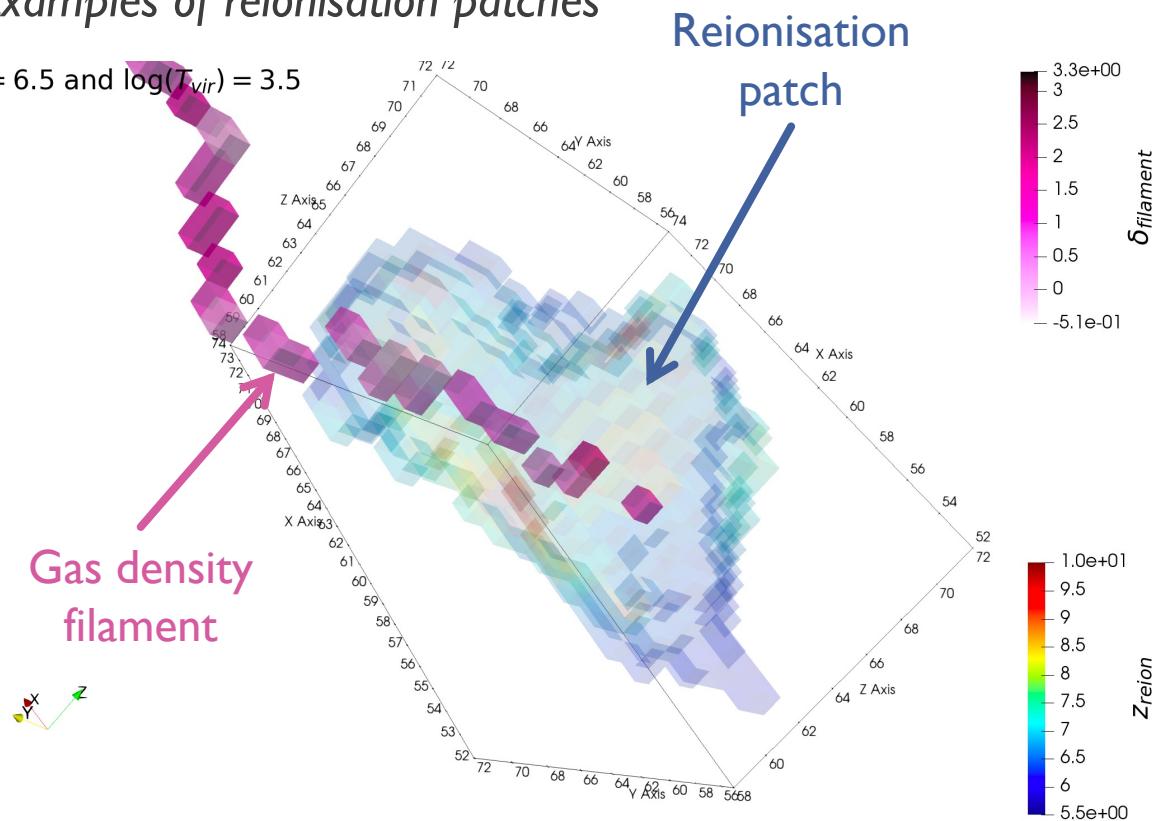
$z_{reion}$

- Maxima = first places to reionise
- Propagation of ionisation fronts along gradient lines
- Reionisation patches = extension of radiative influence of the sources

## B. REIONISATION PATCHES

Examples of reionisation patches

$$\zeta = 6.5 \text{ and } \log(T_{vir}) = 3.5$$

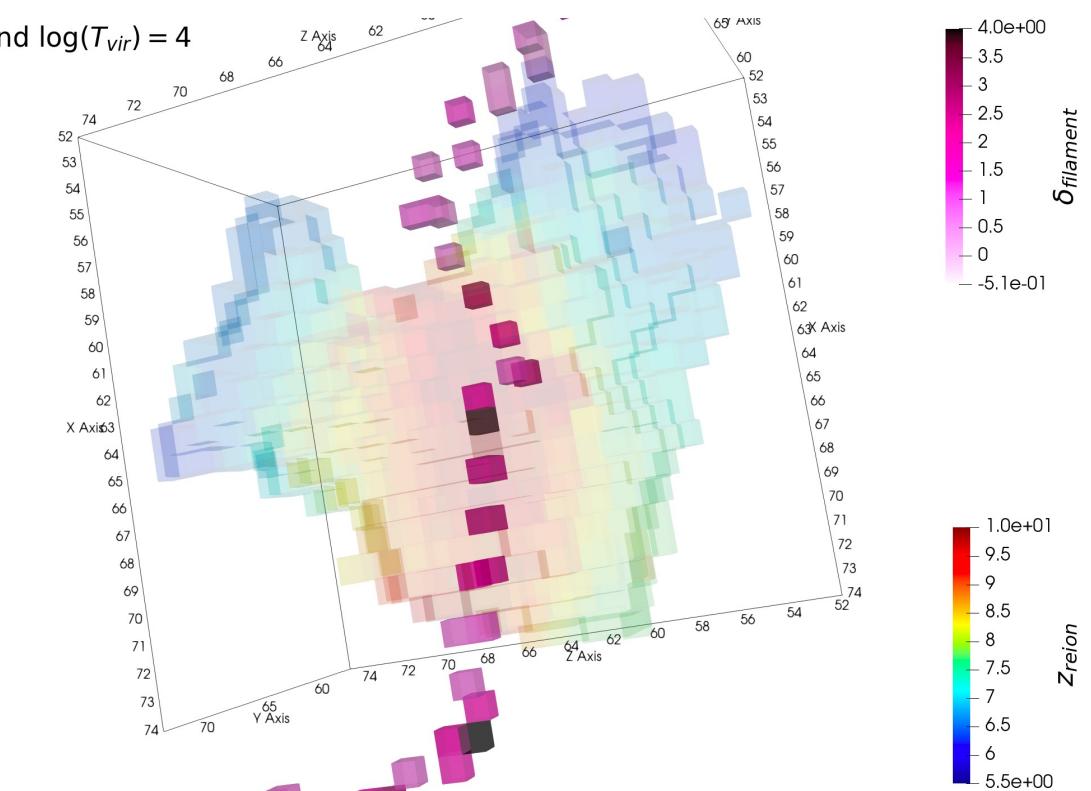


Reionisation  
patch

$$\delta_{\text{filament}}$$

filament

$$\zeta = 12 \text{ and } \log(T_{vir}) = 4$$



$$\delta_{\text{filament}}$$

Zreion

From a 21cmFAST semi-analytical model (128 cMpc/h)

Thélie+ (2022)

Reionisation patch aligned with the filament

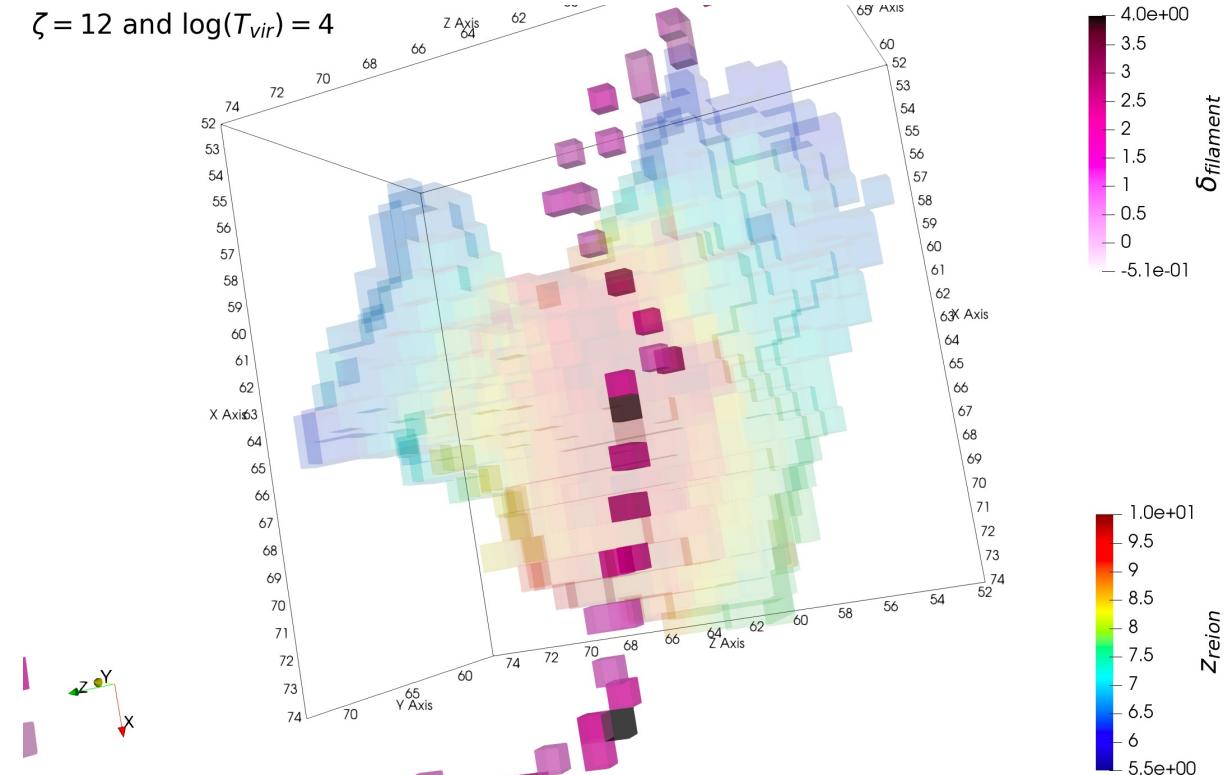
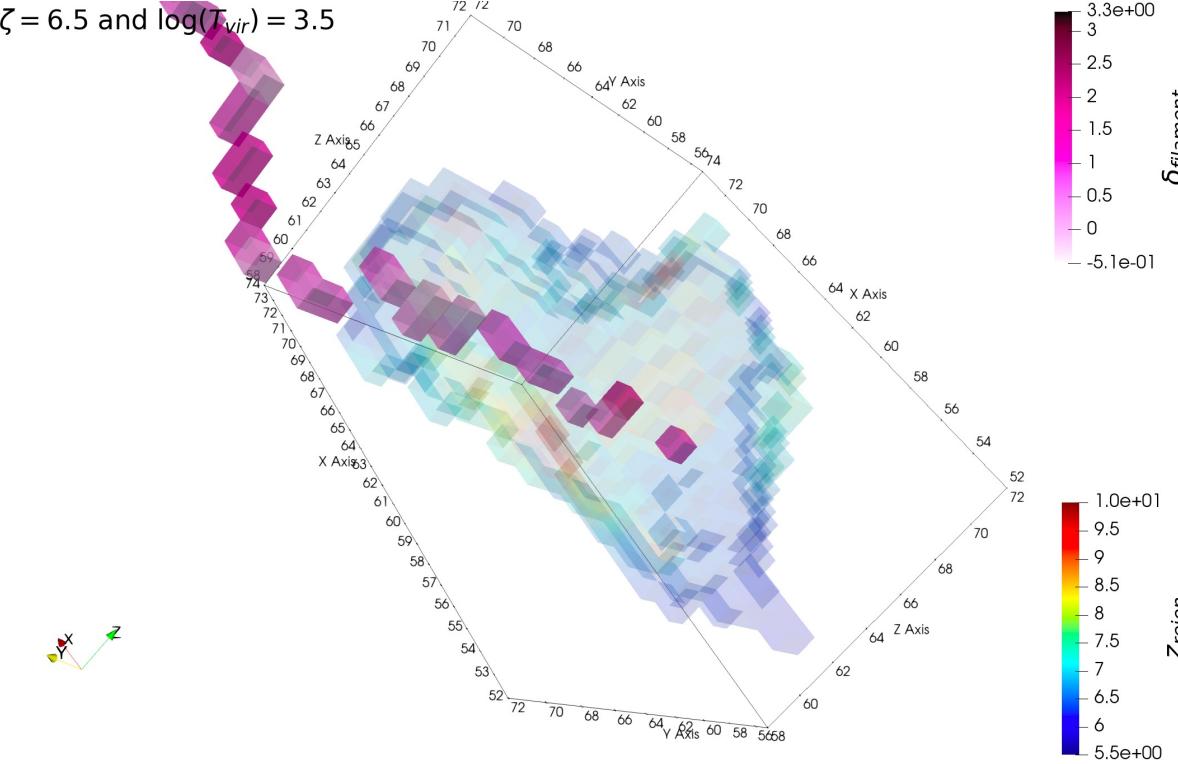
= radiations follow the matter filament

Reionisation patch in a butterfly shape

= radiations follow the path of least resistance

# B. REIONISATION PATCHES

**Results:** orientation of patches with respect to filaments



From a 21cmFAST semi-analytical model (128 cMpc/h)  
Thélie+ (2022)

## DOMINANT CASE (47%)

### Prolate-aligned patch

= beaded sources along the matter filament  
(that supposedly have same properties of emissivity, age...)

## UNDER-REPRESENTED CASE (2%)

### “Butterfly” patch

= isolated and/or strong emitter driving reionisation

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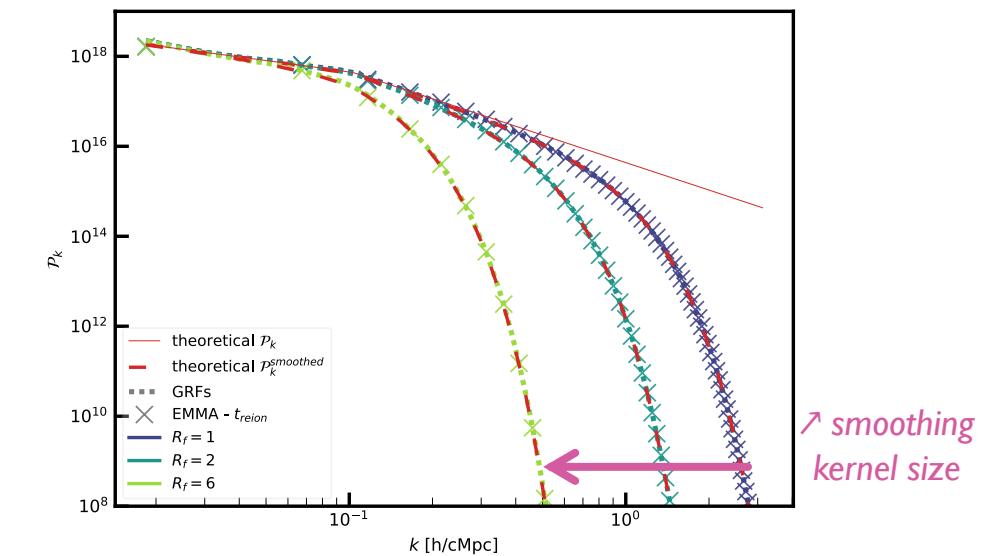
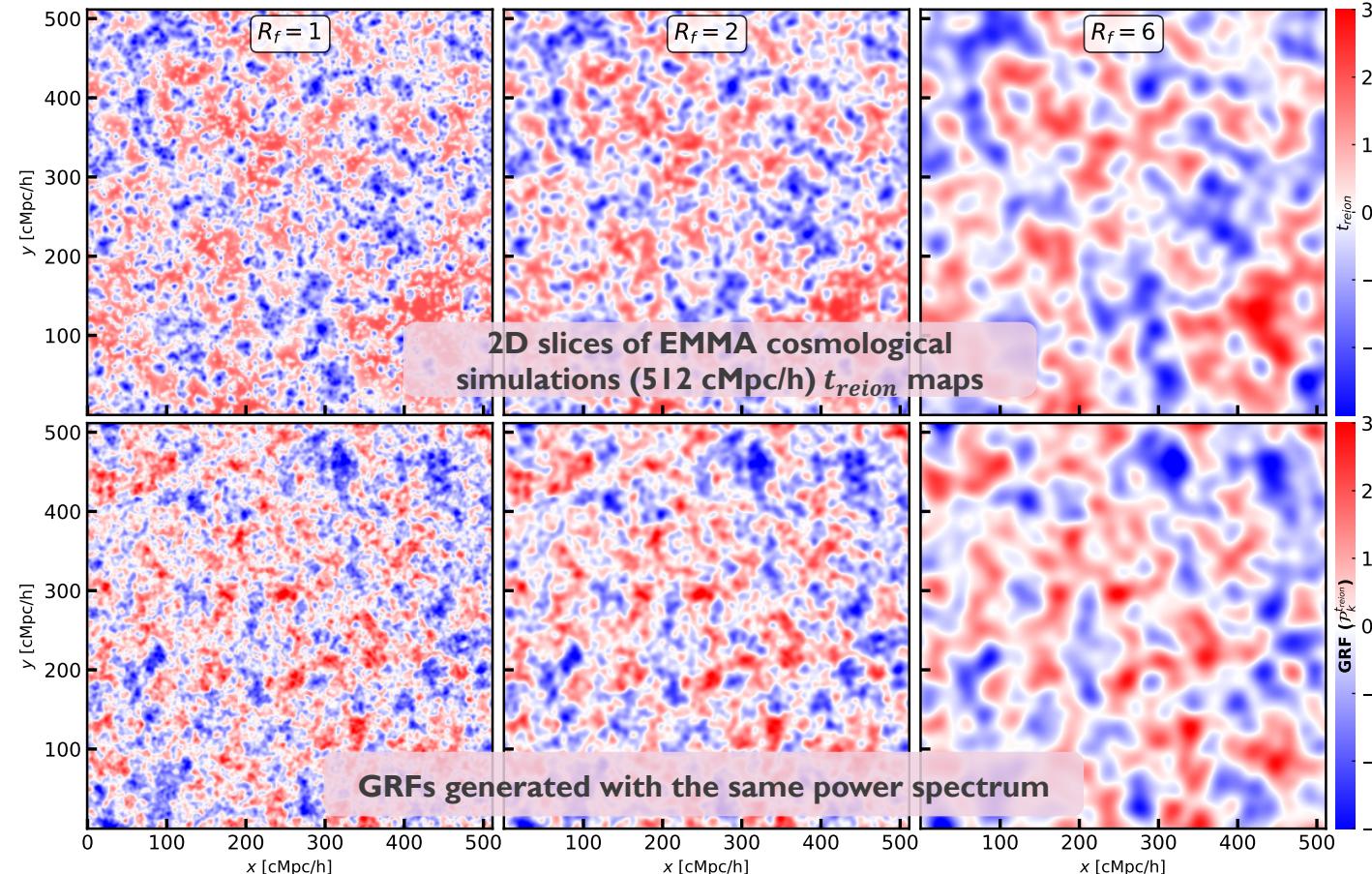
# C.TOPOLOGY AND GRF THEORY

Characterisation of the evolving topology of the EoR

**How to characterise the evolution of reionisation?**

By using topological statistics that are...

- ... measurable in reionisation times maps,
- ... predictable with GRF theory,
- ... and entirely defined with the power spectrum of the gaussian field



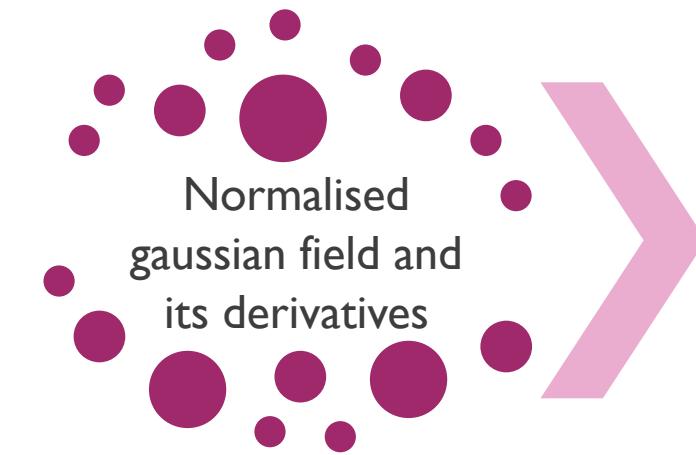
Thélie+ (submitted to A&A)

**6 data sets**

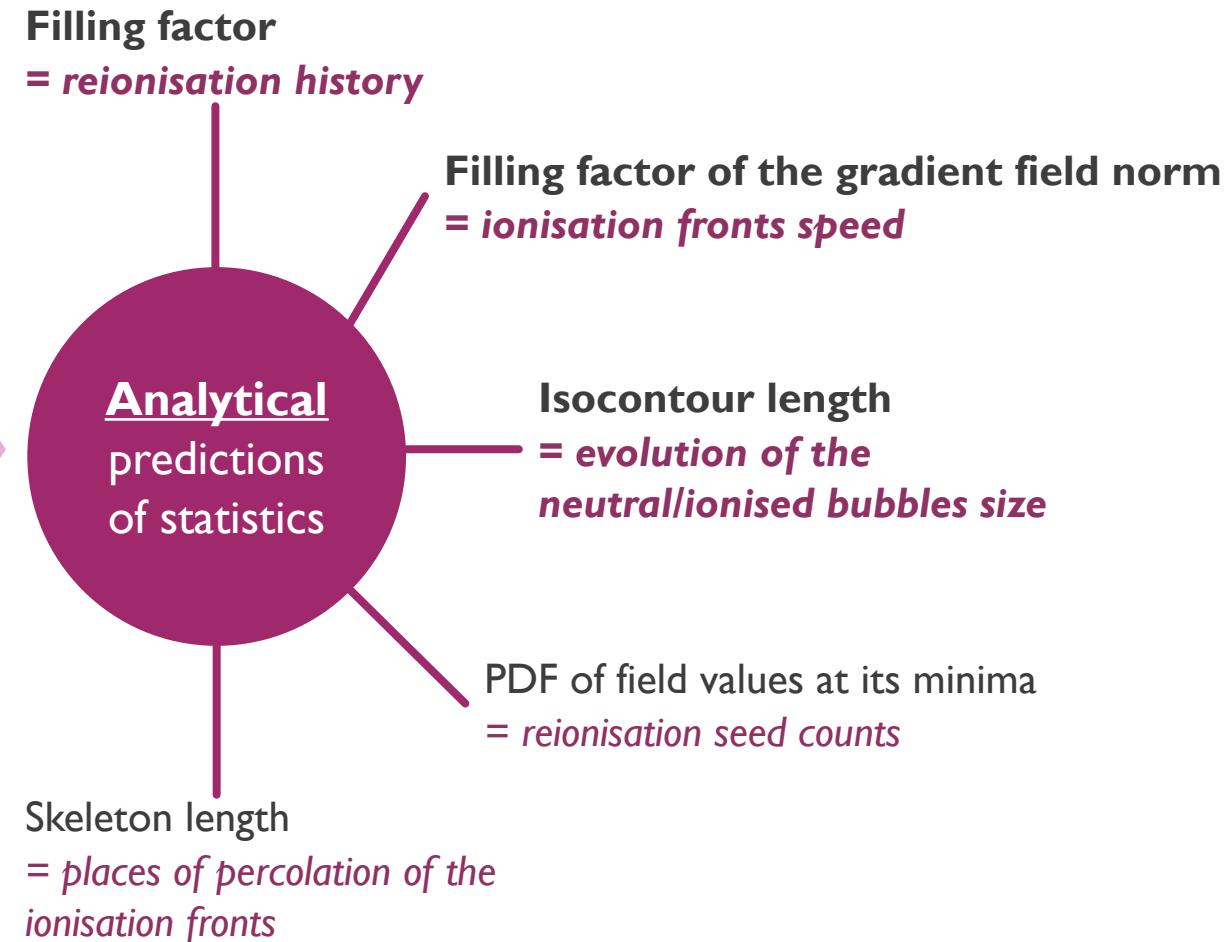
- 100 EMMA maps of  $t_{reion}(x, y)$
- 100 GRFs
- 3 different smoothings

# C.TOPOLOGY AND GRF THEORY

Gaussian random fields theory (Rice+44, Longuet-Higgins+57, Bardeen+86, Gay+12)



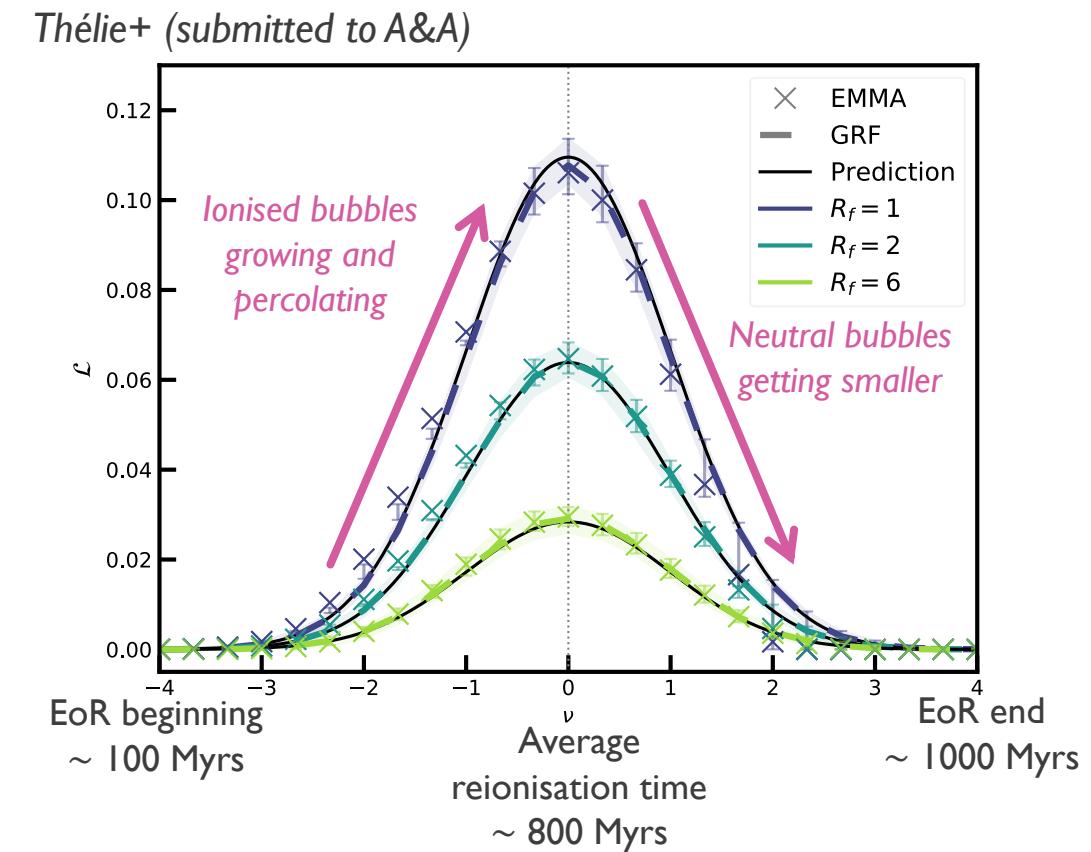
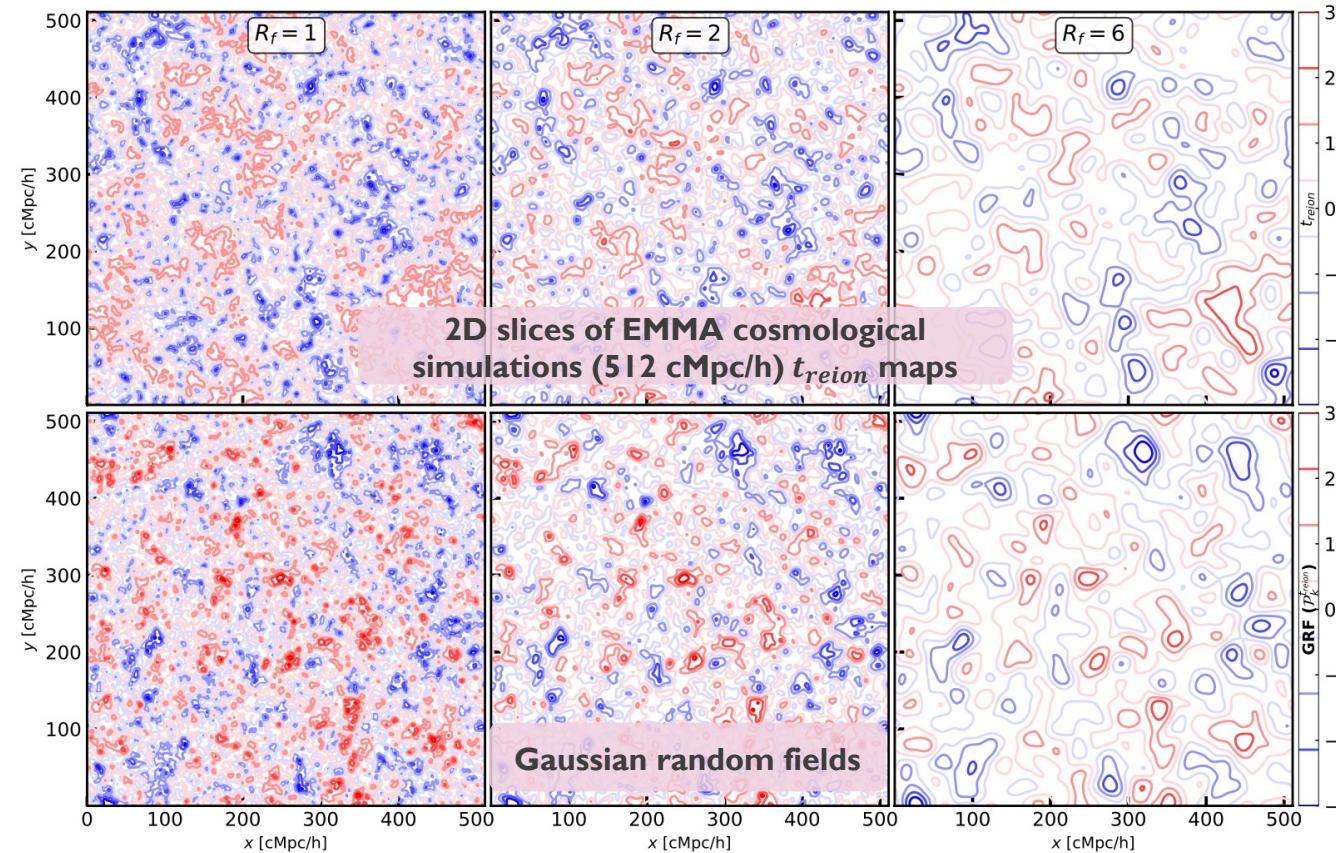
Probability distribution functions  
(joint PDFs of the field/its derivatives)



# C.TOPOLOGY AND GRF THEORY

Simulation measurements & GRFs predictions

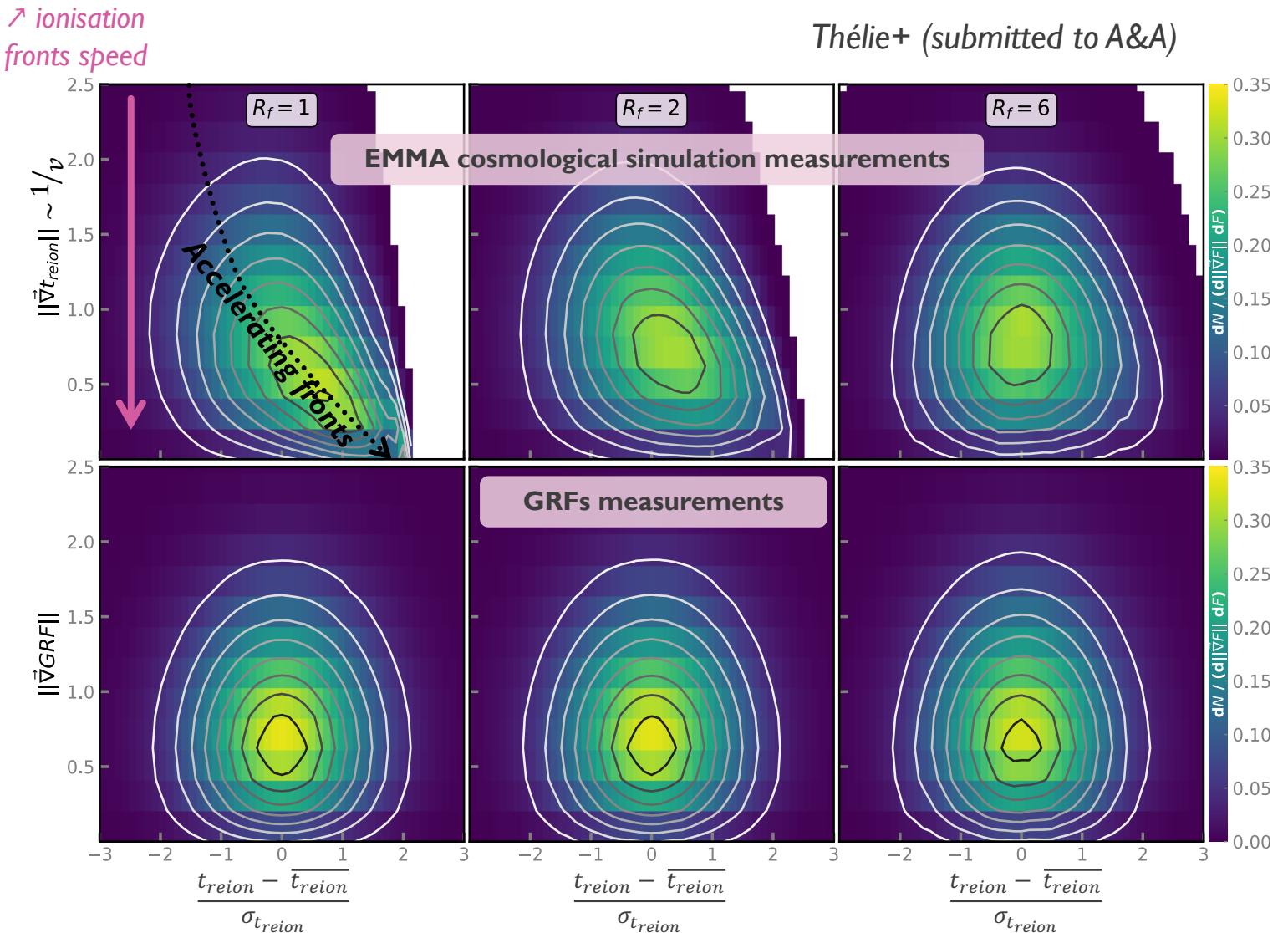
- Isocontours = regions reached by ionisation fronts at the same time  
→ follow the size evolution of ionised/neutral bubbles
- EMMA measurements close to gaussian predictions



# C.TOPOLOGY AND GRF THEORY

Simulation measurements & GRFs predictions

- EMMA measurements close to gaussian predictions
- When  $R_f \nearrow$  : EMMA measurements more symmetric
- $R_f \in \{1, 2\}$  : asymmetry (non-gaussianity)
  - Slow reionisation before it accelerates
  - Acceleration of ionisation fronts at the end of the EoR



# CONCLUSIONS & PERSPECTIVES

## How does the EoR happen?

Topological studies to analyse:

- *Growth of structures*
- *Ionised/neutral bubbles geometry, distribution, organisation*
- *Percolation, evolution of the process*

## Reionisation times

- Geometrical characterisation of different EoR model
- Comparisons of cosmological and semi-analytical simulations

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### A. 21 CM → $t_{reion}$

- Good reconstruction of reionisation times map, even if they are a little bit smoothed (with and without noise)
- Best reconstructions with observed redshifts  $8 < z < 12$

### B. Reionisation patches

- Beaded sources along the matter filaments
- Minority of butterflies (isolated sources or strong emitters)

### C. Reionisation times: topology and GRFs theory

- Diverse statistics on the evolution of the reionisation process...
- ... that are analytically computable  $t_{reion}$

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- Best reconstructions with observed redshifts  $8 < z < 12$



- Improving the neural network to reconstruct well the small scales
- Topological analyses of the CNN reconstructions  $t_{reion}$  maps

### B. Reionisation patches

- Beaded sources along the matter filaments
- Minority of butterflies (isolated sources or strong emitters)

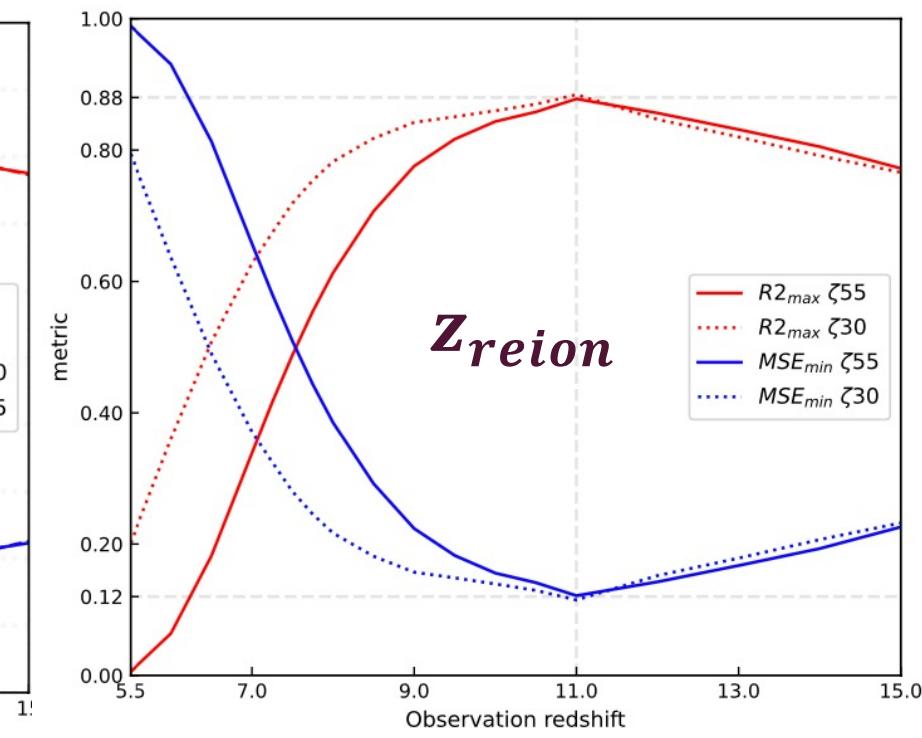
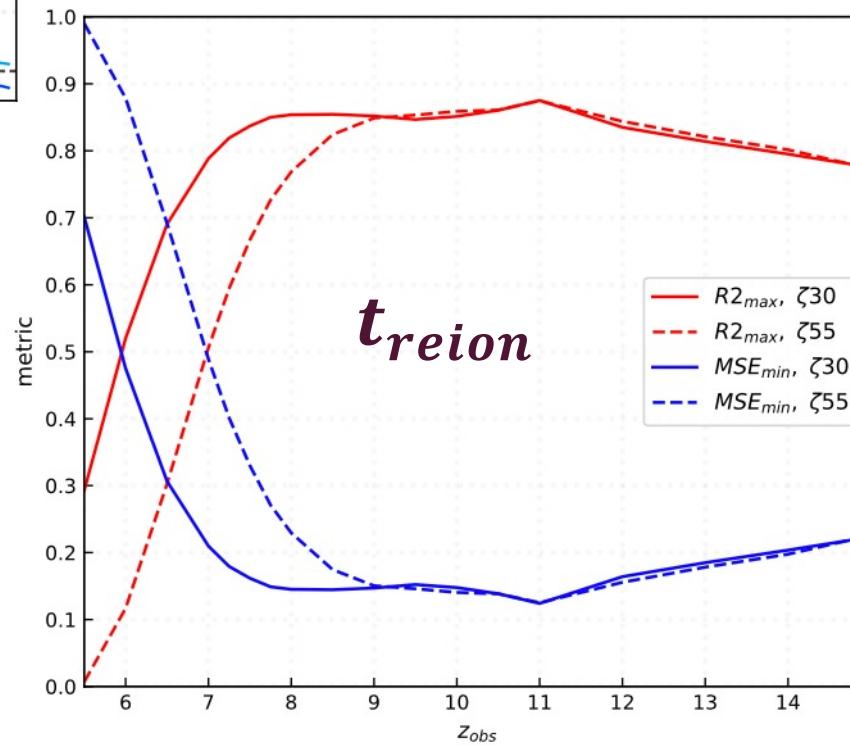
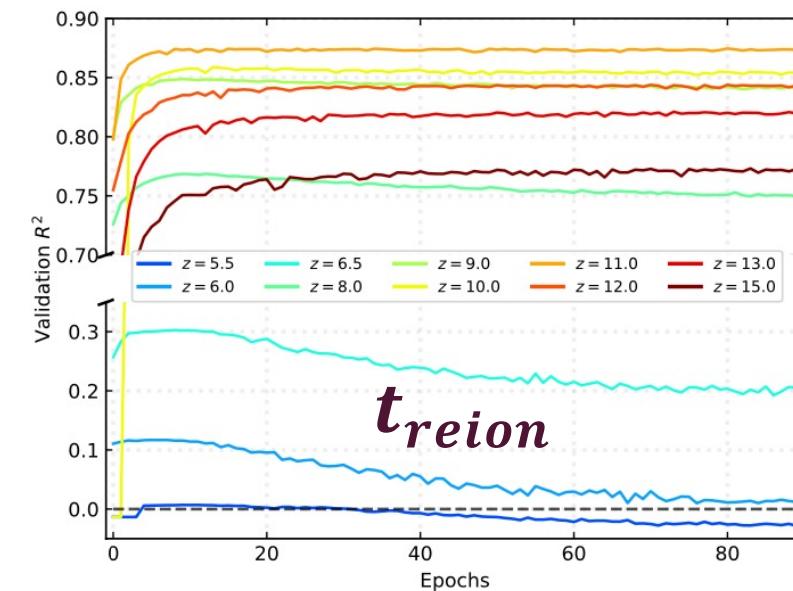


- Take into account the asymmetries in the GRFs predictions (e.g. with Gram-Charlier expansion)
- Same study but with larger or more resolved simulations (e.g. with Dyablo)
- Inference of the power spectrum parameters from topological measurements

Thank you for your attention!

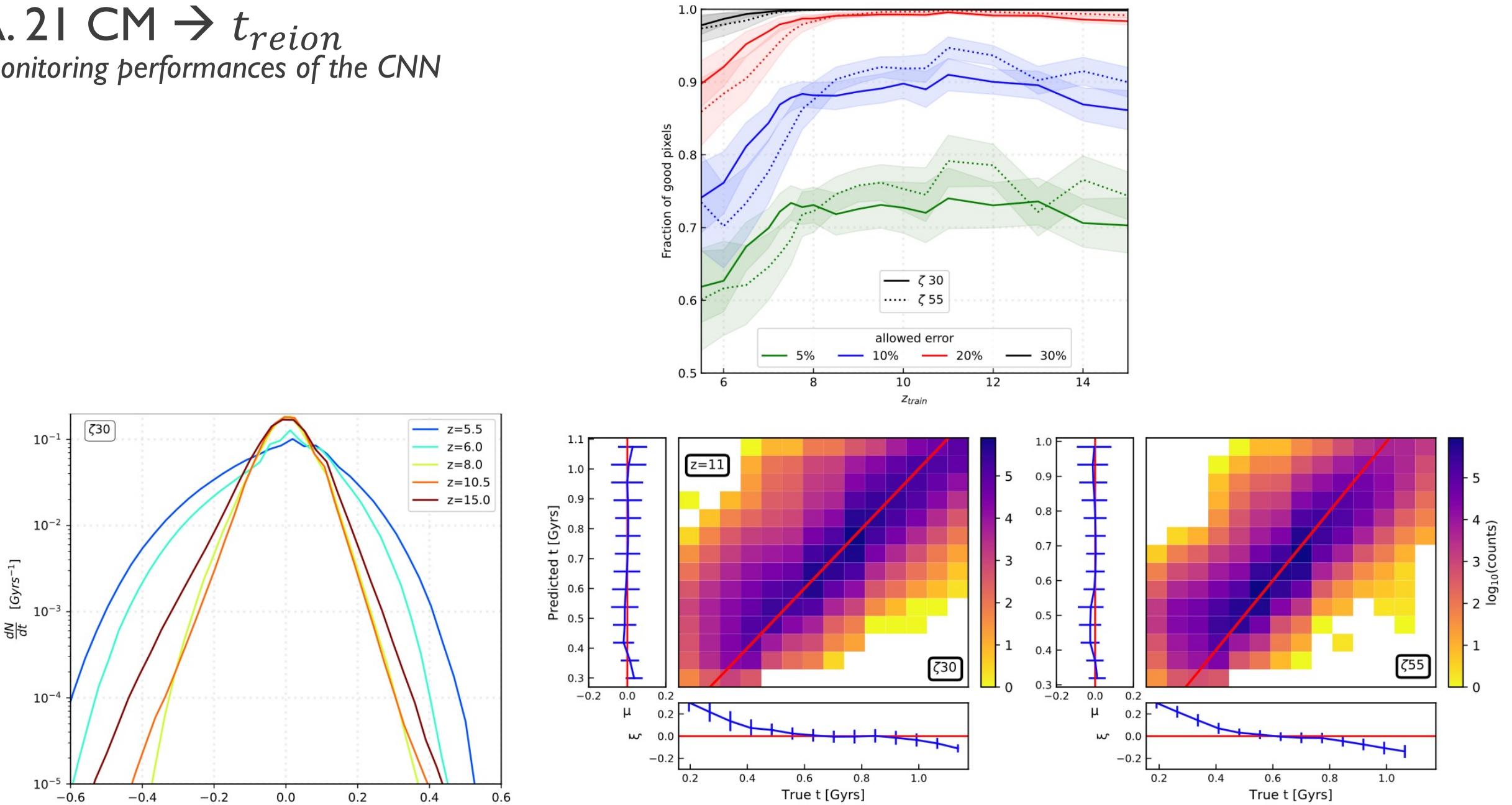
# A.2 | CM $\rightarrow t_{reion}$

## Monitoring performances of the CNN



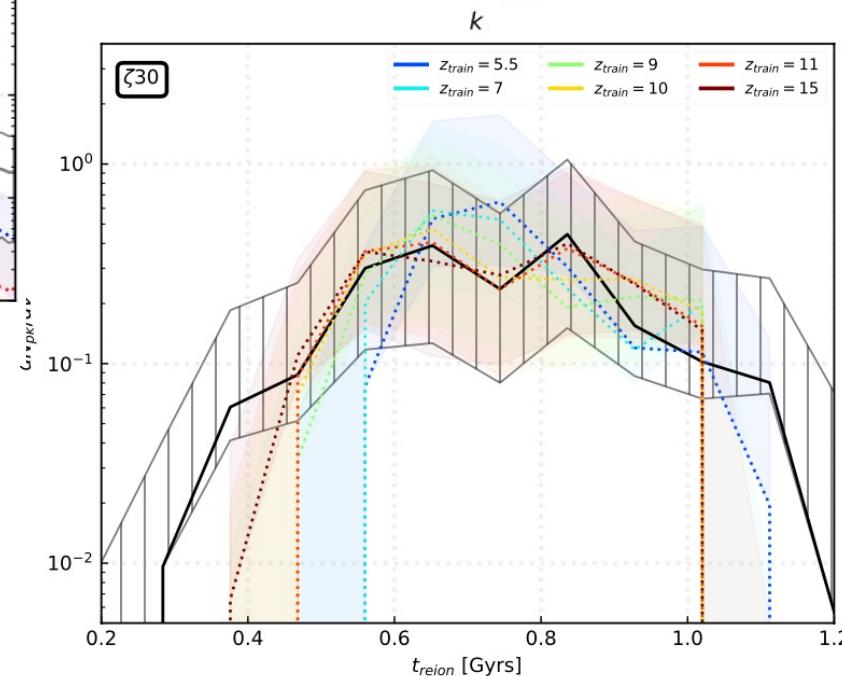
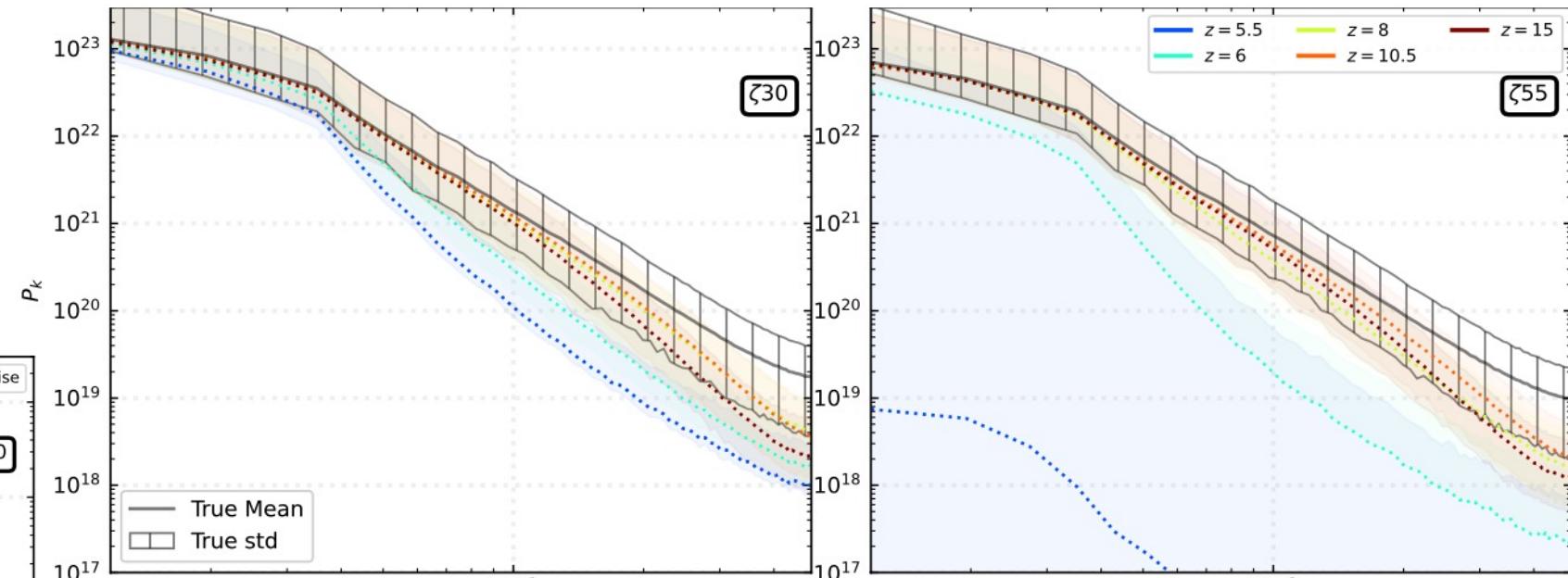
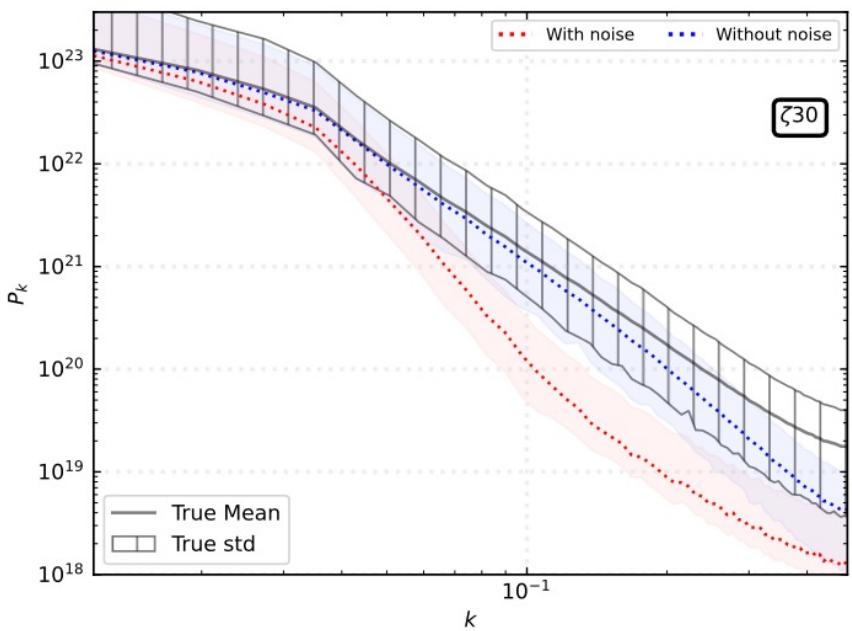
# A.2I CM $\rightarrow t_{reion}$

## Monitoring performances of the CNN



# A.2 | CM $\rightarrow t_{reion}$

## Monitoring performances of the CNN



# B. PATCHS DE REIONISATION

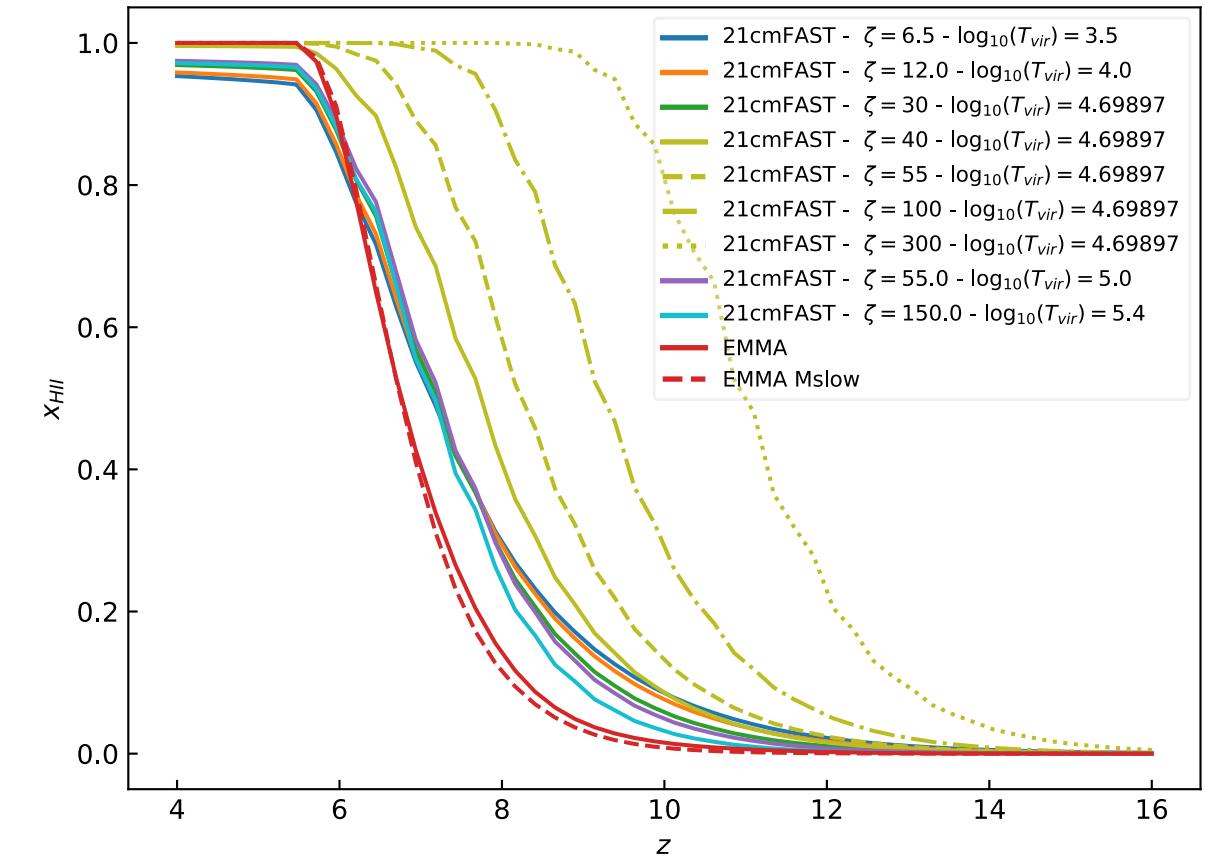
*Simulations*

**21cmFAST semi-analytical simulations ( $128^3$  cellules -  $128^3 \text{ cMpc}^3/\text{h}^3$  ; Mesinger+11) :**

- $\zeta$  : galaxies ionising efficiency
- $T_{vir} \sim M_{min}^3$  : minimal virial temperature so that a halo start to form stars

**EMMA cosmological simulations ( $512^3$  cellules -  $512^3 \text{ cMpc}^3/\text{h}^3$  ; Aubert+15, Gillet+21) :**

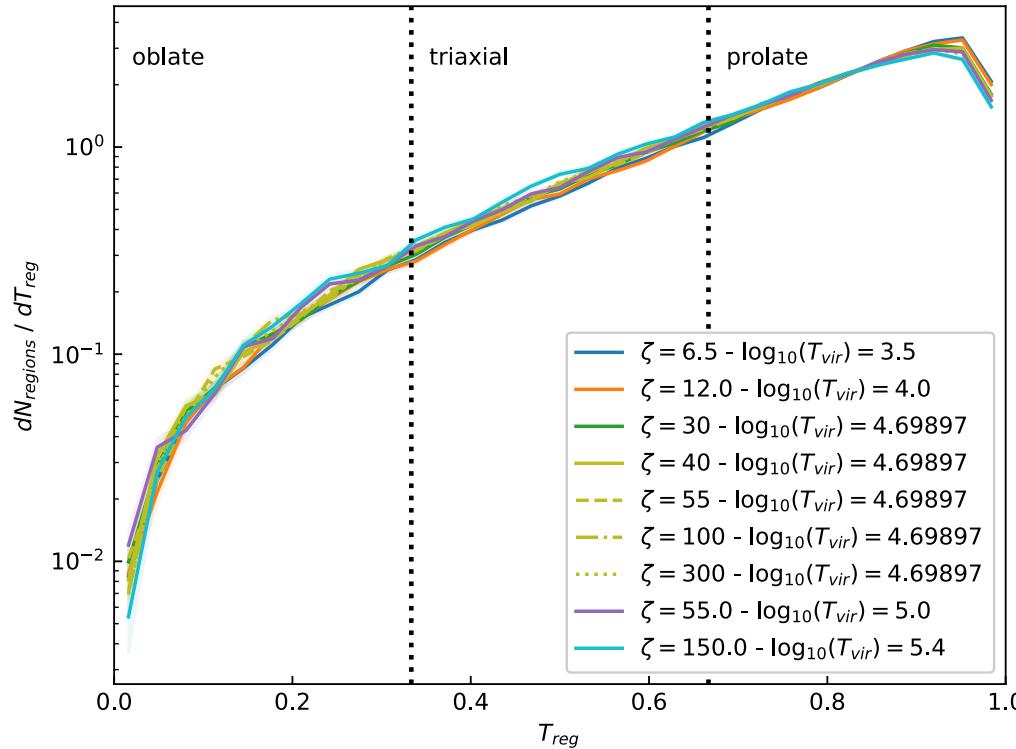
- Mass resolution for the stellar particle ( $10^7 M_\odot$  for the Mslow one and  $10^8 M_\odot$  for the other)



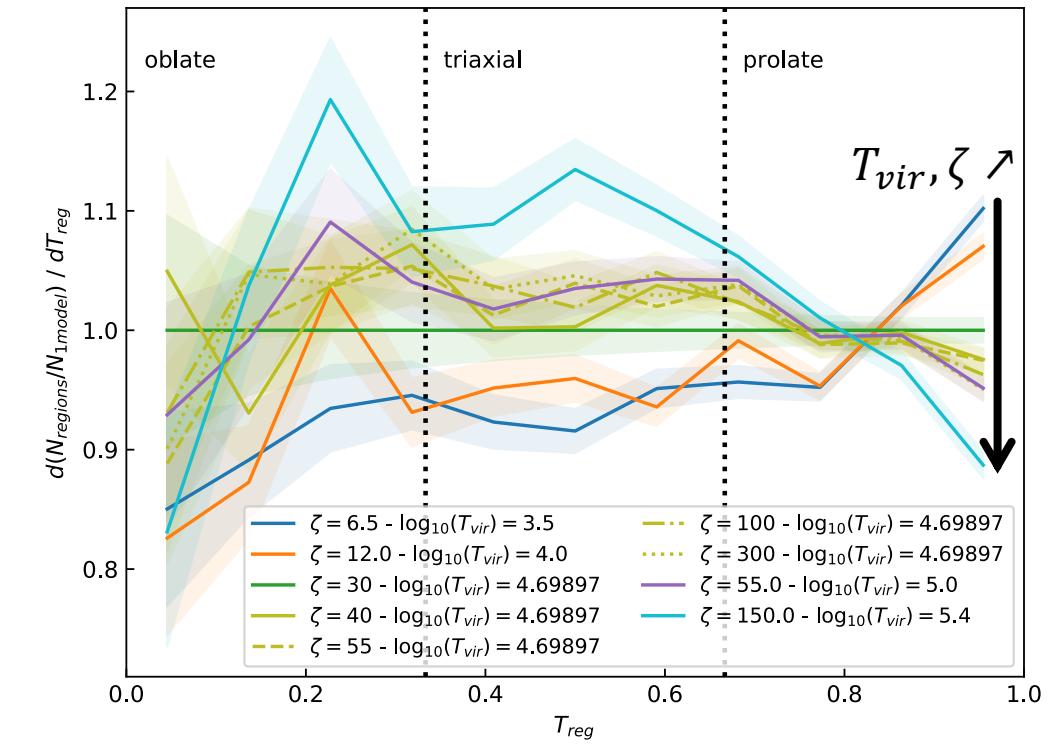
# B. REIONISATION PATCHES

## Patches shape

- Triaxiality parameter:  $T = \frac{\lambda_3^2 - \lambda_2^2}{\lambda_3^2 - \lambda_1^2}$
- Majority of prolate patches
- Less prolate patches for halos that are stronger emitter and more massive

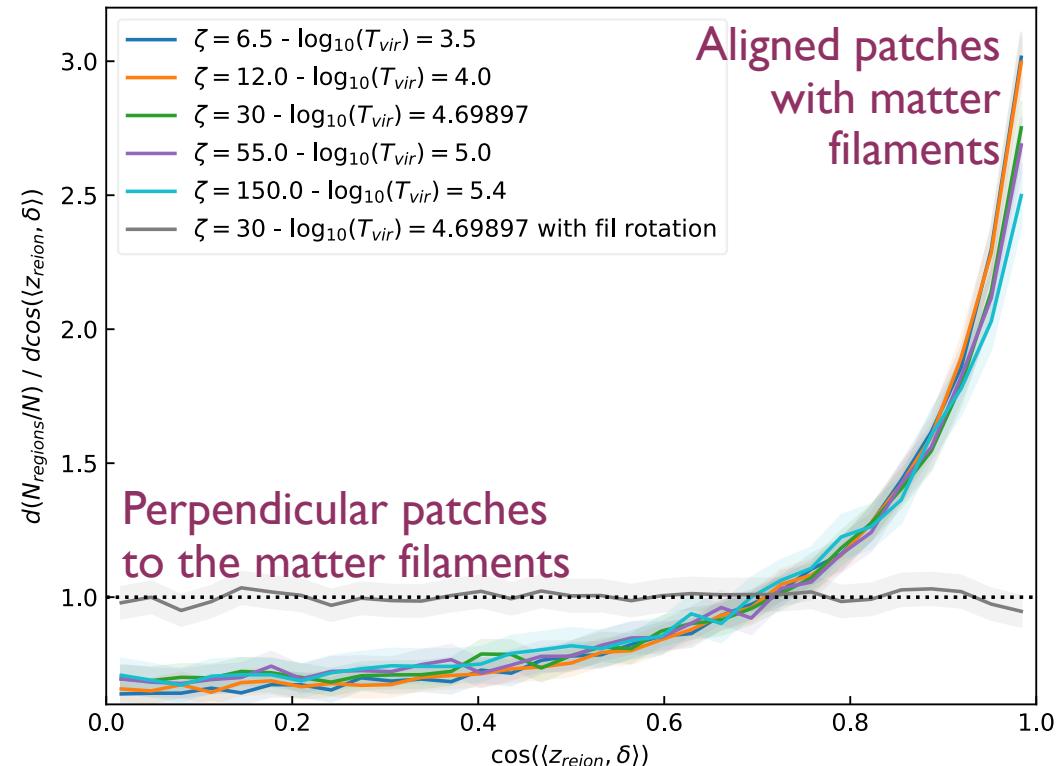


Ratio with model  
 $\zeta = 30$

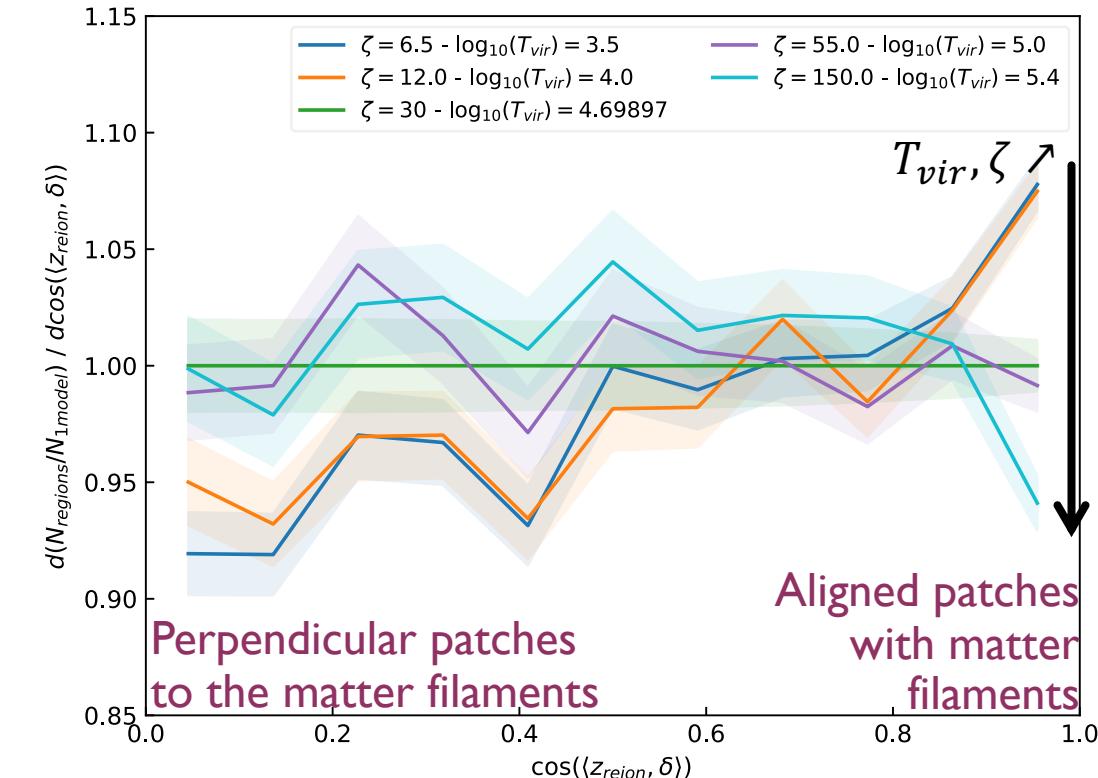


# B. REIONISATION PATCHES

Orientation of patches with respect to the matter filaments



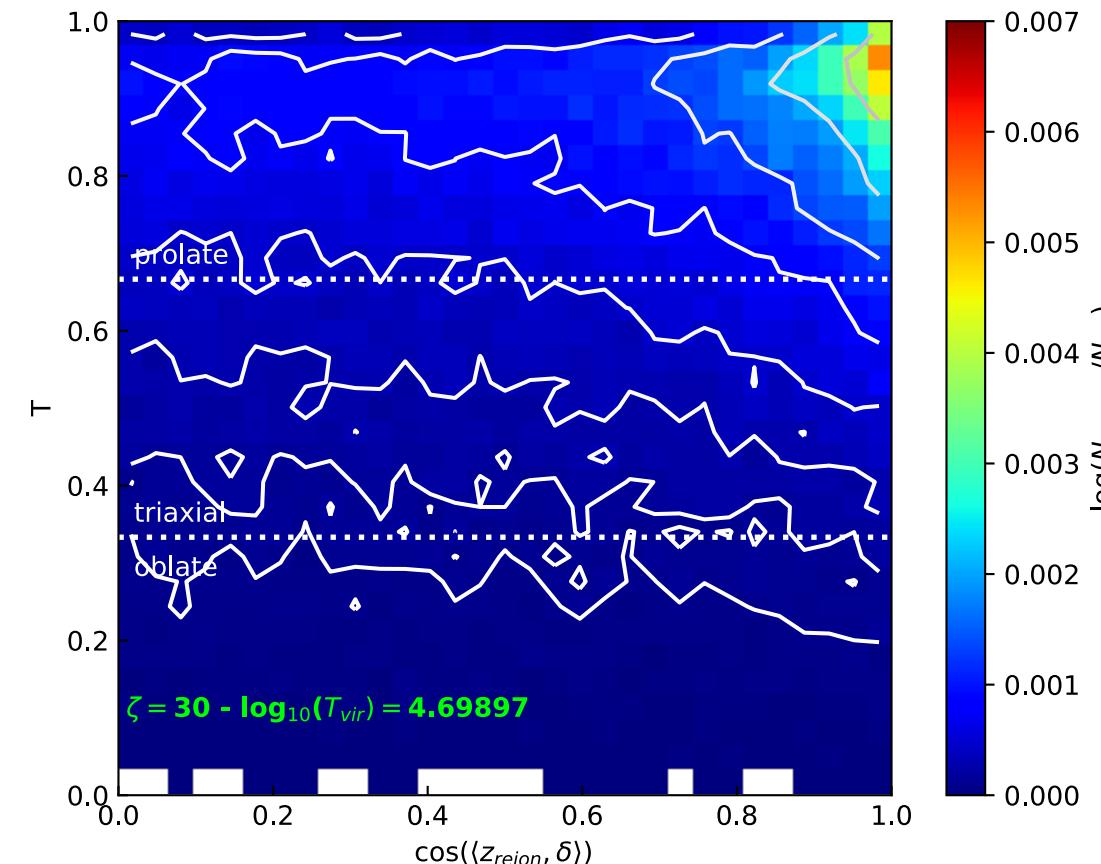
Ratio with model  
 $\zeta = 30$



- Majority of aligned patches to the matter filaments
- Less aligned patches for halos that are stronger emitter and more massive

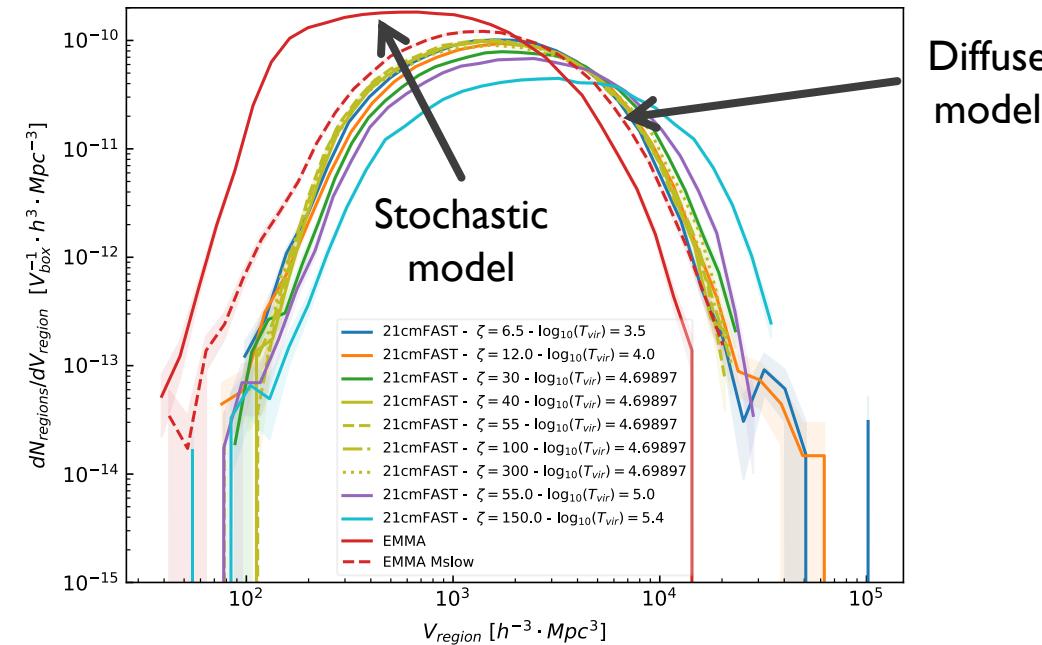
# B. REIONISATION PATCHES

*Shape vs. orientation of patches with respect to the matter filaments*

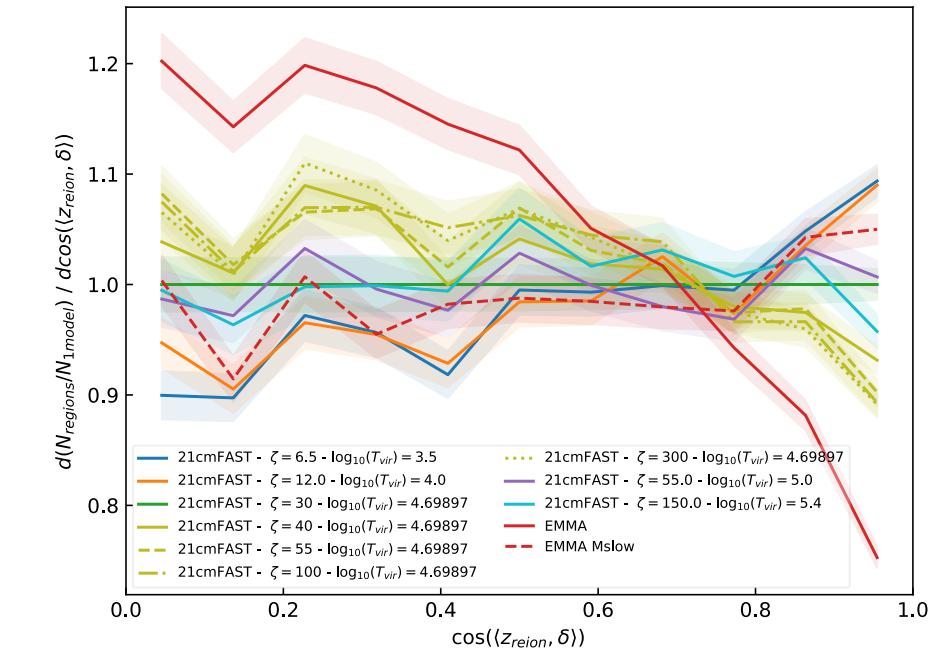


# B. PATCHS DE REIONISATION

Comparaison avec EMMA, une simulation cosmologique



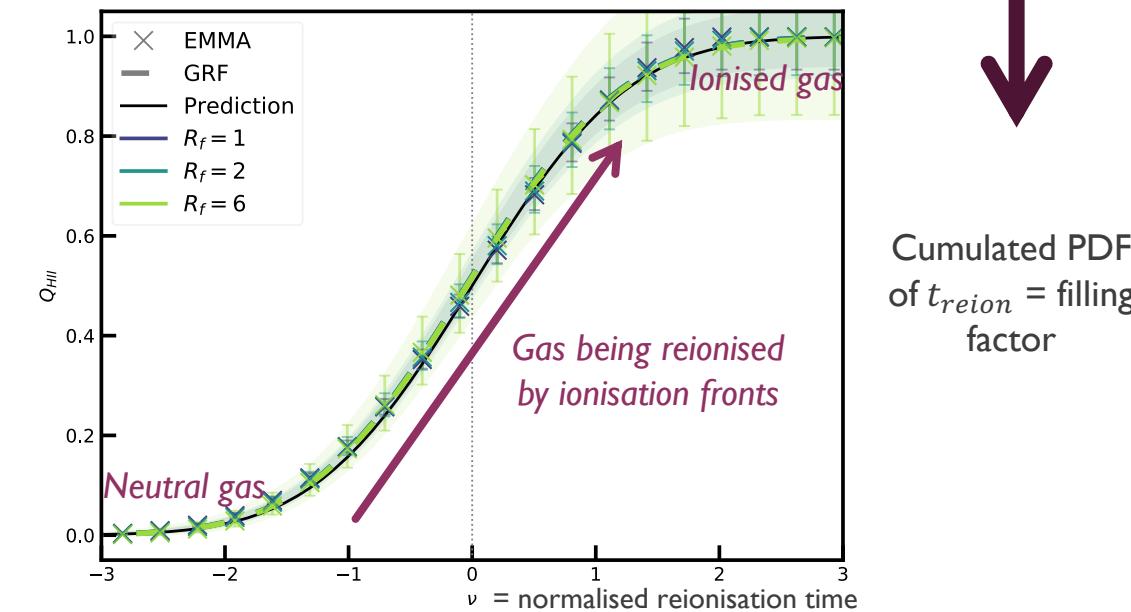
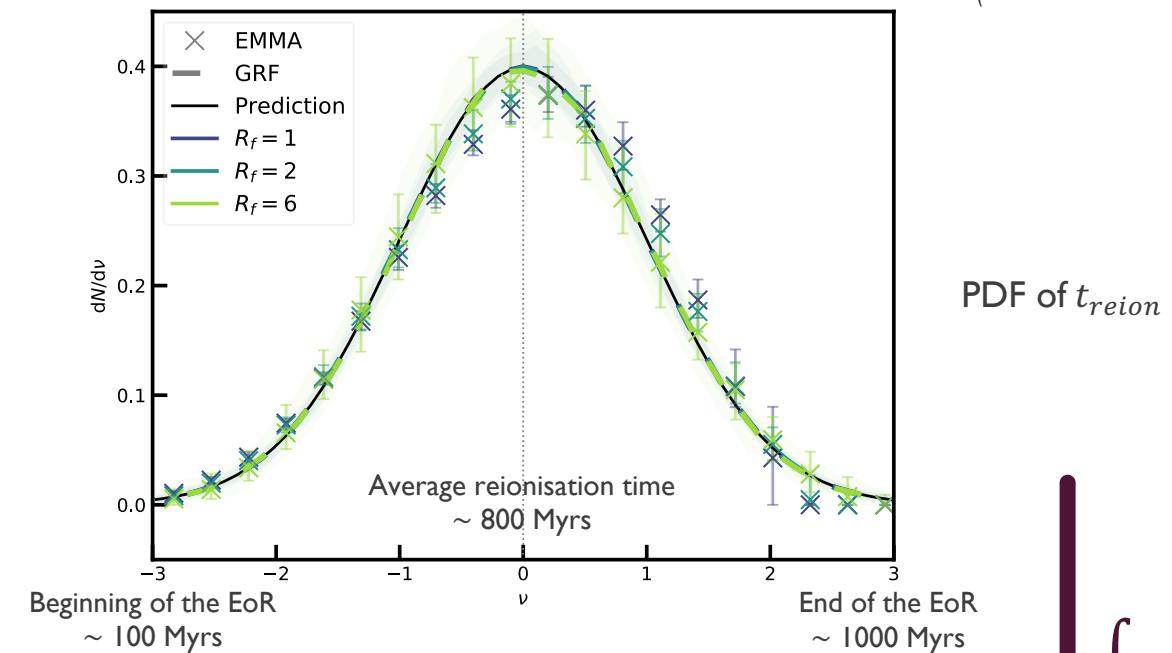
- Same conclusions for both type of simulations
- BUT EMMA can also produce models with rather different topologies (for the same  $x_{HII}(z)$ )



# C.TOPOLOGY AND GRF THEORY

Filling factor: *PDF et reionisation history*

- $R_f \nearrow$  : EMMA measurements more symmetric
- $R_f \in \{1, 2\}$  : imprints of non-gaussianity in the form of an asymmetry  
= slow reionisation before it is accelerated



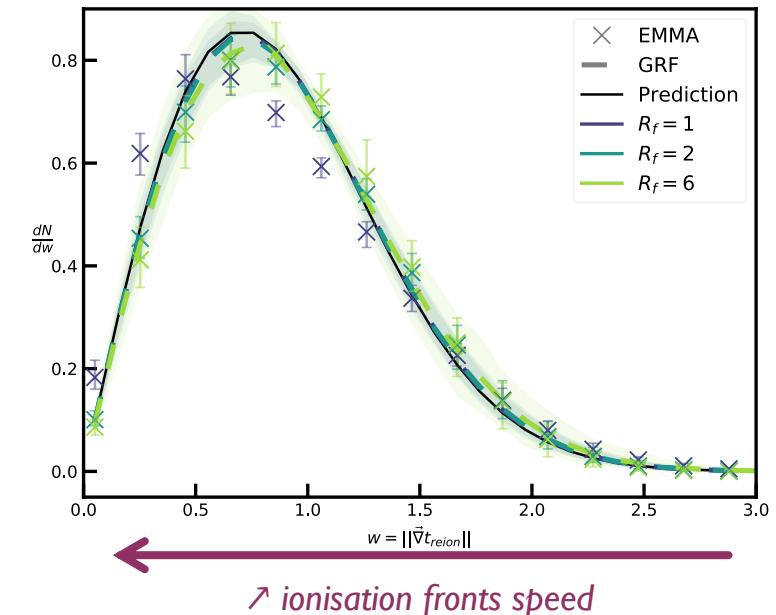
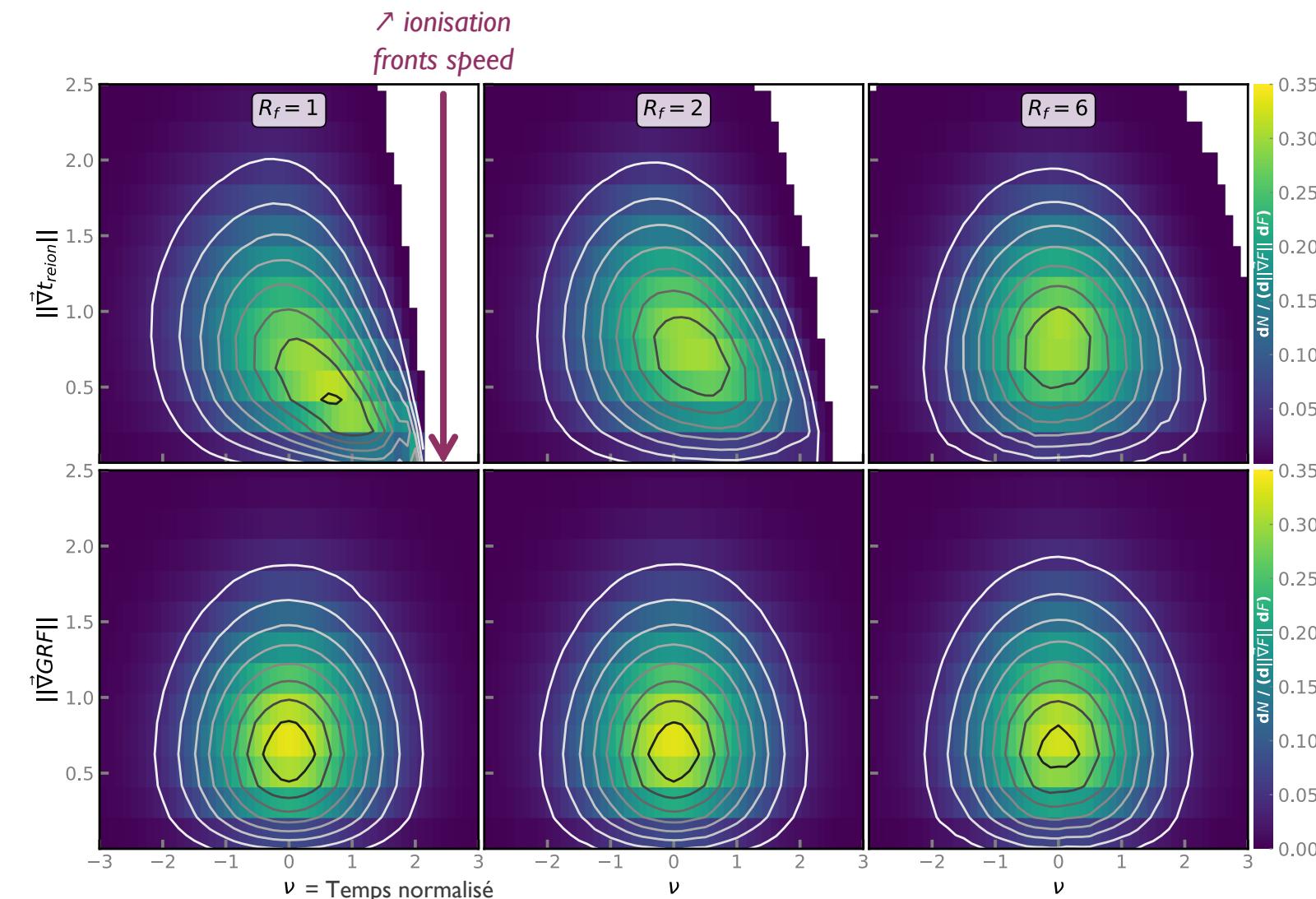
PDF of  $t_{reion}$



Cumulated PDF  
of  $t_{reion}$  = filling  
factor

# C.TOPOLOGY AND GRF THEORY

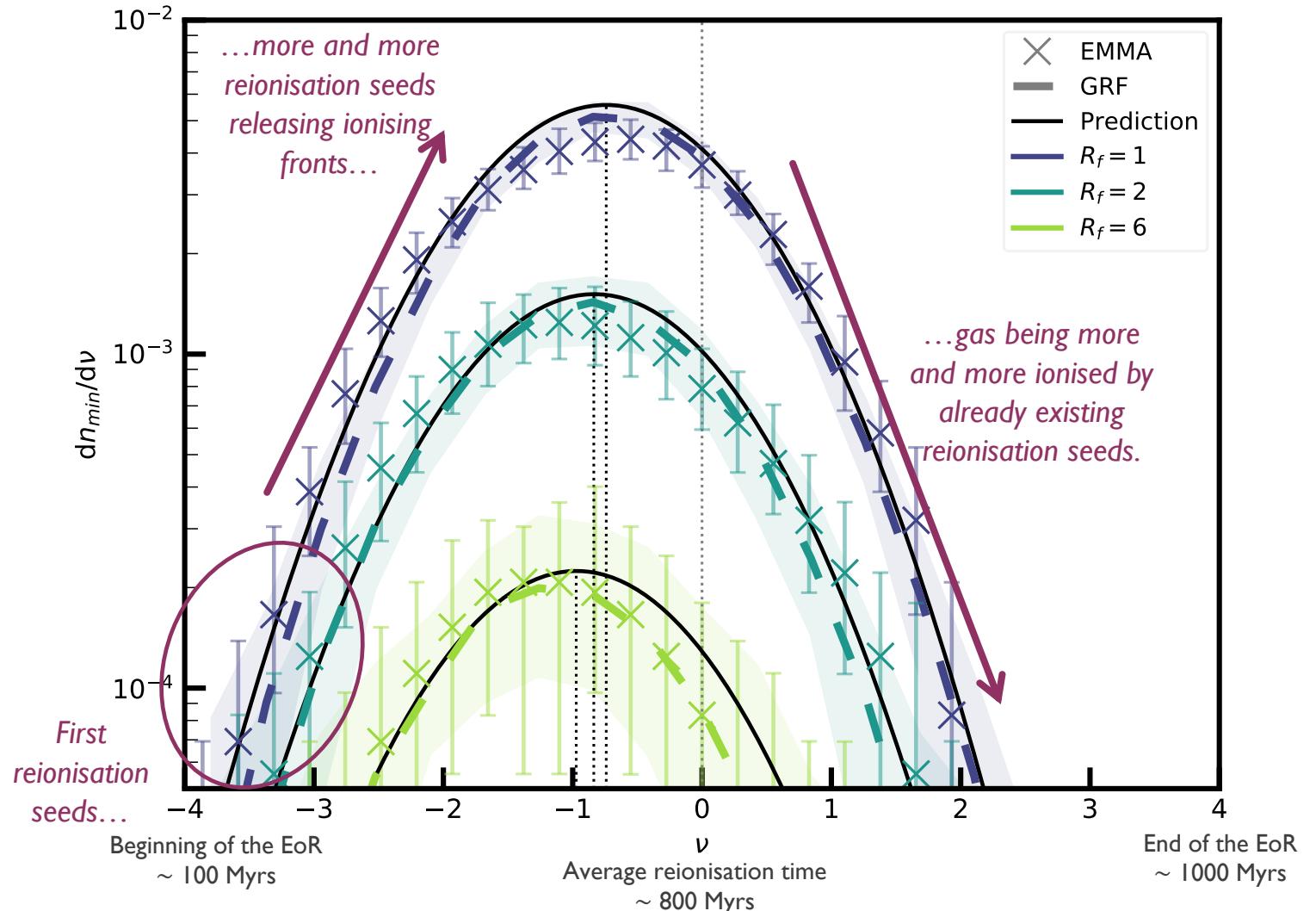
PDF of the gradients field norm: ionisation fronts speed



- $R_f \nearrow$  : EMMA measurements more symmetric
- $R_f \in \{1, 2\}$  : imprints of non-gaussianity in the form of an asymmetry  
= acceleration of the ionisation fronts at the end of the EoR

# C.TOPOLOGY AND GRF THEORY

PDF of the field value at its minima: *reionisation seed counts*

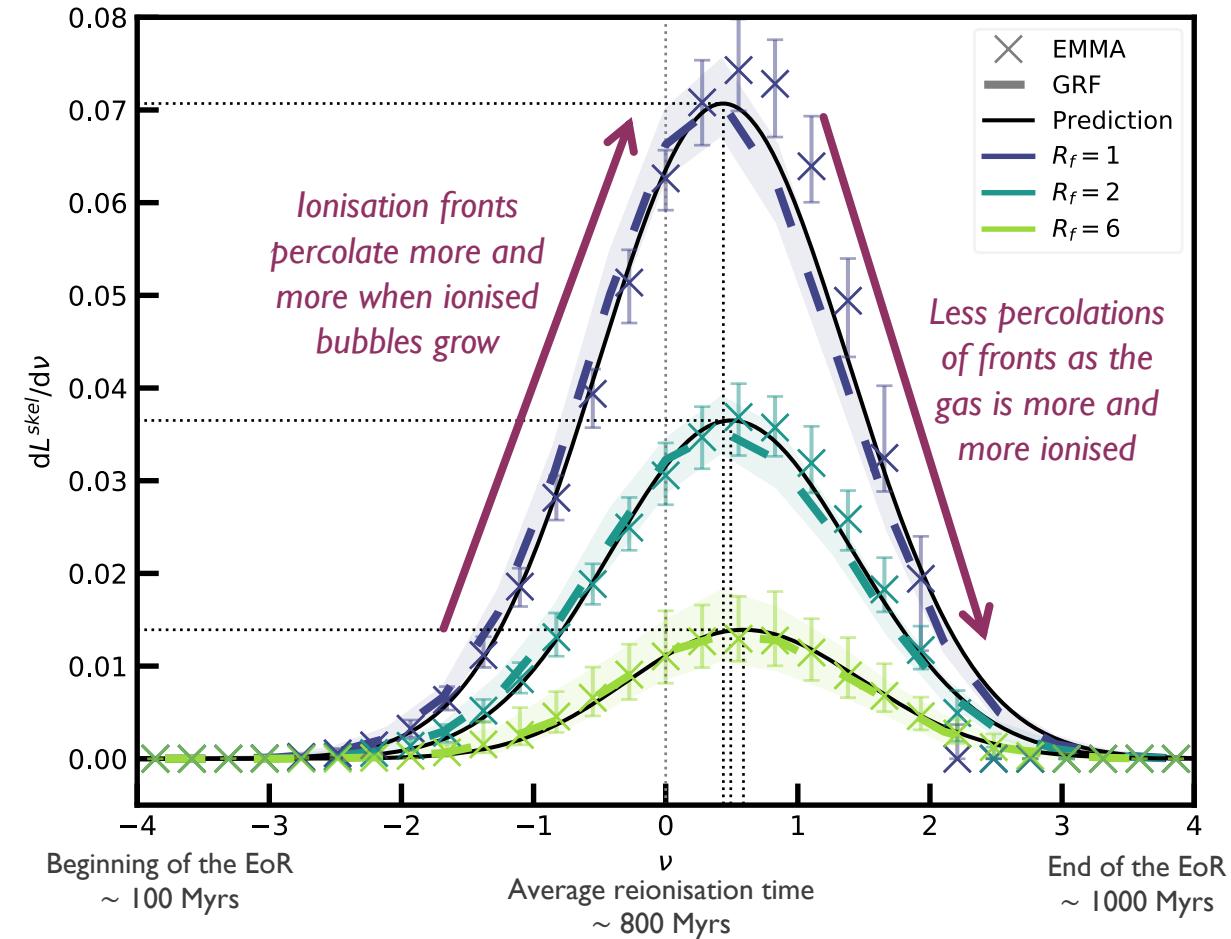


# C.TOPOLOGY AND GRF THEORY

*Skeleton length: places where the ionisation fronts percolate*

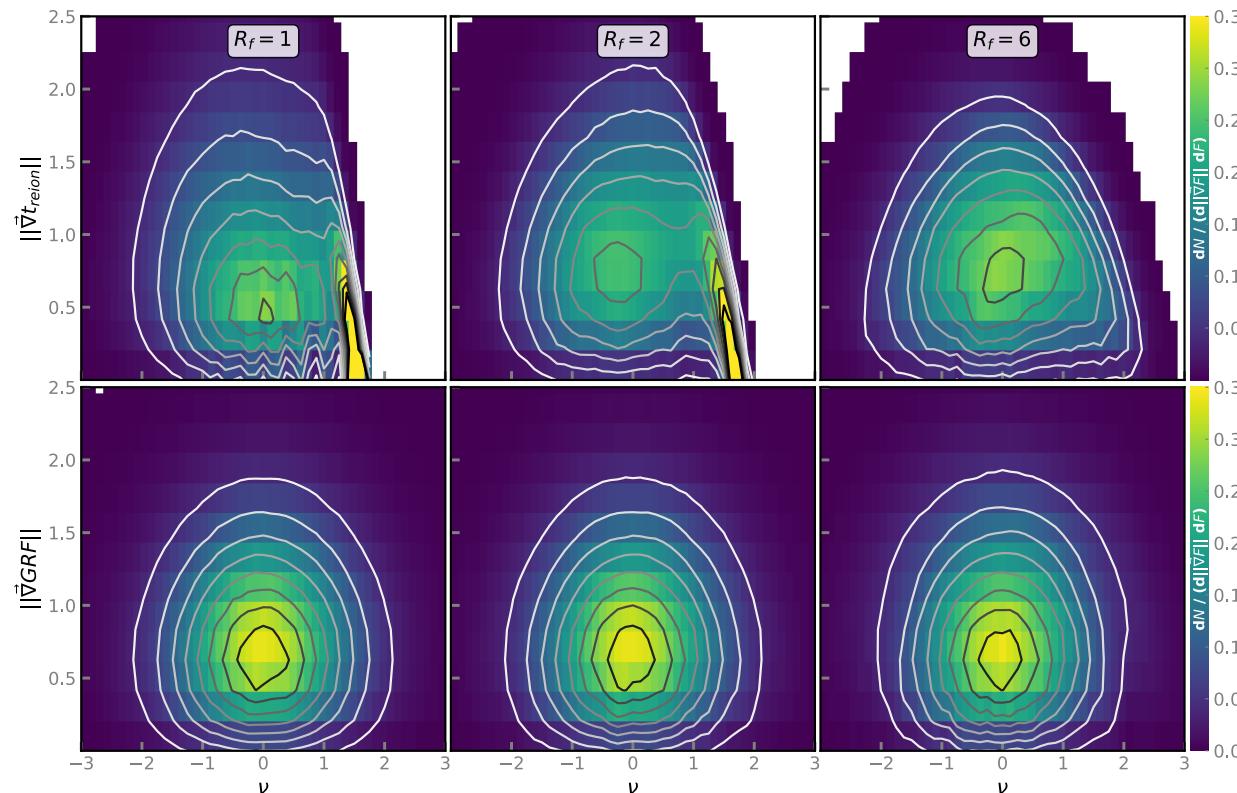
$$L^{\text{tot}} = \left( \frac{1}{8} + \frac{\sqrt{2}}{4\pi} \right) \frac{1}{R_*}$$

- "Stiff" approximation + global (measurements) or local (GRFs) calculations → predictions underestimating the skeleton length: **measurements have to be renormalised**
- $R_f \nearrow$ : EMMA measurements more symmetric
- $R_f \in \{1, 2\}$ : asymmetry  
= *acceleration of ionisation fronts at the end of the EoR*



# C.TOPOLOGY AND GRF THEORY

Comparisons with 21cmFAST



- Same behaviour as the EMMA measurements globally
- $R_f \nearrow$  : 21cmFAST measurements more symmetric
- $R_f \in \{1, 2, 6\}$  : asymmetry because of the absence of modelisation of radiation propagation within 21cmFAST

