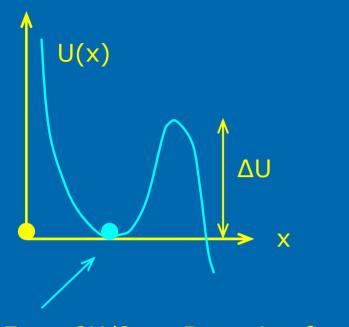
Computation of the mean reversal rate of geodynamo models

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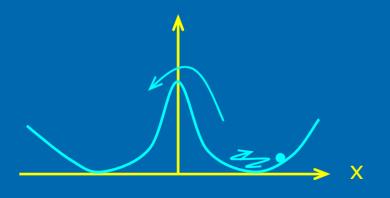


History Kramers' escape problem (1940)

Dissociation of molecule •–• in two parts

 $T_{esc} \sim T_{osc} \exp (\Delta U/kT)$ 





Mean reversal rate problem: bistable, and no thermal kicks

[Hoyng et al. (GAFD 94 2001 (reversals of mean field)

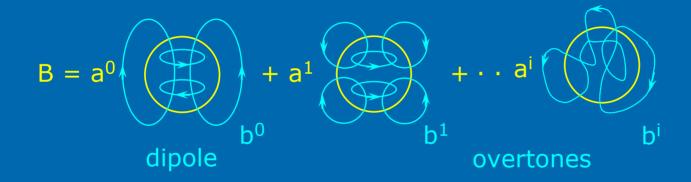
Hoyng & Duistermaat (HD model) (EPL 68 2004; dipole + 2 overtones)]



Mean reversal rate: setup

 $\partial B/\partial t = \nabla x [U \times B - \eta \nabla x B]$ 

- U = v + u(t) v = mean flow, u(t) = convection (given)
- $B = \Sigma_k a^k(t) b^k(r)$  B is decomposed in set of base functions  $b^k$



$$\partial_t a^i = [R^{ik} + V^{ik}(t)] a^k$$

$$\mathsf{R}^{\mathsf{i}\mathsf{k}} = \int_V \, \hat{\mathsf{j}}^{\,\mathsf{i}} \cdot (\mathsf{v} - \mathsf{\eta} \nabla) \, \mathsf{x} \, \mathsf{b}^{\mathsf{k}} \, \, \mathsf{d}^3\mathsf{r}$$

$$V^{ik}(t) = -\int_{V} u(t) \cdot \hat{j}^{i} \times b^{k} d^{3}r$$

- choice of  $b^k$  and adjoints  $\hat{b}^k$ : eigenfunctions of R:  $R^{ik} = \lambda^k \delta^{ik}$
- physics in terms of interaction of global modes
- reversal = sign flip of  $a^{0}(t)$

## S RON

## HD model

nonlinearity V<sup>ik</sup>(t) uncorrelated & 1 complex [must be removed for by hand overtone geodynamo model] equal r.m.s. magnitude  $\partial_t x = (1 - x^2)x + V^{00}x + V^{01}y + V^{02}z$  $\partial_t y = -ay - cz + V^{10}x + V^{11}y + V^{12}z$ У  $\partial_t z = cy - az + V^{20}x + V^{21}y + V^{22}z$ x(t)x = -1separatrix x = 1500 1000 1500 2000 2500 3000 0 3D escape problem time

 $\dot{\mathbf{x}} = [\cdots] \times \rightarrow \partial_t \log \mathbf{x} = \cdots \rightarrow \text{no reversals}$   $\mathbf{x} \downarrow \text{slowly} \rightarrow \mathbf{y}, \mathbf{z} (:) \times \rightarrow \dot{\mathbf{x}} (:) \times \rightarrow \text{no reversals}$   $\mathbf{x} \downarrow \text{fast} \rightarrow \dot{\mathbf{x}} \approx V^{01}\mathbf{y} + V^{02}\mathbf{z} \rightarrow \text{reversal possible}$   $duration reversal \sim \text{decay time first overtone}$   $\mathbf{S}^{\text{ROM}}$ Santa Barbara 18 July 2008
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## Computation mean reversal time in HD model

P(x,y,z,t) = probability distribution

T(x,y,z) = mean time to reach separatrix starting in x,y,z

 $\dot{x} = \cdots$ ;  $\dot{y} = \cdots$ ;  $\dot{z} = \cdots$   $\rightarrow$  Fokker-Planck Eq. for P(x,r,t) :

$$\frac{\partial P}{\partial t} = \begin{bmatrix} -\frac{\partial}{\partial x} (1-x^2)x + \frac{a}{r} \frac{\partial}{\partial r} r^2 + \frac{1}{2} D(x^2+r^2) \left( \frac{\partial^2}{\partial x^2} + \frac{1}{r} \frac{\partial}{\partial r} r \frac{\partial}{\partial r} \right) \end{bmatrix} P$$

$$r = (y^2+z^2)^{1/2}; \quad D = \langle (V^{ij})^2 \rangle \tau_c$$

$$\partial_t P = MP \rightarrow M^{\dagger}T = -1$$

Mean time between reversals  $T_{rev} = 2T(1,0)$ 

Symmetries in the problem allow finding approximate value  $T_{rev}$ Geodynamo: many overtones, nonzero cross correlations between  $V^{ij}$ (symmetries are gone). Need alternative approach: reduce to 1D problem by projecting out the overtones

## S ROH

Generalisation to many overtones + correlations

$$\dot{a} = (1-a^2)a + V^{00}a + \Sigma_{k\geq 1}V^{0k}a^k \quad (a=a^0) \quad \text{variability dipole} \\ \dot{a}^i = \lambda^i a^i + V^{i0}a + \Sigma_{k\geq 1}V^{ik}a^k \quad \text{reversals} \\ \uparrow & \uparrow & \uparrow \\ \text{system attractioner mixing} \\ Take < V^{in}(t)V^{jm}(t-s) > = D \delta^{ij}\delta^{nm}\delta(s) \quad [\text{correlations: } D \to D^{ijnm}] \\ \partial_t P = MP \quad \to \text{ integrate over } a^1, a^2, ... \to \text{ equation for } p(a) \text{ only:} \\ \partial_t p = -\frac{\partial}{\partial a}(1-a^2)ap + \frac{1}{2}D\frac{\partial^2}{\partial a^2}(a^2 + \langle r^2 \rangle)p \qquad r^2 = \Sigma_{k\geq 1}(a^k)^2 \\ \partial_t p = Mp \quad \to M^{\dagger}T = -1 \text{ is ODE for } T: \\ \end{array}$$

$$(-a^{2})a \frac{dT}{da} + \frac{1}{2}D(a^{2} + \langle r^{2} \rangle) \frac{d^{2}T}{da^{2}} + 1 = 0 \qquad \text{find } \langle r^{2} \rangle \text{ near } a = 0 \text{ by}$$
  
ignoring mixing term

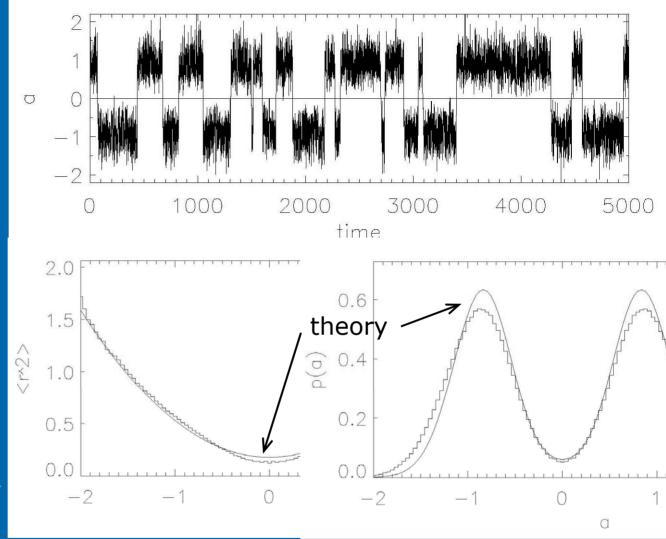
Numerical results

9 overtones no correlations between V<sup>ij</sup>(t)

 $T_{rev} = 146$  (theor), meas: 194 ± 27

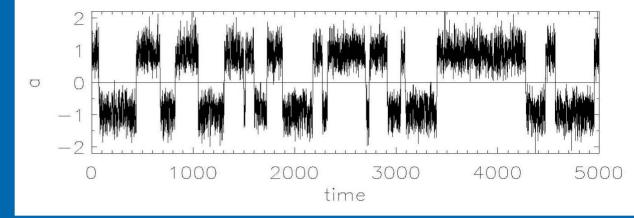
(usually agreement to within 10-40%)

 $T_{rev} \sim exp(K/D)$ K function of D & { $\lambda^k$ }





# What causes a reversal?



 $\dot{a} = (1-a^2)a + V^{00}a + \Sigma_{k\geq 1}V^{0k}a^k$  $\dot{a}^i = \lambda^i a^i + V^{i0}a$  (a = a<sup>0</sup>; mixing term ignored)

#### Three-step process:

1. jump in one or more  $V^{i0}(t)$  makes corresponding  $a^i$  temporarily large; 2. jump in  $V^{00}(t)$  makes the dipole amplitude a(t) collapse; 3.  $V^{0k}(t)a^k$  pushes a(t) to the other basin of attraction.

Step 1 and 2 need to be cotemporal within overtone decay time  $\rightarrow$  explains why reversals are rare, despite large fluctuations in a(t)

In a numerical model the jumps may be connected to properties of u(t) via

$$V^{ik}(t) = -\int_V u(t) \cdot \hat{j}^i x b^k d^3r$$

## Implementation

Collaboration with Dieter Schmitt, Martin Schrinner, Johannes Wicht MPS / Lindau

Numerical model

Velocity averages U = v + u(t)

Find eigenmodes  $b^{n}(r)$ , adjoints  $\hat{b}^{n}(r)$  & eigenvalues  $\lambda^{n}$  of R

Compute  $V^{ik}(t) = -\int_{V} u(t) \cdot \hat{j}^{i} \times b^{k} d^{3}r$ Compute  $\int_{0}^{\infty} \langle V^{ij}(t) V^{kl}(t-\tau) \rangle d\tau$ 

Compute D and  $\langle r^2 \rangle \rightarrow p(a)$  and  $T_{rev}$ 

### Discussion

Numerical tests of theory extended HD model encouraging But: quenching of fundamental mode and diffusion coefficients

