

Turbulence in quantum gases

scale invariance, synthetic dissipation and cascade fluxes

Nir Navon

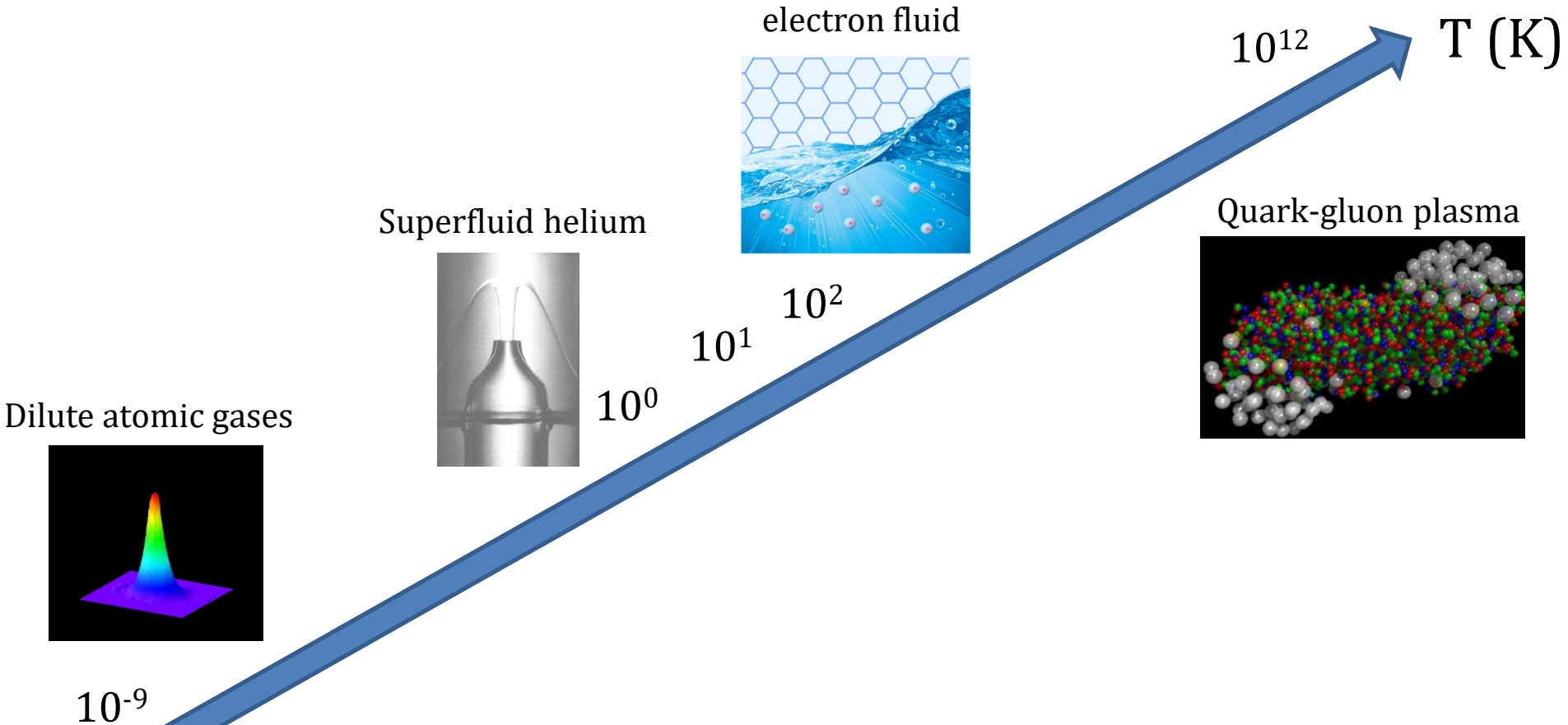
- Navon et al, *Nature* **539**, 72 (2016)
- Navon et al, *arXiv:1807.07564* (2018)
- Garratt et al, *arXiv:1810.08195* (2018)

KITP, 24th October 2018

Outline

- Turbulence in quantum gases
- The optical box trap: a new tool for ultracold atoms
- Low-energy excitations of a box-trapped BEC
- A turbulent steady state
- A synthetic dissipation scale and turbulent-cascade fluxes

Hydrodynamics of Quantum fluids

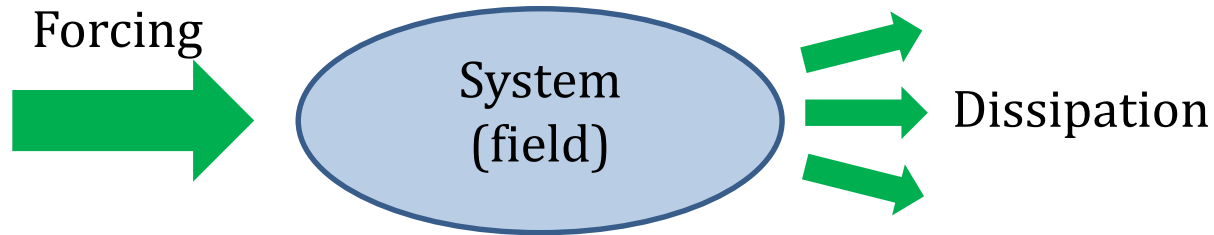


What is the fate of a periodically (and resonantly) driven quantum fluid?

- Existence and properties of far-from-equilibrium steady states for continuously driven-dissipative quantum fluid?
- Emergence of *universal* behavior?

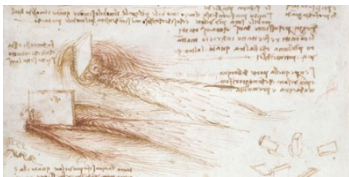
What is Turbulence ?

Quintessential phenomenon of out-of-equilibrium physics

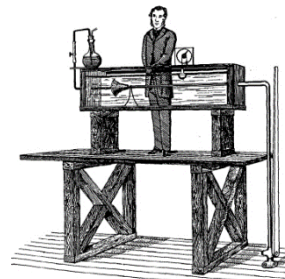


Far-from-equilibrium states that (usually) involves

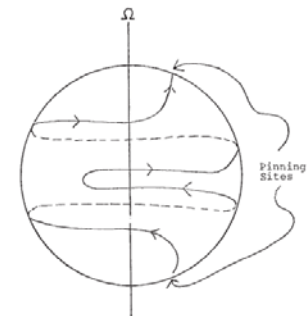
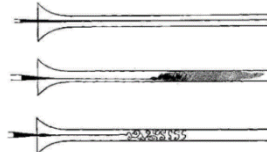
- drive + dissipation
- steady in a statistical sense
- local restoration of symmetries (isotropy/homogeneity)
- many interacting degrees of freedom
- chaotic properties



Leonardo da Vinci (1509)



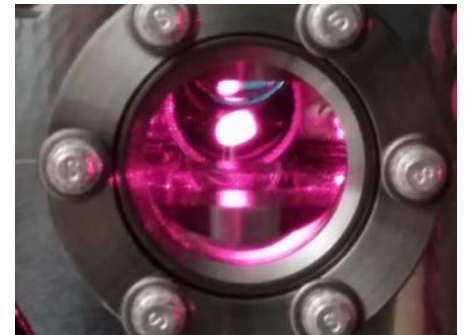
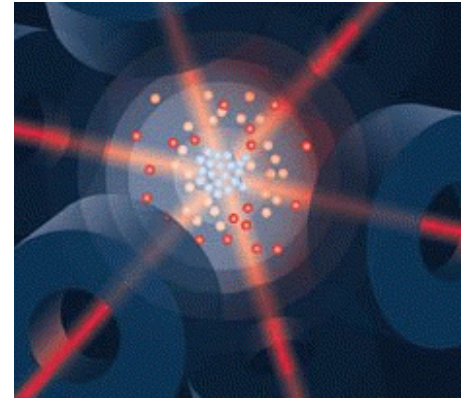
O. Reynolds
(1883)



SF turbulence neutron
stars / glitches? (1969)

Quantum gases as toy fluids

- Perfectly pure
- Resolvable from system size (typ. 100 μm) to interparticle spacing (typ. 1 μm)
- Easy to set out of equilibrium
- Real-time dynamics
- Highly controllable
 - (i) Interparticle interactions
 - (ii) Trapping geometry
 - (iii) Custom forcing
 - (iv) Tunable dissipation



→ Good candidates for simple turbulence experiments, « new knobs »

Described from first principles possible, e.g. GPE for Bose gas
$$i\hbar\frac{\partial\psi}{\partial t} = \left(-\frac{\hbar^2}{2m}\Delta + g|\psi|^2 \right) \psi$$

Experimental activity: V. Bagnato, B. P. Anderson, Y. Shin, P. Engels, K. Helmerson and T. Neely

Turbulent cascades

A central phenomenology of turbulence:
cascades of *conserved quantities, gradually*
transported across different length scales

→ Generic prediction: scale invariant fluxes
and power-law spectra

Famous example: “Kolmogorov 5/3” (hydro)

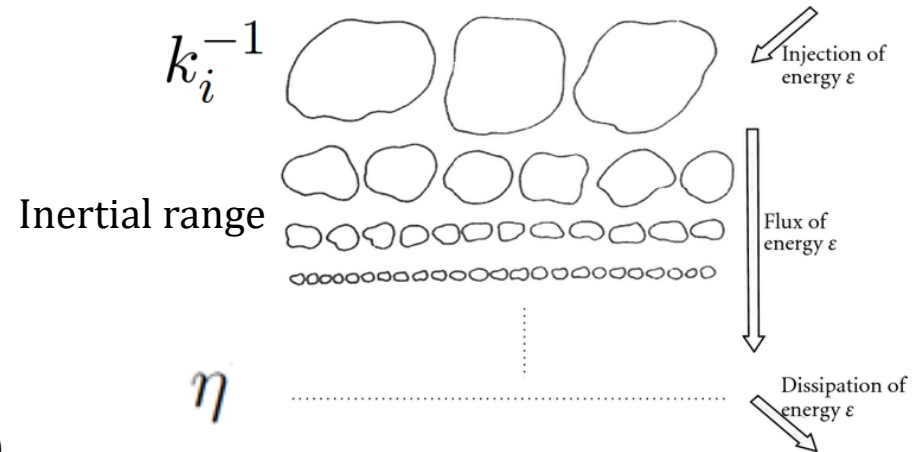
$$E_k = C_K \epsilon^{2/3} k^{-5/3}$$

kinetic energy cascade of incompressible flow

ϵ scale-invariant energy flux

Power-law turbulent spectra measured in many types of flows

Fluxes more challenging to measure (often indirect)



Can we observe in real time a turbulent cascade?

Can we directly measure a turbulent flux in momentum space?

How does turbulence look like in the quantum realm?

How does a turbulent cascade end in a dissipationless quantum fluid?

Is there a generic scenario for a turbulent cascade of excitations?

Hydrodynamic vs. **Wave** Turbulence

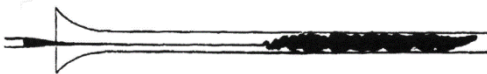
Hydrodynamic turbulence

Hydrodynamic (HD) equations for $\mathbf{v}(\mathbf{r},t)$

- Static solutions (e.g. Poiseuille)
- Static solutions can be unstable
 - time invariance is broken
 - chaotic, turbulent state can emerge



Laminar flow



Turbulent flow

Wave turbulence

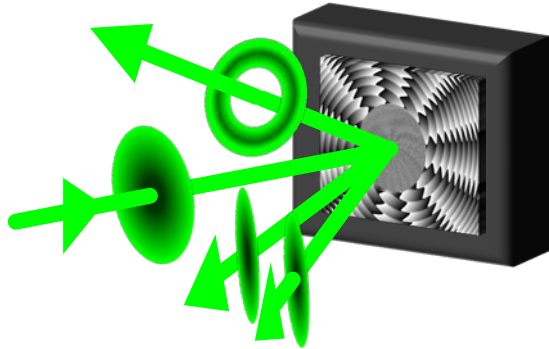
Eq of motion for **wave** occupations
e.g. $\mathbf{n}(\mathbf{k},t)$

Nonlinearity: waves **eigenmodes**
Analytic insights for weak interactions
(weak wave turbulence)

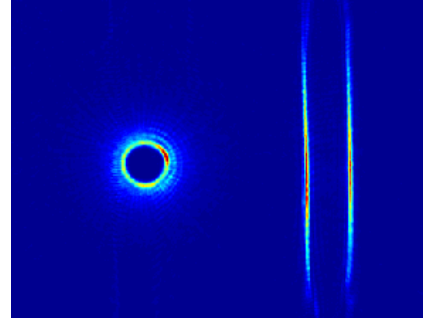


Experimental protocol I

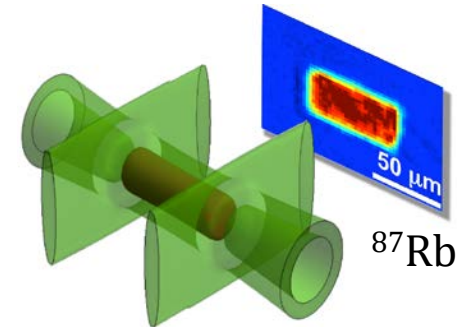
Spatial Light Modulator



repulsive light



Uniform density gas in hard-wall potential

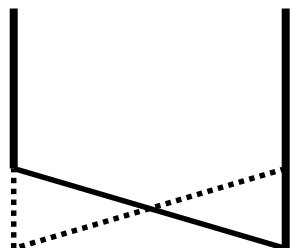


$N \sim 10^5$ $\mu \sim 2 \text{ nK}$
 $T \sim 5 \text{ nK}$ $\xi \sim 1 \mu\text{m}$
 $L \sim 30 \mu\text{m}$

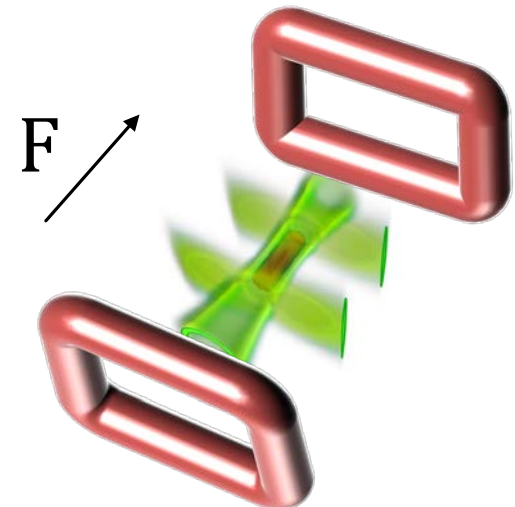
Clean setting for many-body physics, e.g.:

- Thermodynamics of uniform Bose gas, (Gaunt et al PRL 2013, Schmidutz et al PRL 2014)
- Kibble-Zurek mechanism in quenched gas, (NN et al, Science 2015)
- Bogoliubov quantum depletion of a Bose gas (Lopes et al, PRL 2017)
- Unitary Bose gas (Eigen et al PRX 2017, Eigen et al arXiv 2018)
- ...

Excite gas on system-size scale
with spatially uniform force



$$V_{\text{shake}}(\mathbf{r}, t) = \frac{\Delta U}{L} \sin(\omega t)x$$

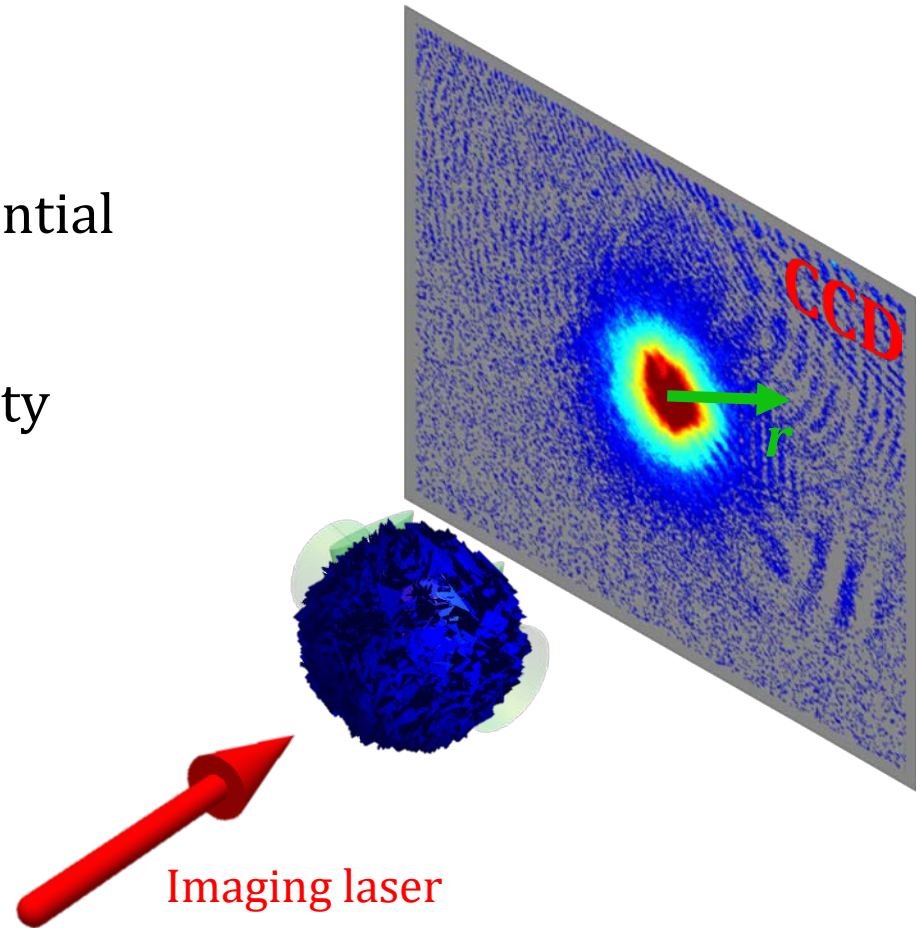


Experimental protocol II

For MWT, crucial quantity: $n(\mathbf{k})$

- 1) Switch off *abruptly* trapping potential
- 2) Let gas expand for time t_{TOF}
- 3) Take image of the resulting density by absorption imaging along \mathbf{z} axis

Position $\mathbf{r} \rightarrow$ velocity $\mathbf{v} = \mathbf{r}/t_{\text{TOF}}$
Mapping to momentum space
 \rightarrow 2D distribution $\tilde{n}(k_x, k_z)$

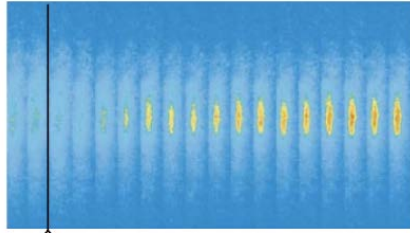
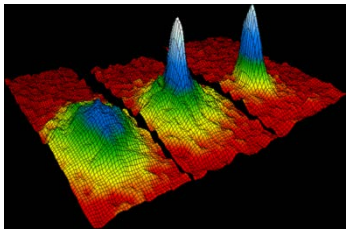


Important caveats:

- (i) *free* flight: no residual forces nor interactions in flight
- (ii) long flights to avoid in-trap density convolution

Harmonic vs. Uniform BEC Equilibrium

BEC occurs both in real and momentum space



Time of Flight (JILA)

In-situ (MIT)



2001



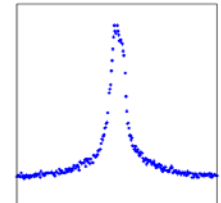
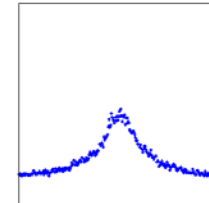
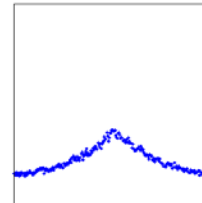
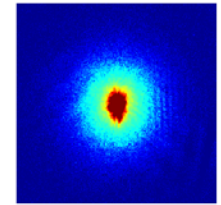
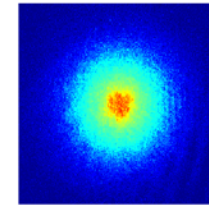
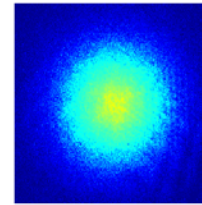
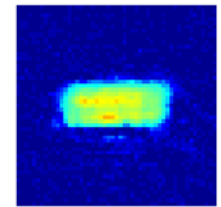
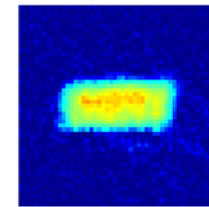
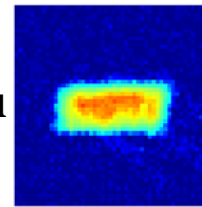
C. Wieman, E. Cornell

W. Ketterle

Reduce Temperature T_c



In-situ



Only in momentum space !

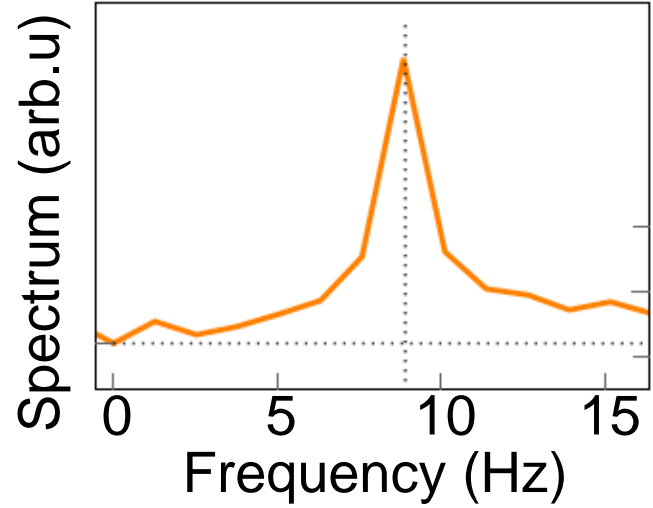
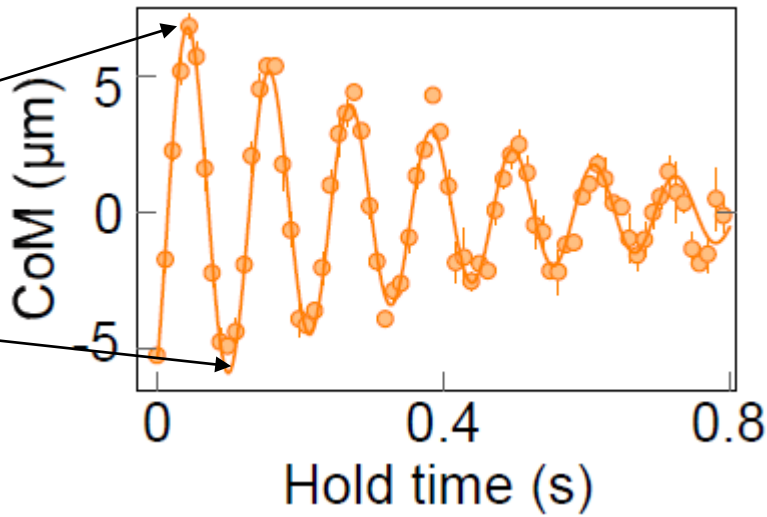
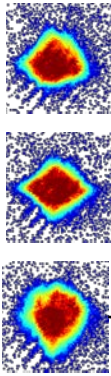
Linear regime: lowest-lying phonon

Near-equilibrium

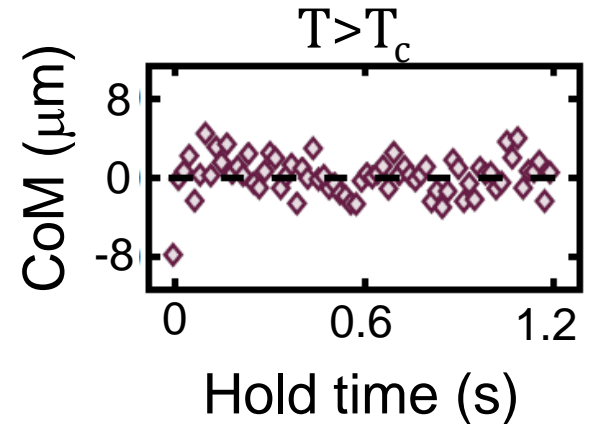
Single kick



ToF



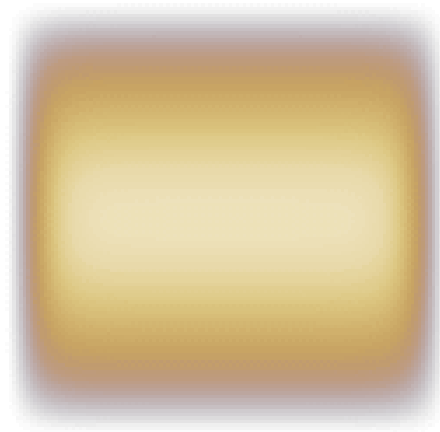
Note: this mode exists only below T_c ($mfp \sim 10L$ for $T > T_c$)
Ideal to start a direct matter-wave cascade (largest scale)



See Garratt et al, arXiv:1810.08195 (2018)

Emergence of Turbulence

(numerical simulations)

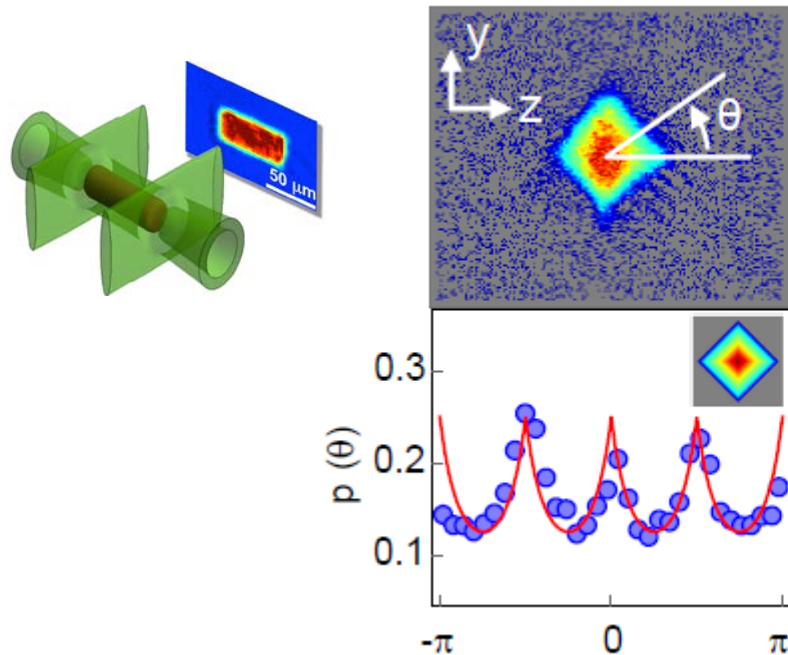


In situ density

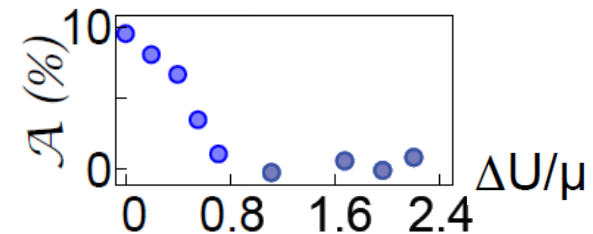
GPE in 256^3 grid with $10 \mu\text{s}$ steps for up to 4s (GPU)
30 min computation time per 1 sec experimental time

Emergence of turbulence

Initial BEC



Anisotropy parameter $\mathcal{A} = \frac{1}{2} \int d\theta \left| p(\theta) - \frac{1}{2\pi} \right|$

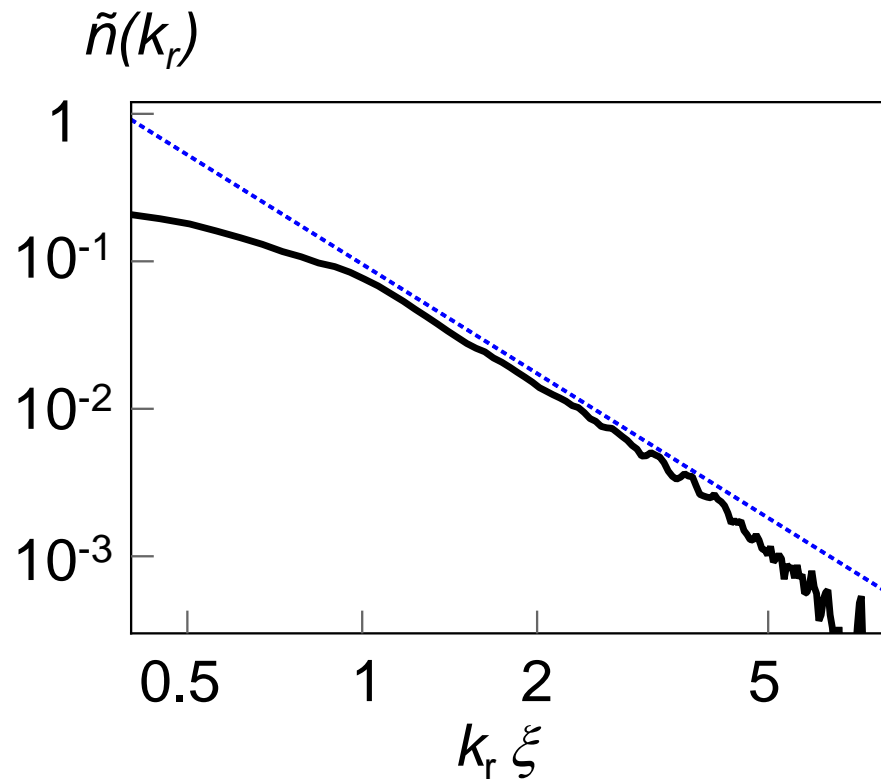
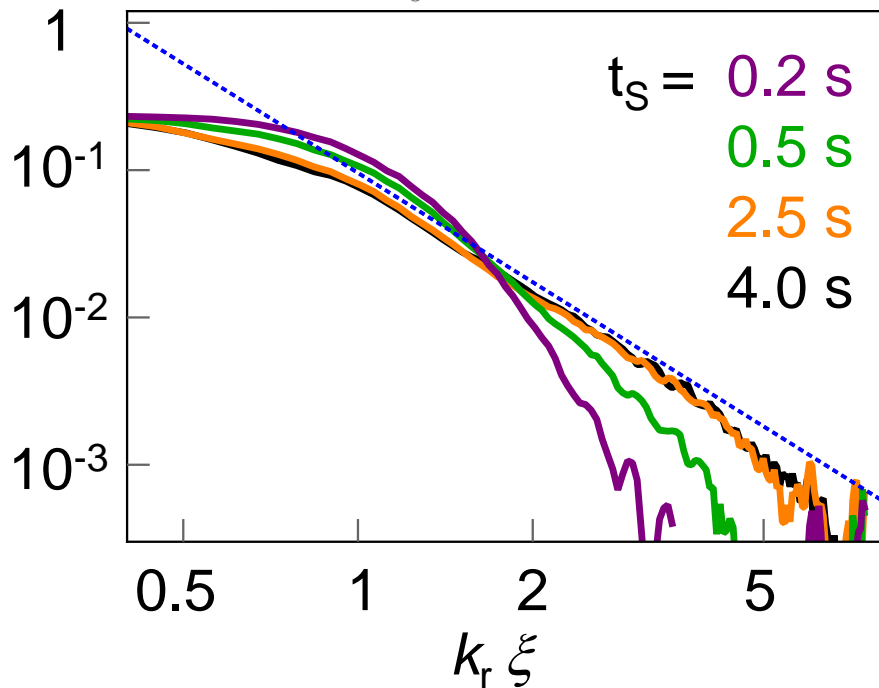


Isotropic state yet highly degenerate (not thermal!)

Matter-wave Turbulent Cascade

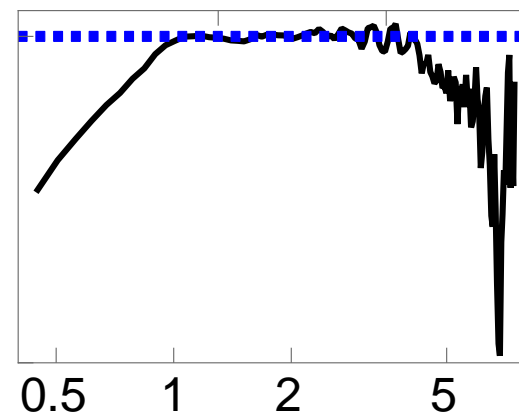
N. Navon *et al*, *Nature* **539** 72 (2016)

$$\tilde{n}(k_r) \quad \tilde{n}(k_r) = \int dk_z n(\sqrt{k_r^2 + k_z^2})$$



Compensated spectrum $k_r^{\gamma-1} \tilde{n}(k_r)$

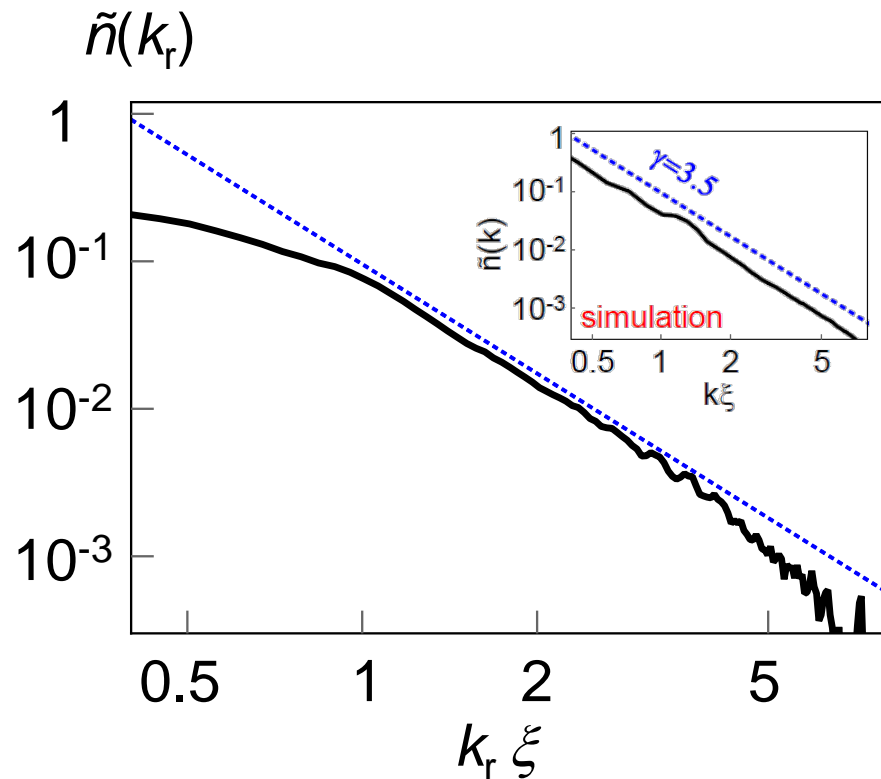
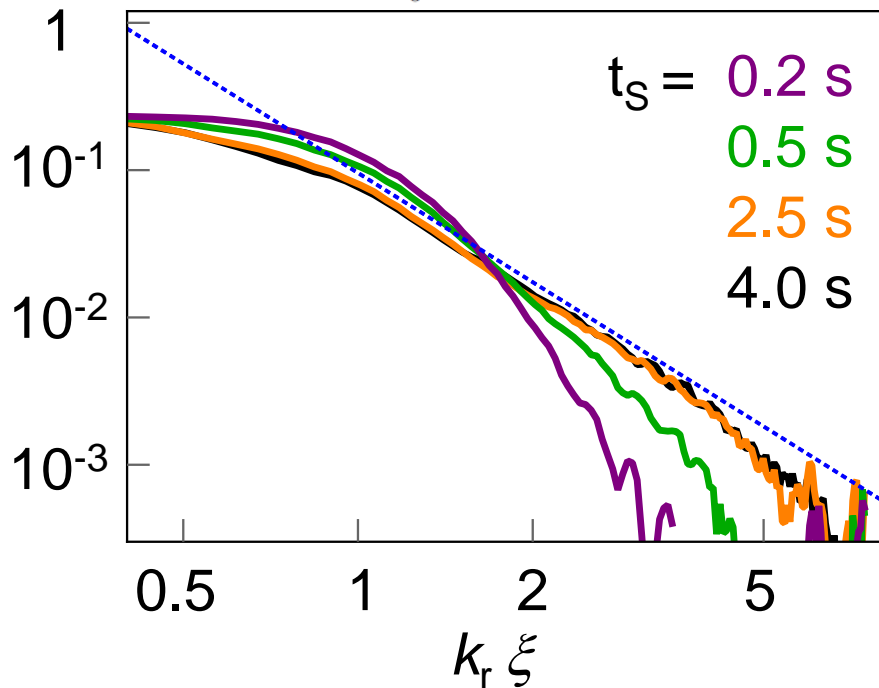
$$n(\mathbf{k}) \propto k^{-\gamma} \quad \text{with} \quad \gamma = 3.5(1)$$



Matter-wave Turbulent Cascade

N. Navon *et al*, *Nature* **539** 72 (2016)

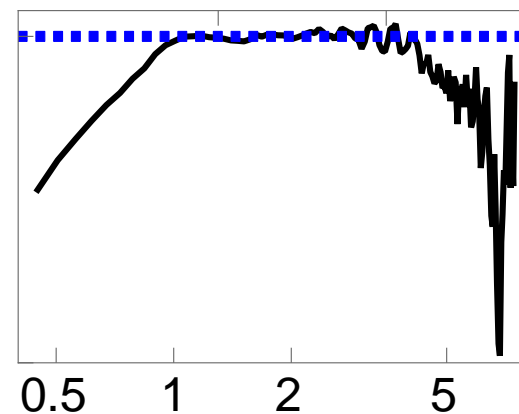
$$\tilde{n}(k_r) \quad \tilde{n}(k_r) = \int dk_z n(\sqrt{k_r^2 + k_z^2})$$



Compensated spectrum $k_r^{\gamma-1} \tilde{n}(k_r)$

$n(\mathbf{k}) \propto k^{-\gamma}$ with $\gamma = 3.5(1)$

Analytical prediction for $\gamma?$



Weak wave turbulence I

from Peierls to Zakharov
(a two-slide crash course)

- Starting from GP, analytical theory of WT in the weak coupling limit

$$i\hbar \frac{\partial \psi}{\partial t} = \left(-\frac{\hbar^2}{2m} \Delta + g|\psi|^2 \right) \psi$$

(classical Bose field, not BEC wavefunction!)

- With assumption of weak coupling and ‘random phase approximation’:

$$\frac{\partial n(\mathbf{k}, t)}{\partial t} = g^2 \mathcal{I}_{\text{coll}}[n(\mathbf{k}, t)]$$

where $\psi(\mathbf{r}, t) \propto \int d\mathbf{k} c(\mathbf{k}, t) e^{-i\mathbf{k}\cdot\mathbf{r}}$ and $\langle c(\mathbf{k}, t) c^*(\mathbf{k}', t) \rangle \approx n(\mathbf{k}, t) \delta(\mathbf{k} - \mathbf{k}')$
higher order Gaussian correlators (Wick)

- ‘Trivial’ steady states: thermodynamic equilibrium $n_k^{\text{eq}} = \frac{T}{\frac{k^2}{k_L^2} - \mu}$

- Exact turbulent cascade solutions $n_k \sim k^{-\gamma}$

Weak wave turbulence II

- If collision integral converges: KE is continuity equation (in \mathbf{k} space)

$$\frac{\partial n(\mathbf{k}, t)}{\partial t} = -\nabla_{\mathbf{k}} \cdot \Pi_n$$

- Continuity equation for energy $\mathcal{E}(\mathbf{k}, t) = \hbar\omega_k n(\mathbf{k}, t)$

$$\frac{\partial \mathcal{E}(\mathbf{k}, t)}{\partial t} = -\nabla_{\mathbf{k}} \cdot \Pi_{\mathcal{E}}$$

- WWT theory of GP predicts two cascade solutions

(i) Direct energy cascade $n(k) \sim k^{-3}$

$$\begin{aligned} \Pi_{\mathcal{E}} &= \epsilon \\ \Pi_n &= 0 \end{aligned}$$

?!?

(ii) Inverse particle cascade $n(k) \sim k^{-7/3}$

$$\Pi_{\mathcal{E}} = 0$$

$$\Pi_n = Q$$

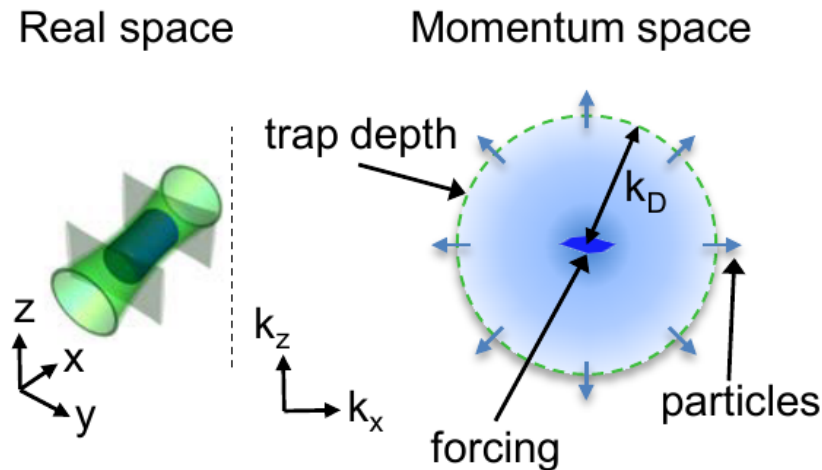
experimental relevance not obvious! source and sink terms...

Link to 0th law of turbulence (phenomenological): keeping everything else the same, dissipation rate $\epsilon \rightarrow$ constant when $\nu \rightarrow 0$ (anomalous limit)

“independence of forcing and dissipation”

Probing fluxes in a turbulent quantum gas

- Exotic feature: box depth introduces a momentum cutoff k_D
a *synthetic* dissipation scale



- Direct measurement of particle flux through shell of *tunable* radius k_D

$$\frac{\partial N}{\partial t} = \frac{\partial}{\partial t} \int_0^{k_D} n^{1D}(k) dk = -\Pi_n(k_D, t)$$

We measure total E (not just HD) \rightarrow dissipation = particle loss

- Note: in HD turbulence, dissipation (Kolmogorov) scale $\eta \propto \nu^{3/4}$

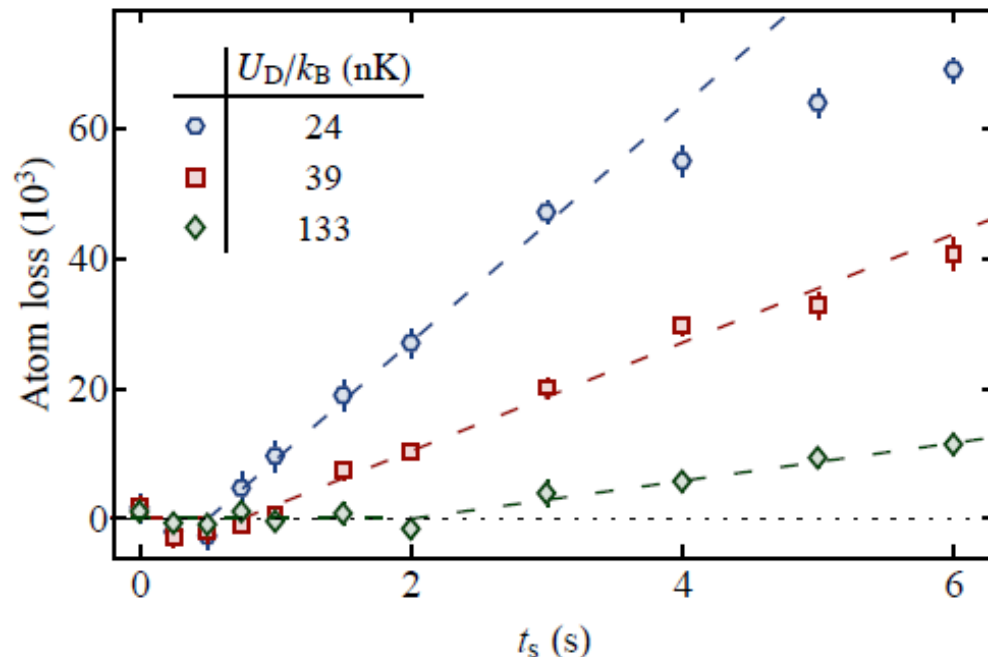
Measurement of remaining atoms

Navon et al, *arXiv:1807.07564* (2018)

- We measure the atom lost during the shaking

$$N_{\text{lost}} = N(t_s)|_{\Delta U} - N(t_s)|_{\Delta U=0}$$

Differential measurement factors out important systematic effects:
residual background gas induced losses, slow drifts



$$\mu \approx 2 \text{ nK}$$

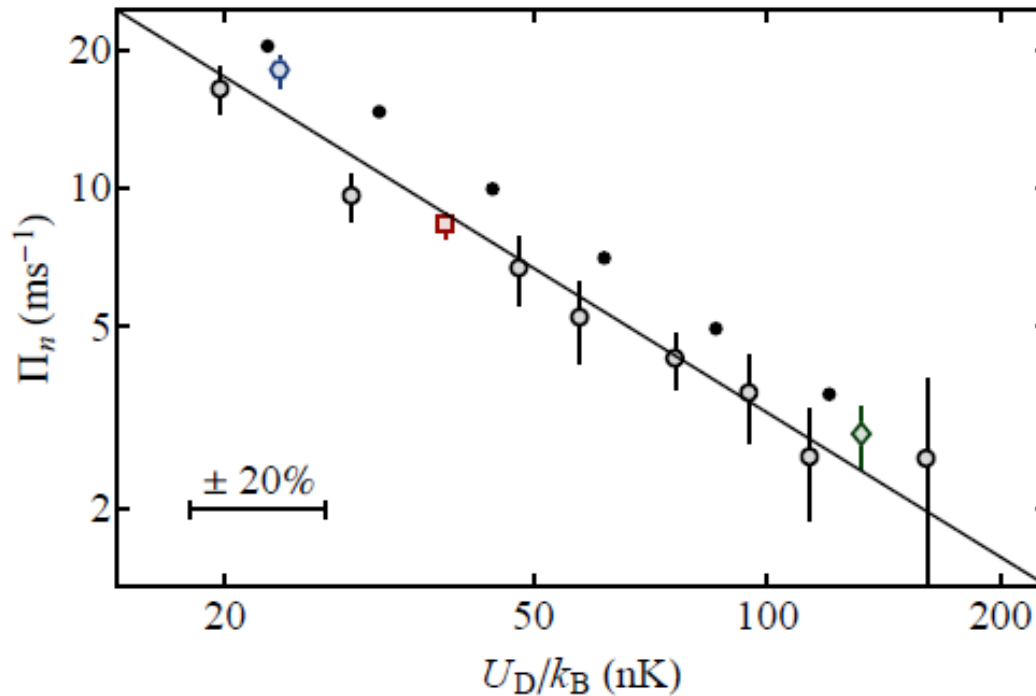
$$\Delta U = 1.3\mu$$

$$U_D = \frac{\hbar^2 k_D^2}{2m}$$

Two features: (1) delayed onset of losses
(2) lost rate constant afterwards
Both depend on the dissipation scale

The particle flux

Navon et al, *arXiv:1807.07564* (2018)



$$\Pi_n \propto U_D^\alpha$$

$$\alpha = -1.05(8)$$

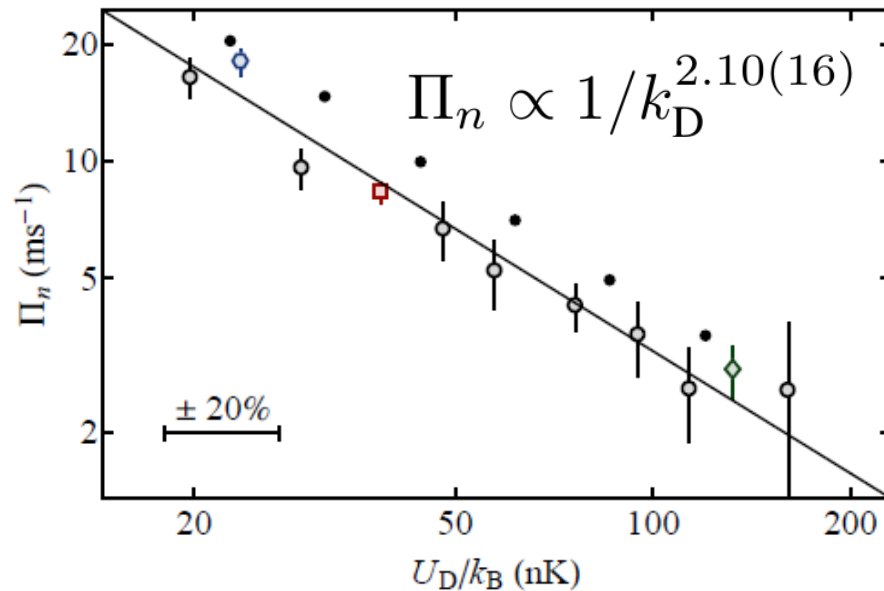
$$\Rightarrow \Pi_n \propto 1/k_D^{2.10(16)}$$

Simulation of GPE + dissipation
(no free parameters)

$$\alpha_{\text{sim}} = -1.06(1)$$

→ Particle flux vanishes in the limit of zero dissipation scale

Relation between particle and energy fluxes



For weak interactions $\mathcal{E}(k) = \hbar\omega(k)n(k)$

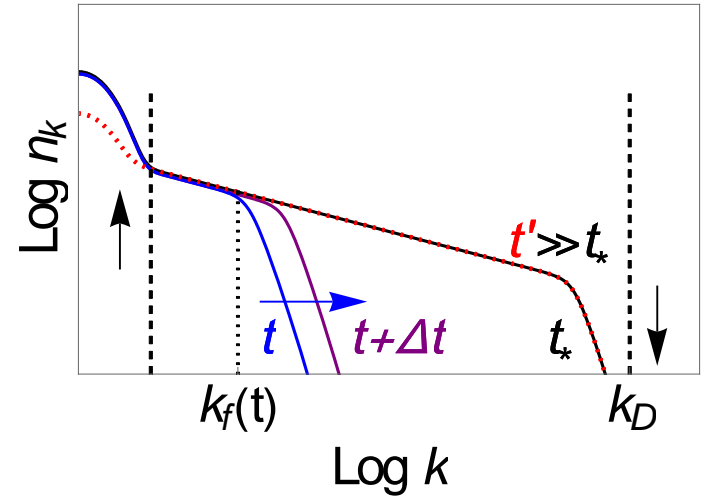
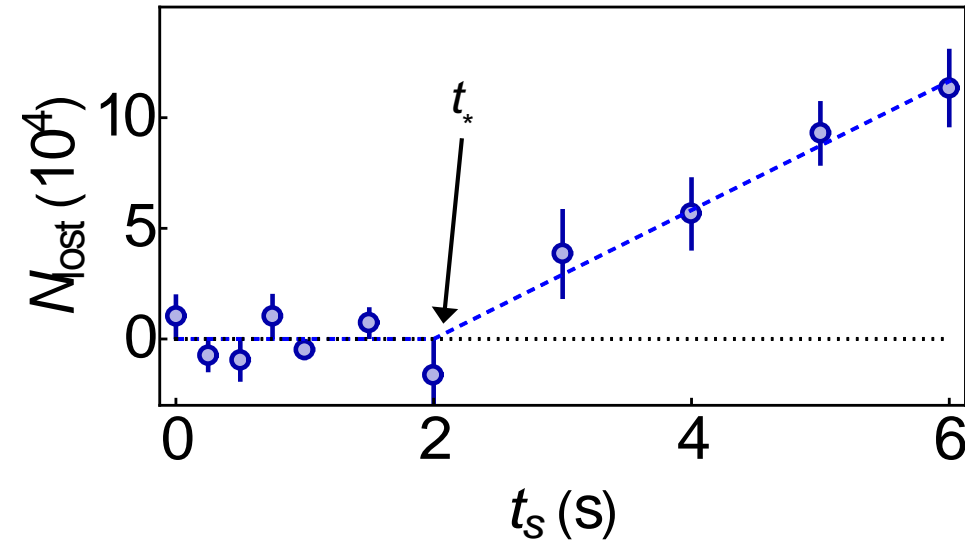
Then, naively $\Pi_{\mathcal{E}}(k) = \hbar\omega(k)\Pi_n(k)$ *can't be true in steady state!*

From continuity equations $\Pi_{\mathcal{E}}(k_D) = \hbar\omega(k_D)\Pi_n(k_D)$

We thus find $\Pi_{\mathcal{E}}(k_D) = \Pi_{\mathcal{E}} (= \epsilon)$ energy cascade

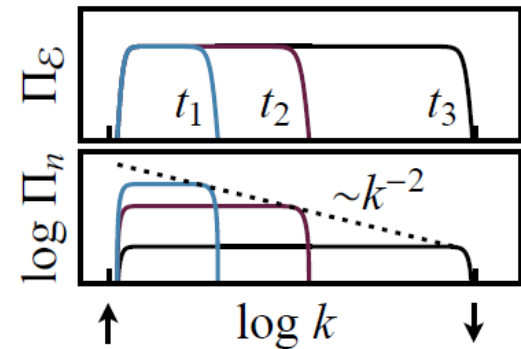
« 0th law » of turbulence: for fixed forcing, energy input/dissipation rate tends to a constant when $\nu \rightarrow 0$

Consistent picture of the energy cascade (and its fluxes)



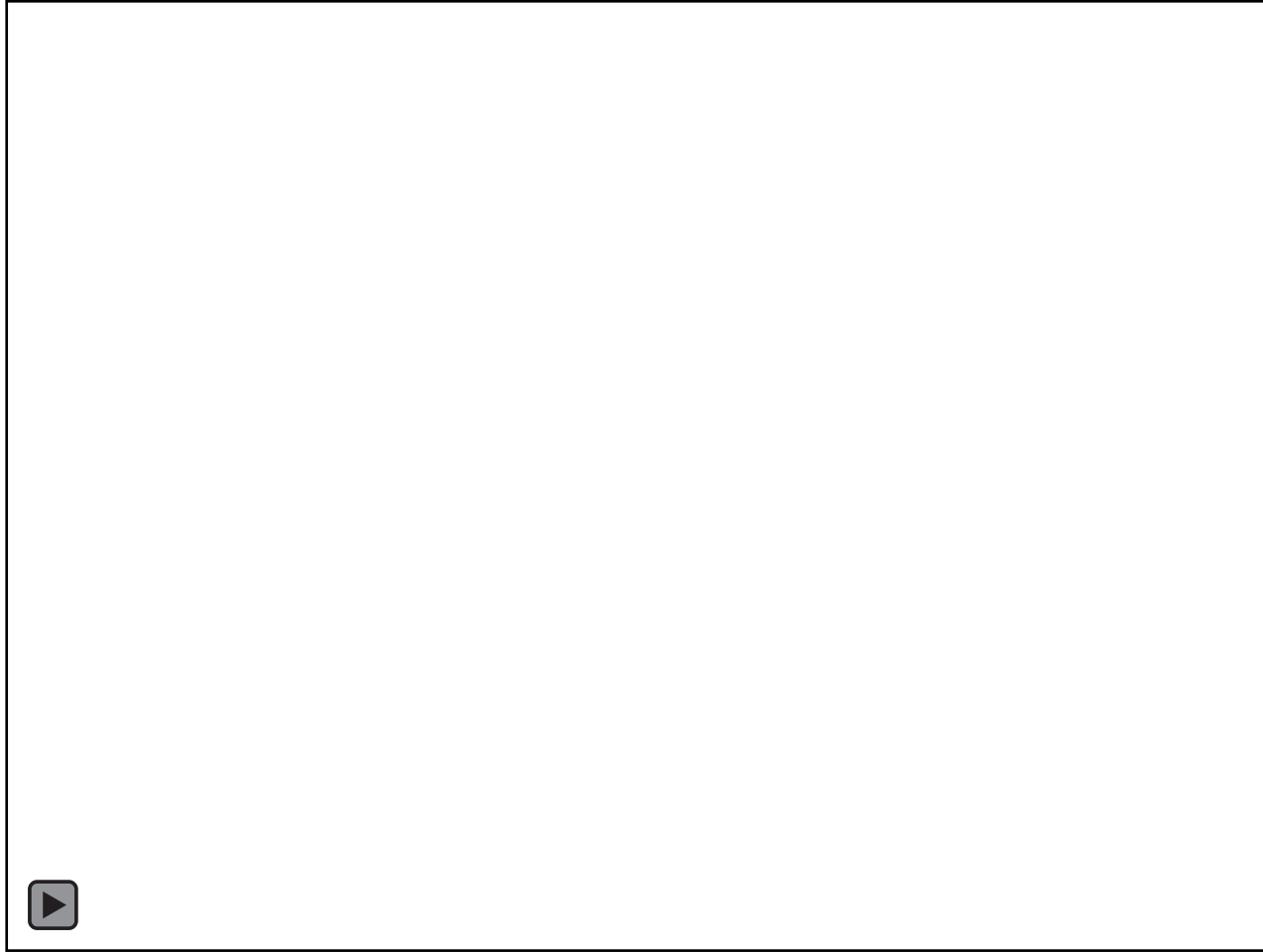
Real-time observation of turbulent cascade

Particle flux has to “adjust” to sustain the direct energy cascade



Testable prediction of cascade front dynamics (pre-steady state)

Momentum distribution dynamics (numerical simulations)



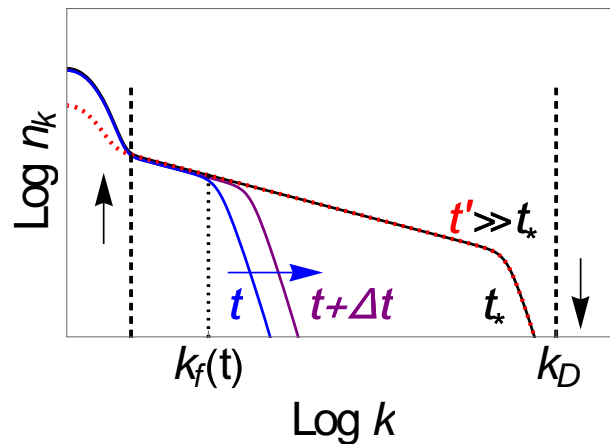
and M. Tsubota

Amazing that such complex real-space dynamics is captured by
simple Fourier space dynamics

Cascade front propagation

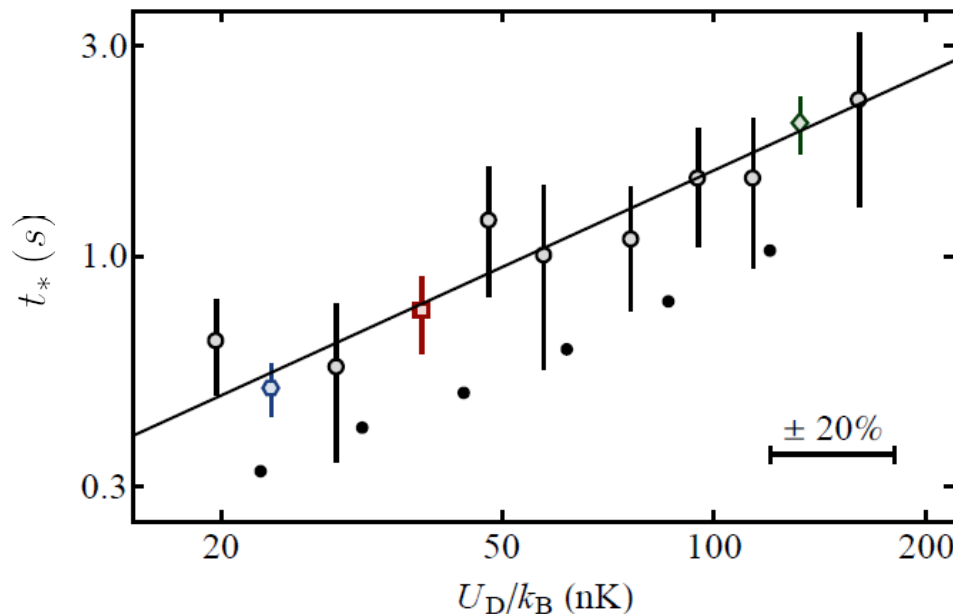
Energy in inertial range up to the cascade front:

$$E_{\text{inert}}(t) = \int_{k_{\text{low}}}^{k_f(t)} \mathcal{E}(k) dk$$



Energy injected pushes cascade front forward

$$\epsilon \equiv \frac{dE_{\text{inert}}}{dt} = \frac{dk_f}{dt} k_f^2 \cdot (k_f^2 n_{k_f}) \quad \text{with} \quad n_k \sim k^{-\gamma} \quad \text{leads to} \quad t_* \propto U_D^\beta$$



$$\beta = \frac{5 - \gamma}{2} = 0.75(5)$$

$$\beta_{\text{exp}} = 0.73(6)$$

$$\beta_{\text{sim}} = 0.68(2)$$

Summary/Outlook

- Turbulent steady-state cascade in quantum gas
- Measurement of the fluxes
 - Particle flux vanishes in the zero-dissipation limit
 - Energy flux is constant
 - Quantitative understanding of cascade front propagation
- Simple technique specific to ultracold gases
- No-free parameters GPE simulations including realistic dissipation
- Open questions:
 - Cascade exponent value from analytics?
(related: Quantify deviation from WWT theory?)
 - Ultimate dissipation mechanism for cascade? (effective dissipation beyond the classical regime?)
 - Establish the validity of classical field methods
 - Interplay of compressible and incompressible-fluid turbulence

The Team(s)

Color scale: Prof, Dr, Ph.D student

Cambridge



Zoran Hadzibabic



Rob Smith



Alex Gaunt



Rapha Lopes



Chris Eigen



Jinyi Zhang

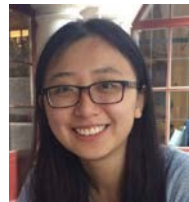
Yale



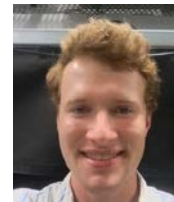
NN



Franklin Vivanco



Yunpeng Ji



Grant Schumacher



Gabriel Assumpcao



Jingjing Pan

Japan



Makoto
Tsubota



Kazuya
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