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Causality & Positivity with Gravity











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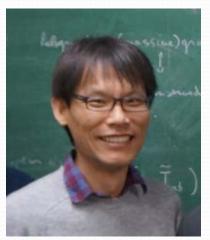
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Within low-energy gravitational EFTs,

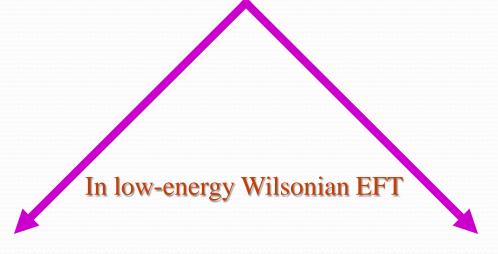
- Constraints from standard UV completion?
- Constraints from causality?



Non-Gravitational EFT

UV completion

- ✓ Local (Froissart Bound)
- ✓ Unitary (optical theorem)
- ✓ Lorentz invariant (crossing symmetry)
- ✓ CAUSAL (analyticity)



(sub)luminal sound speed

positivity bounds

$$\left. \frac{\mathrm{d}^2 \mathcal{A}(s,t)}{\mathrm{d}s^2} \right|_{t=0} > 0$$

A: 2-2 elastic scattering amplitude

Adding Gravity?

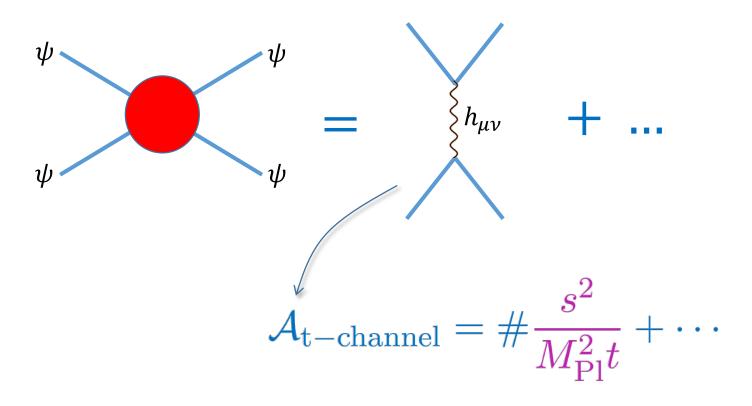
Both of these requirements are more subtle for gravitational EFTs

(sub)luminal sound speed

positivity bounds

Justified for completions of string/Regge higher spin type Hamada, Noumi & Shiu, 1810.03637

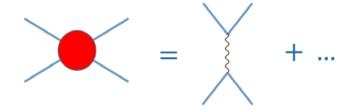
Positivity Bounds in Gravitational LEEFT



t-channel pole from gravity exchange compromises positivity bound

$$\left. \frac{\mathrm{d}^2 \mathcal{A}(s,t)}{\mathrm{d}s^2} \right|_{t=0} > 0$$

Positivity Bounds in Gravitational LEEFT



Gravity is non-dynamical in 3d, upon compactifying $4d\rightarrow 3d\times S^1$, contribution from t-channel pole should disappear Are bounds simply applicable to rest of amplitude?

$$4d \rightarrow 3d \times S^1 \rightarrow 4d$$

$$1902.03250$$

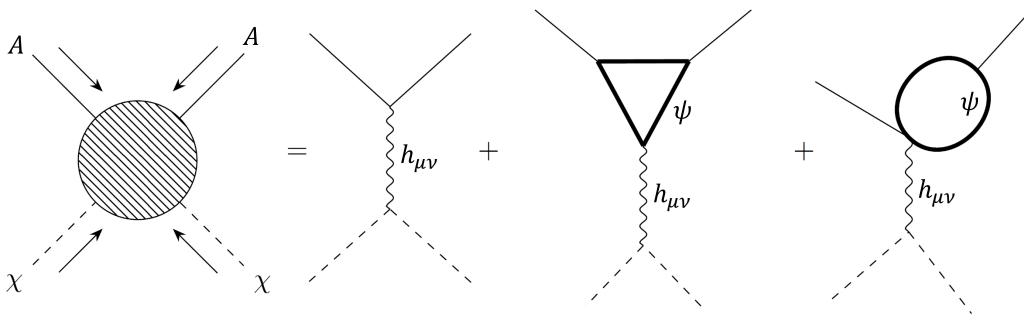
$$\frac{d^2}{ds^2} \mathcal{A}_{\text{t-pole subtracted}}^{\text{(4d)}}(s,t) \Big|_{t=0} > 0$$

Potential caveats pointed out in Loges, Noumi & Shiu, 1909.01352

Let's explore the validity of this bound in a specific example with known partial UV completion

Scalar QED with gravity

$$\mathcal{L}_{\text{sQED}} = \frac{M_{\text{Pl}}^2}{2} R - \frac{1}{2} (\partial A)^2 - \frac{1}{2} (\partial \psi)^2 - \frac{1}{2} M^2 \psi^2 - \alpha M A \psi^2 - \frac{1}{2} (\partial \chi)^2 - \frac{1}{2} m^2 \chi^2$$



Scalar QED with gravity

$$\mathcal{A}_{\text{sQED}}(s,t) = -\frac{s^2}{M_{\text{Pl}}^2 t} - \frac{\alpha^2 s^2}{90(4\pi)^2 M^2 M_{\text{Pl}}^2} + \mathcal{O}(t^0)$$

$$\frac{d^{2}}{ds^{2}} A_{\text{t-pole subtracted}}(s,t) \Big|_{t=0} > 0$$

$$\frac{d^{2} \mathcal{A}_{\text{sQED,no pole}}(s,0)}{ds^{2}} = -\frac{2\alpha^{2}}{90(4\pi)^{2} M^{2} M_{\text{Pl}}^{2}} > 0$$

in contradiction...

Same contradiction for QED minimally coupled to gravity

Compactified bounds & Scalar QED

$$\left(\frac{\mathrm{d}^2}{\mathrm{d}s^2} \mathcal{A}_{\text{t-pole subtracted}}(s,t) \Big|_{t=0} > 0 \right)$$



(Scalar) QED minimally coupled with QED

• **Either** QED minimally coupled with gravity is not consistent...

would require new interactions between any massive particles (eg. DM) and the photon at the scale

$$\Lambda \leq (M_{\rm Pl}M)^{1/2}$$

Compactified bounds & Scalar QED

$$\left(\frac{\mathrm{d}^2}{\mathrm{d}s^2} \mathcal{A}_{\text{t-pole subtracted}}(s,t) \Big|_{t=0} > 0 \right)$$

(Scalar) QED minimally coupled with QED

• Or 3d compactified bounds are **not** justified

Even though gravity is not dynamical in 3d, the t-channel pole only disappears after Eikonal resummation —— leading to an overall delta function

The delta function is the 3d manifestation of 4d pole albeit in a different form

Removing delta function leads to a resulting amplitude $\tilde{\mathcal{A}}$ with $\lim \tilde{\mathcal{A}} \not > 0$ Ciafaloni (1992)

Alternatively amplitude $\tilde{\tilde{\mathcal{A}}}$ can be defined with gravity-redressed states \longrightarrow compromises crossing symmetry

with Alberte, Jaitly and Tolley 2007.12667

Compactified bounds & Scalar QED

$$\left(\frac{\mathrm{d}^2}{\mathrm{d}s^2} \mathcal{A}_{\text{t-pole subtracted}}(s,t) \Big|_{t=0} > 0 \right)$$

(Scalar) QED minimally coupled with QED

- **Either** QED minimally coupled with gravity is not consistent...
- Or 3d compactified bounds are **not** justified

There is no properly defined 3d amplitude which is simultaneously:

- Finite and Analytic
- Has positive Imaginary part
- Enjoys manifest crossing symmetry

Essential for the derivation of the positivity bounds

--- t-channel pole affects positivity bounds

Approximate Positivity

• Or 3d compactified bounds are **not** justified

The best we can then argue is that the Positivity bounds ought to be satisfied in a limit $M_{\rm Pl} \rightarrow \infty$ where gravity decouples

More precisely, if a 2-2 low-energy elastic scattering amplitude is of the form:

$$A(s,t) \sim -\frac{s^2}{M_{\rm Pl}^2 t} + \frac{c}{M^4} s^2 + \cdots$$

Then the coupling constant needs not be positive but rather

$$c > - \frac{M^2}{M_{\rm Dl}^2} imes {\cal O}(1)$$
 Not assuming specific UV behavior

$$\mathcal{L}_{\text{EFT GR}} = \sqrt{-g} \left[\frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}^{(\text{light})}(g, \psi) + C_{R^2} R^2 + C_{W^2} W_{\mu\nu\alpha\beta}^2 + \dots \right]$$

$$\mathcal{L}_{\text{EFT GR}} = \sqrt{-g} \left[\frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}^{(\text{light})}(g, \psi) + C_{R^2} R^2 + C_{W^2} W_{\mu\nu\alpha\beta}^2 + \dots \right]$$

All the light fields at low-energy (e.g. including photon)

Consider these fields to be minimally coupled In this frame, light travels at the speed of light c = 1 in the vacuum

Curvature-square operators can be removed by field redefinition at the price of including non-minimal couplings to light fields

Respective causal structure remains the same, just shifts the question somewhere else

$$\mathcal{L}_{\text{EFT GR}} = \sqrt{-g} \left[\frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}^{(\text{light})}(g, \psi) + C_{R^2} R^2 + C_{W^2} W_{\mu\nu\alpha\beta}^2 + \dots \right]$$

Consider tensor fluctuations on FLRW,

$$ds^{2} = a^{2}(\eta)\eta_{\mu\nu}dx^{\mu}dx^{\nu} + a h_{ij}dx^{i}dx^{j}$$

$$\left[-\partial_{\eta}^2 + \left(1 - \frac{16C_{W^2}\dot{H}}{M_{\rm Pl}^2} \right) \nabla^2 \right] \tilde{h} = m_0^2 \tilde{h}$$

$$\mathcal{L}_{\text{EFT GR}} = \sqrt{-g} \left[\frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}^{(\text{light})}(g, \psi) + C_{R^2} R^2 + C_{W^2} W_{\mu\nu\alpha\beta}^2 + \dots \right]$$

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$$c_s^2 = 1 + \frac{16C_{W^2}(-\dot{H})}{M_{\rm Pl}^2} + \mathcal{O}\left(\frac{H^4}{M_{\rm Pl}^4}, \frac{k^4H^2}{M_{\rm Pl}^6}\right)$$

with Tolley 1909.00881

Speed of Gravity

Within the regime of validity of the EFT,

$$c_s^2 = 1 + \frac{16C_{W^2}(-\dot{H})}{M_{\text{Pl}}^2} + \mathcal{O}\left(\frac{H^4}{M_{\text{Pl}}^4}, \frac{k^4H^2}{M_{\text{Pl}}^6}\right)$$

• For a maximally symmetric spacetime $\dot{H} = 0$, modes are luminal

• We expect
$$C_{W^2} \sim \mathcal{O}(1)$$
 \Rightarrow $|\Delta c_S| \ll \frac{H^2}{M_{\rm Pl}^2} \sim 10^{-120}$

Speed of Gravity

Within the regime of validity of the EFT,

$$c_s^2 = 1 + \frac{16C_{W^2}(-\dot{H})}{M_{\rm Pl}^2} + \mathcal{O}\left(\frac{H^4}{M_{\rm Pl}^4}, \frac{k^4H^2}{M_{\rm Pl}^6}\right)$$

If $\dot{H} \neq 0$ and NEC is satisfied, $\dot{H} < 0$ modes are

$$\begin{array}{ccc} \text{subluminal} & \Leftrightarrow & C_{W^2} < 0 \\ super \text{luminal} & \Leftrightarrow & C_{W^2} > 0 \end{array}$$

Does it mean that the low-energy EFT is only consistent if $C_{W^2} < 0$??

Speed of Gravity

$$\mathcal{L}_{\text{EFT GR}} = \sqrt{-g} \left[\frac{M_{\text{Pl}}^2}{2} R + \mathcal{L}^{(\text{light})}(g, \psi) + C_{R^2} R^2 + C_{W^2} W_{\mu\nu\alpha\beta}^2 + \dots \right]$$

$$\begin{array}{lll} \text{subluminal} & \Leftrightarrow & C_{W^2} < 0 \\ super \text{luminal} & \Leftrightarrow & C_{W^2} > 0 \end{array}$$

From a field theory perspective the constraints on enjoying a standard causal high energy completion are (so far) simply

$$C_{W^2} > -\mathcal{O}\left(\frac{M^2}{M_{\mathrm{Pl}}^2}\right)$$

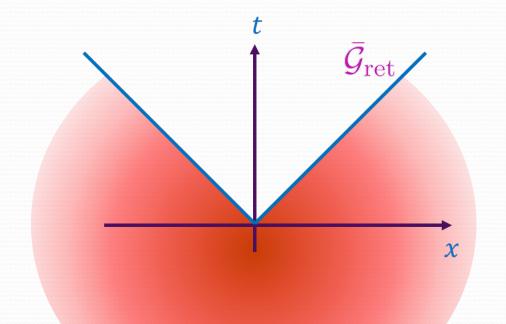
How is this consistent with causality within the low-energy EFT???

Causality

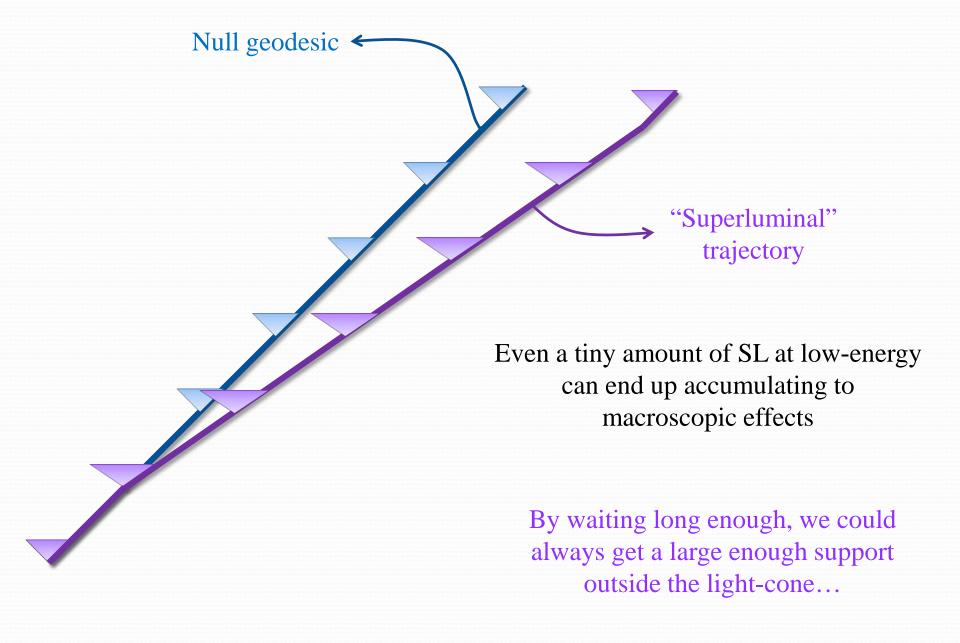
The physical speed of propagation is given by the front velocity:

$$v_{\text{front}} = \lim_{k \to \infty} v_{\text{phase}}(k)$$

But causality itself requires that the retarded propagator vanishes outside the light-cone which typically requires (sub)luminality even at low-energy



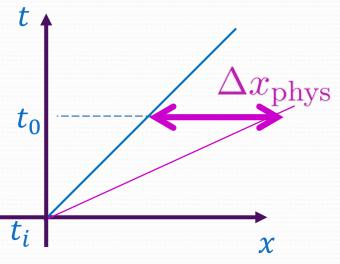
Support Outside Light-Cone



Support Outside Light-Cone

EFT has a cutoff $M \leq M_{\rm Pl}$

For any mode, with physical frequency k, one can only trust EFT so long as $\Box_{\text{FLRW}} \sim \Box_{\text{Minkowski}} + \frac{kH}{a} \ll M_{\text{Pl}}^2$



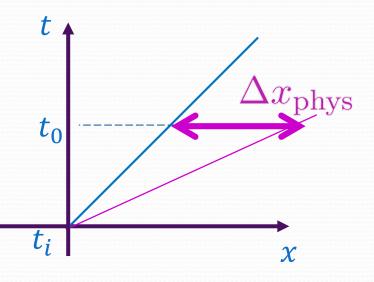
Cannot send a mode with arbitrarily small wavelength

$$\lambda_i \gg \frac{H_i}{M_{
m Pl}^2}$$

$$\Delta x_{\rm phys} = a_0 \int_{\eta_i}^{\eta_0} \Delta c_S d\eta = a_0 \int_{t_i}^{t_0} \frac{-\dot{H}}{aM_{\rm Pl}^2} dt < \frac{a_0}{a_i} \frac{H_i}{M_{\rm Pl}^2} \ll \lambda_{\rm phys}$$

Support Outside Light-Cone

$$\Delta x_{\rm phys} \ll \lambda_{\rm phys}(t_0)$$



There is never support outside the light cone by a resolvable amount within the regime of validity of the EFT

→ No violation of causality

The amount of superluminality is so small that it can never build up to lead to macroscopic violation of causality.

QED on curved spacetime

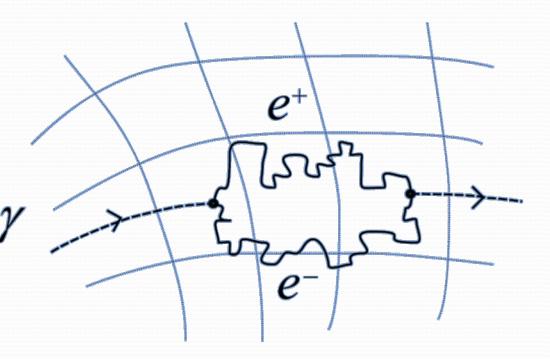
Drummond & Hathrell, PRD 1980
Hollowood & Shore 0707.2302, 0707.2303, 0806.1019, 0905.0771, 1006.0145, 1006.1238, 1111.3174, 1205.3291, 1512.04952
Goon & Hinterbichler, 1609.00723

M: electron mass

$$\mathcal{L} = \sqrt{-g} \left(-\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{\alpha}{M^2} R_{abcd} F^{ab} F^{cd} + \dots \right)$$

As the photon propagates, it interacts with virtual electron pairs

feels the curvature in region around its geodesic



From Hollowood & Shore

QED on curved spacetime

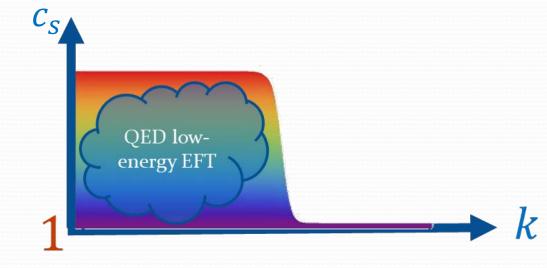
There are many space-time backgrounds for which the low-energy group velocity is superluminal. Eg. Schwarzschild, Type I & II conformally flat backgrounds, ...

E.g. on Schwarzschild,
$$c_s^2=1+rac{eta_P}{M^2}rac{r_g}{r^3}+\mathcal{O}\left(rac{r_g^2}{M^4r^6}
ight)+\mathcal{O}\left(rac{k^4}{M^4}
ight)$$

M: electron mass

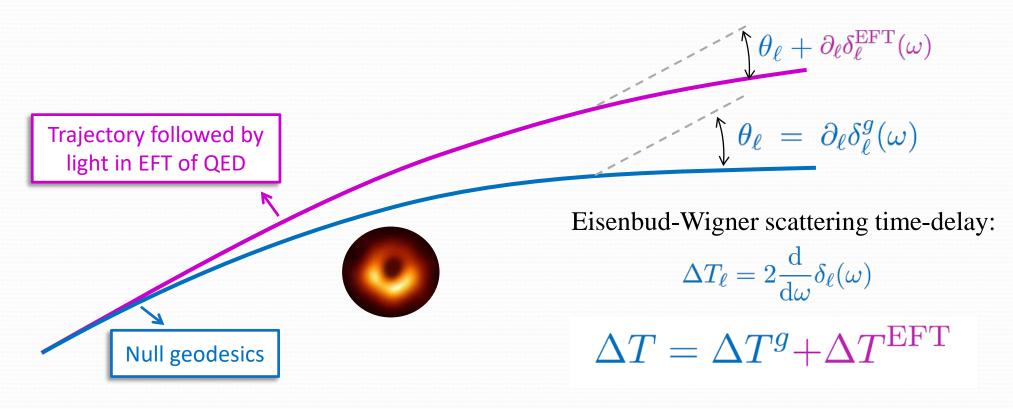
 $\beta_P \sim {\it O}(1)$ - polarization dependent constant

 $\beta_P > 0$ for radially polarized light



Low-energy `superluminality' is precisely related to (non)-positivity bounds

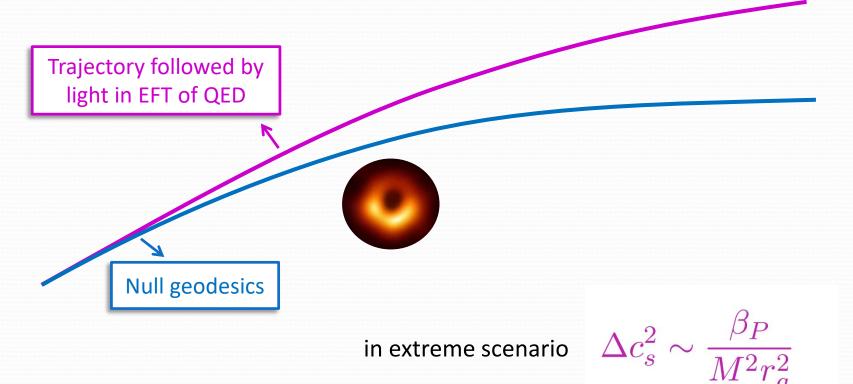
Time-delay/advance



It was previously argued that while $\Delta T^{\rm EFT} < 0$ we still have $\Delta T > 0$ in the regime of validity of the EFT

but that's not enough...

Time-delay/advance



$$\Delta T^{\rm EFT} \sim -\frac{\beta_P}{M^2 r_g} < 0 \quad {\rm for} \quad \beta_P > 0$$

Regime of Validity of EFT

The low-energy EFT is only valid below the scale *M*Above that scale one should go back to the microscopic description

wave of null momentum $k_{\mu} = (\omega, ...)$ in a curved background with Weyl tensor $W_{\mu\nu\alpha\beta}$

$$|W^p| \ll M^{2p}$$

For null momenta this cannot put any direct bound on k_{μ} nor on ω (ω is not a Lorentz scalar)

However to be within the regime of validity of the EFT other invariants ought to be bounded E.g. any invariant of the form

$$\left| \left(W^a{}_{bcd} k^c k^d \right)^p \right| \ll M^{4p}$$

Regime of Validity of EFT

The low-energy EFT is only valid below the scale M Above that scale one should go back to the microscopic description

$$\left| \left(W^a{}_{bcd} k^c k^d \right)^p \right| \ll M^{4p}$$

In the extreme scenario
$$\left[\begin{array}{ccc} \dfrac{\omega^2}{r_g^2} \ll M^4 & \Rightarrow & \dfrac{1}{M^2 r_g} \ll \lambda \end{array} \right]$$

 ω : asymptotic energy of the scattering particle

Causality in Gravitational Theories

$$\Delta T^{\rm EFT} \sim -\frac{\beta_P}{M^2 r_g} < 0 \quad {\rm for} \quad \beta_P > 0$$

$$\left(\begin{array}{ccc} \frac{\omega^2}{r_g^2} \ll M^4 & \Rightarrow & \frac{1}{M^2 r_g} \ll \lambda \end{array}\right)$$

$$\left|\Delta T^{
m EFT}
ight|\sim rac{\mathcal{O}(1)}{M^2r_g}\ll \pmb{\lambda}$$

Amount of SL is small enough not to lead to any resolvable time advance (as it should be)

Causality in Gravitational Theories

Conjecture: In a frame where gravity can be decoupled, a small amount of SL at low-energy is still consistent with causality so long as

$$\lim_{M_{
m Pl} o \infty} |c_s^2 - 1| \sim M_{
m Pl}^{-lpha} \quad ext{with} \quad lpha \geq 2$$

Time-advance

In the EFT of gravity or QED, the time advance due to SL is always unresolvable

$$\left|\Delta T_{\ell}^{\mathrm{EFT}}\right| \ll \omega^{-1}$$

 $\left|\Delta T_{\ell}^{\mathrm{EFT}}\right| \ll \omega^{-1}$ The time advance is smaller than the geometric optics resolution scale it is not resolvable

This is a very different statement than

$$\left|\Delta T_{\ell}^{\mathrm{EFT}}\right| \ll M^{-1}$$

While this relation is also true it is not relevant:

- The low-energy EFT is only used to determine the trajectory, Nothing demands that the time advance should be measured with apparatus that live in the low-energy EFT
- The time delay is not a Lorentz invariant quantity so one cannot use M^{-1} as its cutoff

No support outside the light-cone

If $|\Delta T^{\rm EFT}| \ll \omega^{-1}$, its sign cannot be directly linked with causality

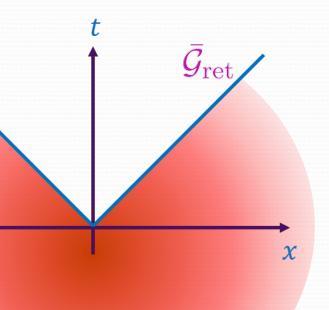
No support outside the light-cone

If $|\Delta T^{\rm EFT}| \ll \omega^{-1}$, there are is no dangerous growth of secular effects

Retarded Green's function can be computed perturbatively



There can be no support outside the light-cone



$$ar{\Box}ar{\mathcal{G}}_{ ext{ret}} \ = \ \delta$$
 $ar{\left(ar{\Box} + \Delta\mathcal{O}_{ ext{EFT}}
ight)\mathcal{G}}_{ ext{EFT}} \ = \ \delta$

A perturbative approach $\mathcal{G}_{EFT} = \overline{\mathcal{G}}_{ret} (1 + \Delta \mathcal{G} + \cdots)$ is justified if the secular effects are bounded

$$|\Delta \mathcal{G}| \ll 1 ext{ with } \Delta \mathcal{G} \sim \int \Delta \mathcal{O}_{ ext{EFT}} ar{\mathcal{G}}_{ ext{ret}}$$

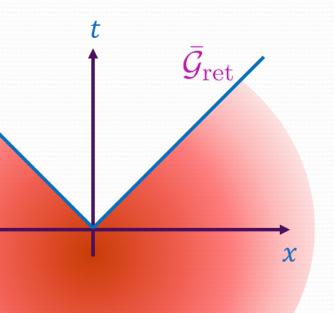
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There can be no support outside the light-cone



Note: in practice we can replace

$$\Delta T^{\mathrm{EFT}} \longleftrightarrow \partial_{\omega} \delta^{\mathrm{EFT}}(\omega)$$

$$|\Delta T^{\mathrm{EFT}}| \ll \omega^{-1} \iff |\delta^{\mathrm{EFT}}| \ll 1$$

Living with Superluminality

- Gravitational Waves are luminal to a (VERY) good accuracy at LIGO frequencies $-\mathcal{O}\left(10^{-15}\right) < c_T 1 < \mathcal{O}\left(10^{-16}\right)$
- Within the standard EFT of gravity, GWs are no longer perfectly luminal on backgrounds that spontaneously break Lorentz invariance (eg Schwarzschild, FLRW, the real world,...)

Lesson 1:

- In an arbitrary frame, GWs may be superluminal
- Imposing subluminality priors only makes sense in a frame where gravity can be decoupled
- In the original frame this may correspond to GWs being superluminal by a 'large' amount (not suppressed by Mpl⁻²)

Living with Superluminality

Lesson 2:

- Even in the frame where matter and gravity can decouple,
 a tiny amount of SL or a negative phase shift be it for GWs or
 other fields is not in conflict with causality.
 It may even follow from consistent causal and Lorentz invariant
 UV completions.
- In the frame where matter and gravity can decouple, superluminality is consistent with causality so long as

$$\lim_{M_{
m Pl} o \infty} |c_s^2 - 1| \sim M_{
m Pl}^{-lpha} \quad ext{with} \quad lpha \geq 2$$

Living with Negativity

Lesson 3: Conjecture

• For a 2 - 2 scattering amplitude of the form

$$A(s,t) \sim -\frac{s^2}{M_{\rm Pl}^2 t} + \frac{c}{M^4} s^2 + \dots$$

c needs not be positive so long as

$$c > -rac{M^2}{M_{
m Pl}^2} imes {\cal O}(1)$$
 Not assuming specific UV behavior

Amount of "positivity"-violation directly connected to "allowed" amount of SL