

# ENTANGLEMENT IN HOLOGRAPHY

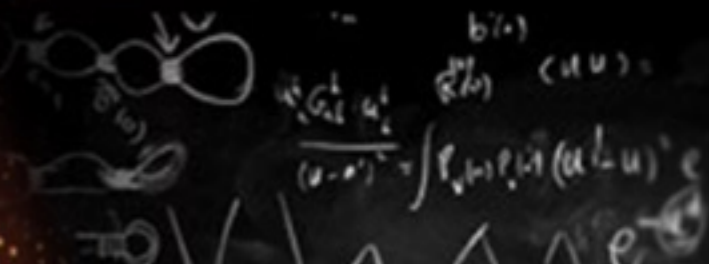
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Durham University

June 1, 2015

**Closing the entanglement gap:**  
Quantum information, quantum matter, and quantum fields



The Kavli Institute for  
Theoretical Physics  
University of California, Santa Barbara



# Entanglement

- Most non-classical manifestation of quantum mechanics
  - “Best possible knowledge of a whole does not include best possible knowledge of its parts — and this is what keeps coming back to haunt us” [Schrodinger '35]
- New quantum resource for tasks which cannot be performed using classical resources [Bennet '98]
- Plays a central role in wide-ranging fields
  - quantum information (e.g. cryptography, teleportation, ...)
  - quantum many body systems
  - quantum field theory
- Hints at profound connections to geometry...

# Entanglement Entropy (EE)

Suppose we only have access to a subsystem  $A$  of the full system  $= A + B$ . The amount of entanglement is characterized by Entanglement Entropy  $S_A$ :

- reduced density matrix  $\rho_A = \text{Tr}_B |\psi\rangle\langle\psi|$   
(more generally, for a mixed total state,  $\rho_A = \text{Tr}_B \rho$ )
- EE = von Neumann entropy  $S_A = -\text{Tr} \rho_A \log \rho_A$

Defined if we can divide a quantum system into a subsystem  $A$  and its complement  $B$ , such that the Hilbert space decomposes:

$$\mathcal{H} = \mathcal{H}_A \otimes \mathcal{H}_B$$



# Entanglement Entropy (EE)

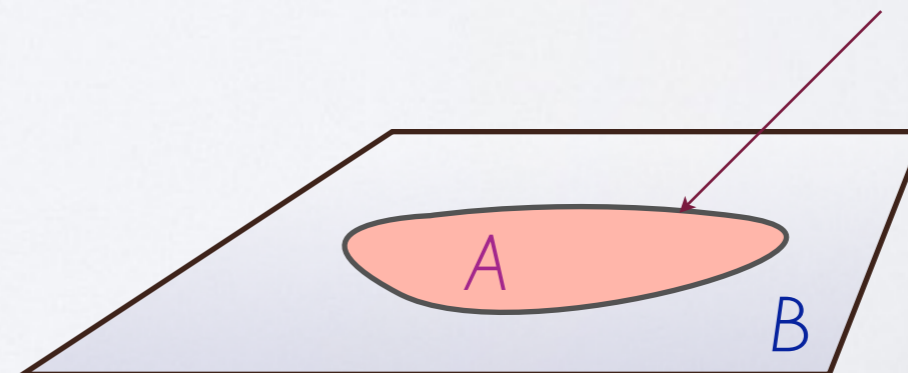
Suppose we only have access to a subsystem  $A$  of the full system  $= A + B$ . The amount of entanglement is characterized by **Entanglement Entropy  $S_A$** :

- reduced density matrix  $\rho_A = \text{Tr}_B |\psi\rangle\langle\psi|$   
(more generally, for a mixed total state,  $\rho_A = \text{Tr}_B \rho$ )

- **EE** = von Neumann entropy  $S_A = -\text{Tr} \rho_A \log \rho_A$

- e.g. in local QFT:

$A$  and  $B$  can be spatial regions, separated by a smooth entangling surface



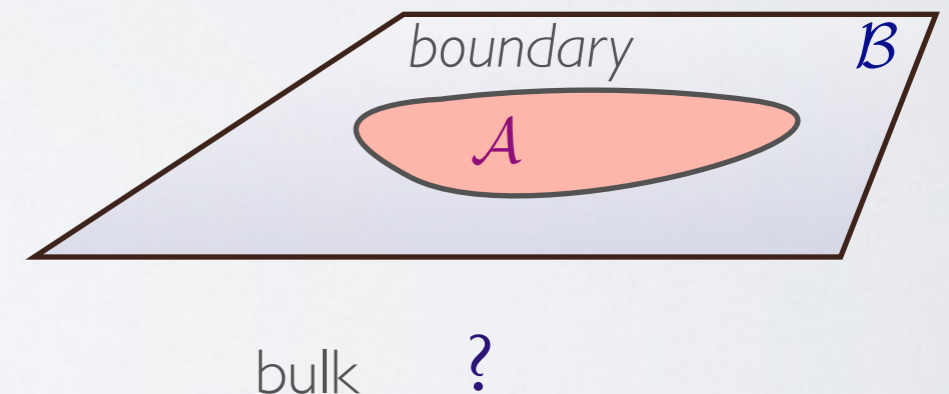


# The good news & the bad news

- **But** EE is hard to deal with...
  - non-local quantity, intricate & sensitive to environment
  - difficult to measure
  - difficult to calculate... especially in strongly-coupled quantum systems

- **AdS/CFT to the rescue?**

- ~ Is there a natural bulk dual of EE?  
(= “Holographic EE”)



Yes! - described geometrically...

# OUTLINE

- Entanglement ✓
- Holography
  - AdS/CFT Correspondence
  - Essential elements of the AdS/CFT dictionary
- Holographic Entanglement Entropy
  - RT & HRT prescriptions
- Utility of Geometry
  - Easy-to-prove properties of EE
  - Entanglement plateaux
  - Dynamics (example: quench & thermalization)
  - Curious features of EE
- Summary & Outlook

# AdS/CFT correspondence

String theory ( $\ni$  gravity)  $\iff$  gauge theory (CFT)

“in bulk” asymp.  $\text{AdS} \times S$

“on boundary”

[Maldacena, '97]

‘soup can’ diagram of AdS:



here label is everything...

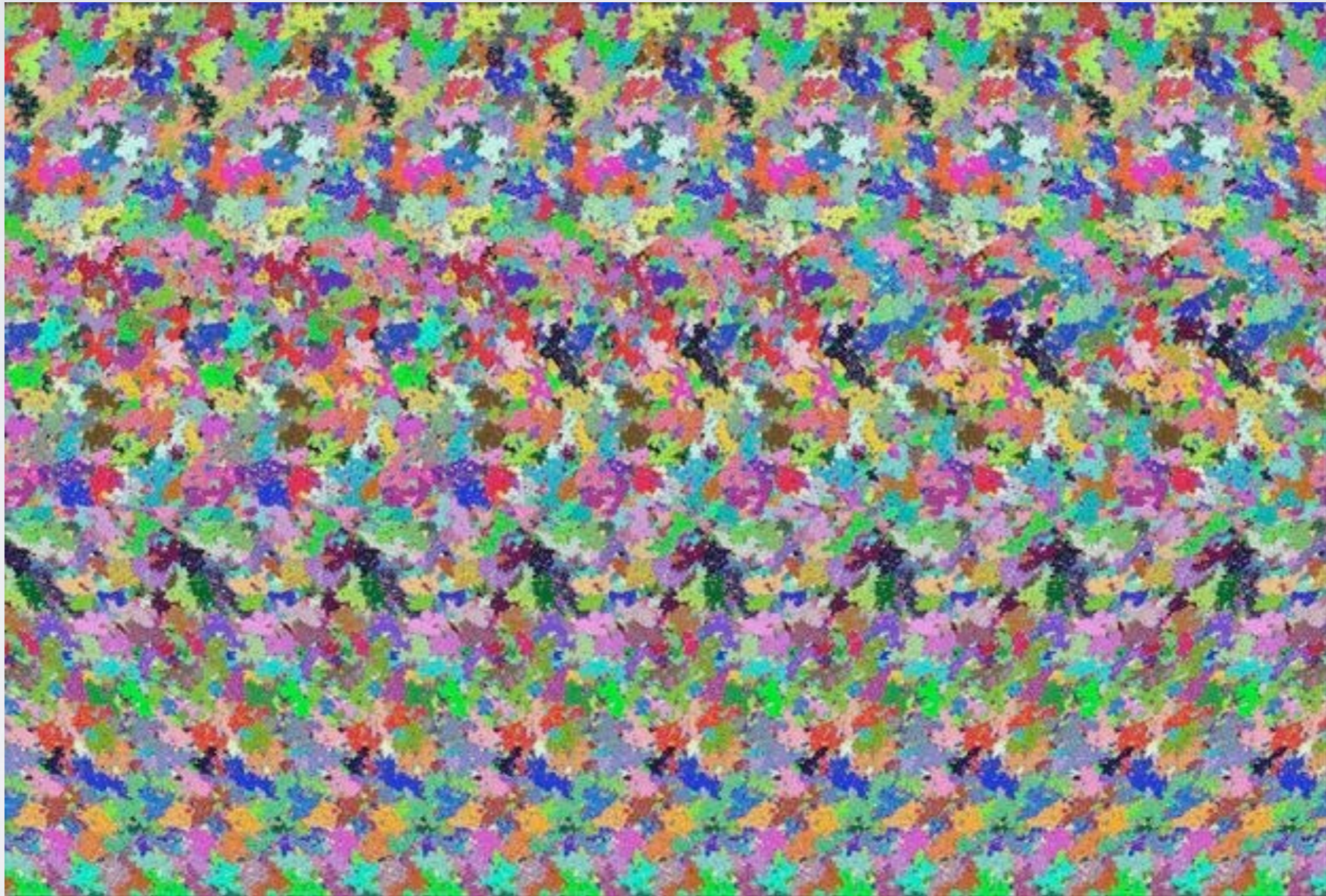
## Key aspects:

- \* Gravitational theory maps to non-gravitational one!
- \* *Holographic*: gauge theory lives in fewer dimensions.



# AdS/CFT correspondence

\* better analogy: stereogram...



...but infinitely more complicated



# AdS/CFT correspondence

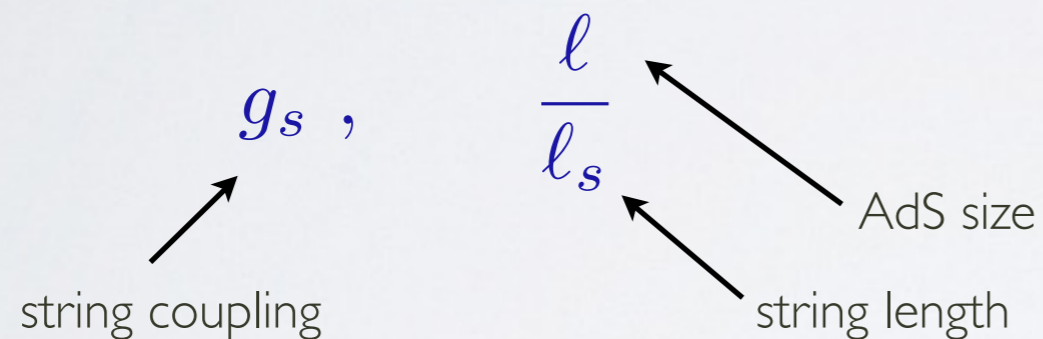
String theory ( $\ni$  gravity)  $\iff$  gauge theory (CFT)

“in bulk” asymp.  $\text{AdS} \times \text{K}$

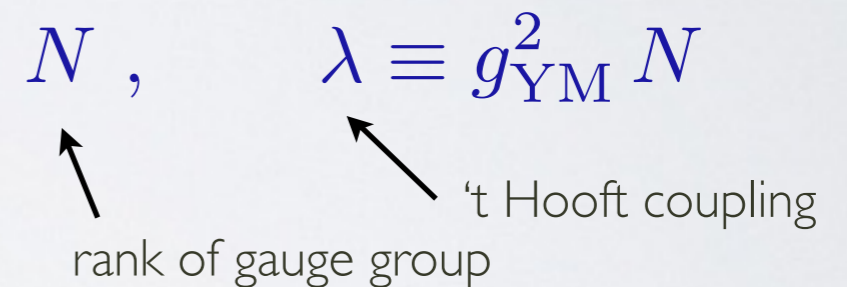
“on boundary”

## Specific example:

IIB string theory on  $\text{AdS}_5 \times S^5$ :



$\mathcal{N} = 4$   $\text{SU}(N)$  SYM:



where  $g_s \sim \frac{\lambda}{N}$  and  $\frac{\ell}{\ell_s} \sim \lambda^{1/4}$

- large  $\lambda \implies$  small stringy corrections
- large  $N \implies$  small quantum corrections
- Hence  $N \gg \lambda \gg 1 \implies$  classical gravity on  $\text{AdS}_5 \times S^5$

# AdS/CFT correspondence

String theory ( $\ni$  gravity)  $\iff$  gauge theory (CFT)

*“in bulk”* asymp. AdS  $\times$  K

*“on boundary”*

## Key aspects:

- \* Gravitational theory maps to non-gravitational one!
- \* *Holographic*: gauge theory lives in fewer dimensions.
- \* Strong/weak coupling duality.

## Invaluable tool to:

- ~ Use gravity on AdS to learn about strongly coupled field theory  
(as successfully implemented in e.g. AdS/QCD & AdS/CMT programs)
- ~ Use the gauge theory to define & study quantum gravity in AdS

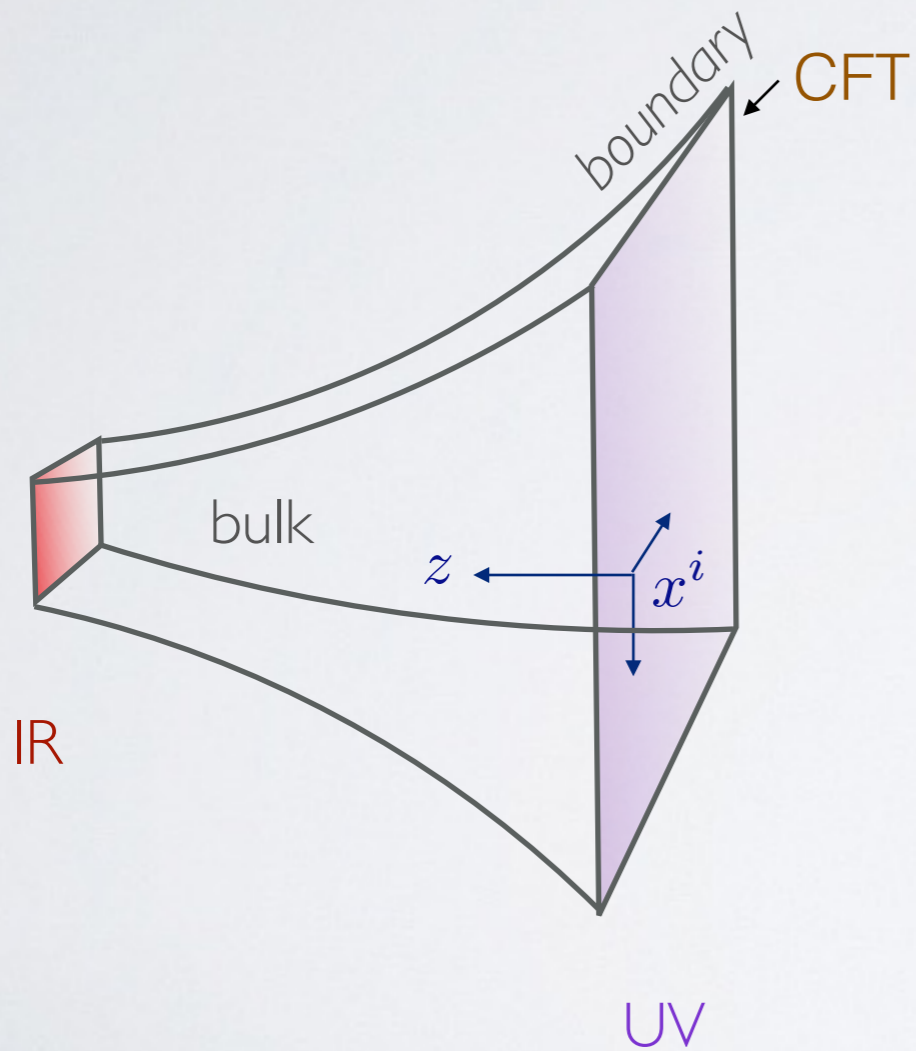
**Pre-requisite:** Understand the AdS/CFT ‘dictionary’...



# Geometry of AdS

Poincare AdS:

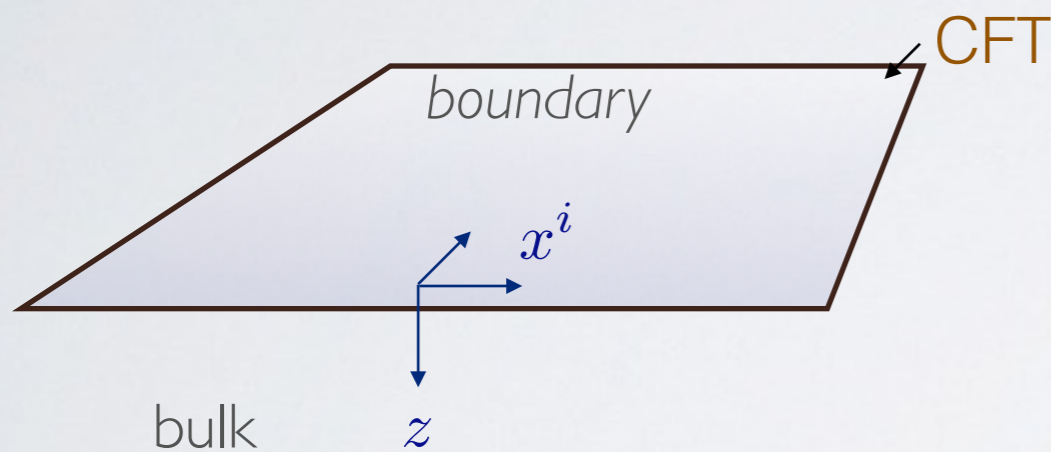
$$ds^2 = \frac{\ell^2}{z^2} (-dt^2 + dx_i dx^i + dz^2)$$



# Geometry of AdS

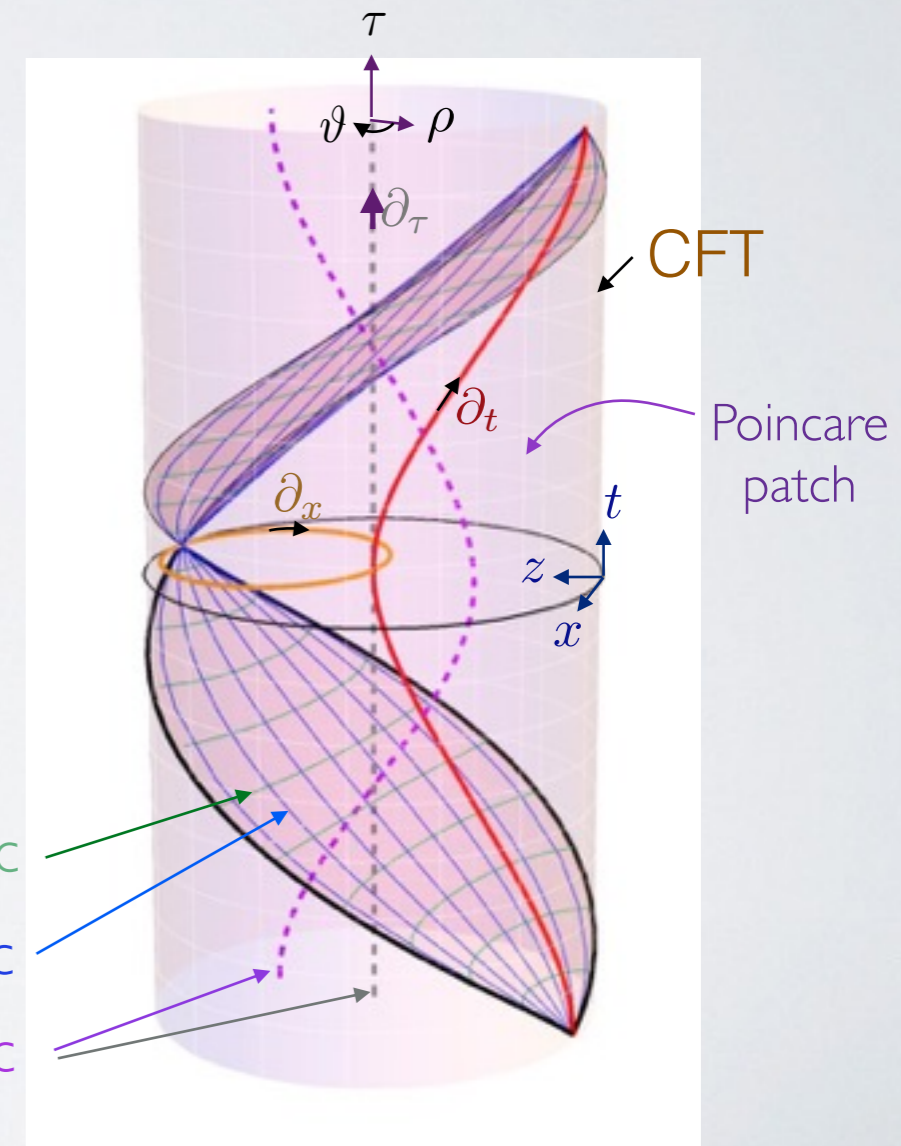
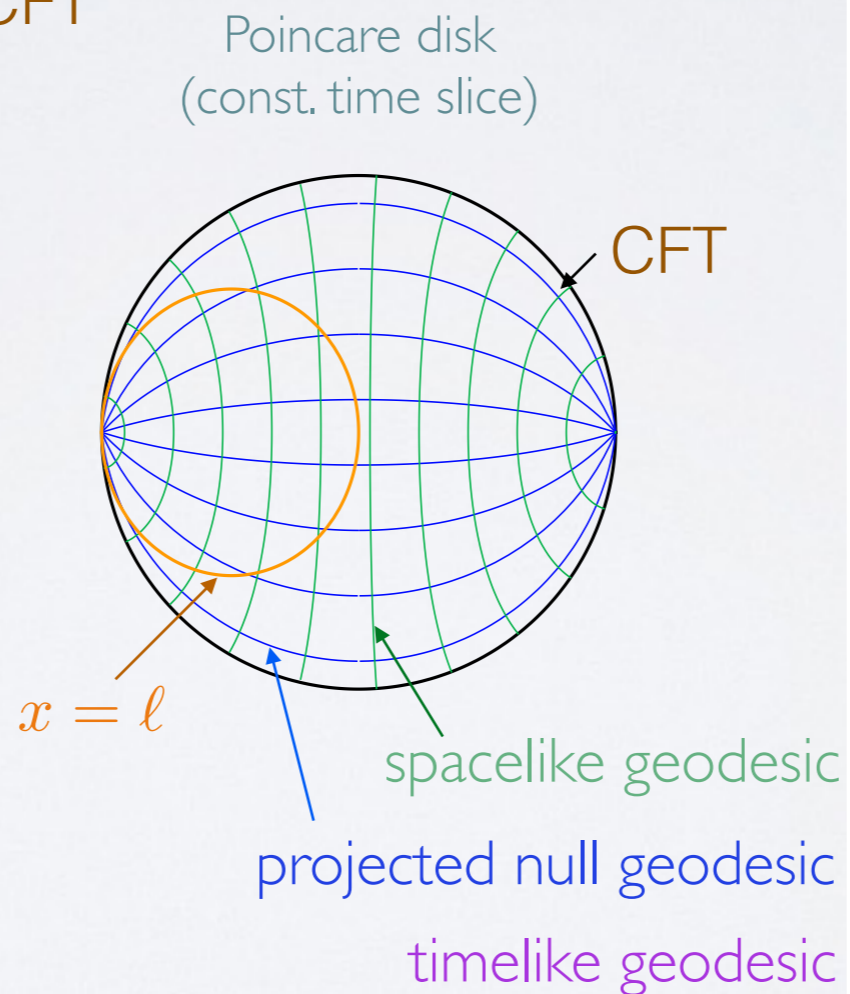
Poincare AdS:

$$ds^2 = \frac{\ell^2}{z^2} (-dt^2 + dx_i dx^i + dz^2)$$



Global AdS:

$$ds^2 = - \left( \frac{\rho^2}{\ell^2} + 1 \right) d\tau^2 + \frac{d\rho^2}{\left( \frac{\rho^2}{\ell^2} + 1 \right)} + \rho^2 d\Omega_3^2$$



# Scale/radius duality

What CFT quantity encodes the extra bulk direction?

- Scale/radius (or UV/IR) duality:

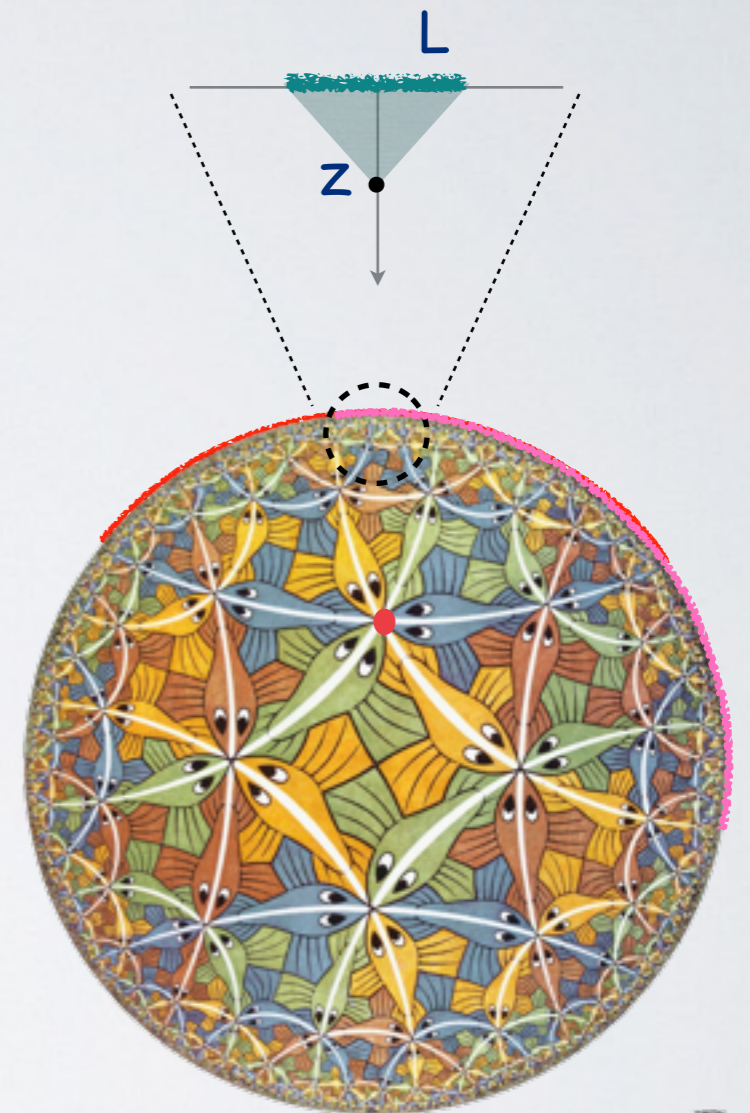
- UV (small scale) in CFT  $\leftrightarrow$  IR (large radius) in AdS
- Local bulk excitation at radial position  $z$  in AdS is manifested by CFT excitation at scale  $L \sim z$ .

[Susskind & Witten]

- Follows from AdS geometry...

- Provides useful intuition: e.g. object falling into a black hole  $\leftrightarrow$  CFT excitation spreads & thermalizes [Banks, Douglas, Horowitz, Martinec]

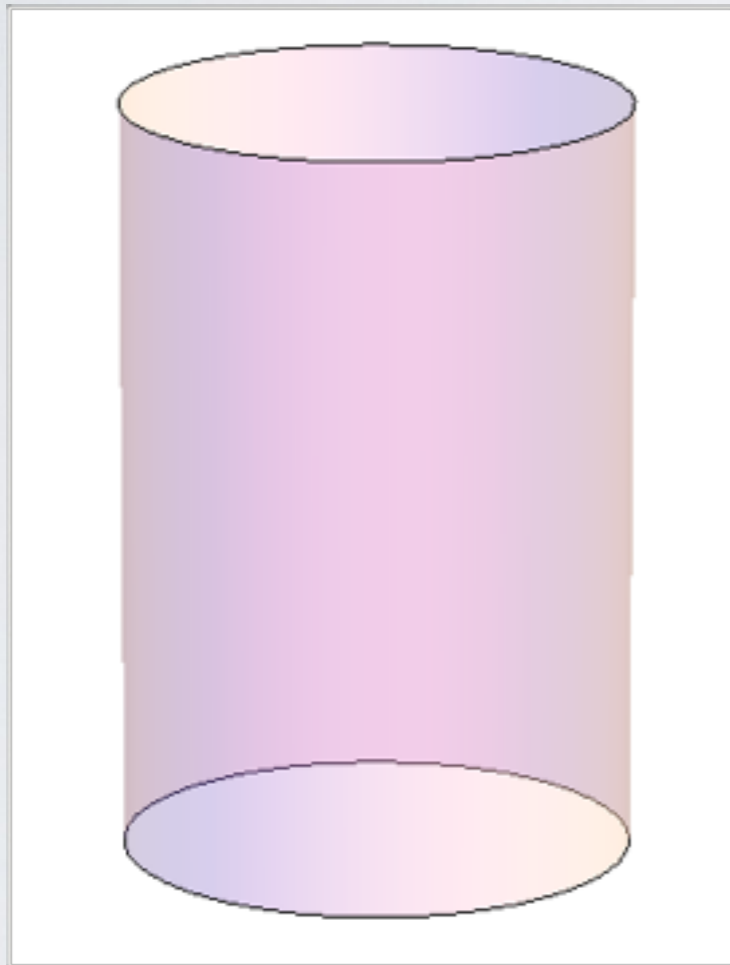
- Asymptotic fall-off of bulk fields  $\leftrightarrow$  Expectation values of local gauge-invariant operators in CFT





# Bulk geometries and CFT states

different bulk geometries  $\leftrightarrow$  different states in CFT  
(asymptotically AdS)

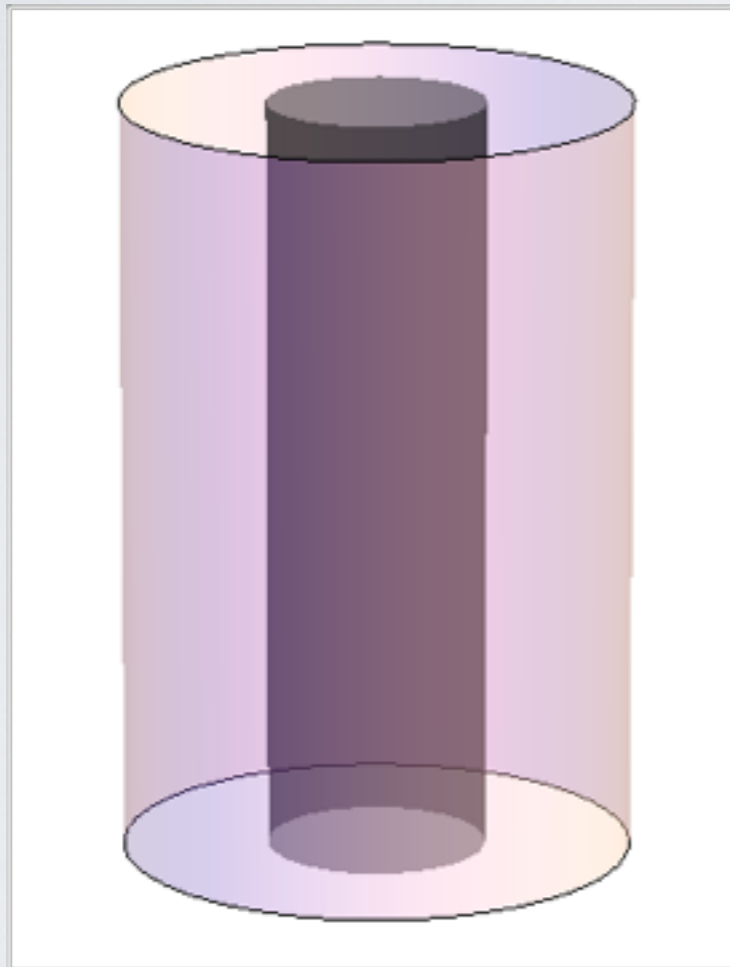


- Pure AdS  $\leftrightarrow$  vacuum state in CFT

Finite-mass deformations  
of the bulk geometry result in  
non-zero boundary stress tensor

# Bulk geometries and CFT states

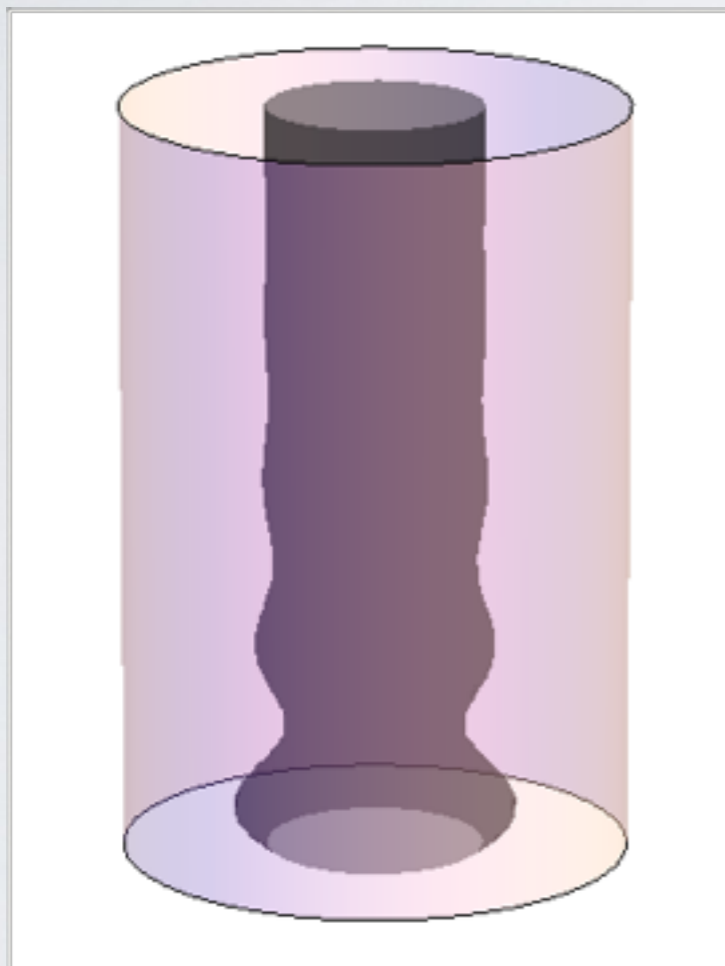
different bulk geometries  $\leftrightarrow$  different states in CFT  
(asymptotically AdS)



- Pure AdS  $\leftrightarrow$  vacuum state in CFT
- Black hole  $\leftrightarrow$  thermal state in CFT  
(large BH  $\leftrightarrow$  high temperature)

# Bulk geometries and CFT states

evolving bulk geometries  $\leftrightarrow$  corresponding dynamics



- Pure AdS  $\leftrightarrow$  vacuum state in CFT
- Black hole  $\leftrightarrow$  thermal state in CFT

- quasinormal modes of perturbed black hole  $\leftrightarrow$  approach to thermal equilibrium  
[Horowitz & VH]

- horizon response properties  $\leftrightarrow$  CFT transport coefficients  
[Kovtun, Son, Starinets]

- at non-linear level, in hydro regime (large BHs)  $\rightarrow$  fluid/gravity correspondence  
[Bhattacharyya, VH, Minwalla, Rangamani]



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  - RT & HRT prescriptions
- Utility of Geometry
  - Easy-to-prove properties of EE
  - Entanglement plateaux
  - Dynamics (example: quench & thermalization)
  - Curious features of EE
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# Paths to Holographic EE

String theory ( $\ni$  gravity)  $\iff$  gauge theory (CFT)

*“in bulk”* asymp.  $\text{AdS} \times K$

*“on boundary”*

Applied AdS/CFT:

- study specific system via its dual
- e.g. AdS/QCD, AdS/CMT, ...

Fundamentals of AdS/CFT:

- why/how does the duality work
- map between the 2 sides

Holographic Entanglement Entropy

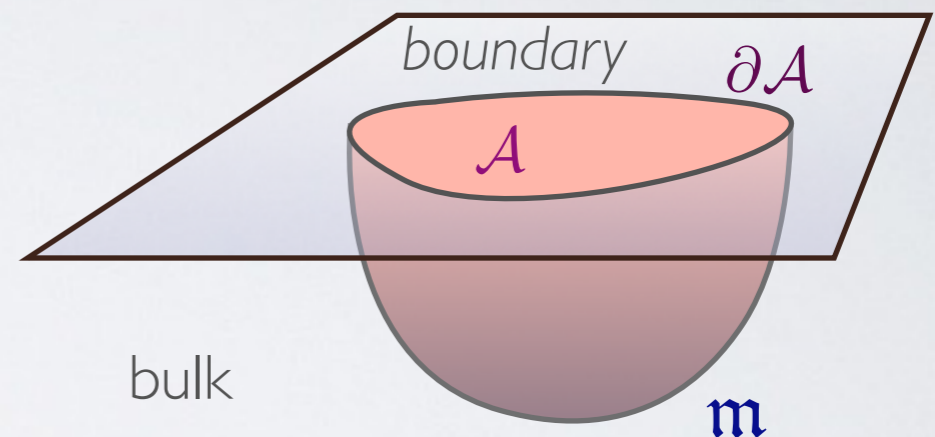
Quantum Gravity

# Holographic Entanglement Entropy

Proposal [RT = Ryu & Takayanagi, '06] for *static* configurations:

In the bulk, EE  $S_{\mathcal{A}}$  is captured by the area of minimal co-dimension 2 bulk surface  $\mathfrak{m}$  (at constant  $t$ ) anchored on  $\partial\mathcal{A}$ .

$$S_{\mathcal{A}} = \min_{\partial\mathfrak{m}=\partial\mathcal{A}} \frac{\text{Area}(\mathfrak{m})}{4G_N}$$



Remarks:

- Large body of evidence, culminating in [Lewkowycz & Maldacena]
- cf. black hole entropy...
- Minimal surface “hangs” into the bulk due to large distances near bdy.
- Note that both LHS and RHS are in fact infinite...

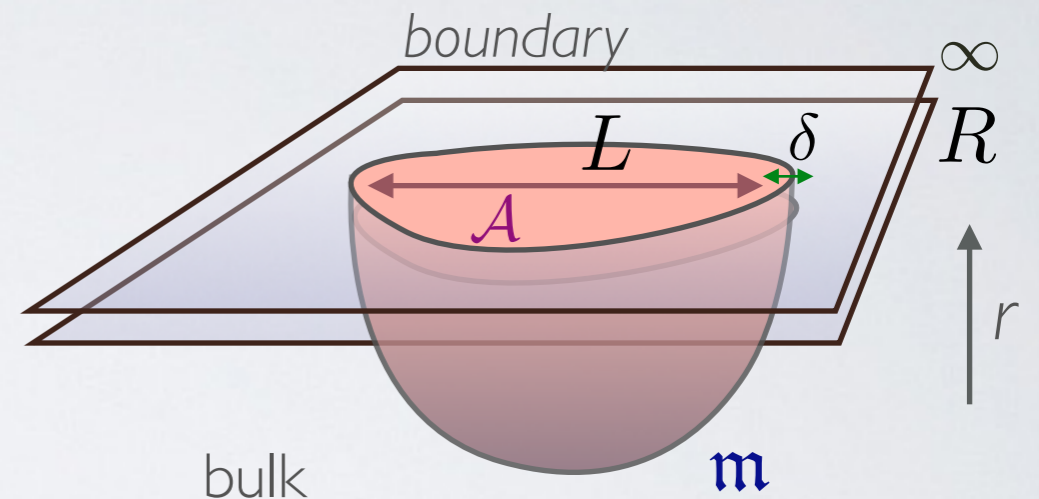


# Area-law divergence of HEE

Short-distance cutoff  $\delta$  in the CFT translates to large-radius cutoff  $R$  in  $\text{AdS}_{d+1}$

with  $\delta = \frac{\ell^2}{R}$  (cf. UV/IR duality)

Bulk area reproduces the correct divergence structure:



$$S_{\mathcal{A}} = c_0 \left(\frac{L}{\delta}\right)^{d-2} + c_1 \left(\frac{L}{\delta}\right)^{d-4} + \dots$$

↑ cutoff-dependent coefficients     
 ↑ universal coefficients

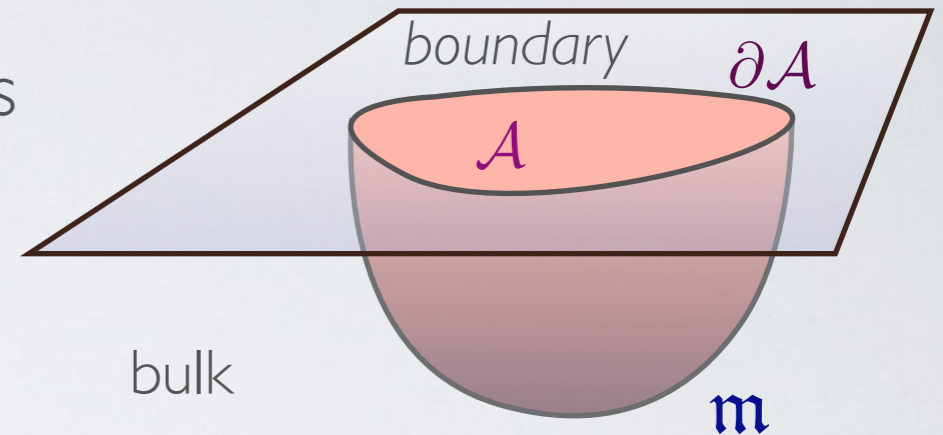
$$+ \begin{cases} c_{d-2} \log\left(\frac{L}{\delta}\right) + \dots & , & d \text{ even} \\ c_{d-2} + \dots & , & d \text{ odd} \end{cases}$$

We can regulate EE by e.g. background subtraction.

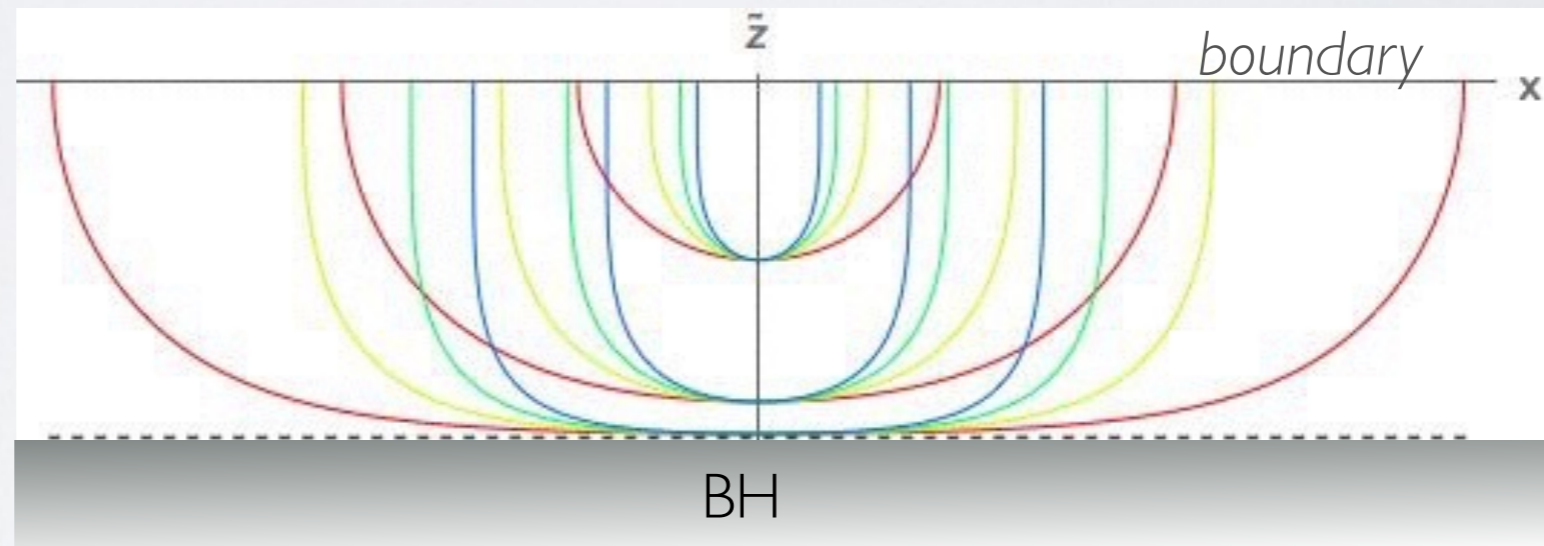
# Area-law vs. Volume-law

The leading divergence of  $S_A$  necessarily scales with the area of the entangling surface  $\partial A$

→ Area-law

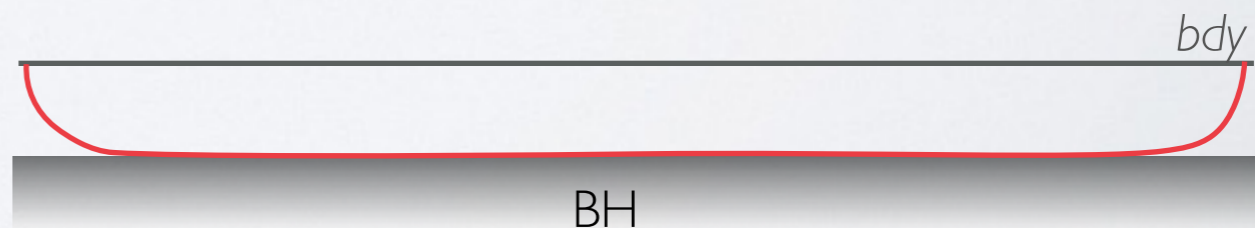


But in the presence of a (static) black hole, extremal surface get repelled by the horizon.



At high temperature ( $T L \gg 1$ ), the finite part of  $S_A$  scales with the volume of the region  $A$

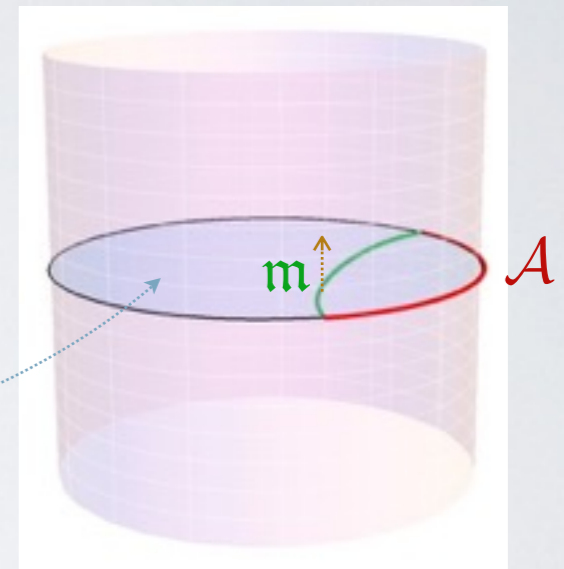
→ Volume-law



# Covariant Holographic EE

But the RT prescription is not well-defined outside the context of static configurations:

- In Lorentzian geometry, we can decrease the area arbitrarily by timelike deformations
- In time-dependent context, no natural notion of “const.  $t$ ” slice...



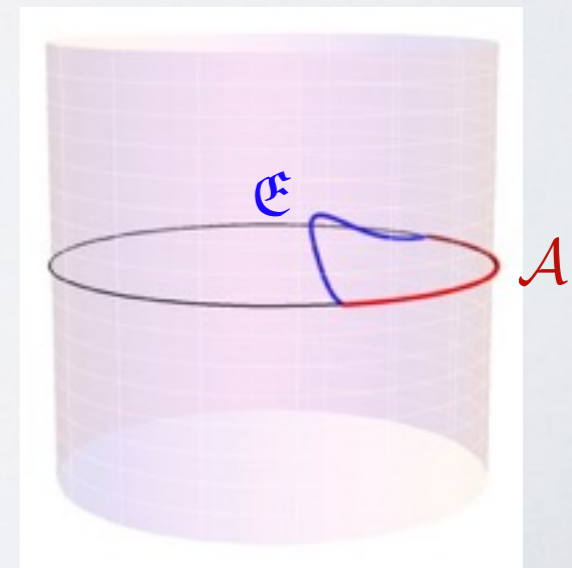
In *time-dependent* situations, RT prescription must be covariantized:

Simplest candidate:

minimal surface  $\mathfrak{m}$   
at constant time



extremal surface  $\mathfrak{E}$   
in the full bulk



[HRT = VH, Rangamani, Takayanagi '07]



# Covariant Holographic EE

HRT Prescription:

In the bulk EE  $S_{\mathcal{A}}$  is captured by the area of extremal co-dimension 2 bulk surface  $\mathcal{E}$  anchored on  $\partial\mathcal{A}$  & homologous to  $\mathcal{A}$

$$S_{\mathcal{A}} = \min_{\partial\mathcal{E}=\partial\mathcal{A}} \frac{\text{Area}(\mathcal{E})}{4G_N}$$

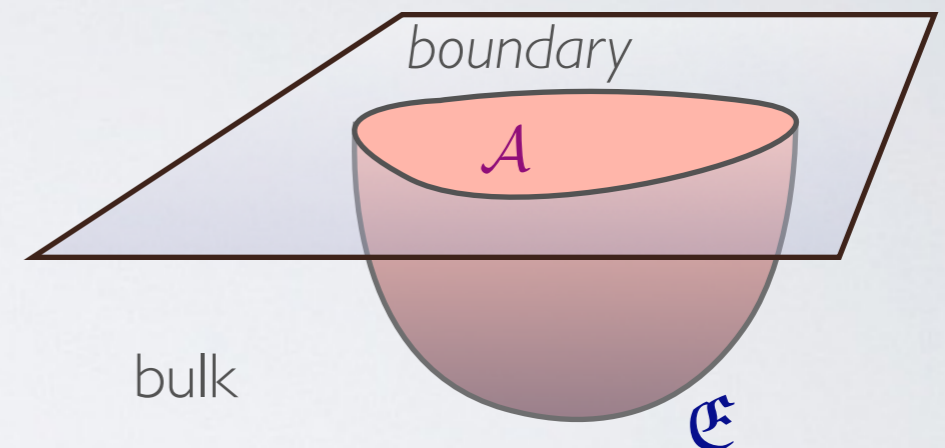
Equivalently:

- $\mathcal{E}$  is the surface with zero null expansions; (cf. light sheet construction [Bousso] )
- maximin construction: maximize over minimal-area surface on a spacelike slice [Wall]

This gives a well-defined quantity in any (arbitrarily time-dependent asymptotically AdS) spacetime  $\Rightarrow$  equally robust as in CFT

But we can't use Euclidean techniques / minimization for proofs...

[VH, Rangamani, Takayanagi]

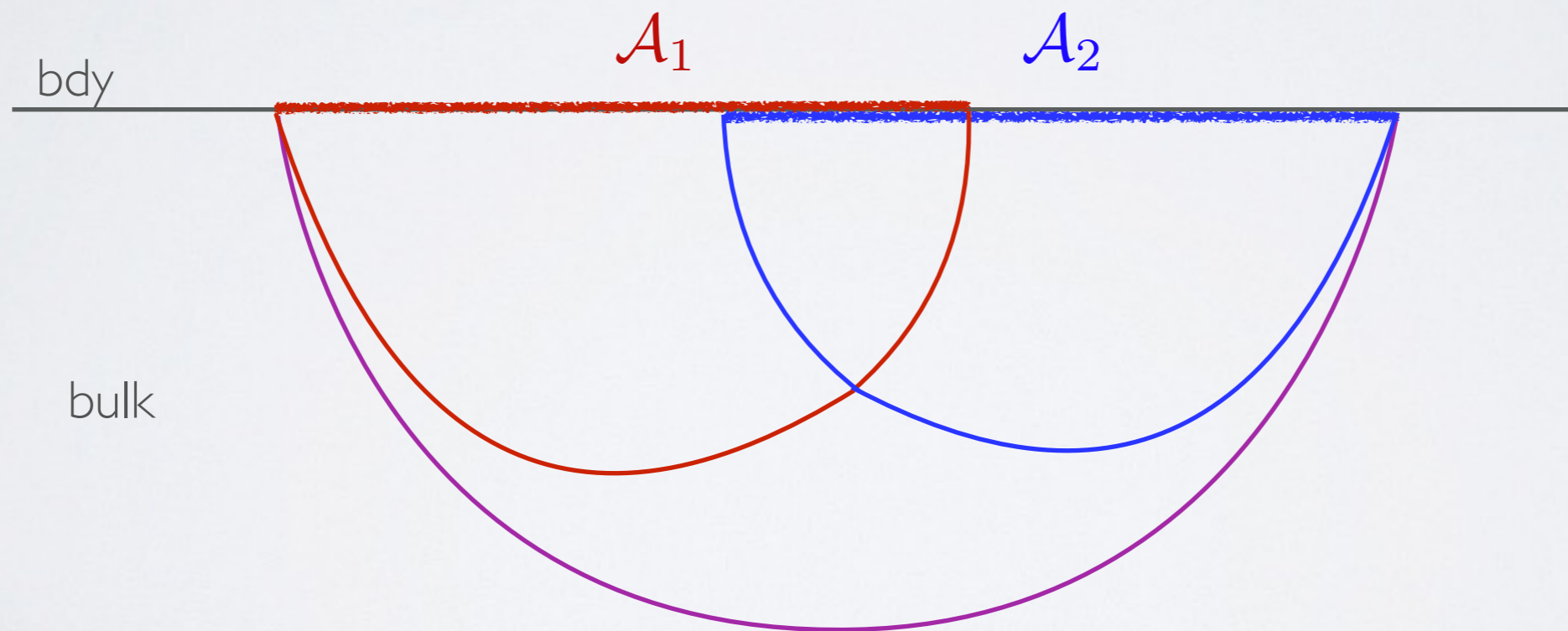
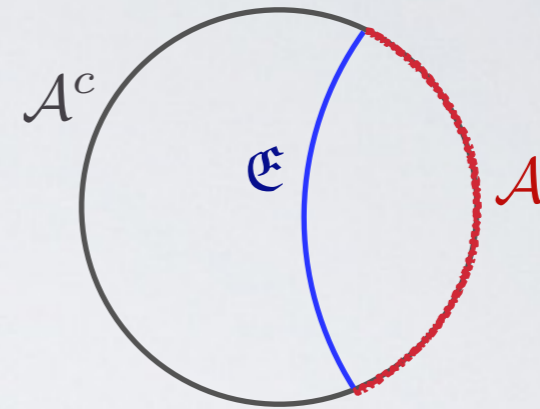


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- Entanglement ✓
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  - AdS/CFT Correspondence
  - Essential elements of the AdS/CFT dictionary
- Holographic Entanglement Entropy ✓
  - RT & HRT prescriptions
- **Utility of Geometry**
  - Easy-to-prove properties of EE
  - Entanglement plateaux
  - Dynamics (example: quench & thermalization)
  - Curious features of EE
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# Manifest properties of EE

- For pure states  $S_{\mathcal{A}} = S_{\mathcal{A}^c}$
- Positivity:  $S_{\mathcal{A}} \geq 0$
- Subadditivity:  $S_{\mathcal{A}_1} + S_{\mathcal{A}_2} \geq S_{\mathcal{A}_1 \cup \mathcal{A}_2}$



- Implies positivity of mutual information:  $I(\mathcal{A}_1, \mathcal{A}_2) = S_{\mathcal{A}_1} + S_{\mathcal{A}_2} - S_{\mathcal{A}_1 \cup \mathcal{A}_2}$

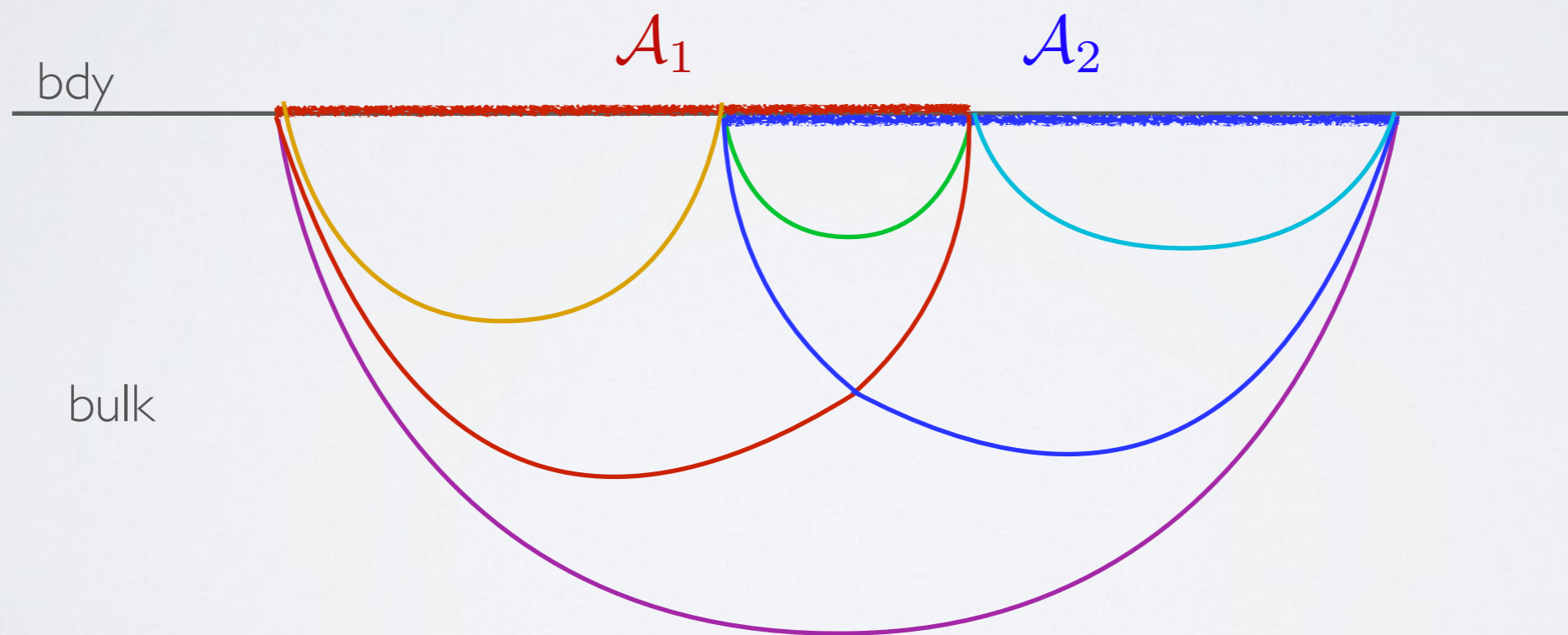


# Strong Subadditivity

- strong subadditivity:

$$S_{\mathcal{A}_1} + S_{\mathcal{A}_2} \geq S_{\mathcal{A}_1 \cup \mathcal{A}_2} + S_{\mathcal{A}_1 \cap \mathcal{A}_2}$$

$$S_{\mathcal{A}_1} + S_{\mathcal{A}_2} \geq S_{\mathcal{A}_1 \setminus \mathcal{A}_2} + S_{\mathcal{A}_2 \setminus \mathcal{A}_1}$$

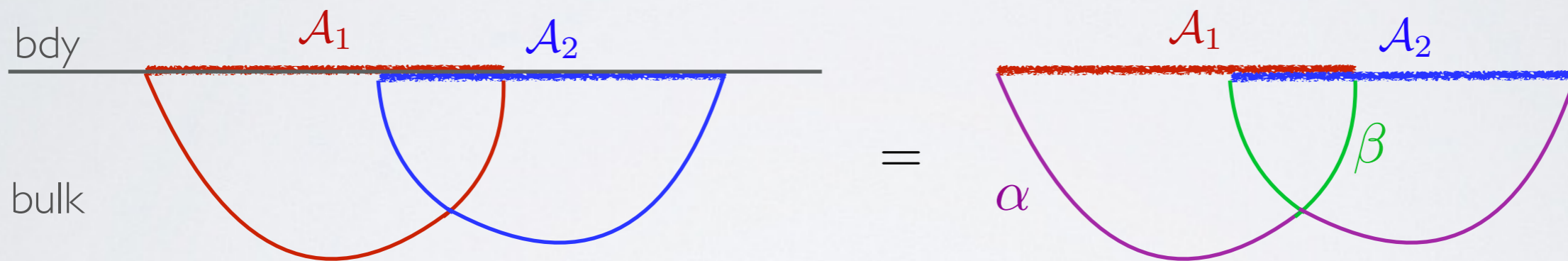


# Proof of Strong Subadditivity

- strong subadditivity:

$$S_{A_1} + S_{A_2} \geq S_{A_1 \cup A_2} + S_{A_1 \cap A_2}$$

- proof in static configurations [Headrick & Takayanagi]



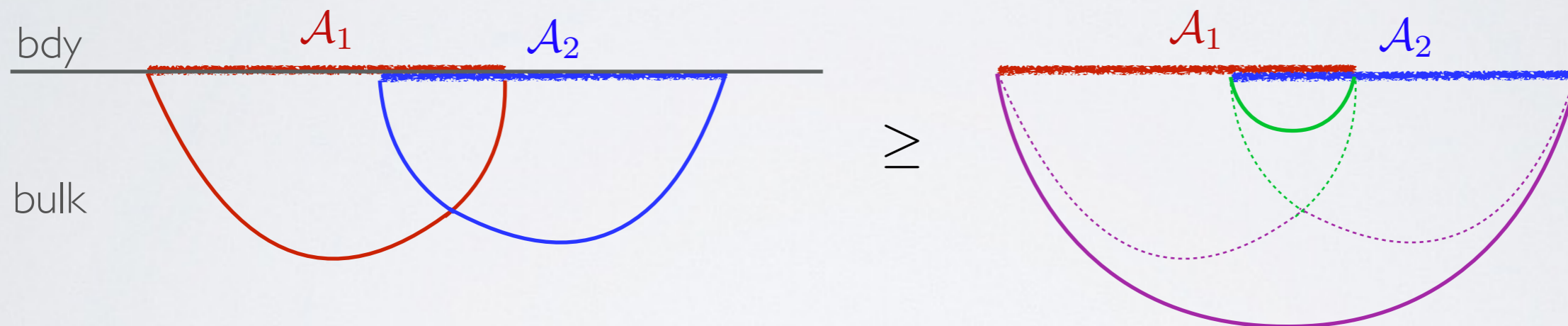
$$S_{A_1} + S_{A_2} = \alpha + \beta$$

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$$S_{A_1} + S_{A_2} \geq S_{A_1 \cup A_2} + S_{A_1 \cap A_2}$$

- proof in static configurations [Headrick & Takayanagi]



$$S_{A_1} + S_{A_2} = \alpha + \beta \geq S_{A_1 \cup A_2} + S_{A_1 \cap A_2}$$

- Similarly prove monogamy of mutual information [Hayden, Headrick, Maloney] valid in holography but not in general:  $S_A + S_B + S_C + S_{ABC} \leq S_{AB} + S_{BC} + S_{AC}$
- general proof uses properties of null geodesics + energy conditions [Wall]

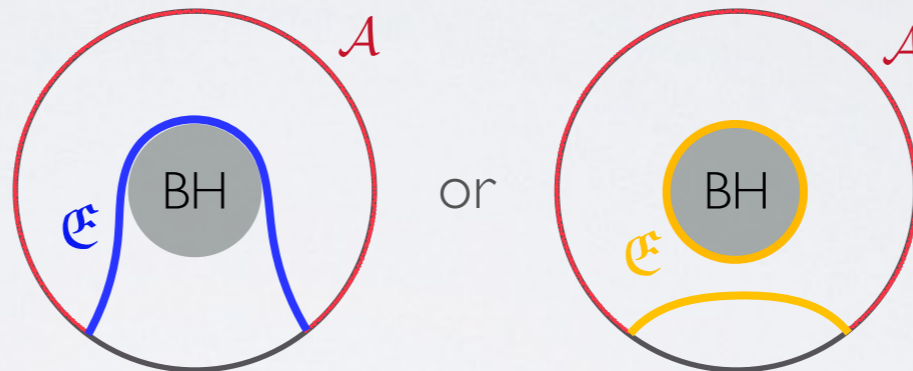


# Araki-Lieb inequality

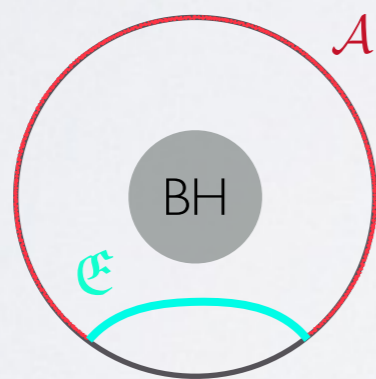
- for system in a mixed state (density matrix)  $\rho_\Sigma$  on spatial slice  $\Sigma = \mathcal{A} \cup \mathcal{A}^c$
- in general  $S_{\mathcal{A}} \neq S_{\mathcal{A}^c}$ , due to the homology constraint

[Headrick&Takayanagi]

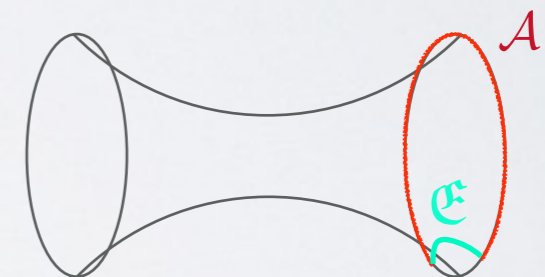
E.g. for thermal state on  $S^1$ , dual to eternal (BTZ) BH, both of the following satisfy the homology constraint:



but NOT



since  $\nexists$  interpolating region whose only boundaries are  $\mathcal{A}$  and  $\mathcal{E}$ :



- but EE satisfies: 
$$\underbrace{|S_{\mathcal{A}} - S_{\mathcal{A}^c}|}_{\delta S_{\mathcal{A}}} \leq S_{\rho_\Sigma} \leq S_{\mathcal{A}} + S_{\mathcal{A}^c}$$

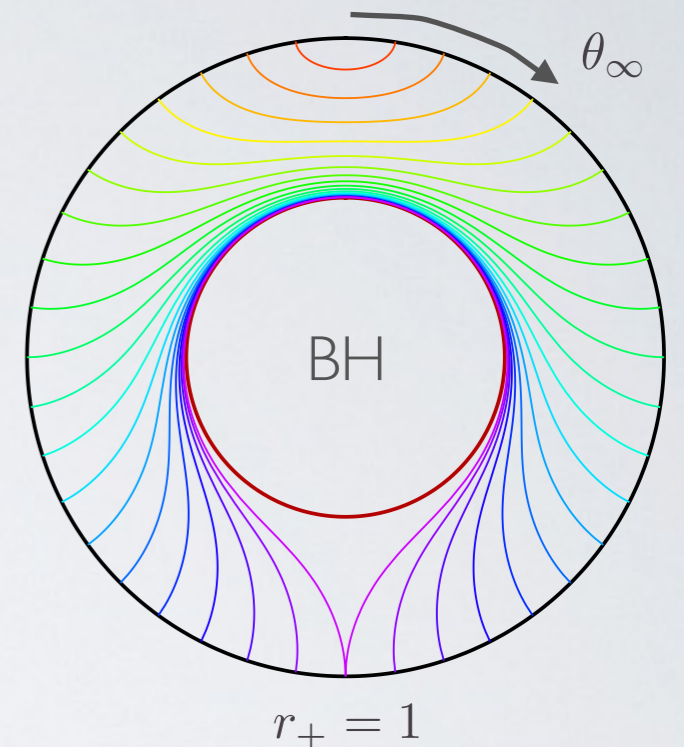
Araki-Lieb                      subadditivity ✓

# Entanglement plateaux

- Naively, Araki-Lieb appears in danger of being violated;  
E.g. for thermal state on compact space  $S^1$

Bulk dual: BTZ black hole:

$$(S_{\mathcal{A}})_{\text{naive}} = \frac{c}{3} \log \left( \frac{2r_{\infty}}{r_+} \sinh(r_+ \theta_{\infty}) \right)$$



- Resolution is supplied by the **minimality** condition:

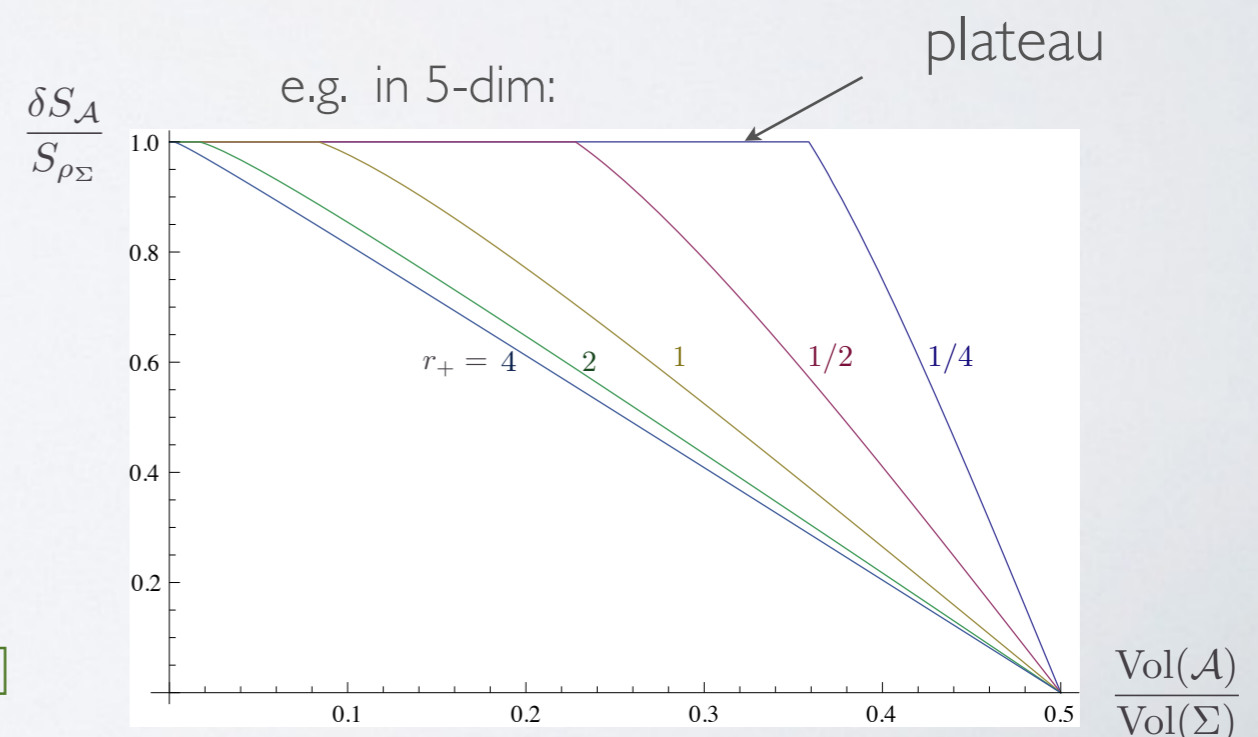
$$S_{\mathcal{A}} = \min_{\partial \mathcal{E} = \partial \mathcal{A}} \frac{\text{Area}(\mathcal{E})}{4G_N}$$

for  $\mathcal{A}$  large enough, the disjoint configuration dominates:

$$\Rightarrow S_{\mathcal{A}} = S_{\mathcal{A}^c} + S_{\text{BH}}$$

→ entanglement plateau

[VH, Maxfield, Rangamani, Tonni]



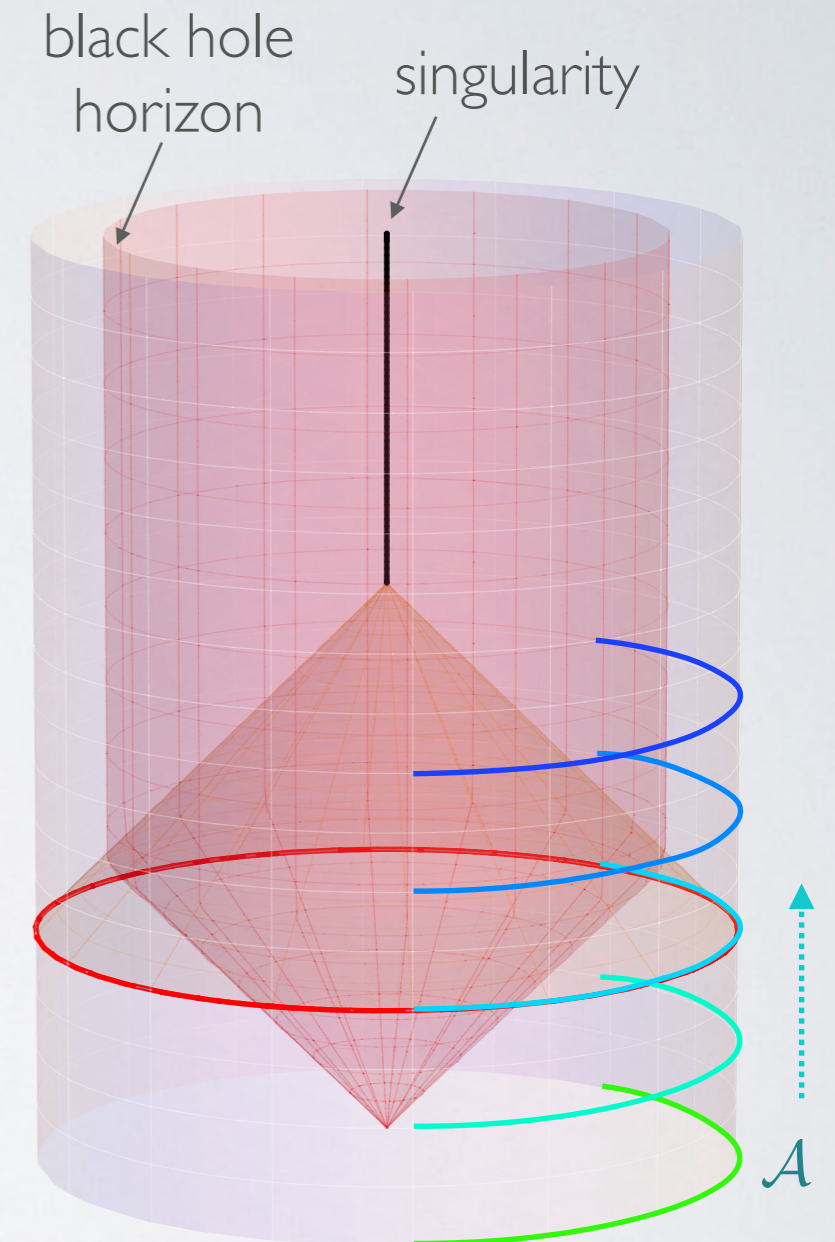
# Utility for dynamics

- For strongly time-dependent situations, holography is often the best tool presently available...
- E.g. consider quantum quench & thermalization
  - prepare a system in ground state of Hamiltonian  $H_0$
  - at  $t=0$  deform the Hamiltonian to a new Hamiltonian  $H$  by sudden change in some parameter
  - let the system evolve ( $\sim$  thermalize) with the new  $H$
- Global quench: change to  $H$  is homogeneous
  - in the bulk dual, we can model this by Vaidya-AdS corresponding to collapse of a null shell which forms a black hole.



# Building up Vaidya-AdS

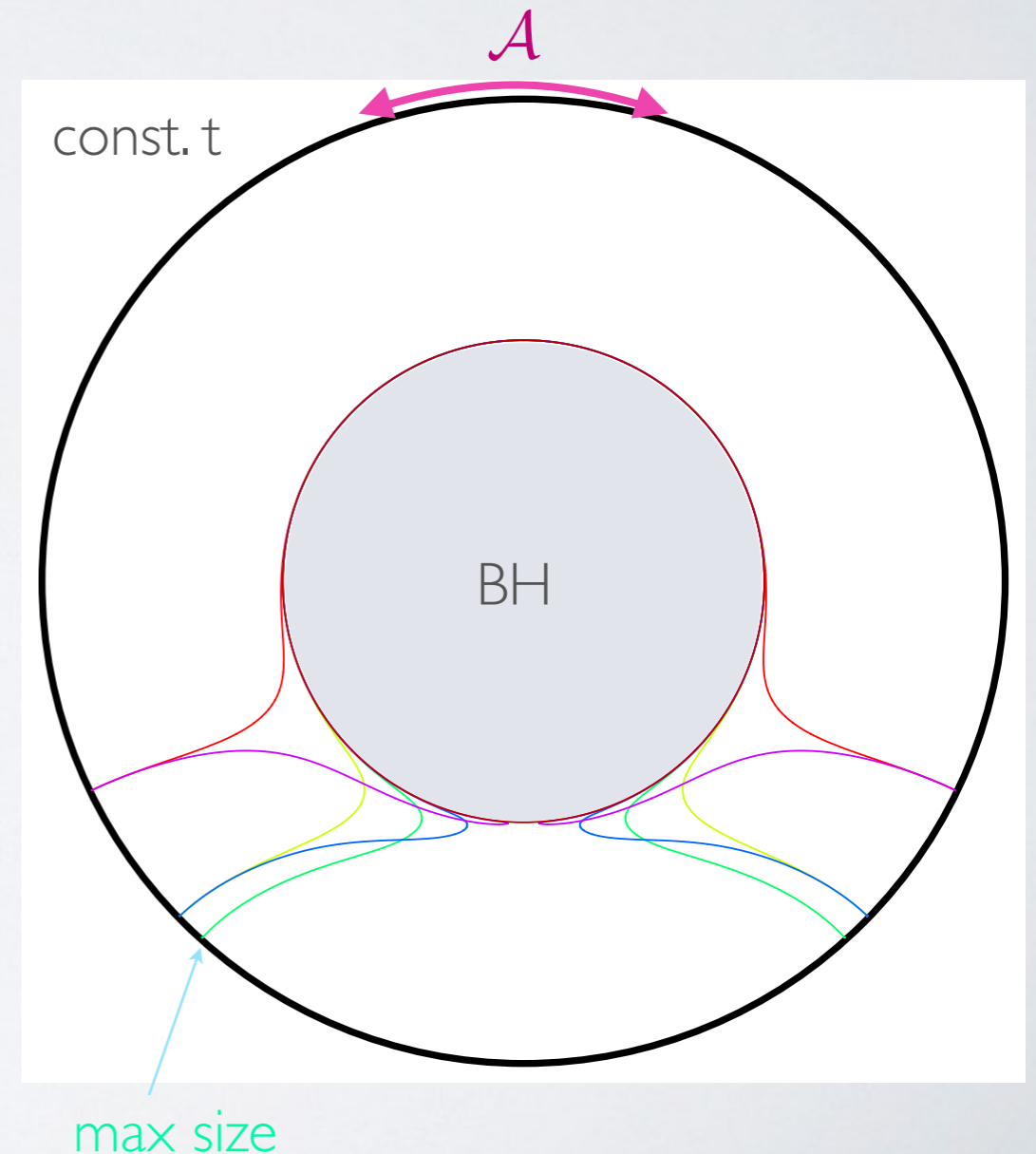
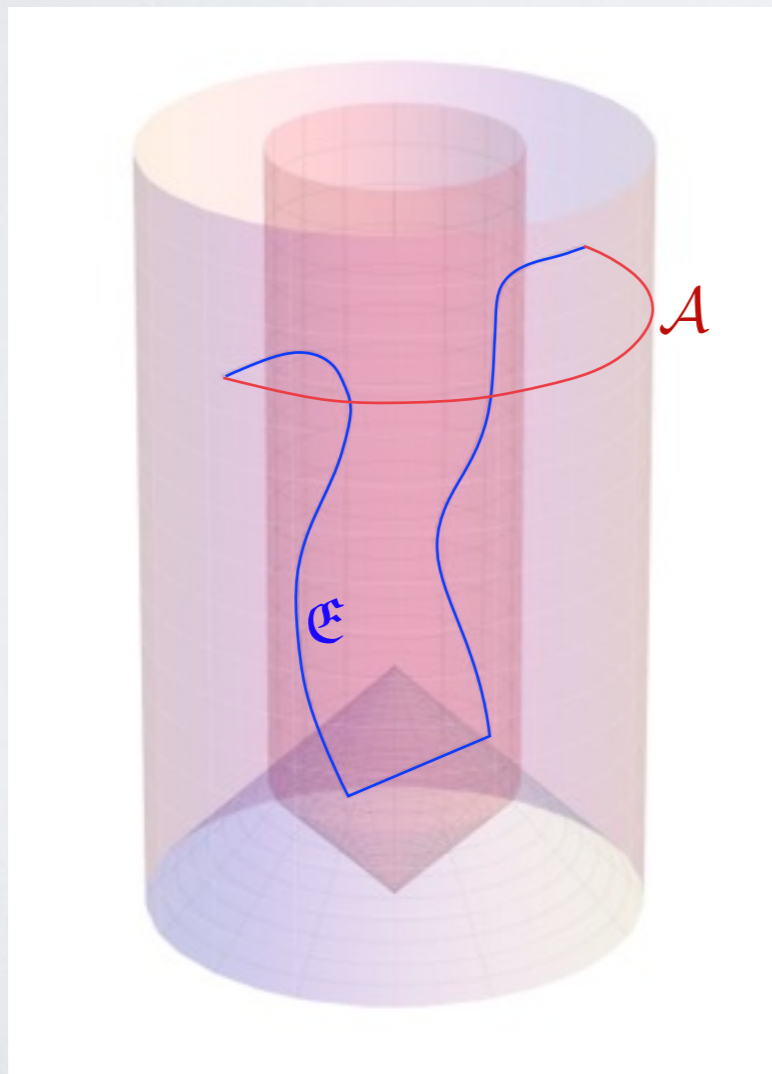
- start with vacuum state in CFT = pure AdS in bulk
- at  $t=0$ , create a short-duration disturbance in the CFT (global quench)
- this will excite a pulse of matter (shell) in AdS which implodes under evolution
- gravitational backreaction: collapse to a black hole  $\Rightarrow$  CFT 'thermalizes'
- large CFT energy  $\Rightarrow$  large BH



- To study thermalization of  $S_{\mathcal{A}}$ , fix spatial extent of  $\mathcal{A}$  and vary the time...

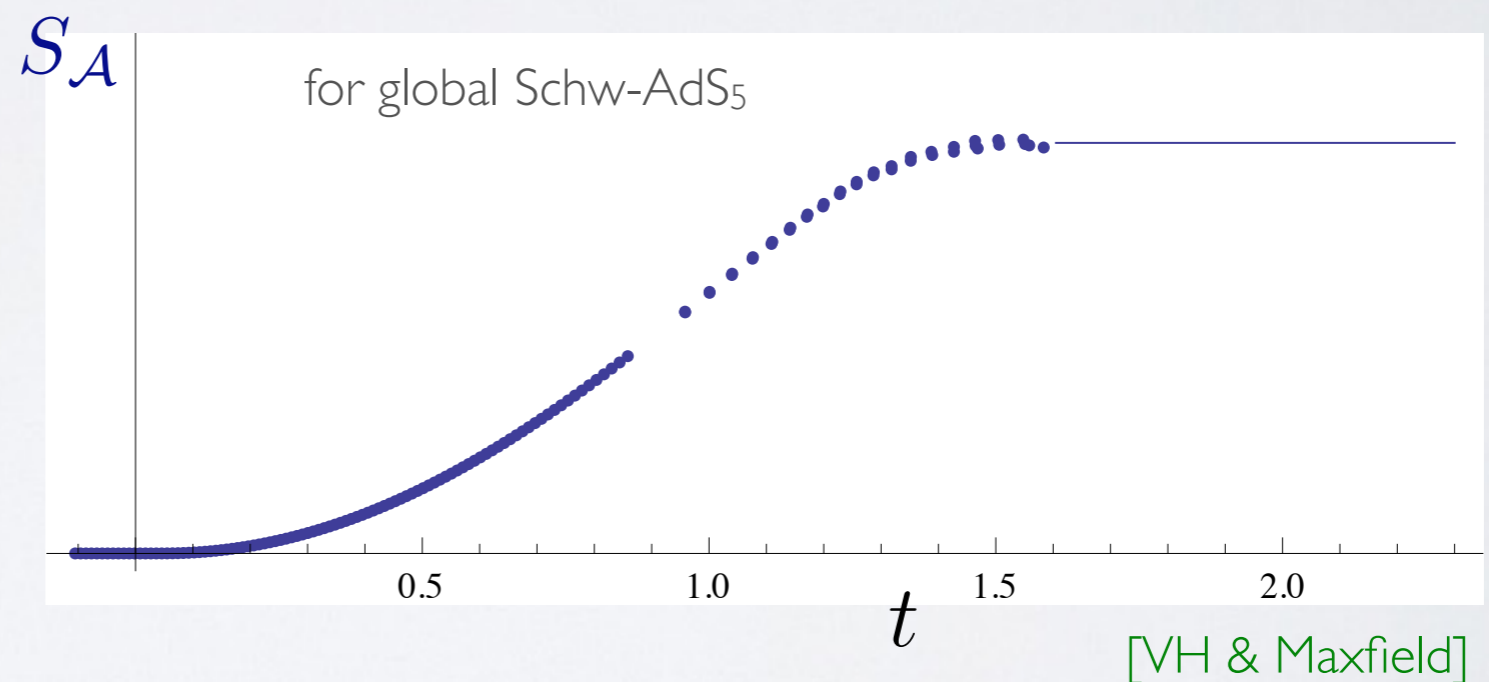
# Thermalization

- Results:  $\exists$  rather exotic extremal surfaces for  $d > 3$ , e.g. ones
  - penetrating the BH from arbitrarily late boundary time
  - having arbitrarily many 'folds' (even for the static BH)



# Thermalization

- Results:  $\exists$  rather exotic extremal surfaces for  $d > 3$ , e.g. ones
  - penetrating the BH from arbitrarily late boundary time
  - having arbitrarily many ‘folds’ (even for the static BH)
- nevertheless, minimality seems to ensure that  $S_{\mathcal{A}}$  ‘thermalizes’ continuously and monotonically:



- cf. ‘entanglement tsunami’ [Liu & Suh] for planar Vaidya-AdS
- (Note: a-priori, monotonicity & continuity wasn’t guaranteed)



# Power of covariant constructs

- ‘Natural’ geometrical constructs (defined for general bulk spacetimes, independent of coordinates) provide useful candidates for dual of ‘natural’ quantities in CFT
- e.g. dual of  $\rho_{\mathcal{A}}$ ? [Bousso, Leichenauer, Rosenhaus; Czech, Karczmarek, Nogueira, Van Raamsdonk;...]
- In generic Lorentzian spacetime, null congruences which define a causal set provide useful characterization of ‘natural’ bulk regions.

2 options:

...starting from bdy:

$D[\mathcal{A}] \rightsquigarrow$  Causal Wedge:  $\blacklozenge_{\mathcal{A}}$

= future and past causally-separated from bdy region determined by  $\rho_{\mathcal{A}}$

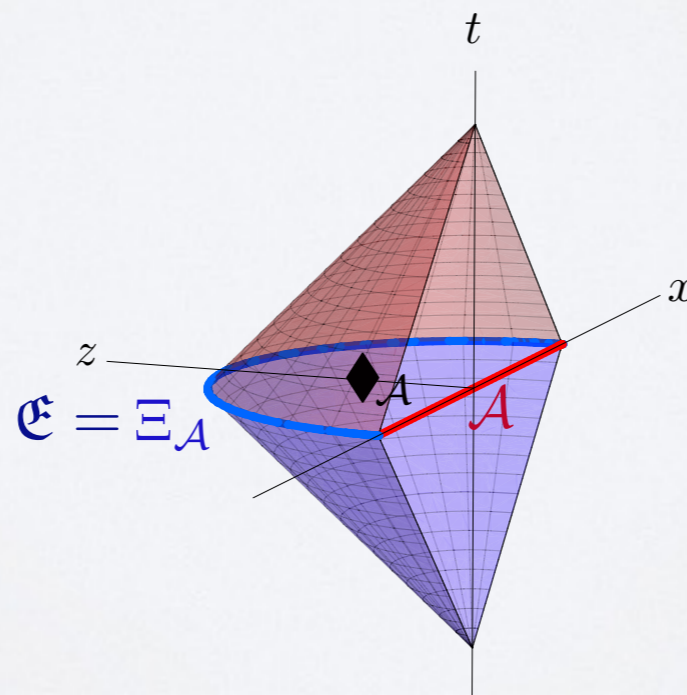
[VH & Rangamani]

...starting from bulk:

$\mathfrak{E} \rightsquigarrow$  Entanglement Wedge:  $\mathcal{W}_E[\mathcal{A}]$

= spacelike-separated (toward  $\mathcal{A}$ ) from  $\mathfrak{E}$

[Headrick, VH, Lawrence, Rangamani]

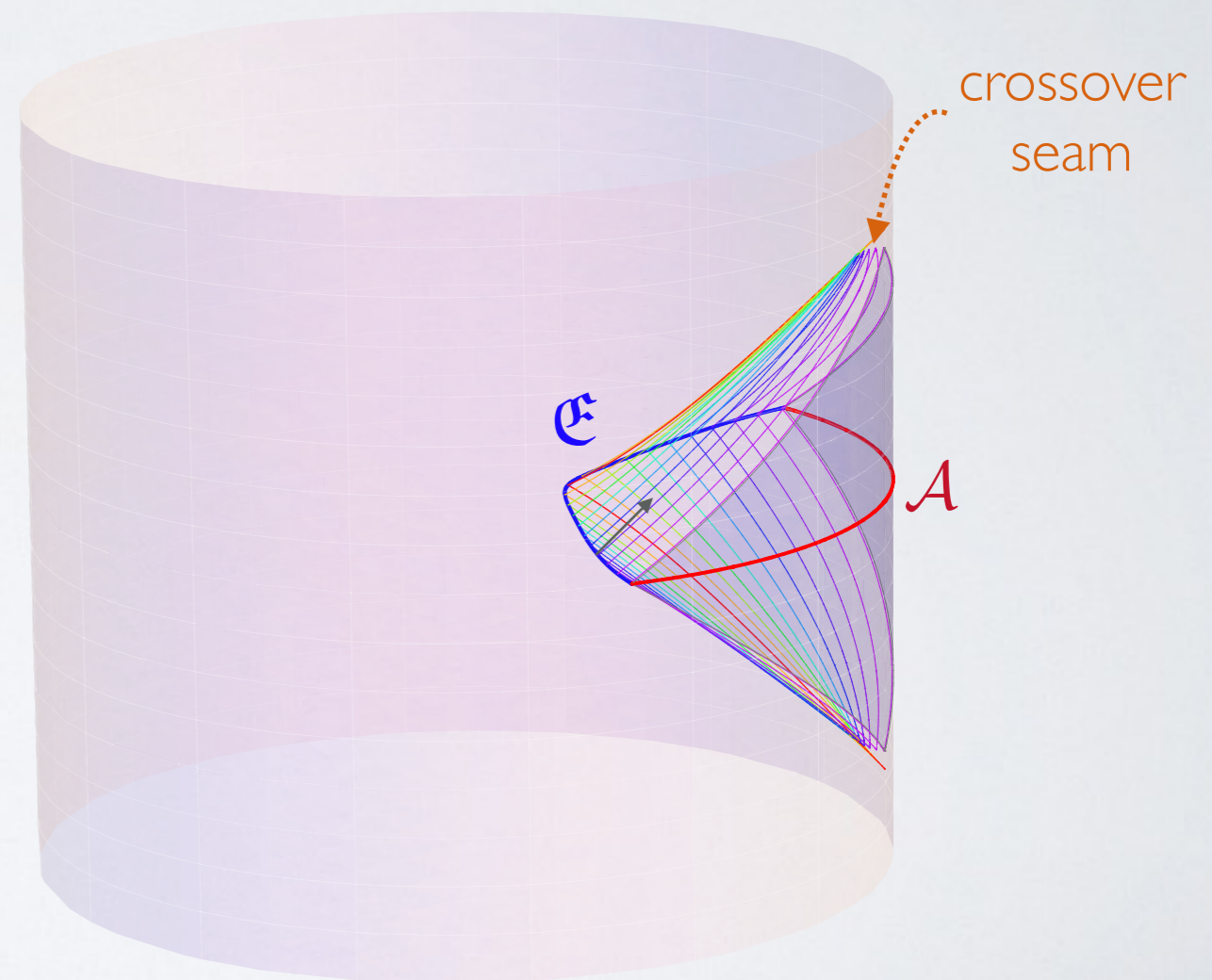
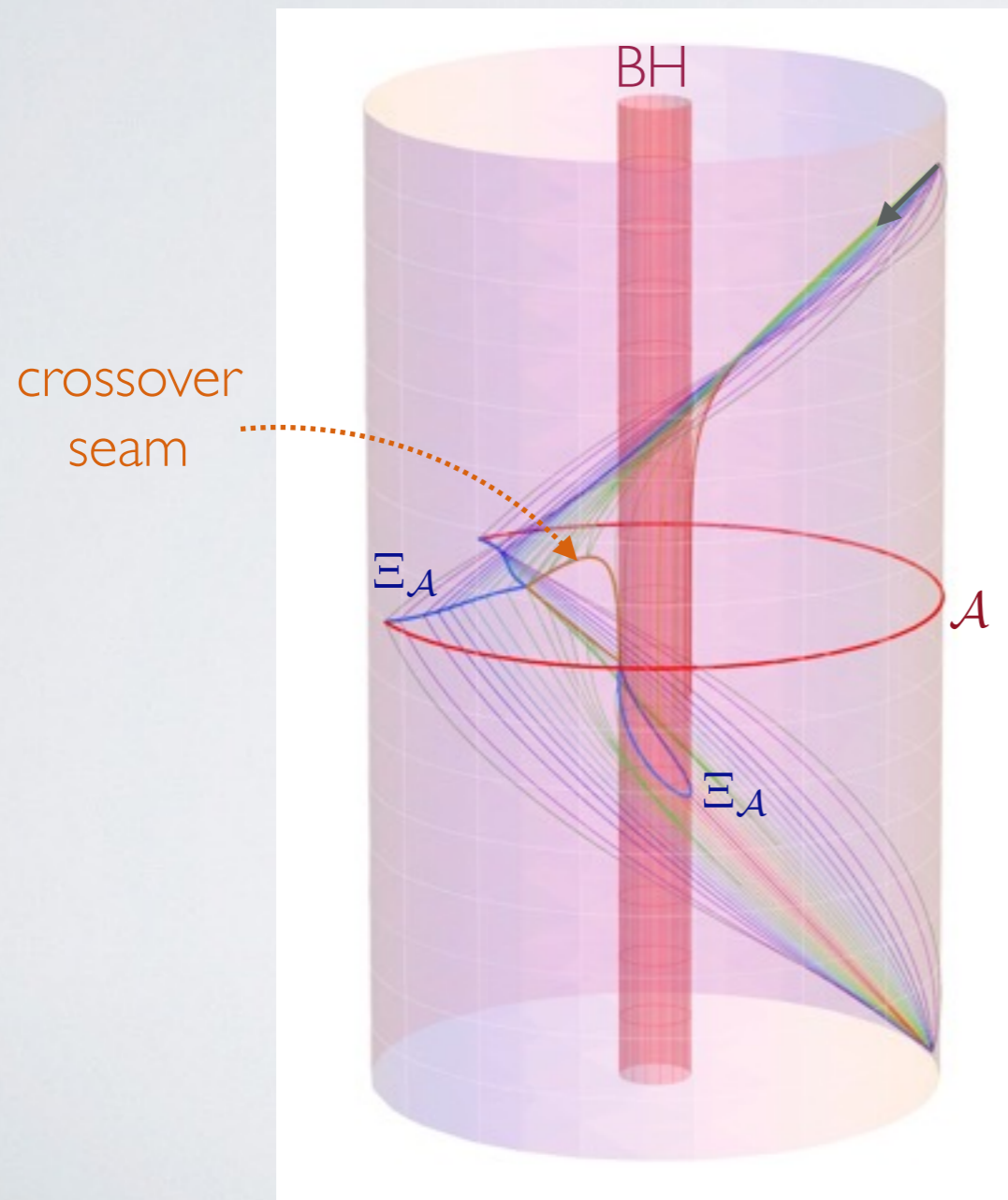


NB: in pure AdS,  
& for spherical  $\mathcal{A}$ ,  
these coincide:  $\blacklozenge_{\mathcal{A}} = \mathcal{W}_E[\mathcal{A}]$   
(but not in general)

# Causal wedge vs. Entanglement wedge

$D[\mathcal{A}] \rightsquigarrow$  Causal Wedge:  $\blacklozenge_{\mathcal{A}}$

$\mathcal{E} \rightsquigarrow$  Entanglement Wedge:  $\mathcal{W}_E[\mathcal{A}]$





# Power of covariant constructs

$D[\mathcal{A}] \rightsquigarrow$  Causal Wedge:  $\blacklozenge_{\mathcal{A}}$

$\mathfrak{E} \rightsquigarrow$  Entanglement Wedge:  $\mathcal{W}_E[\mathcal{A}]$

...continued past  $\Xi$ :  $\rightsquigarrow$  Causal Shadow  $\mathcal{Q}_{\partial\mathcal{A}}$

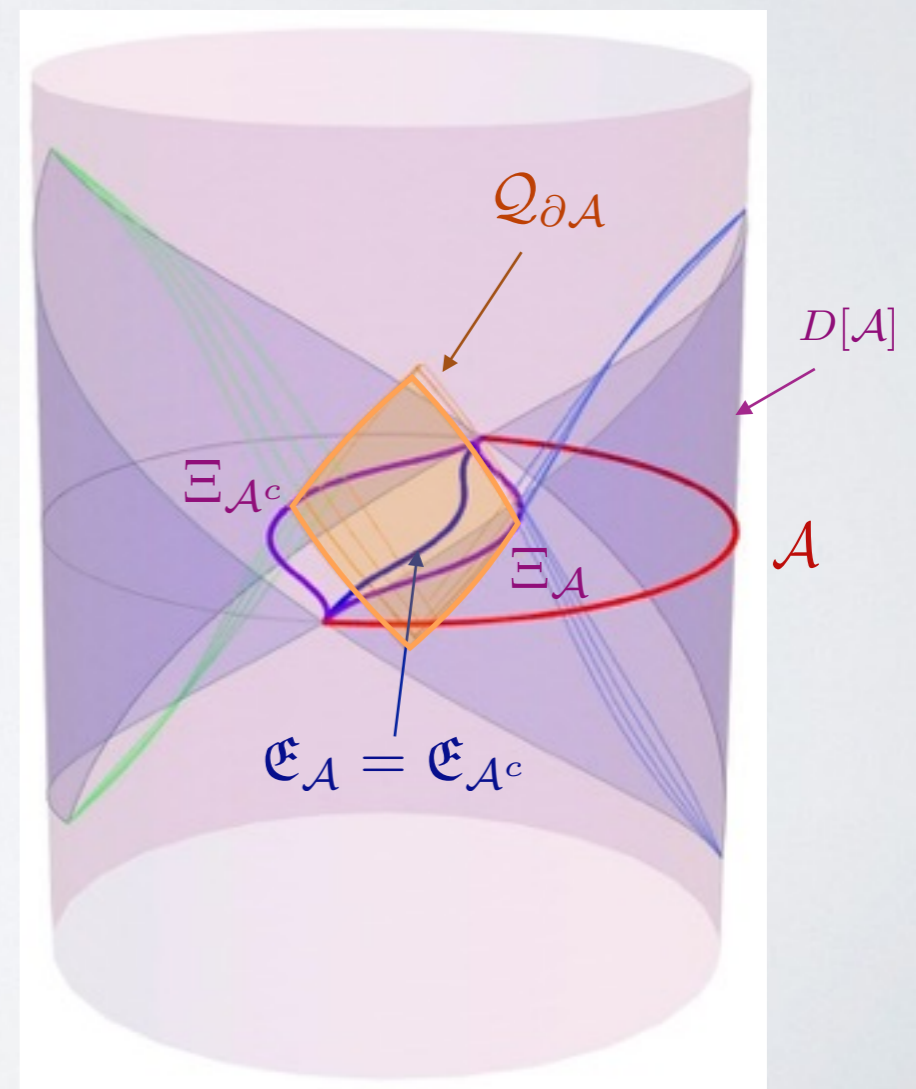
- We can prove the inclusion property [Headrick, VH, Lawrence, Rangamani; Wall]

$$\blacklozenge_{\mathcal{A}} \subset \mathcal{W}_E[\mathcal{A}]$$

or equivalently,  $\mathfrak{E} \subset \mathcal{Q}_{\partial\mathcal{A}}$

- Consequences:

- HRT is consistent with CFT causality (= non-trivial check of HRT)
- Entanglement plateaux
- Entanglement wedge can reach deep inside a black hole!





## Curious features of EE:

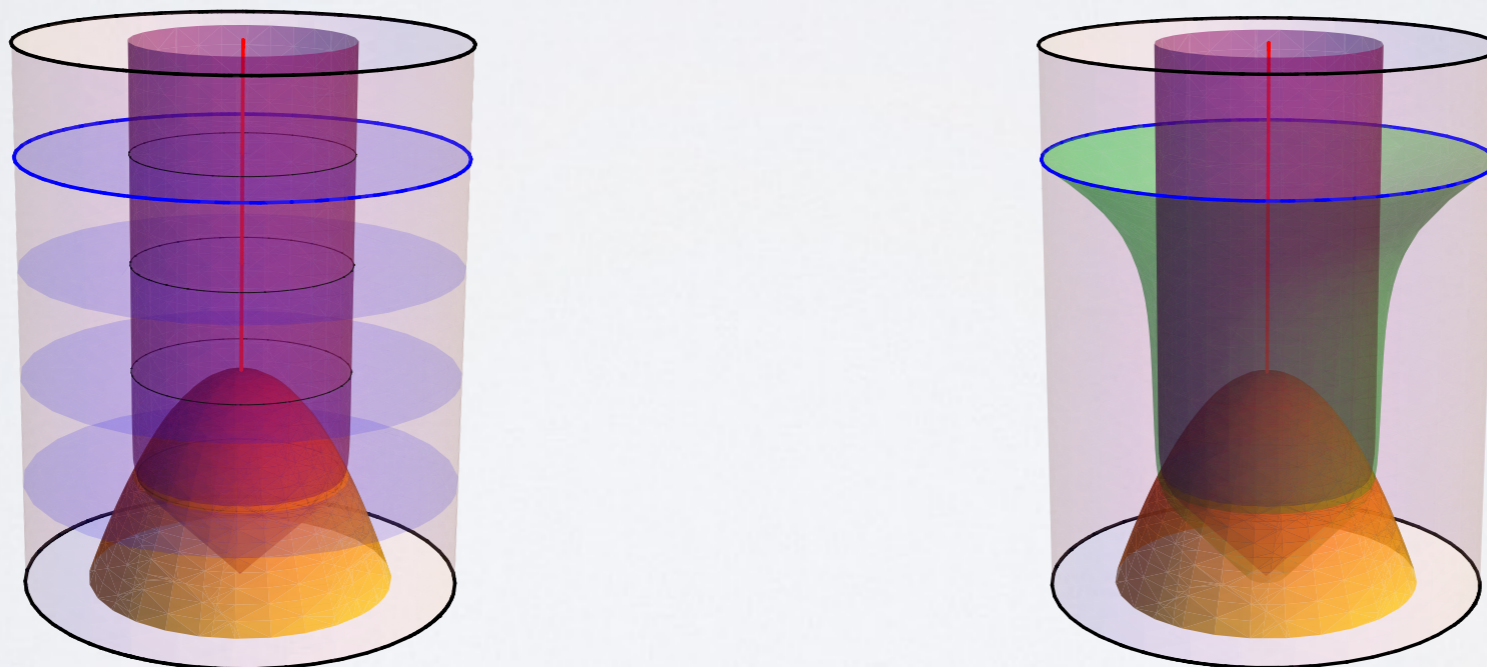
- Extremal surfaces can have intricate behavior:
  - $S_A$  can have discontinuous jumps under smooth variations of  $A$ 
    - phase transitions in EE
  - $\mathcal{E}$  can be topologically nontrivial even for simply-connected regions  $A$
- Holographic EE seems too local:
  - sharply-specified both on boundary **and** in bulk
  - but: → we can reconstruct the bulk metric (modulo caveats) solely from the set  $\{S_A\}$  for a suitable set of  $\{A\}$
- Holographic EE seems too **non**-local:
  - global minimization condition + homology constraint makes  $S_A$  sensitive to arbitrarily distant regions in the bulk...

# EE is fine-grained observable!

Example: black hole formed from a collapse

- In contrast to the static (i.e. eternal) black hole, for a collapsed black hole, there is no non-trivial homology constraint on extremal surfaces.

[cf. Takayanagi & Ugajin]



- Hence we always have  $S_{\mathcal{A}} = S_{\mathcal{A}^c}$  as for a pure state.

# Summary & Outlook

- Holography conveniently geometrizes entanglement
  - Finding bulk extremal surfaces and their area is (relatively) easy!
  - Useful in proving important properties!
  - Can we prove HRT directly?
  - Why is EE related to geometry so simply?
  - Duals of other measures of entanglement?
- General covariance is a powerful guiding principle
  - Motivated entanglement wedge, causal wedge, ...
  - How is bulk geometry encoded in  $\rho_{\mathcal{A}}$ ?
  - (In what sense) is entanglement wedge the 'dual' of  $\rho_{\mathcal{A}}$ ?
  - What is the CFT dual of causal wedge (from first principles)?
- Relation between spacetime (gravity) and entanglement?





*Thank you*