## Structure of Solid Matter Under Super-Earth Conditions

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### **Physics of Exoplanets: From Earth-sized to Mini-Neptunes**

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2

## We have built a broad base of academic collaborators

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  - D. Klug, Y. Yao
- Additional Collaborators / Consultants
  - Andrew Comley, Brian Maddox, Chris Wehrenberg, Hye-Sook Park, and Bruce Remington
- Plus farget fabrication, Omega and NIF facility and diagnostic teams.

\*Currently, Washington State University \*\*Currently, University of Edinburgh \*\*\*Currently, Harvard University \*\*\*\*Currently, DESY



- Introduction
- Ramp-Compression Stress-Density EOS
- X-Ray Diffraction at Omega and NIF
- Future Experiments

## Matter at High Energy Density (HED) is found <sup>1 Mbar = 100 GPa = 0.1 TPa</sup> throughout our universe







# High pressures have particular relevance to the interiors of planets



As of February 21, 2015: 1890 planets 1189 planetary systems 477 multiple planet systems (http://exoplanet.eu)

1189 transiting exo-planets have been confirmed ~405 with known mass and radius.

The mass-radius plot is our best way to constrain the interior make-up of the planets.

# We can study materials at rocky-planet cores pressures for 0.5–20 earth masses in the laboratory

Pressures from, Mass-radius relationships for exoplanets, D. C. Swift, et al., Astrophys. J., 744:59 (2012).

NIF is a football stadiumsized facility containing 192 of the biggest lasers in the world

NIF can deliver 2MJ of laser energy in up to 30 ns into a cm-sized gold hohlraum that then reaches radiation temperatures of 100s of eV

30 3 3

NIF-2008-Aerial L2





## **Inside the Target Chamber**







# High pressures result from the rocket effect generated by laser-induced ablation.



Time-integrated photos of shots at the Omega laser (60 beams, 30 kJ) at University of Rochester

## A uniform photon field supplied by a gold radiation cavity or *Hohlraum* ensures good drive planarity



# Typical *T<sub>rad</sub>* of 200 eV will generate pressures of about 40 Mbar



### How hot is 200 eV?





- 200 eV = 2 million degrees
- 200 eV = 400 x temperature at the surface of sun
- 200 eV = the temperature about half way to the center of the sun, in the radiative zone



## We want to measure: Stress-Density Temperature Melting **Solid-Solid Phase Transitions** Structure Texture Strength

# Equilibrium thermodynamics is described by an equation-of-state (EOS) surface







Ramp Compression -> Lower Temperature



## Ramp-compression EOS of nano-crystalline diamond to 50 Mbar.



(Smith, et al., Nature, 2014) 18

# Recent shots on NIF have measured the equation of state of Fe at 8 Mbar, >2x pressure of Earth's core



R. Smith, J. Eggert, D, Braun, R. Patterson, G. Collins et al. (LLNL), R. Jeanloz (UCB), T. Duffy (Princeton) 19

But, we wanted to measure... Stress-Density Temperature Melting Solid-Solid Phase Transitions Structure Texture Strength

So we developed an X-ray diffraction platform for Omega and NIF

# Thermo-mechanical properties can vary significantly between structures









**Physical Properties** 

	Graphite	Diamond
Density (g/cm³)	2.2	3.51
Bulk Modulus (GPa)	34	440
Coefficient of Thermal Expansion (K <sup>-1</sup> )	7.8 x 10 <sup>-6</sup>	1.1 x 10 <sup>-6</sup>
Strength (GPa)	< 0.2	>110

Structure difference leads to: 1.6x difference in density 10x difference in compressibility 500x difference in strength

# Surprising phases have been discovered at high

Traditional view that all materials become simple at high pressure is incorrect!

"... what the present results most assuredly demonstrate is the importance of pressure in revealing the limitations of previously hallowed models of solids" -Neil Ashcroft (2009).



#### 1 Mbar = 100 GPa = 0.1 TPa

## High pressure phases of aluminum are also predicted to be complex





At even higher pressures carbon is also predicted to adopt electride phases.



Canales, Pickard, Needs, Phys. Rev. Lett. 108, 045704 (2012)



# In situ diffraction gives critical data:

- -Crystal structure / phase diagram -EOS
- -Phase-transition mechanisms

X ray beam

- -Deformation texture / microstructure
- -Potential to determine liquid structure
- Target and diagnostics are simple
- We have done 22 NIF TARDIS shots

Crystal atoms





**PXRDIP** (Powder X-Ray Diffraction Image Plates)



1 Mbar = 100 GPa = 0.1 TPa

#### 1 Mbar = 100 GPa = 0.1 TPa We use diamond-sandwich targets to obtain uniform stress density conditions after ramp compression X-ray Collimation 20 1.0 (a) laser pulse Ramp drive: 197J 15 TW/cm<sup>2</sup> PW/cm<sup>2</sup> $(t_{BL} = 5.0 \text{ ns})$ P = 3.2Mbar ←target drive 10 XRS drive → Pinhole 5 0.0 0 (b) visar data Diamond Diamond (c) $\sigma_{ave}$ , U<sub>fs</sub> 10 VISAR 300 Stress (GPa) 001 002 002 002 Iron U<sub>fs</sub> (km/s) Drive ← sample stress free surface velocity $\rightarrow$ 5 0 0 300 50 (d) Stress Lagrangian position (µm) 0 0 0 0 0 0 0

We determine stress by backward propagation of the diamond free-surface velocity, assuming that we know the EOS of diamond.

diamond iron diamond

2

0

0

200 Stress (GPa)

0

8

XRS

6

4 t (ns)

# Quasi-monochromatic beam by thermal emission of Cu He-a line.



3.5% bandwidth

1 Mbar = 100 GPa = 0.1 TPa

# *In situ* diffraction provides crystal structure. Demonstrated at Omega



We observed the B1-B2 phase transition in MgO



1 Mbar = 100 GPa = 0.1 TPa

# We have observed Shock-Melt – Refreeze for first time ever in Sn at Omega



Rick Kraus, Federica Coppari

## Shock-Melt, Refreeze was observed for first time ever in Sn at Omega



NIF Discovery Science campaign to measure Fe melt line from 5-20 Mbar (~4-20  $M_F$ ). Led by R. Hemley and R. Kraus, beginning in Oct.

#### <u>1</u> Mbar = 100 GPa = 0.1 TPa TARDIS: TARget Diffraction In-Situ diagnostic for xray diffraction experiments on NIF 19 months 22 shots • 19 HED X-ray source: Drive: • 4 FS 24 beams Te beams Sample TARDIS, inserted into the Farget Alignment Sensor (TAS) in the NIF target chamber. 2 Diffraction 3 deaks Sample: 1: 40 µm diamond VISAR: 2: 2 µm gold Determines 3: 40 µm diamond pressure 4: 40 $\mu$ m diamond slurry 'Catcher' Four SiO<sub>2</sub> windows 5: 150 µm diamond

(1<sup>st</sup> two shatter)

# This year we began angle-dispersive x-ray diffraction

### **TARDIS = TARget Diffraction In Situ**



The TARDIS is a modified version of the PXRDIP used on Omega.

# Target sustains significant damage, but the support does not



1 Mbar = 100 GPa = 0.1 TPa

### At least 5 BCC diffraction lines for Pb at 2.0 Mbar



### **Preliminary Stress-Density Results**



!!! Stress, Density, and Phase
are very preliminary !!!



## Current Fundamental Science on NIF campaign to try to locate BC8 phase of carbon



# We had only seen highly-structured diffraction from diamond. On Monday we saw power diffraction.





### Validate the accuracy of our stress determination

- Effect of strength in diamond is uncertain
- $\circ~$  Use different sandwich materials, e.g.: Diamond, LiF, MgO

### Experimentally constrain the temperatures we reach

- Heating by: thermal conduction, plastic work, shock formation, etc.
- o Debye-Waller factor in diffraction
- Debye-Waller factor in EXAFS
- Thermal-expansion in diffraction

### **Explore systematics of rate-dependent phase transitions**

- Simple structures may form more rapidly than complex structures
- Explore phase diagrams on multiple platforms

### Improve signal to noise ratios

- $\circ~$  Lower ablation plasma and other background
- $\circ~$  Use higher energy x-rays

### **Develop cheaper targets**