

# An Infalling Observer and the Black Hole Information Paradox in AdS-CFT

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**Complementarity**, Fuzz or Fire

26 August 2013

Based on [arXiv:1211.6767](https://arxiv.org/abs/1211.6767) (with Kyriakos Papadodimas)

# Overview

- Construction of local operators outside the Black Hole.
- Construction of local operators behind the horizon.
- Complementarity and State Dependence.
- Addressing various counter-arguments:
  - 1 Suppressing commutators by system size. (AMPSS)
  - 2 Lack of a Left Inverse (AMPSS)
  - 3  $\langle N_a \rangle \neq 0$  (MP)
  - 4 Restoring unitarity with small corrections. (Mathur)
  - 5 Unfreezing the vacuum. (Bousso, van Raamsdonk)

# Generalized Free Fields

- The boundary theory has **generalized free fields**,  $\mathcal{O}(t, x)$  of low dimension.
- Correlators of these fields **factorize**:

$$\begin{aligned} & \langle \mathcal{O}(t_1, x_1) \dots \mathcal{O}(t_{2n}, x_{2n}) \rangle \\ &= \langle \mathcal{O}(t_1, x_1) \mathcal{O}(t_2, x_2) \rangle \dots \langle \mathcal{O}(t_{2n-1}, x_{2n-1}) \mathcal{O}(t_{2n}, x_{2n}) \rangle \\ & \quad + \text{permutations} + \frac{1}{N} \dots, \end{aligned}$$



$$\langle 0 | \mathcal{O}(t, x), \mathcal{O}(0, 0) | 0 \rangle = \left( \frac{-1}{t^2 - x^2 - i\epsilon t} \right)^\Delta = \left( \frac{-1}{(t - i\epsilon)^2 - x^2} \right)^\Delta.$$

- However,  $\mathcal{O}$  does **not obey an equation of motion**.

# Local Observables in empty AdS

- We can recast dynamics of  $\mathcal{O}$  using a **one-to-one** map to another operator  $\phi_{\text{CFT}}$

[Banks et. al., Bena, Kabat et al., 1998–2012]

$$\mathcal{O} \Leftrightarrow \phi_{\text{CFT}}$$

- The precise definition is:

$$\phi_{\text{CFT}}^{\text{vac}}(t, x, z) = \int_{\omega>0} \frac{d\omega d^{d-1}k}{(2\pi)^d} [\mathcal{O}_{\omega,k} \xi_{\omega,k}(t, x, z) + \text{h.c.}]$$

where  $\xi$  are appropriately chosen functions.

- $\phi_{\text{CFT}}^{\text{vac}}$  behaves like a **free-field** in AdS. For example:

$$[\phi_{\text{CFT}}^{\text{vac}}(t, x, z), \dot{\phi}_{\text{CFT}}^{\text{vac}}(t, x', z')] = \frac{i}{(2\pi)^d} \delta^{d-1}(x - x') \delta(z - z') z^{d-1}.$$

# Emergent AdS

- This is the **clearest way** to understand the emergence of the bulk from the CFT.
- If we did not know AdS/CFT, this map would look **miraculous**.
- For example, no normalizable modes in the bulk with spacelike momenta. Correspondingly, the spacelike modes of generalized free-fields can be discarded at large  $N$ .
- So, the map is **quite constrained** even at infinite  $N$ .

# Local Observables outside a Black Hole

- Consider the CFT in a pure state  $|\Psi\rangle$  that is “close” to a thermal state.
- The same generalized free-fields  $\mathcal{O}$  have different correlators in the state  $|\Psi\rangle$ .
- However, we can still construct perturbative local fields

$$\phi_{\text{CFT}}^\beta(t, \mathbf{x}, \mathbf{z}) = \int_{\omega>0} \frac{d\omega d^{d-1}k}{(2\pi)^d} [\mathcal{O}_{\omega,k} f_{\omega,k}(t, \mathbf{x}, \mathbf{z}) + \text{h.c.}]$$

## More on Operators in front of the horizon

- This is even **more miraculous and constrained**. For example, at large-spacelike momenta, we require:

$$\mathcal{O}_{\omega,k} \rightarrow e^{\frac{-\beta|k|}{4}}$$

for spacelike momenta:  $k \gg \omega$ . True because of **general properties of thermal CFT correlators**.

[Papadodimas, S.R., 2012]

- Moreover, the CFT-bulk map is **state dependent**. Expected, because we do not expect to formulate local observables in a **background-independent** way.

## More on State Dependence

- This is **not** a modification of quantum mechanics: in different states, it is convenient to use **different integral transforms** of operators.
- What if we took the state  $\frac{1}{\sqrt{2}} (|\Omega\rangle + |\Psi\rangle)$ ? **Both**  $\phi_{\text{CFT}}^{\text{vac}}$  and  $\phi_{\text{CFT}}^{\beta}$  are well-defined, but **neither** has a good **interpretation** as a bulk perturbative field.
- If we take  $\sum_E c_E |E\rangle$ , with the right  $c_E$ ,  $\phi_{\text{CFT}}^{\beta}$ , but **not**  $\phi_{\text{CFT}}^{\text{vac}}$ , has a good bulk interpretation.
- The CFT-bulk kernel depends on **more than the energy**. Eg. for a unstable gas, **before it collapses to a black hole**, the kernel is different.



# Operators behind the black hole

- To construct local operators behind the horizon, we need **additional operators** from an operator  $\tilde{\mathcal{O}}(t, x)$  in the CFT.
- This operator should **almost commute** with the original GFF:

$$[\tilde{\mathcal{O}}(t_1, x_1), \mathcal{O}(t_2, x_2)] \sim \epsilon$$

- Results of measurements of  $\mathcal{O}$  and  $\tilde{\mathcal{O}}$  are **correlated**
- i.e. we want a **generic equilibrated CFT state** to look like

$$|\Psi\rangle = \frac{1}{\mathcal{Z}} \sum e^{\frac{-\beta\omega}{2}} |\omega, k\rangle \otimes |\tilde{\omega}, \tilde{k}\rangle$$

# Local operators Behind the Horizon

- If we can find such operators in the CFT, local operators behind the horizon are

$$\phi_{\text{CFT}}^{\text{II}}(t, \mathbf{x}, z) = \int_{\omega > 0} \frac{d\omega d^{d-1}k}{(2\pi)^d} \left[ \mathcal{O}_{\omega, k} g_{\omega, k}^{(1)}(t, \mathbf{x}, z) + \tilde{\mathcal{O}}_{\omega, k} g_{\omega, k}^{(2)}(t, \mathbf{x}, z) + \text{h.c.} \right],$$

where we can **explicitly compute**  $g^{(1)}$  and  $g^{(2)}$ .

[Papadodimas, S.R., 2012]

So, the issue of whether or not we can describe the interior in AdS/CFT is an issue of finding such operators.

# Resolving Various Puzzles

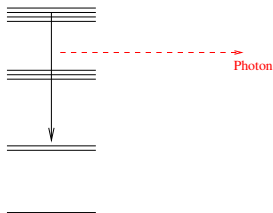
- If the  $\mathcal{O}$  and  $\tilde{\mathcal{O}}$  operators are an overcomplete basis for the same space, then we can resolve the **strong subadditivity puzzle**
- Such a construction would also evade Mathur theorem on **small corrections**, which tacitly does not allow for this possibility of complementarity.
- AMPSS and MP have given several recent arguments in favour of firewalls.
- Later, we will discuss how many of these arguments are **very naturally evaded** if we allow the operators to be state dependent.
- The proposal is strong, precisely where these arguments meet it!

# The Generic Commutator

- But first, let us address the following issue: can we ensure that  $\tilde{\mathcal{O}}$  operators live in the same space as the  $\mathcal{O}$  operators and have a **small commutator** with them.
- We take a basis of operators  $\mathcal{O}_n$  in a large Hilbert space. eg. imagine the spin operators of a spin chain.
- We now scramble it with a generic unitary matrix  $U$  to form  $\tilde{\mathcal{O}}_n = U\mathcal{O}_nU^{-1}$ .
- **Typically**, the commutator  $[\tilde{\mathcal{O}}_m, \mathcal{O}_n]$  has  $\mathcal{O}(1)$  eigenvalues – Bad for locality.

## A more realistic example

- We would like to see if these arguments hold true, in a **more realistic** case.
- Imagine an evaporating black-hole as a system coupled to a radiation field.



- The real puzzle is to find the operators  $\tilde{\mathcal{O}}$  when **most of the energy is in the radiation field**, so that  $\mathcal{O}$  and  $\tilde{\mathcal{O}}$  are acting on the same space.

## Dual operators in a radiation field

- Consider a one-dimensional field, living in a **very long box** of length  $l = \frac{3}{2M_{\text{pl}}} \left( \frac{M}{M_{\text{pl}}} \right)^3$ .
- The frequencies are quantized in units of  $\omega_{IR} = \frac{2}{3} M_{\text{pl}} \left( \frac{M_{\text{pl}}}{M} \right)^3$ .
- Consider putting an energy  $M$  into this box.
- Amusingly, the total entropy of this configuration can be calculated by Cardy's formula and is  $S = \pi M^2$ .
- We can think of this as the **Hawking gas of s-waves** emitted by an old-black hole, living in a **box of the size of the Page time**.

# The natural observable

- The natural observable is the field

$$\phi(t, r) = \sum_n \frac{1}{\sqrt{n}} a_n e^{n\omega_{IR}(t-r)} + \text{h.c.} + \text{left-movers}$$

- And we are interested in correlators  $\langle \phi(r_1, t_1) \phi(r_2, t_2) \dots \phi(r_p, t_p) \rangle$ , where the number of points  $p$  **does not scale with  $N$**

## Coarse Graining the Field

- Consider an observer, with **limited resolving power**: cannot distinguish the frequency  $p\omega_{IR}$  from  $(p+1)\omega_{IR}$ .
- Define the following linear combination of oscillators, where  $m$  can be any  $O(1)$  number.

$$\alpha_p = \frac{1}{\sqrt{m}} \sum_{i=1}^m a_{mp+i}, \quad \beta_p^j = \frac{1}{\sqrt{m}} \sum_{i=1}^m \chi_i^j a_{mp+i},$$

- The only relevant property of the  $\chi_i^j$  coefficients is that

$$\sum_i (\chi_i^{j_2})^* \chi_i^{j_1} = \delta^{j_1 j_2}; \quad \sum_i \chi_i^j = 0, \forall j,$$

so that the  $\beta_p^j$  oscillators are orthogonal to each other and to the  $\alpha_p$  oscillators.



# Coarse and Fine Spaces

- The physical idea is that the **coarse observer sees effectively the space of excitations of the  $\alpha_p$** , but there are fine-grained degrees of freedom in the  $\beta_p^j$  that he cannot access easily.
- Divide the full Hilbert space into

$$\mathcal{H} = \mathcal{H}_c \otimes \mathcal{H}_f$$

where  $\mathcal{H}_c$  consists of excitations of the  $\alpha_p$  oscillators and  $\mathcal{H}_f$  consists of excitations of the  $\beta_p^j$  oscillators.

- If we fix the total energy, then obviously we are restricted to certain entangled states in this product space.

# Entanglement between coarse and fine d.o.f.

- Any state in the CFT can be written

$$|\Psi\rangle = \sum_{i,j} \alpha_{ij} |\Psi_i^c\rangle \otimes |\Psi_j^f\rangle,$$

where  $i$  runs over an orthonormal basis in  $\mathcal{H}_c$  and  $j$  over an orthonormal basis in  $\mathcal{H}_f$ .

- We can perform a **singular value decomposition** of the matrix  $\alpha$

$$\alpha_{ij} = \sum_m U_{im} D_{mm} V_{mj}$$

where  $D$  is a rectangular diagonal matrix

$$D = \begin{pmatrix} D_{11} & 0 & 0 & 0 & \dots \\ 0 & D_{22} & 0 & 0 & \dots \\ 0 & 0 & D_{33} & 0 & \dots \end{pmatrix}$$

# Mirrored Operators

- In this basis, the state becomes

$$|\Psi\rangle = \sum_i D_{ii} |\hat{\Psi}_i^c\rangle \otimes |\hat{\Psi}_i^f\rangle$$

- For some matrix elements  $\omega_{i_1, i_2}$ :

$$\alpha_p = \sum_{i_1, i_2} \omega_{i_1, i_2} |\hat{\Psi}_{i_1}^c\rangle \langle \hat{\Psi}_{i_2}^c|.$$

- **Define** a mirrored operator on the fine-grained space:

$$\widetilde{\alpha}_p = \sum_{i_1, i_2} \omega_{i_1, i_2}^* |\hat{\Psi}_{i_1}^f\rangle \langle \hat{\Psi}_{i_2}^f|,$$

# Mirrored Field

- We can construct a mirrored field

$$\tilde{\phi}(t, x) = \sum_n \frac{1}{\sqrt{nm}} \tilde{\alpha}_n \sum_{j=1}^m e^{i\omega_{IR}(mn+j)(r-t)} + \text{h.c}$$

- This field **commutes up to small corrections** with the original field

$$\begin{aligned} & \langle \phi(r_1, t_1) \phi(r_2, t_2) [\phi(r_3, t_3), \tilde{\phi}(0, 0)] \rangle \\ & \approx \langle \phi(r_1, t_1) \phi(r_2, t_2) [\phi^c(r_3, t_3), \tilde{\phi}(0, 0)] \rangle + \mathcal{O}\left((t_3 - r_3) \frac{M_{\text{pl}}}{S^{\frac{3}{2}}}\right) \\ & = 0 + \mathcal{O}\left((t_3 - r_3) \frac{M_{\text{pl}}}{S^{\frac{3}{2}}}\right) \end{aligned}$$

- In fact, with a little more work, we can argue that only the **square of the commutator** is observable, so the smallest non-local effect is suppressed by  $\frac{1}{S^3}$

# No Left Inverse Argument

- So, this provides an **explicit example** of a model of the interior that can preserve unitarity, with small violations of locality, and small corrections to semi-classical correlators outside and inside the B.H.
- Consider, some other arguments for why the interior operators cannot exist.
- $\tilde{a}_{\omega,k}^\dagger$  **lowers the energy and has a left inverse.**
- There is no such operator in the field theory.
- But, we have an explicit construction of such operators in this toy model. What gives?

[AMPSS, 13]

# State dependence resolves left inverse argument

- Toy model:

coarse space

$|0\rangle$ , with energy = 0

$|1\rangle$ , with energy = 1

fine space

$|0\rangle$ , with energy = 0

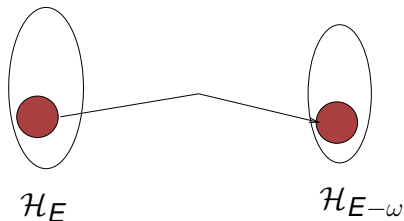
$|1\rangle, |2\rangle, \dots |n\rangle$ , with energy = 1

- Consider some state with total energy = 1:

$$|\Psi\rangle = |0\rangle_{\text{coarse}} \otimes \left( \sum_{j=1}^n c_j |j\rangle_{\text{fine}} \right) + |1\rangle_{\text{coarse}} \otimes |0\rangle_{\text{fine}}$$

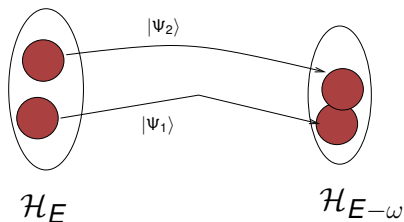
- The mirror of the coarse operator  $a^\dagger = |1\rangle_{\text{coarse}} \langle 0|_{\text{coarse}}$  is  $\tilde{a}^\dagger = c_j |0\rangle_{\text{fine}} \langle j|_{\text{fine}}$

# State Dependent Operators are Sparse



State Dependent Operators are Sparse!: No contradiction with linear algebra!

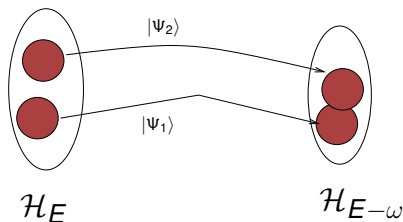
# Union of all constructions?



- Can we use the "union" of all constructions to obtain a contradiction?



# Union of all constructions?



- Can we use the "union" of all constructions to obtain a contradiction?
- NO!
- The explicit construction shows that if we try and cover  $\mathcal{H}_E$  using different states  $|\Psi\rangle_m$  with different structures of entanglement, we **over-cover**  $\mathcal{H}_{E-\omega}$ , as one would expect.

## $\langle N_a \rangle \neq 0$ Argument

- Let  $N_b$  measure the occupation number of some  $b$ -mode outside the horizon.
- **Marolf-Polchinski:** If  $\tilde{N}_b$  is some **fixed operator** in the CFT, then we would expect that a generic state at fixed energy does not look like  $|N_b, \tilde{N}_b\rangle$  and so has  $N_a \neq 0$ .
- Compute the microcanonical expectation value  $\langle N_a \rangle$  by going to the  $N_b$  basis: generically expect this to be non-zero.
- What is the contradiction with our explicit construction, where  $\langle N_a \rangle = 0$ ?

# State dependence resolves the $N_a \neq 0$ argument

- The operator  $N_a$  depends on  $N_b$  and  $\tilde{N}_b$ .
- So, in taking the expectation value  $\sum_{\Psi} \langle \Psi | N_a | \Psi \rangle$ , the **precise operator  $N_a$  varies as we change  $|\Psi\rangle$**
- In fact, **by construction**, for generic states:

$$|\Psi\rangle = \sum_{N_b} e^{-\frac{1}{2}\beta N_b \omega_b + \dots} |N_b, \tilde{N}_b, \dots\rangle$$

So,  $N_a |\Psi\rangle = 0$ .

- So, the Marolf-Polchinski argument should be interpreted as **a strong argument in favour of state-dependence of the interior operator.**

# Frozen Vacuum?

- Another counter-argument (specific to this construction): If the interior operator is **always defined** to be the one chosen by entanglement, then the vacuum is always feature-less (or **frozen**).

[Bousso, van Raamsdonk, 13]

- The rule is **not** that the interior operator is chosen by entanglement.
- Rather it is that the **interior operator is chosen to be the one that gives perturbative local correlators**.
- For a generic equilibrium state, the two rules coincide.

# Interior Operators for a Non-Equilibrium State

- For a slightly out of equilibrium state  $\mathcal{O}(x, t)|\Psi\rangle$ , where  $|\Psi\rangle$  is an equilibrium state, the rule is that we define the operators  $\tilde{\mathcal{O}}$  with respect to  $|\Psi\rangle$ , and then **do not change them**.
- It would be nice to prove that this rule is uniquely fixed by the criterion that we should get local operators.
- This rule **does give local operators with the correct behaviour** and **no frozen vacuum**. So, we conjecture that it is correct.

# Small Corrections

- How do we have a unitary model of evaporation, with only small corrections to correlators?

[Mathur, 2009–13]

- If we look at the full wave-function, it **appears to evolve** as

$$\begin{aligned} |\Psi\rangle &\rightarrow \frac{1}{\sqrt{2}} |\Psi'\rangle \otimes (|0\rangle_c |0\rangle_b + |1\rangle_c |1\rangle_b) \\ &\rightarrow \frac{1}{2} |\Psi''\rangle \otimes (|0\rangle_{c_1} |0\rangle_{b_1} + |1\rangle_{c_1} |1\rangle_{b_1}) \otimes (|0\rangle_{c_2} |0\rangle_{b_2} + |1\rangle_{c_2} |1\rangle_{b_2}) \end{aligned}$$

- The excitations  $b_1, c_1, b_2, c_2$  are all excitations of the same collective modes. They are **not independent excitations**
- This is what allows us to evade the theorem, and obtain unitary evolution with small corrections.

# Summary

- We want to **reorganize boundary correlators** in terms of perturbative fields propagating on a spacetime with an extra dimension.
- Our construction suggests that, to leading order in  $\frac{1}{N}$ , this successfully describes both the interior and the exterior of the black hole.
- The operators, both outside and inside are state-dependent. (At leading order, the exterior operators depend only on the energy, while interior operators seem to have stronger state dependence.)
- This seems to resolve several arguments raised against the existence of such operators.

# Open question 1: size of non-local effects

- The current model has power-law (in entropy) suppression for commutators. It is difficult to extend this to exponential suppression, without making the black-hole too long lived.
- On the other hand, even for points  $r_{\text{out}}$  **outside the black-hole**:

$$[h_{00}(t, \infty), \phi(t, r_{\text{out}})] = \mathcal{O}\left(\frac{1}{N}\right)$$

because the asymptotic metric can measure the energy inside.

- So, we need to understand which operators (if any) should have exponentially suppressed commutators.



## Open questions 2: $\frac{1}{N}$ corrections

- We have made a guess for interior operators. How we know that this is correct?
- Again, the same question exists for exterior operators.
- We conjecture that a careful study of  $\frac{1}{N}$  corrections should allow us to better fix both interior and exterior operators.

## Open questions 3: Bulk Interpretation

- It seems clear that the bulk theory will require some (preferably exponentially small) degree of non-locality to resolve the information paradox.
- If we formulate the theory using a **path integral**, we should expect such corrections since the semi-classical spacetime ceases to make sense at this level of accuracy.
- But, can we find some explicit non-perturbative effect that produces the non-local corrections required to restore unitarity?