

Ages of Giants with High Resolution Spectroscopy

Diane Feuillet

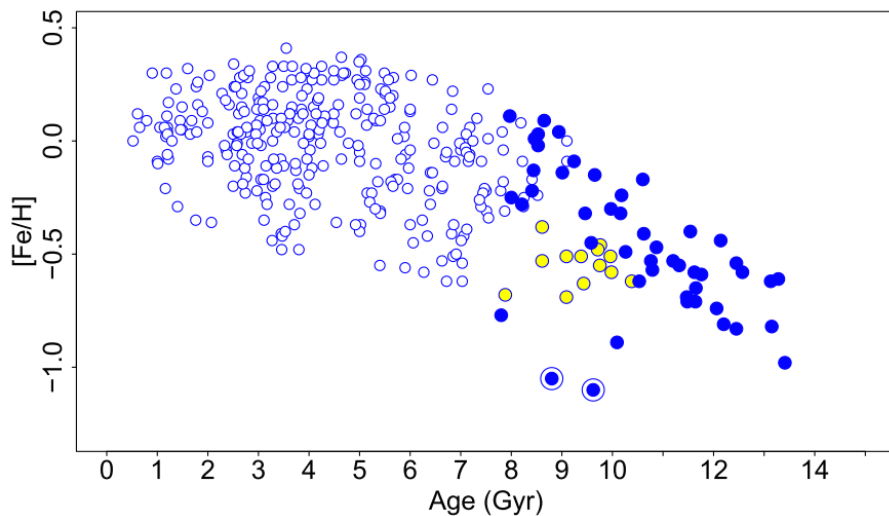
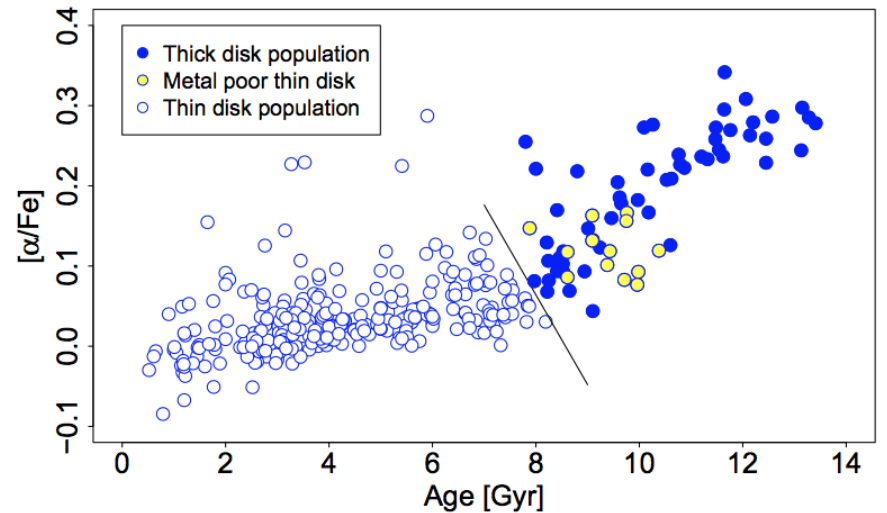
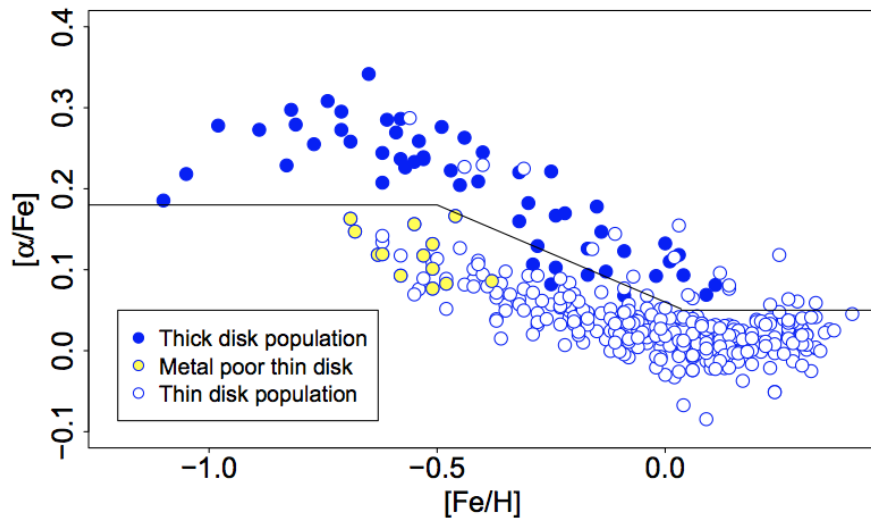
New Mexico State University

February 2, 2015

Jon Holtzman, Leo Girardi
And the APOGEE Team



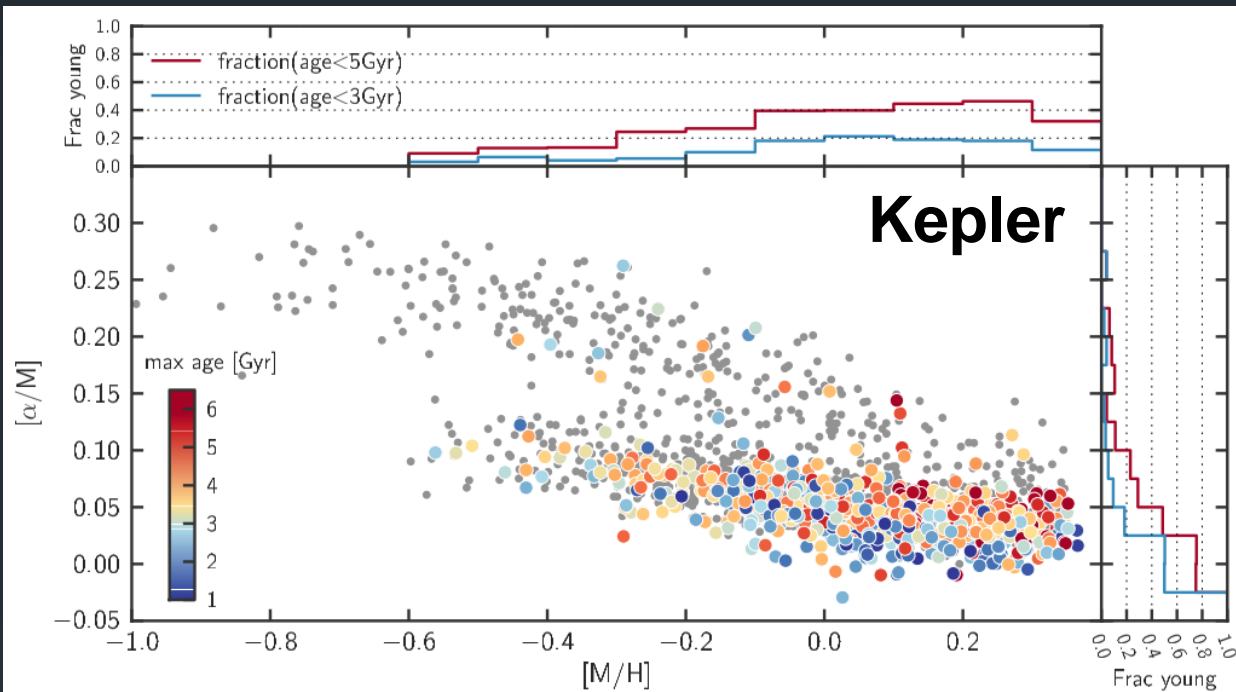
Ages and Abundances



Haywood+ 2013

Age adds crucial evolutionary information and population identification

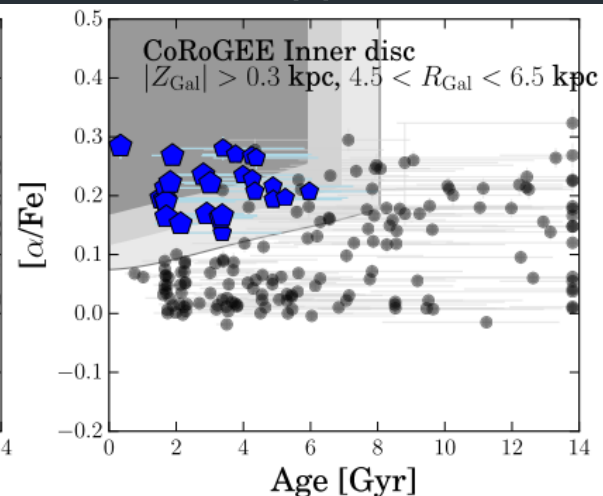
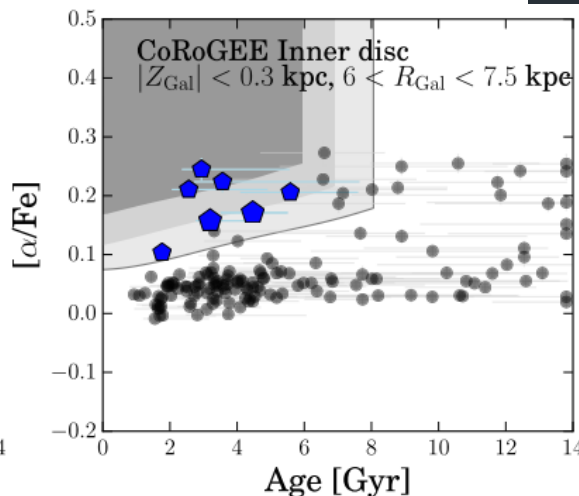
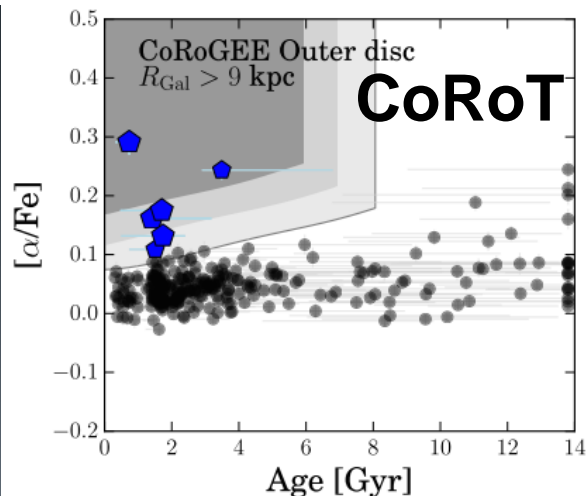
Asteroseismic Ages



Martig+ 2015

- Recent discovery of young alpha-enhanced stars
- Need reliable ages for larger samples of stars.
- See Andrea Miglio talk on Wednesday

Chiappini+ 2015



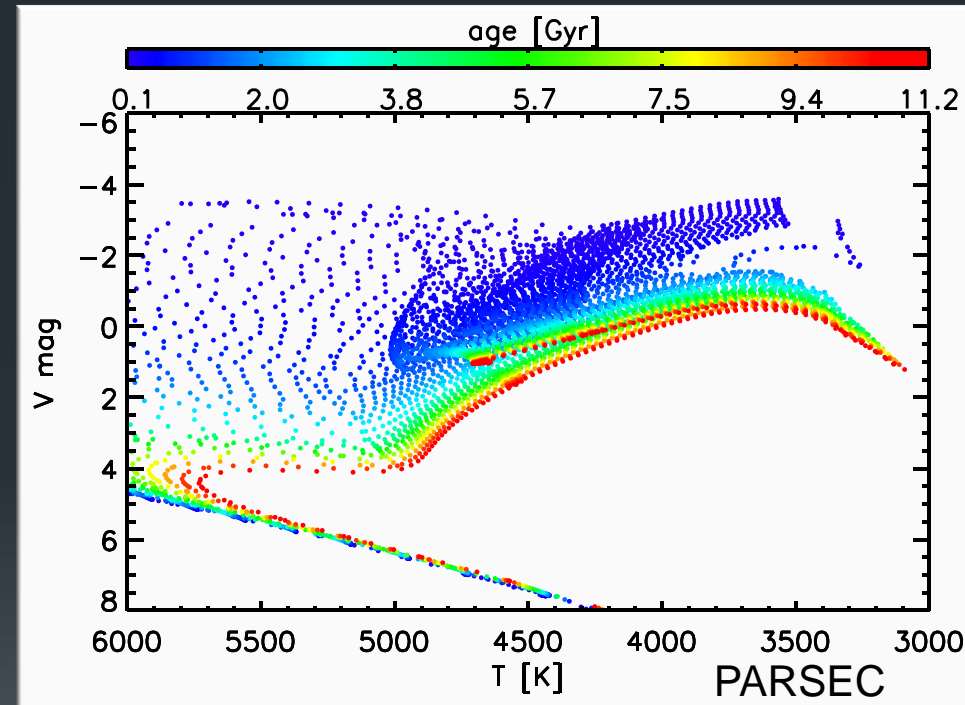
Mass and Age of Giants

Good Distances:

- V mag puts better constraints on age
- Distance + APOGEE T_{eff} & $\log g$ gives mass estimate
- Mass error \approx $\log g$ error = 0.15 dex

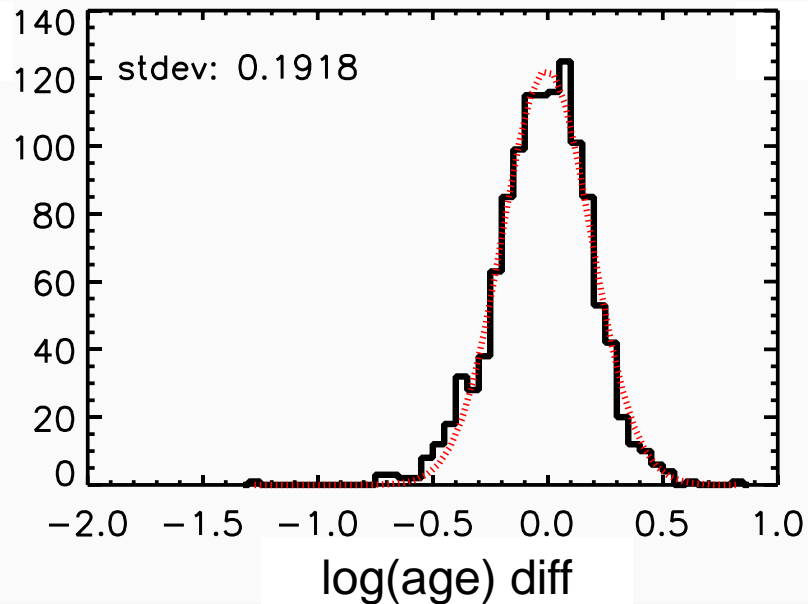
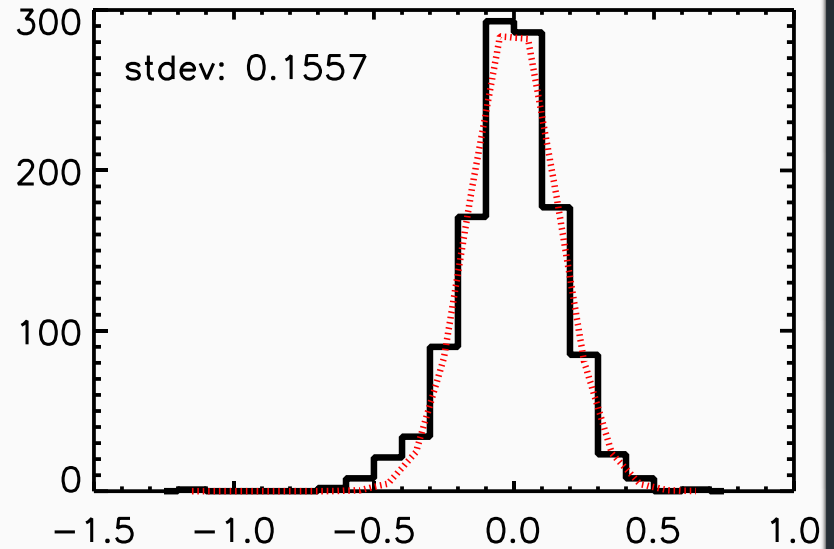
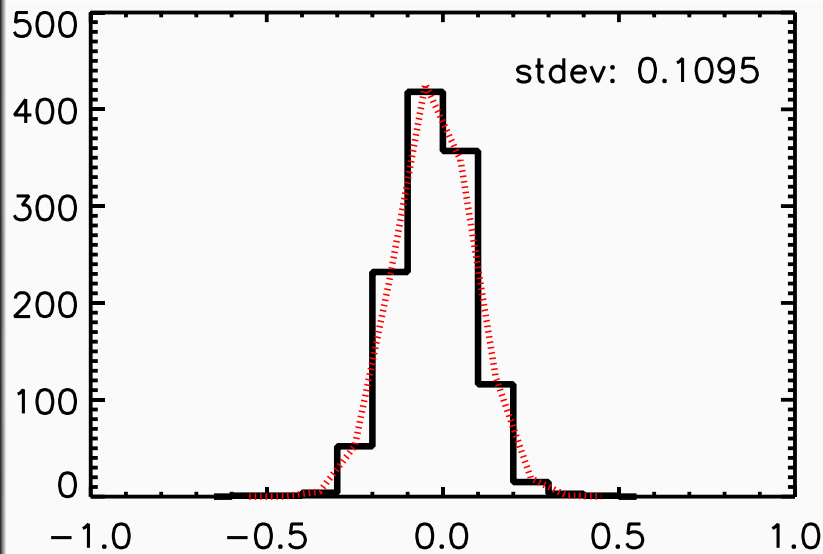
Isochrone Matching:

- $P \propto \exp[-(X_o - X_i)/2\Delta X_o]$
- APOGEE: T_{eff} , $[\text{Fe}/\text{H}]$, $\log g$
- Hipparcos: V mag
- Both: mass



APOGEE + Gaia

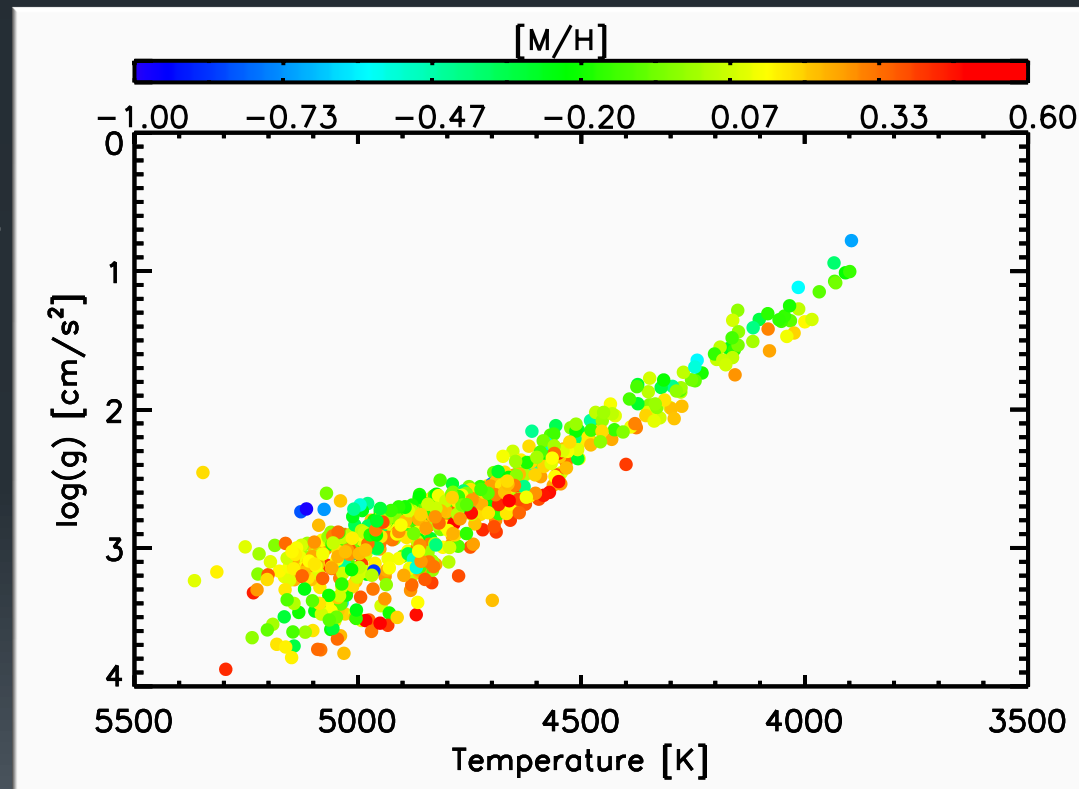
Simulated Sample - PARSEC



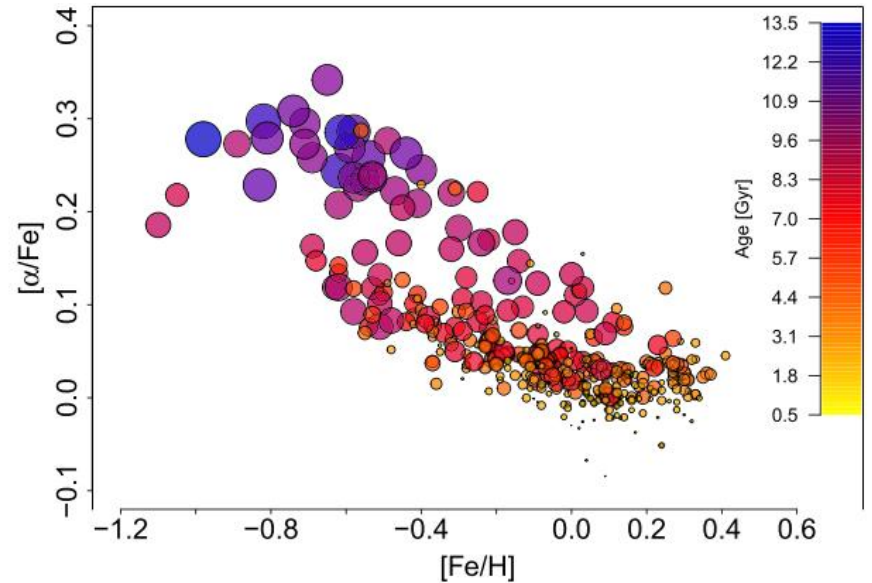
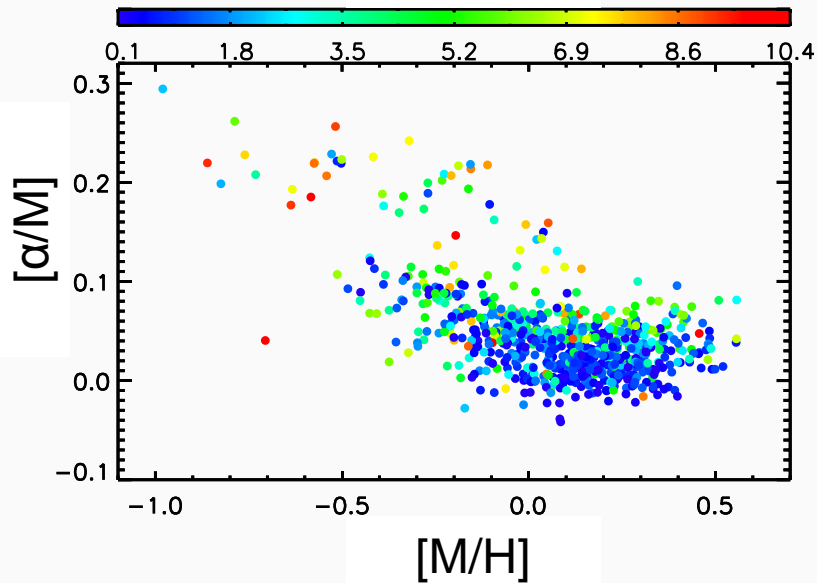
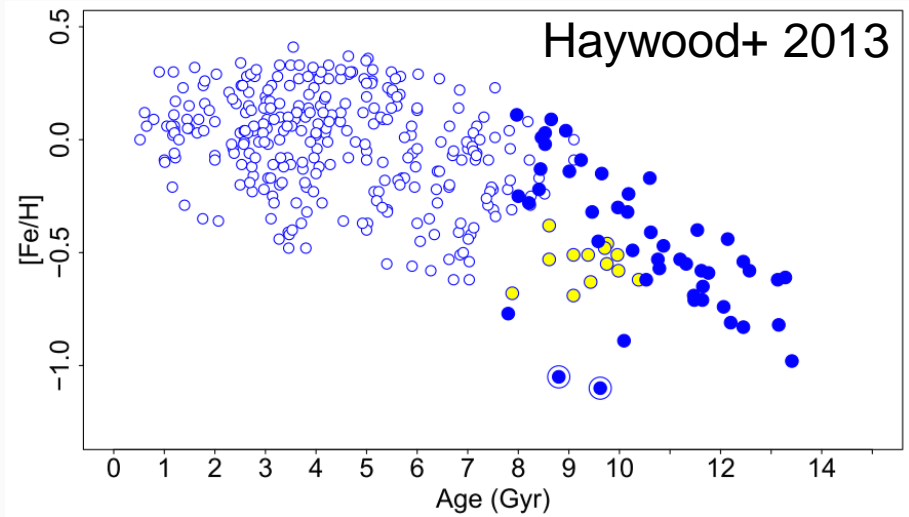
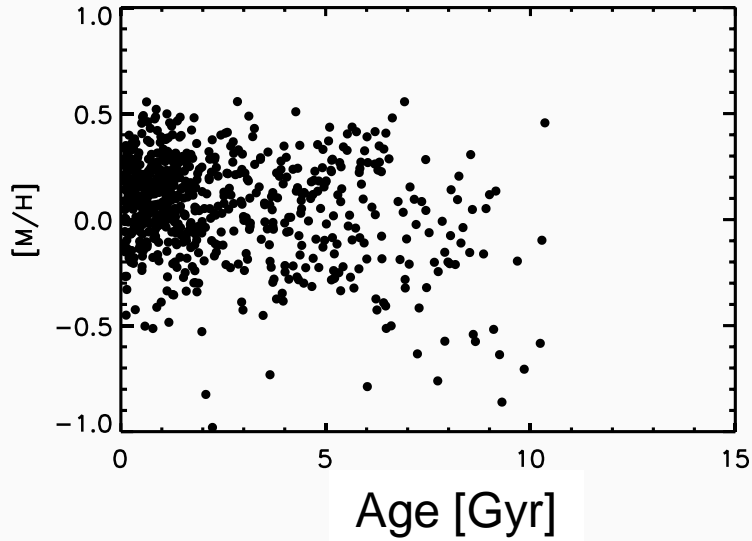
Solar Neighborhood Sample

700 stars observed to SNR 100

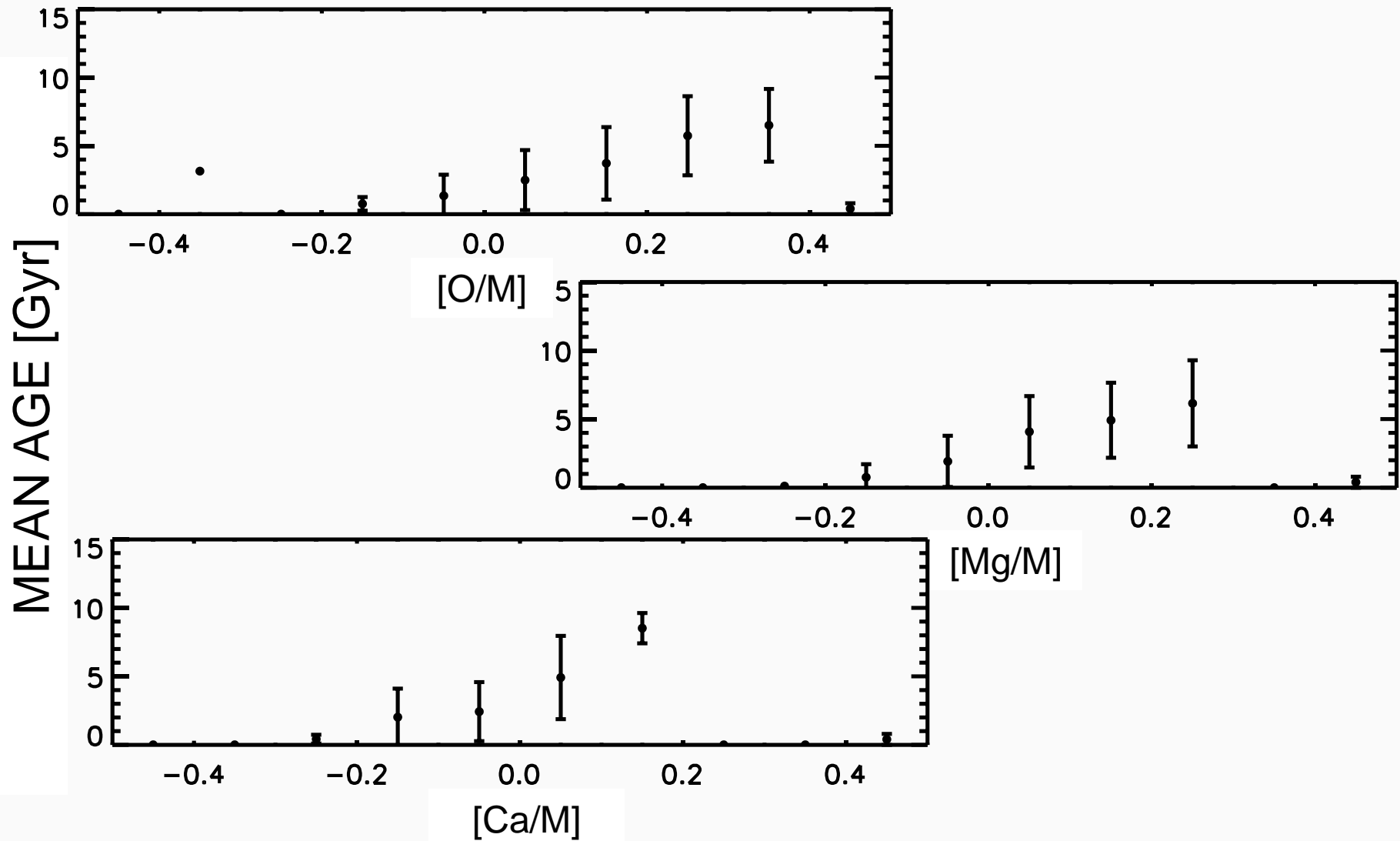
- Sample chosen from Hipparcos catalogue
- Parallax errors less than 10%
- $H < 8$ mag
- $(J-K)_0 > 0.5$
- Observed with NMSU 1m+APOGEE connection



Literature Comparison



Elemental Abundances



Waiting for Gaia



Need a larger sample to statistically examine detailed abundance variations with absolute time on a Galactic scale.

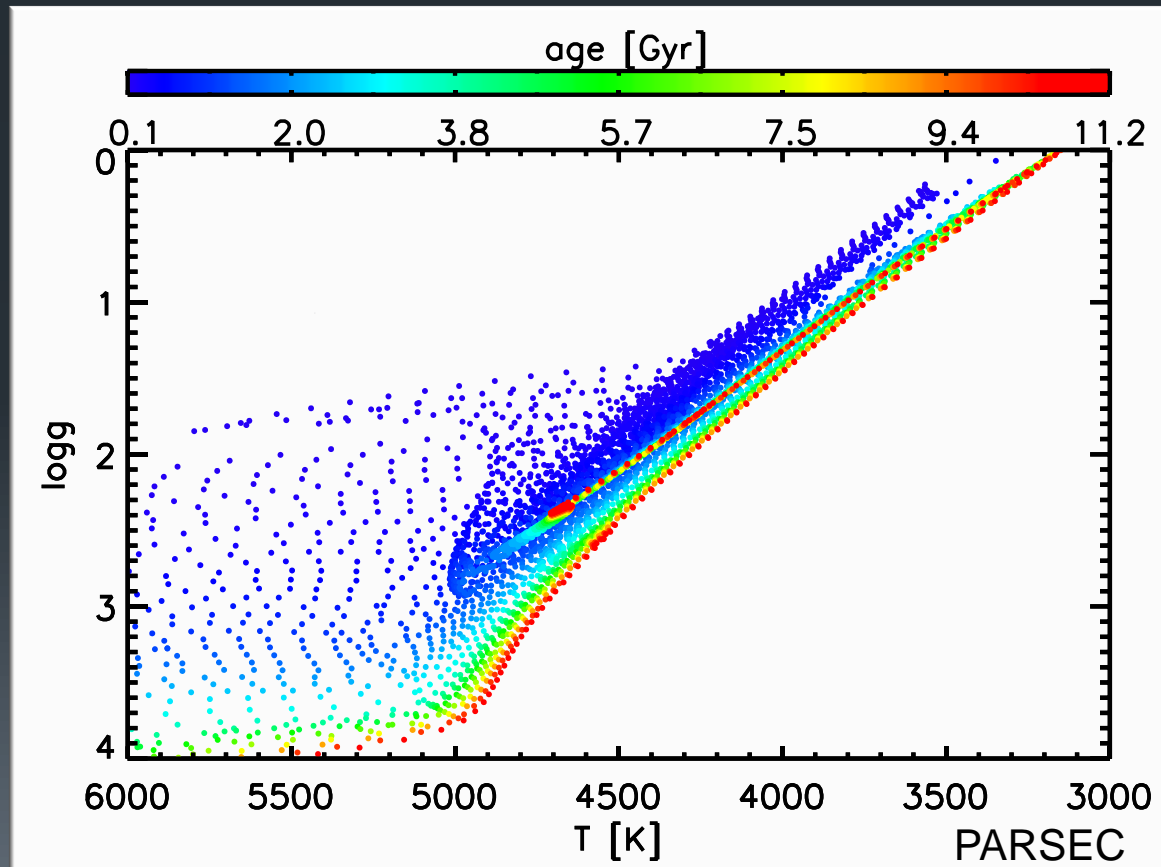
**Developing a method of age determination
with the goal of extending to full APOGEE
sample with Gaia distances**

Until then, get ages at the same time as distances

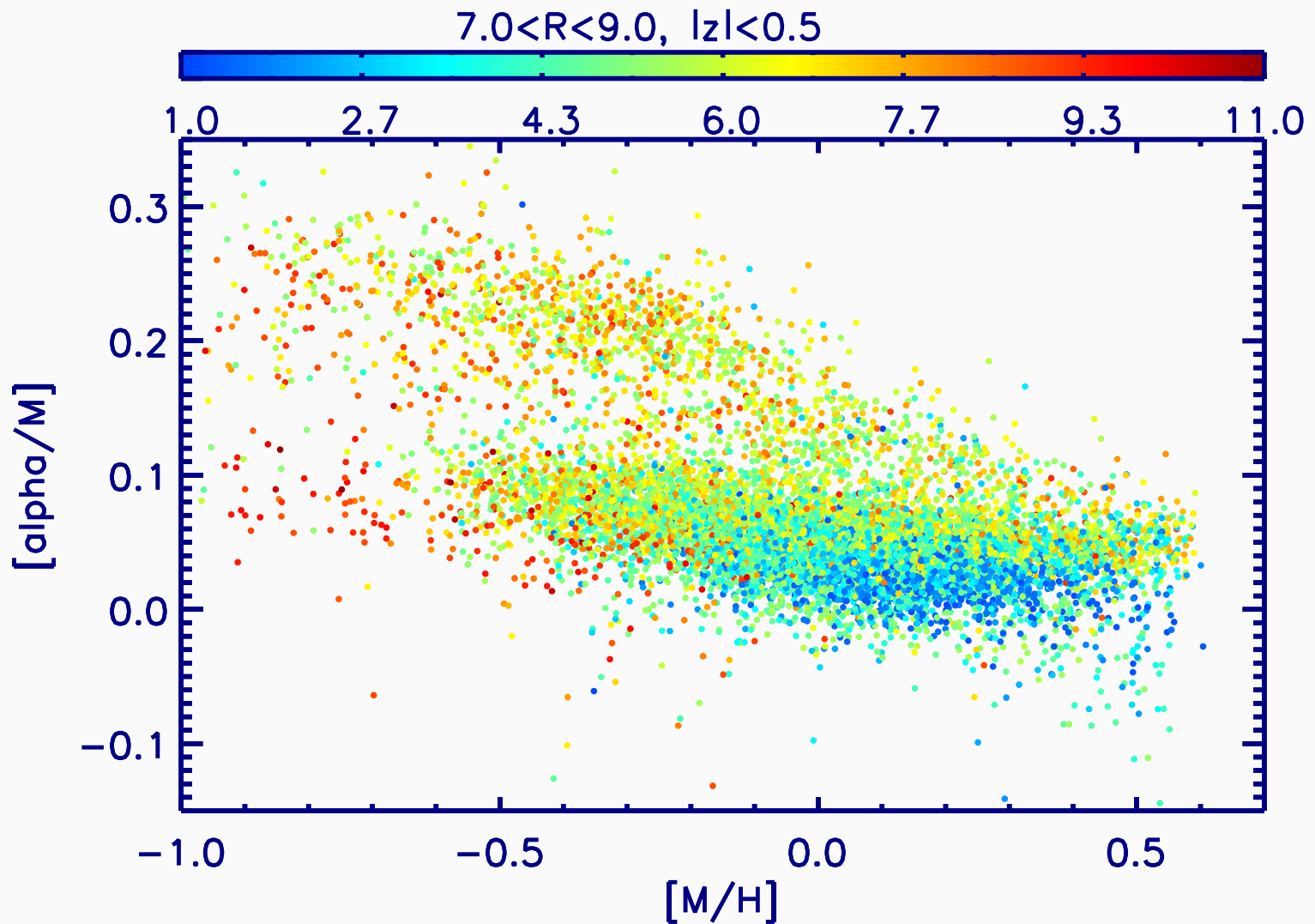
Ages Without Distances

No Distances:

- Ages without distances are possible but crude
- Need very accurate $\log g$ measurements
- Get age and distance simultaneously

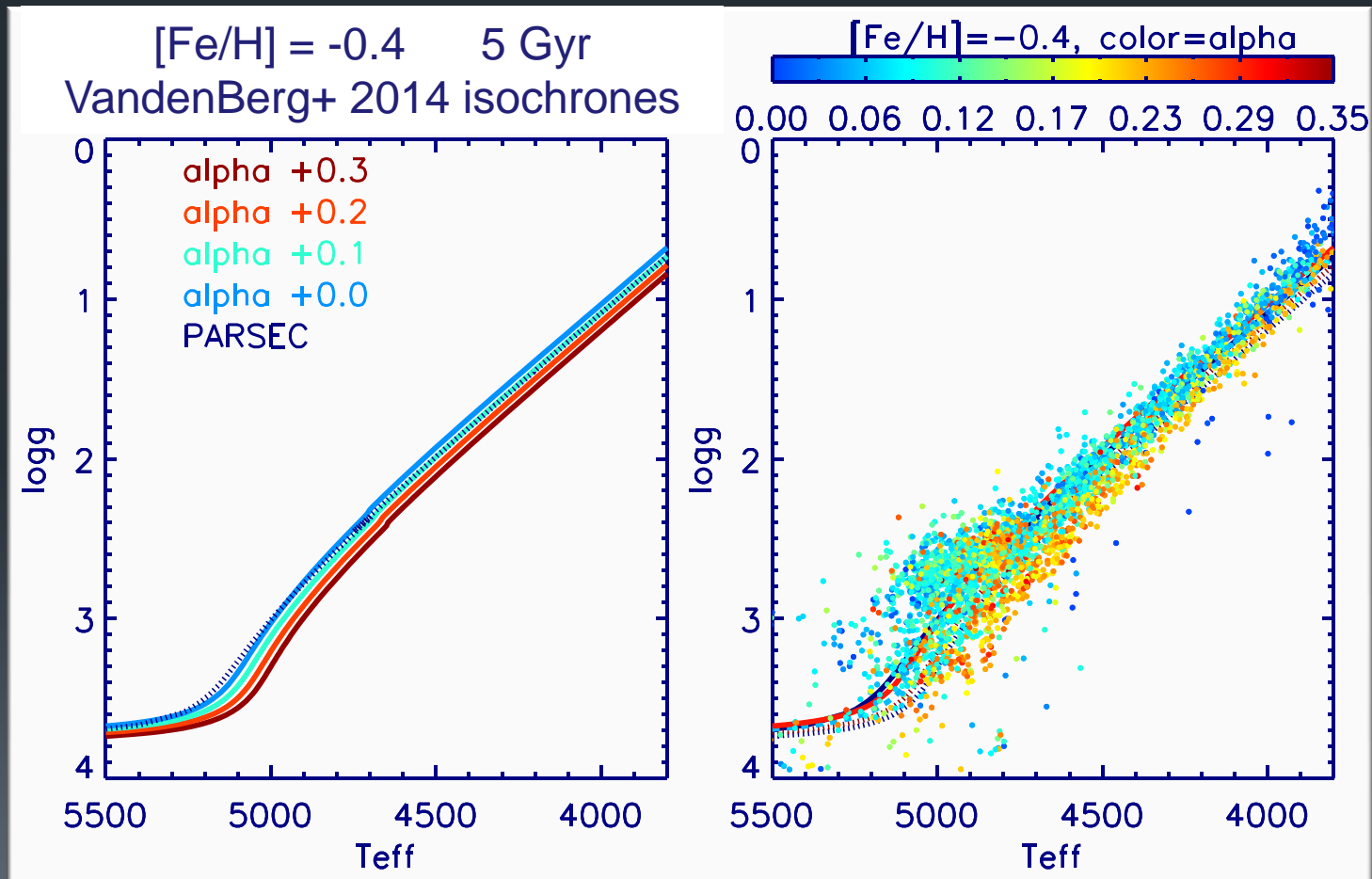


Full APOGEE Sample



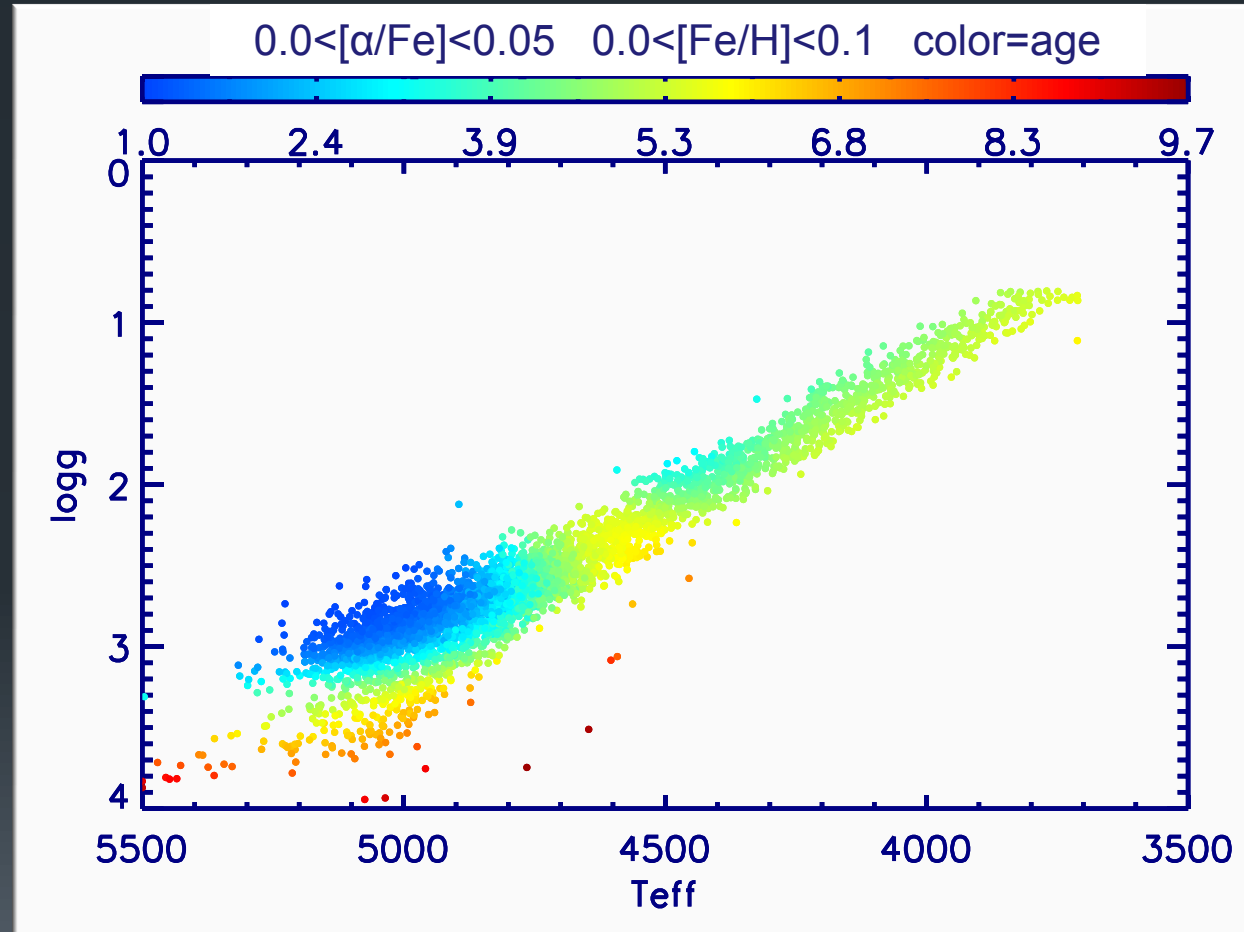
Systematic Effects

Alpha enhancement makes stars cooler...
just as is expected in older stars
We see alpha gradient in APOGEE data



Systematic Effects

- Still see systematics in age at a single metallicity and alpha
- Expect smooth age sequence across the giant branch



Conclusions



Moving towards a Galactic sample with ages and detailed abundance with APOGEE and Gaia

Challenges:

Match observations to isochrones, including alpha enhancements

Applying Precision Astrophysics to Galactic Archaeology