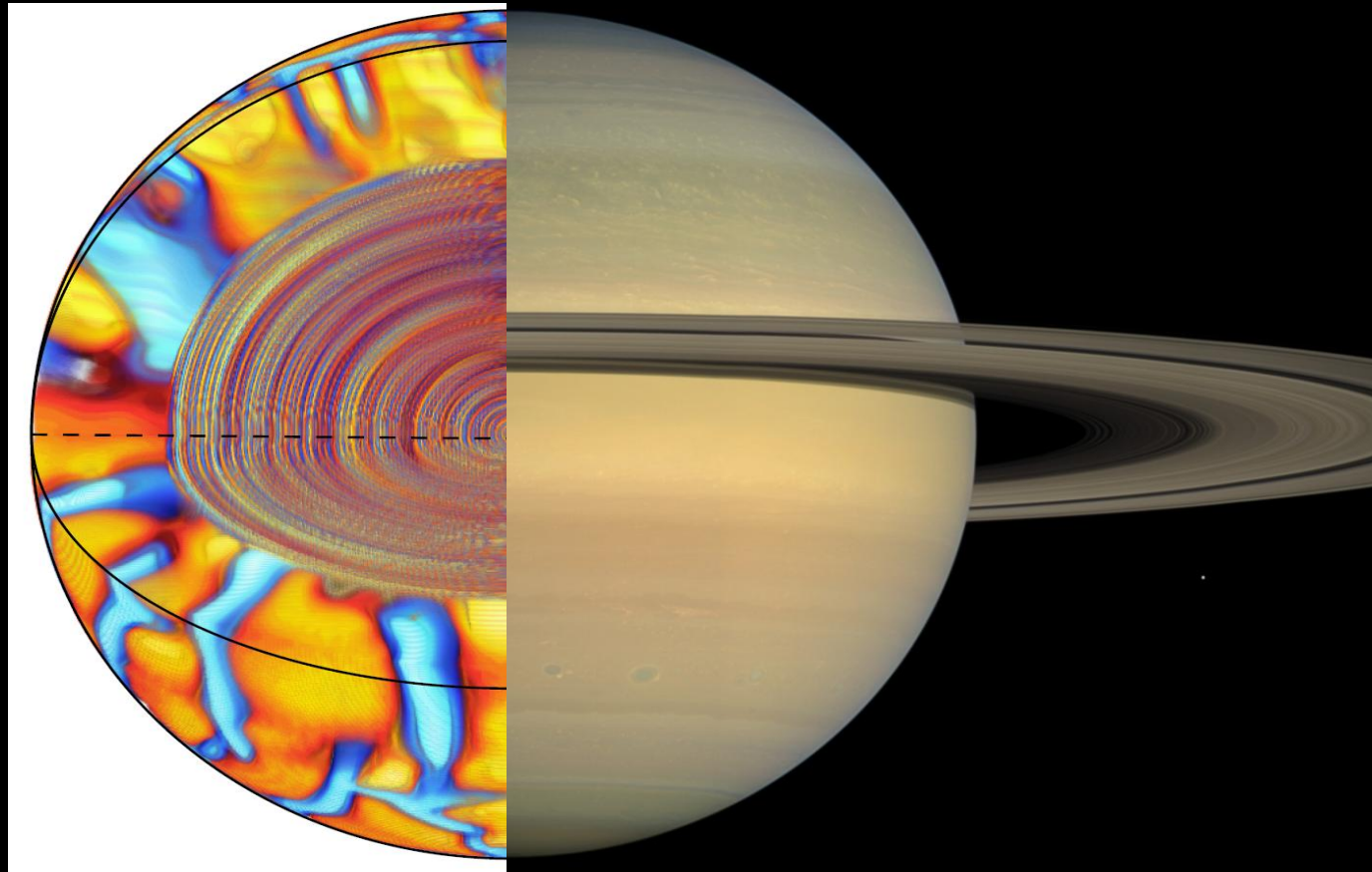


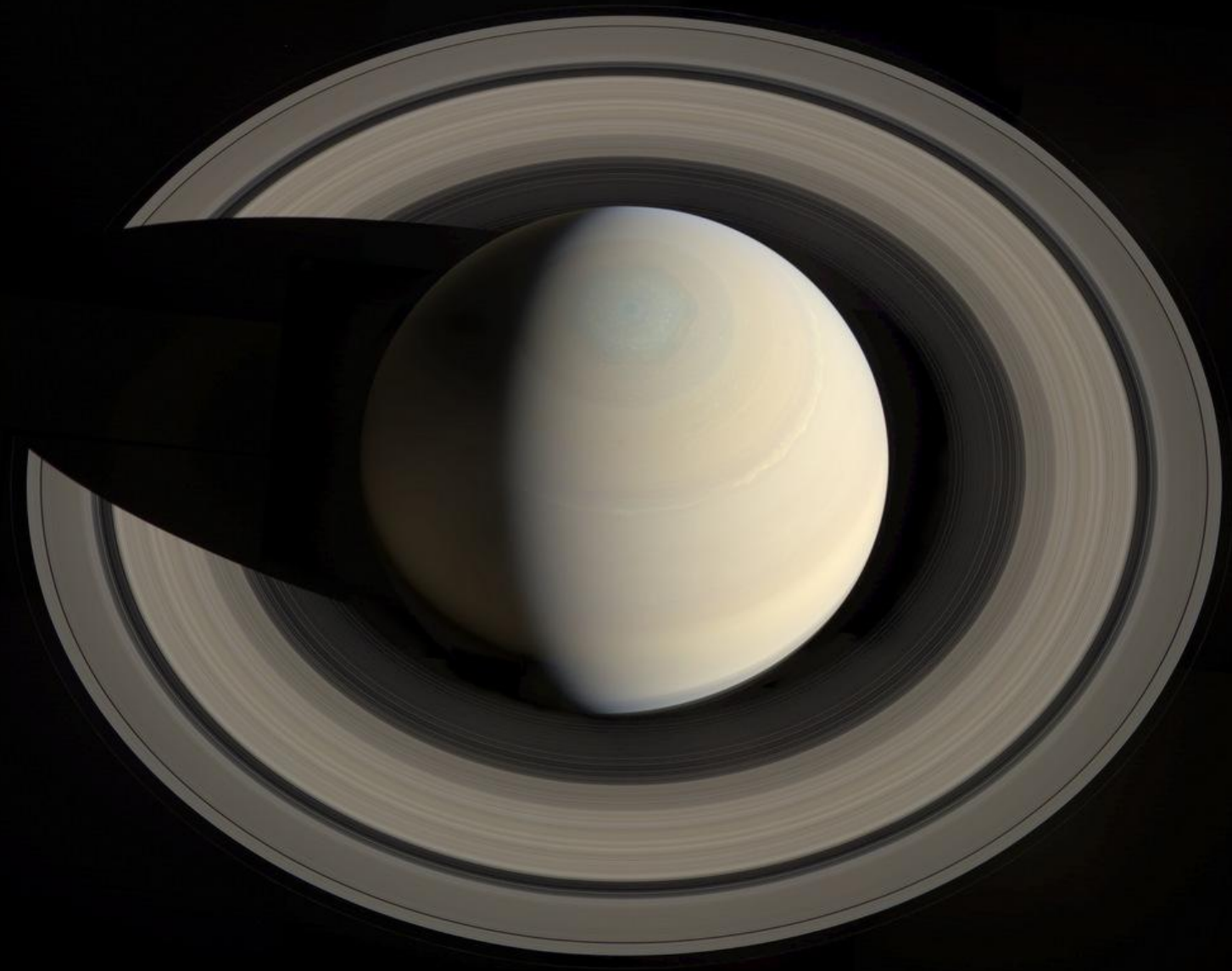
# Seismology

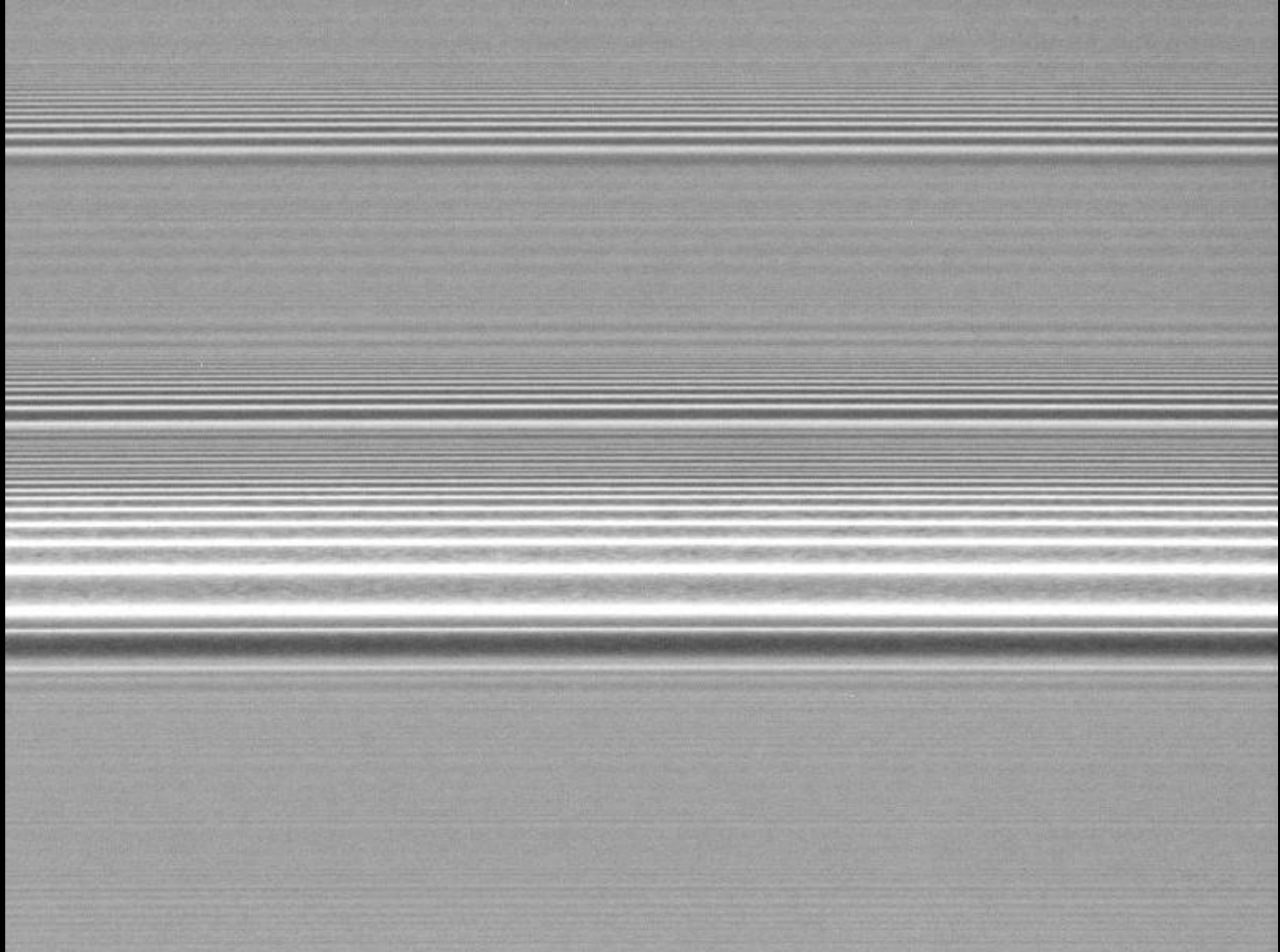
Jim Fuller

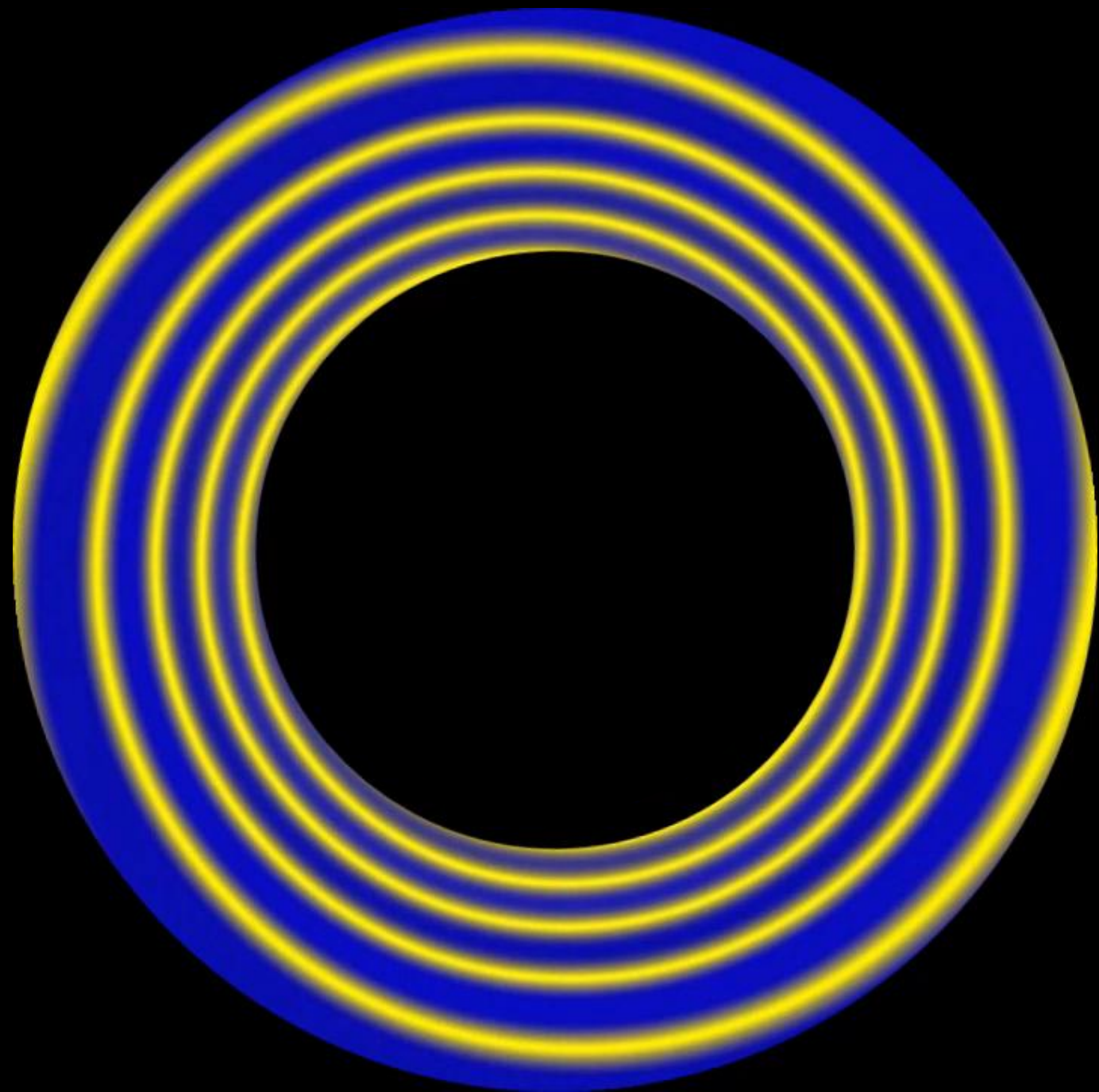
Caltech/KITP



Credits: Matt Hedman, Phil Nicholson, Mark Marley, Carolyn Porco







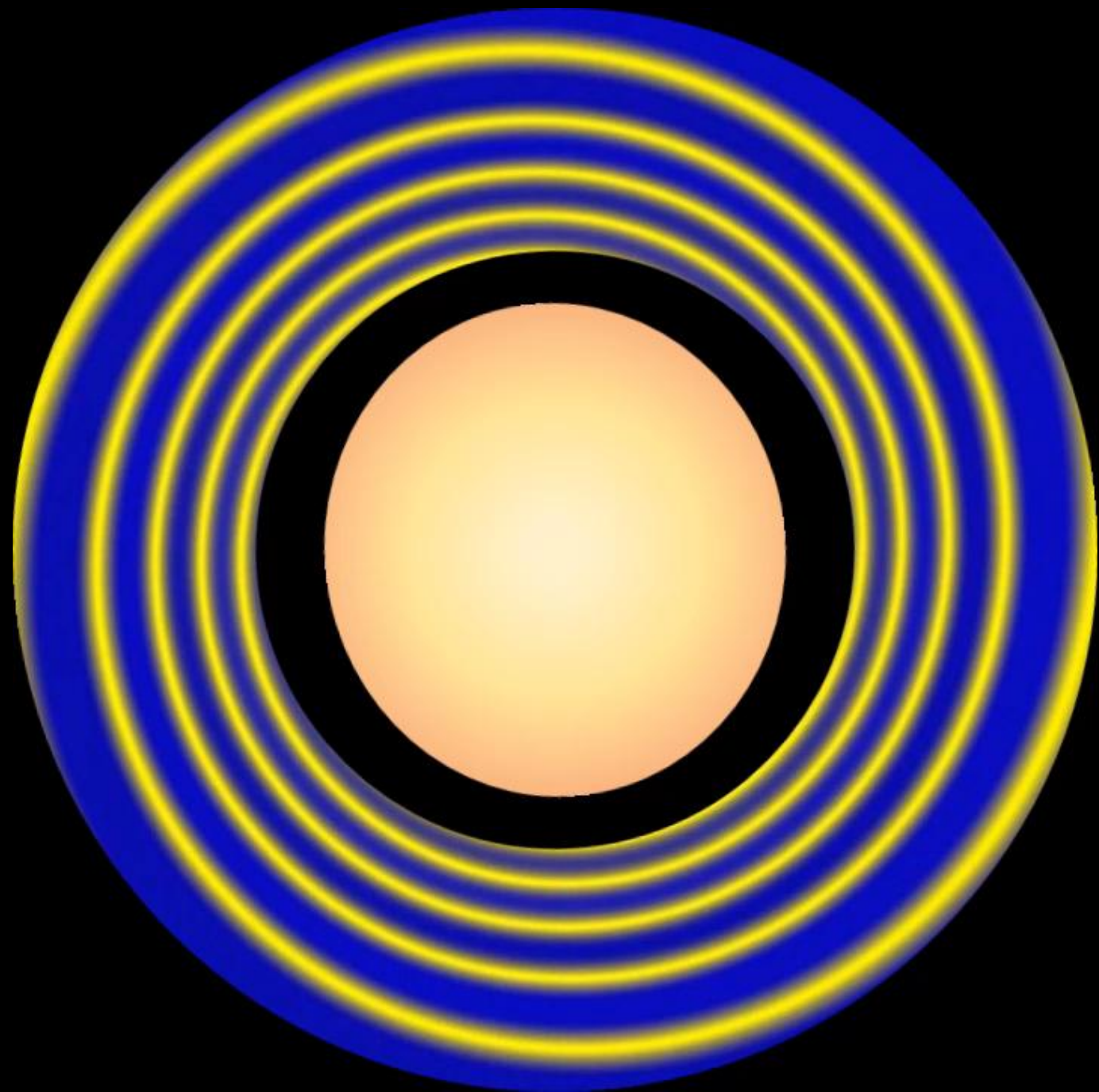
# Wave Excitation

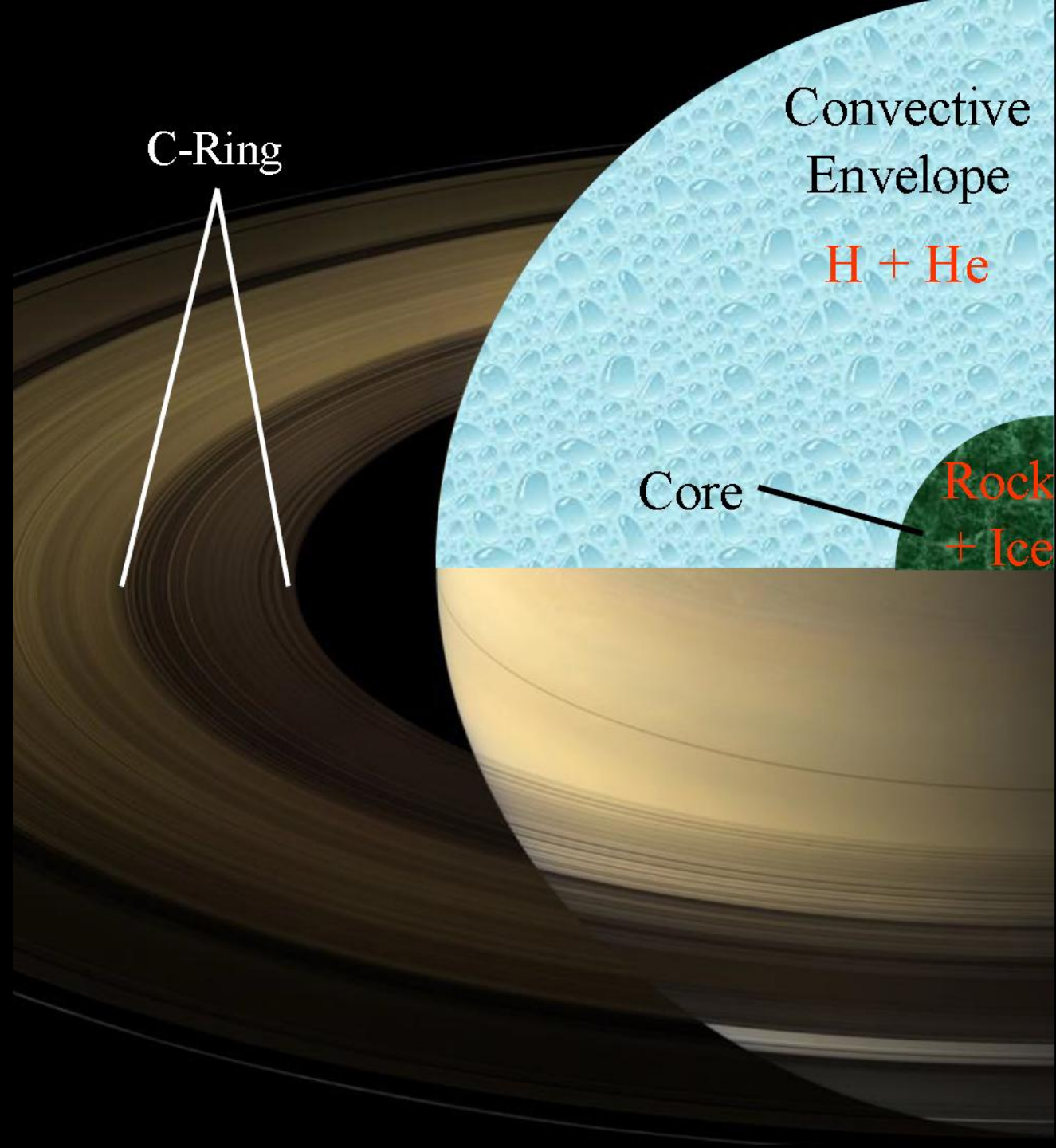
- Waves excited at Lindblad resonances

$$m(\Omega - \Omega_p) = \kappa$$

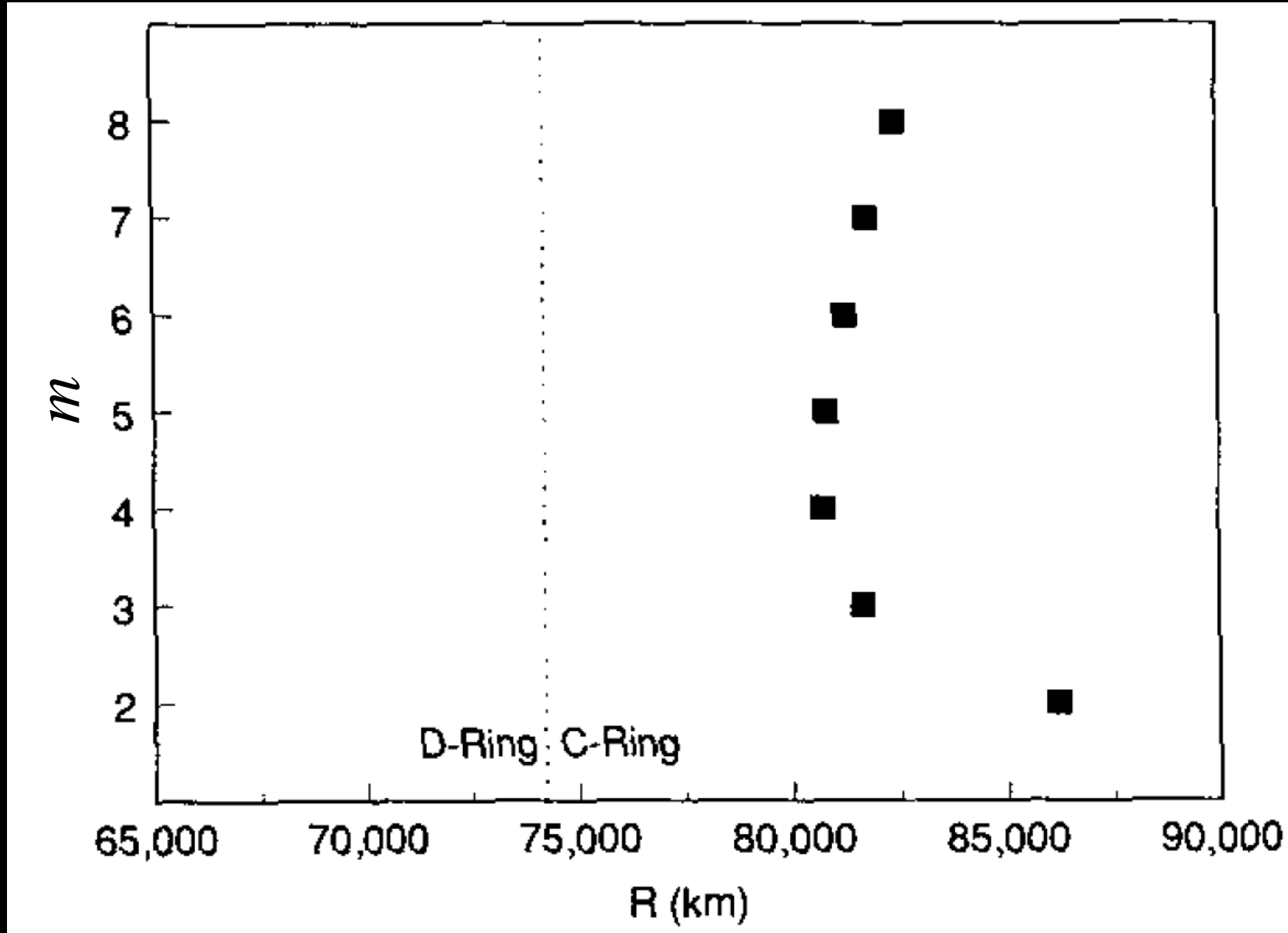
- Wave frequency tells us mode frequency

$$\Omega_p = -\sigma_\alpha / m$$





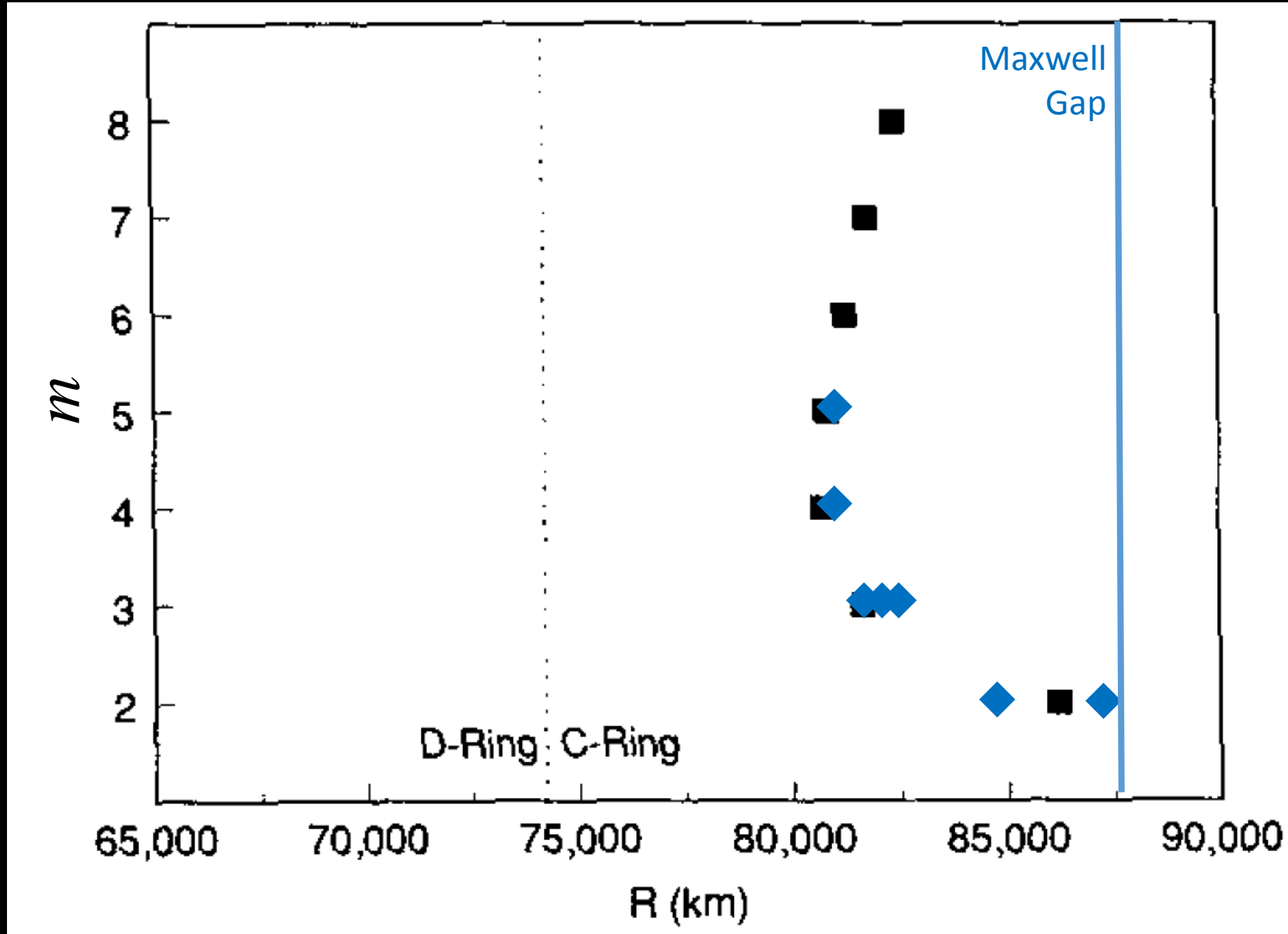
Predicted:

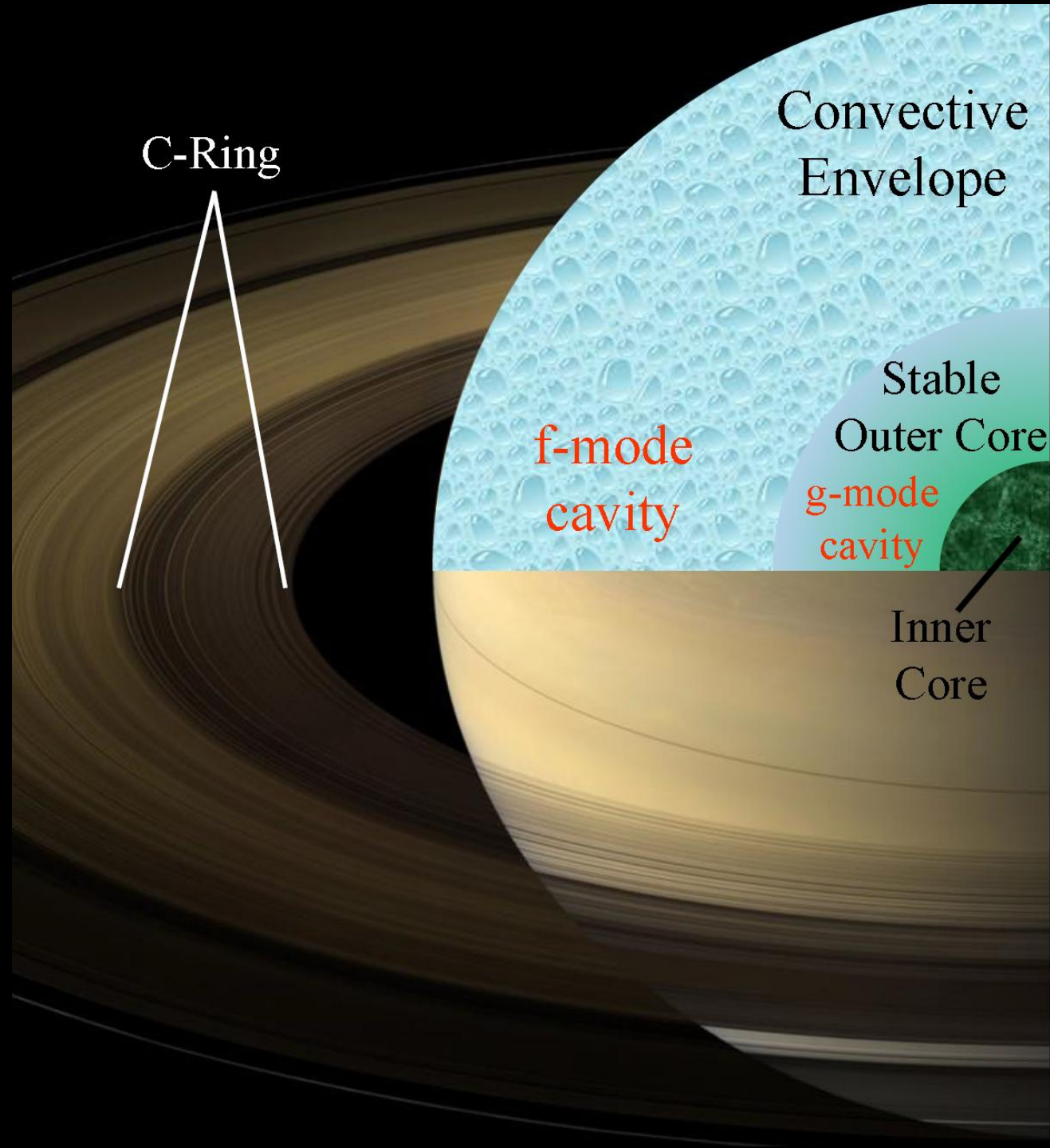


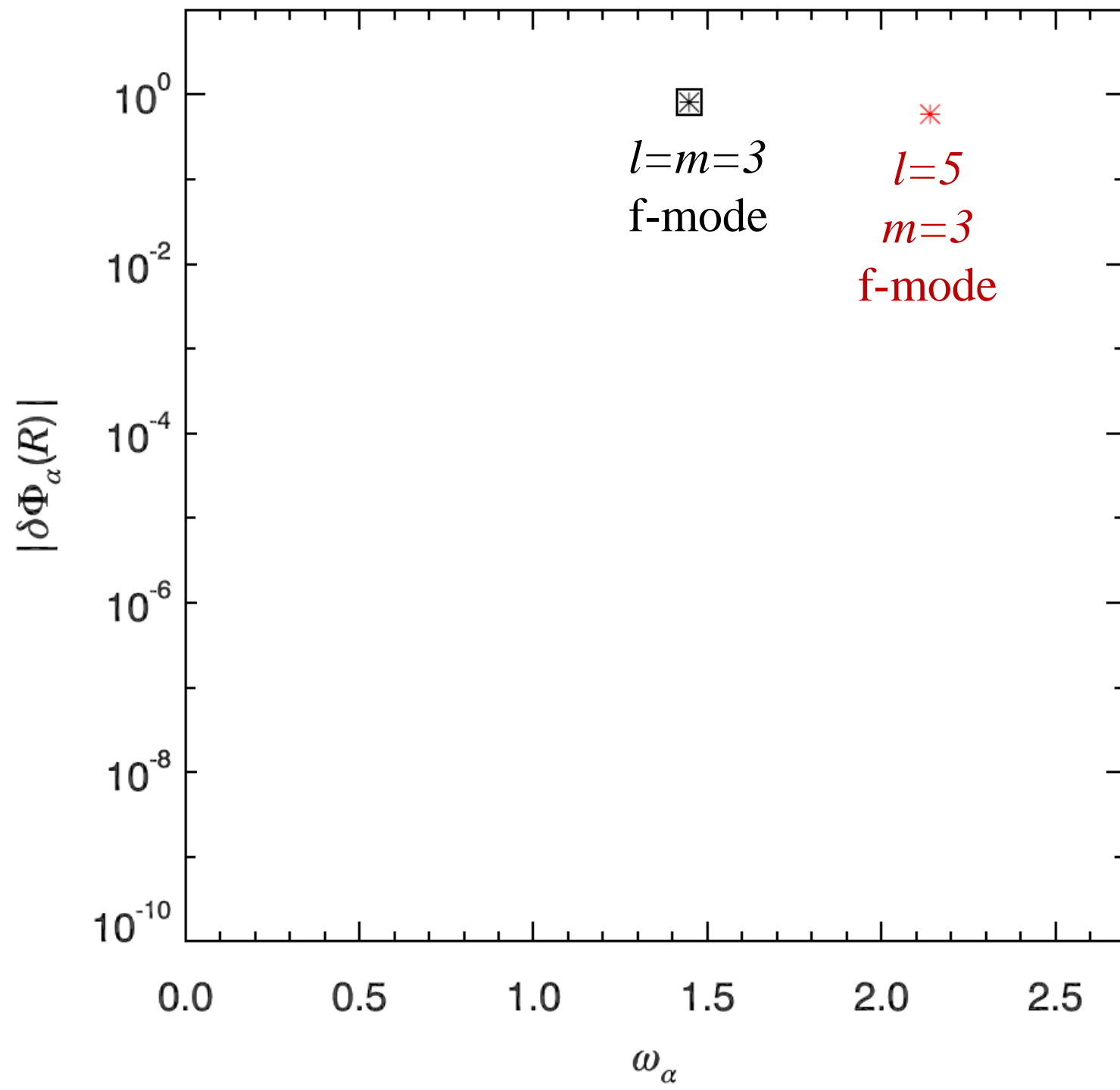
Marley & Porco 1993

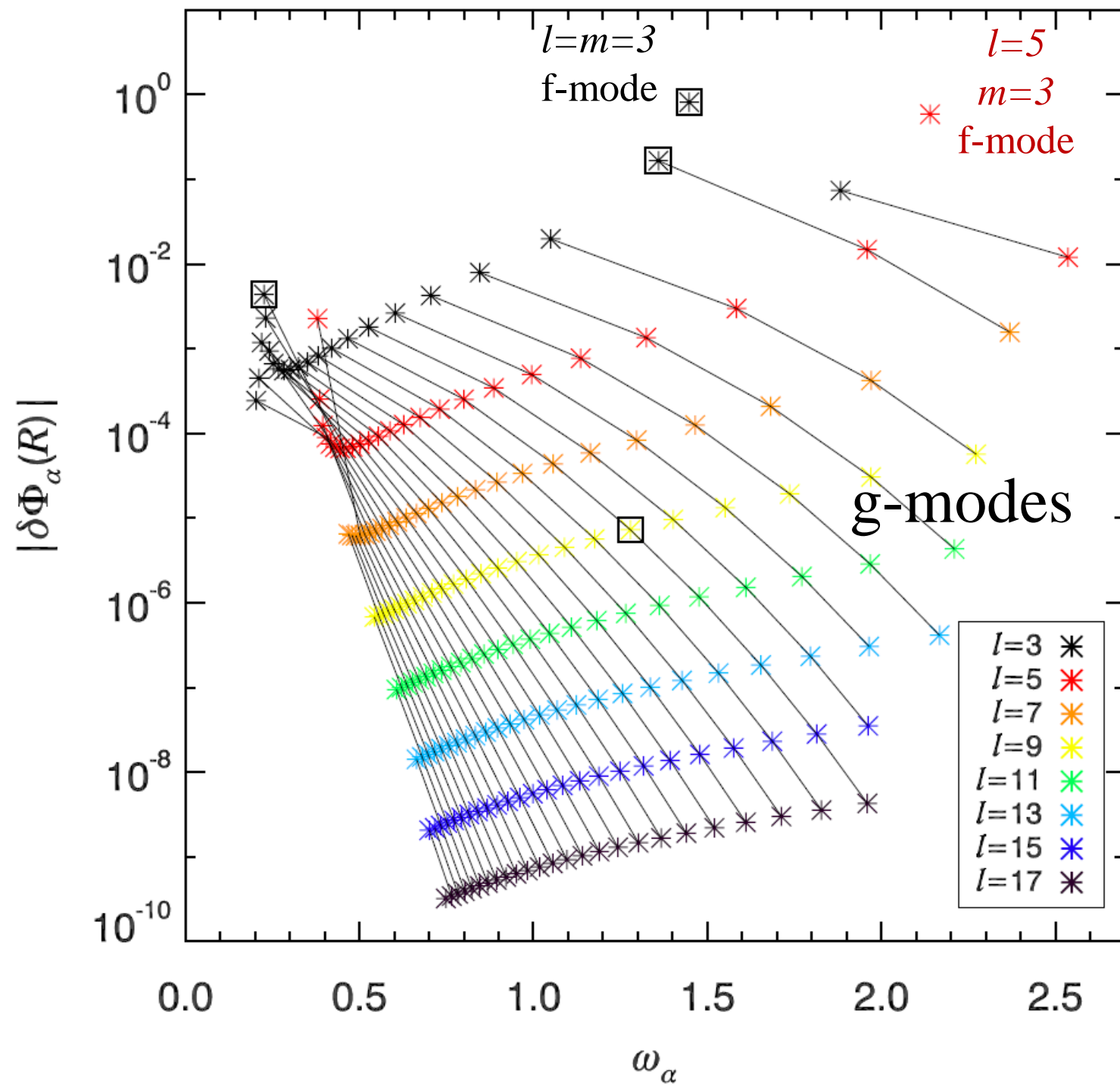


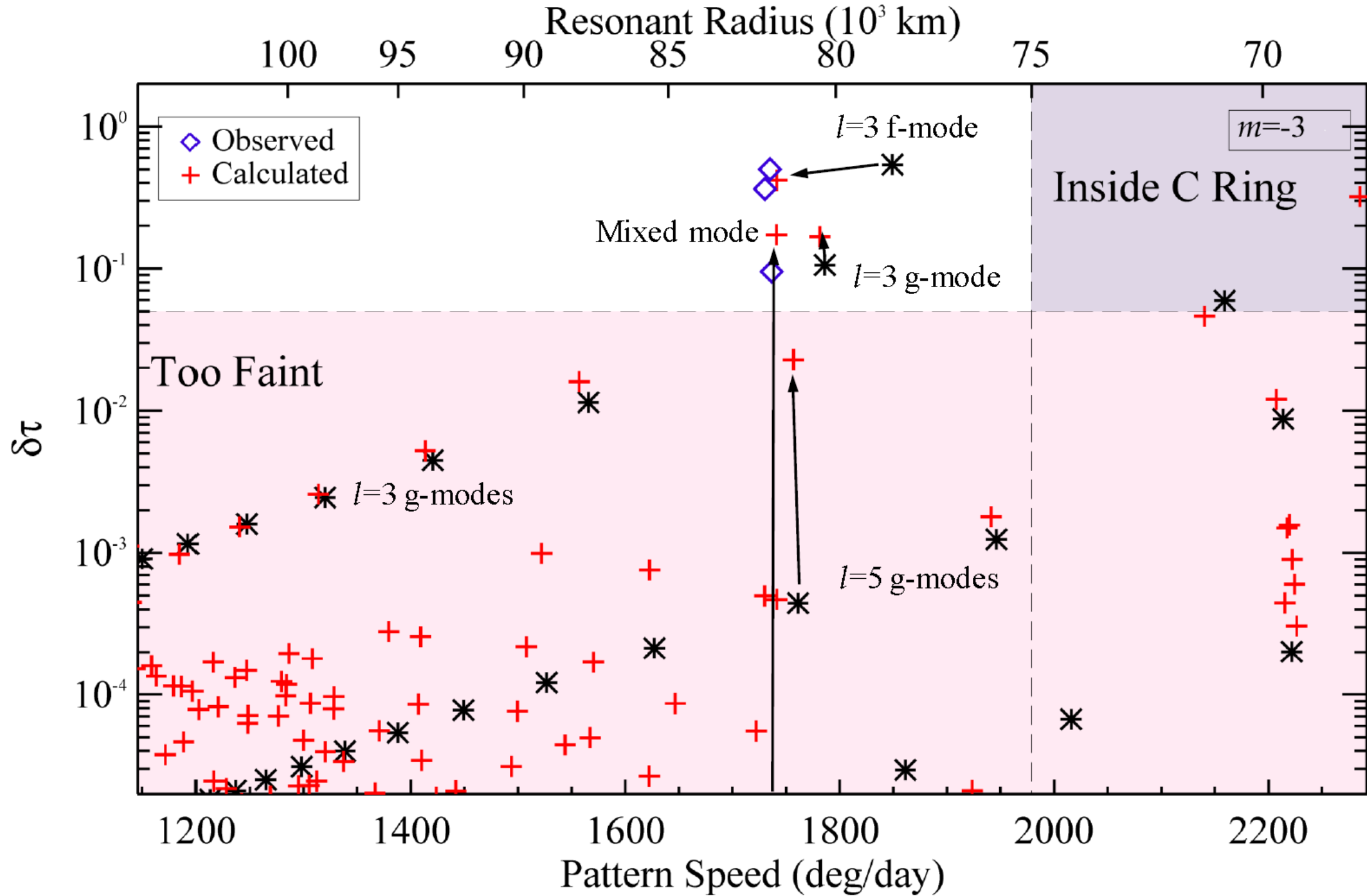
# Observed:

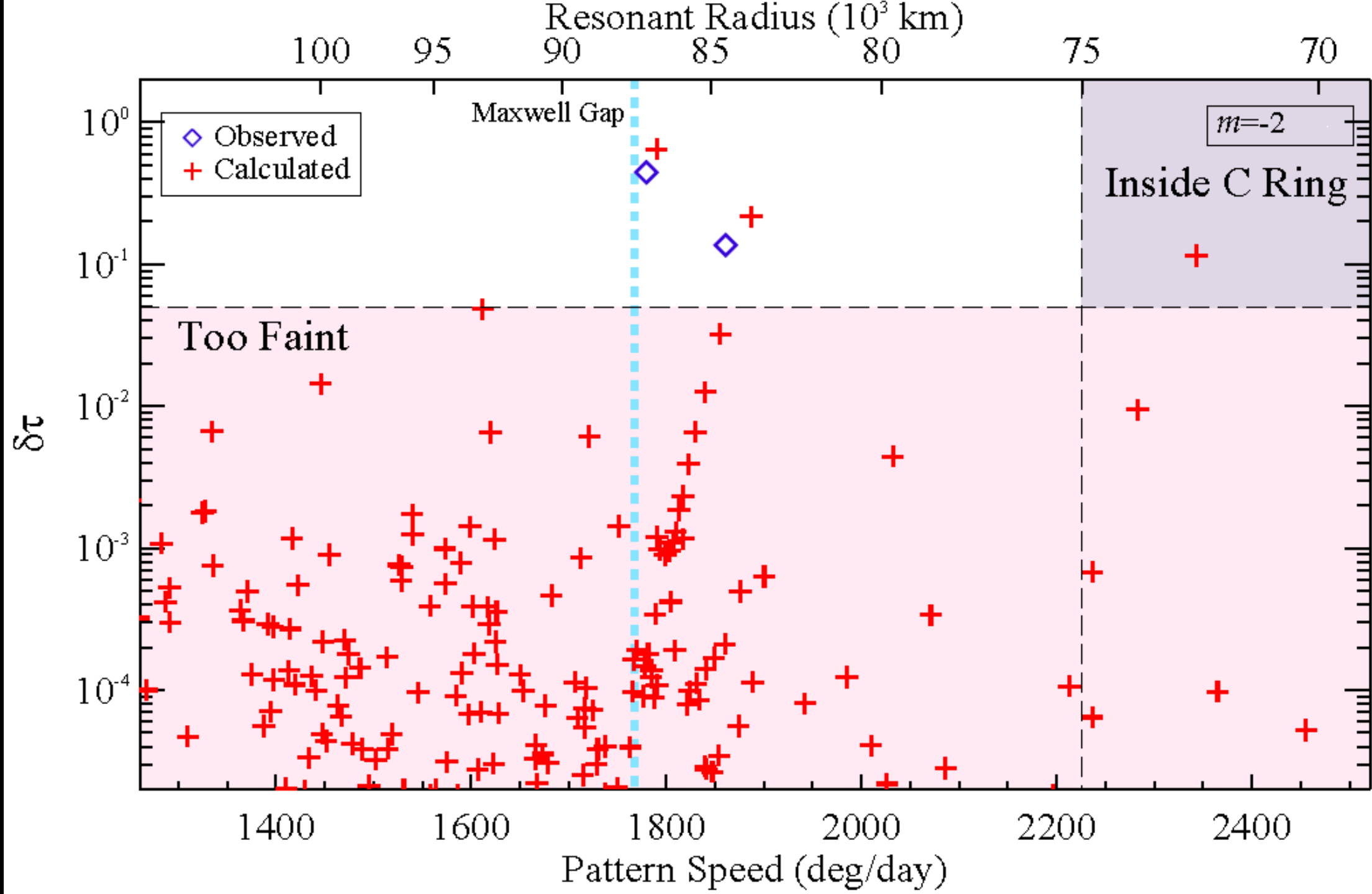








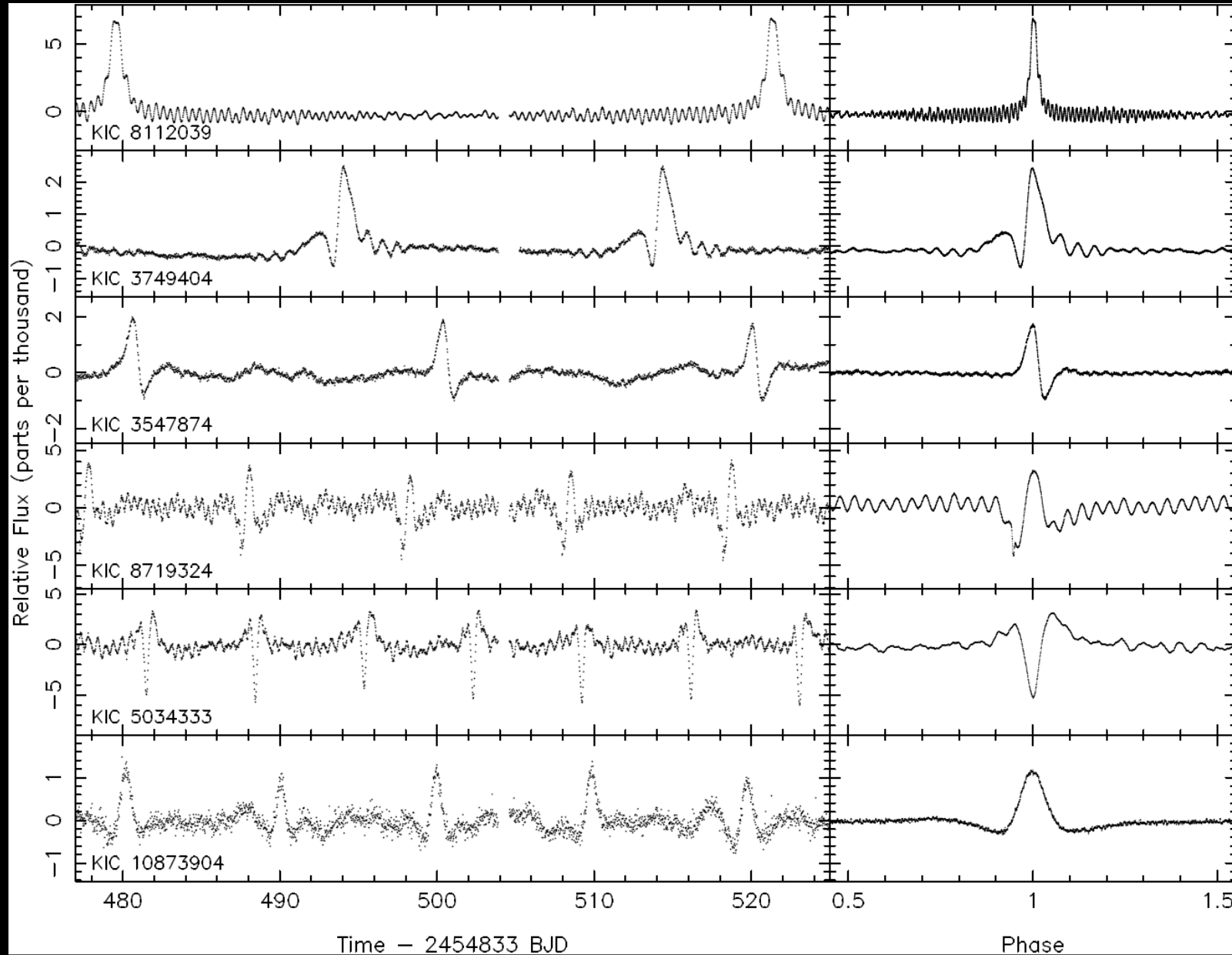




# Conclusions

- Evidence for stable stratification (non-adiabatic interior) of Saturn
- Missing ingredient: Differential rotation? Layered Convection?

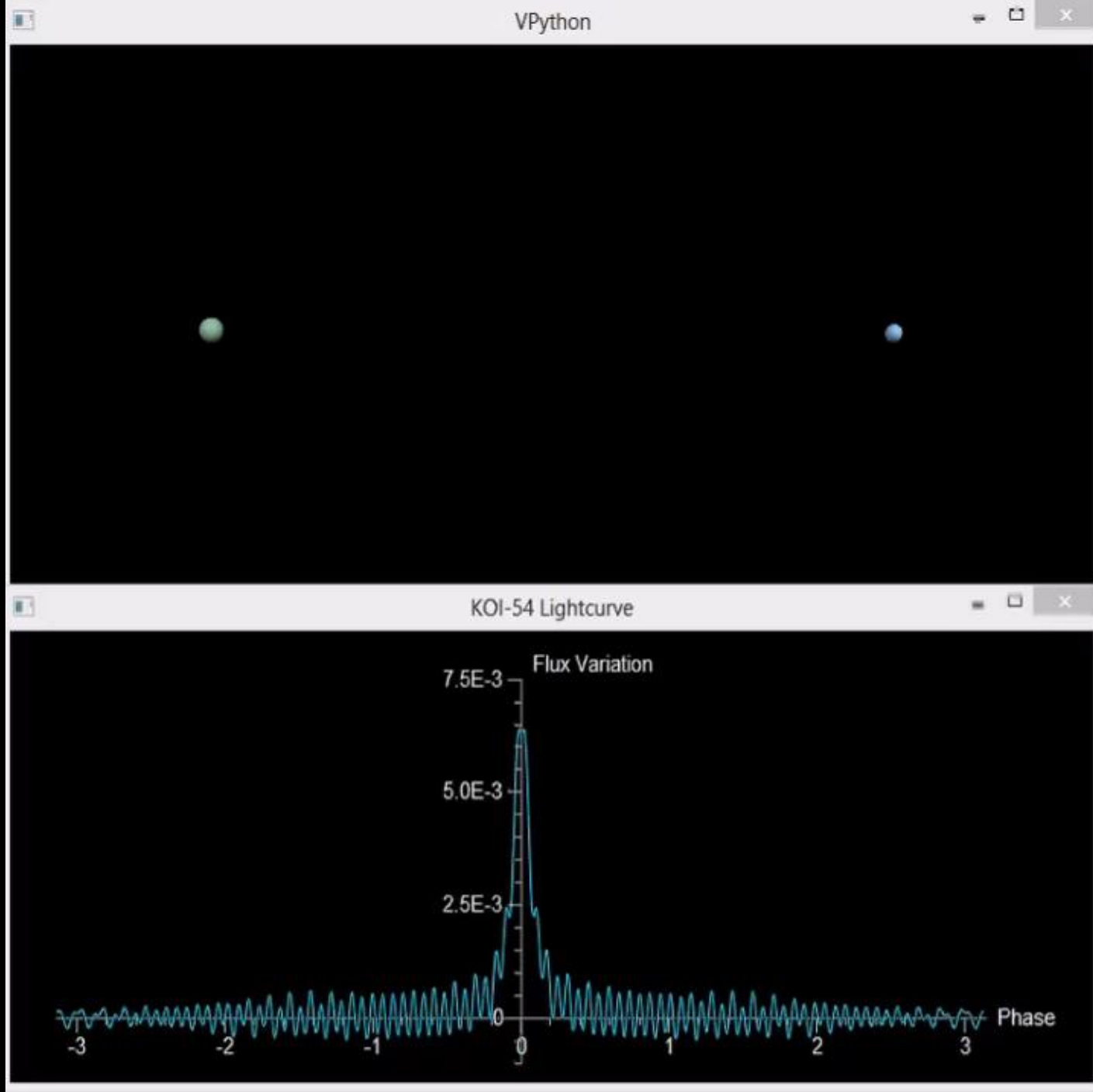
# Heartbeat Stars

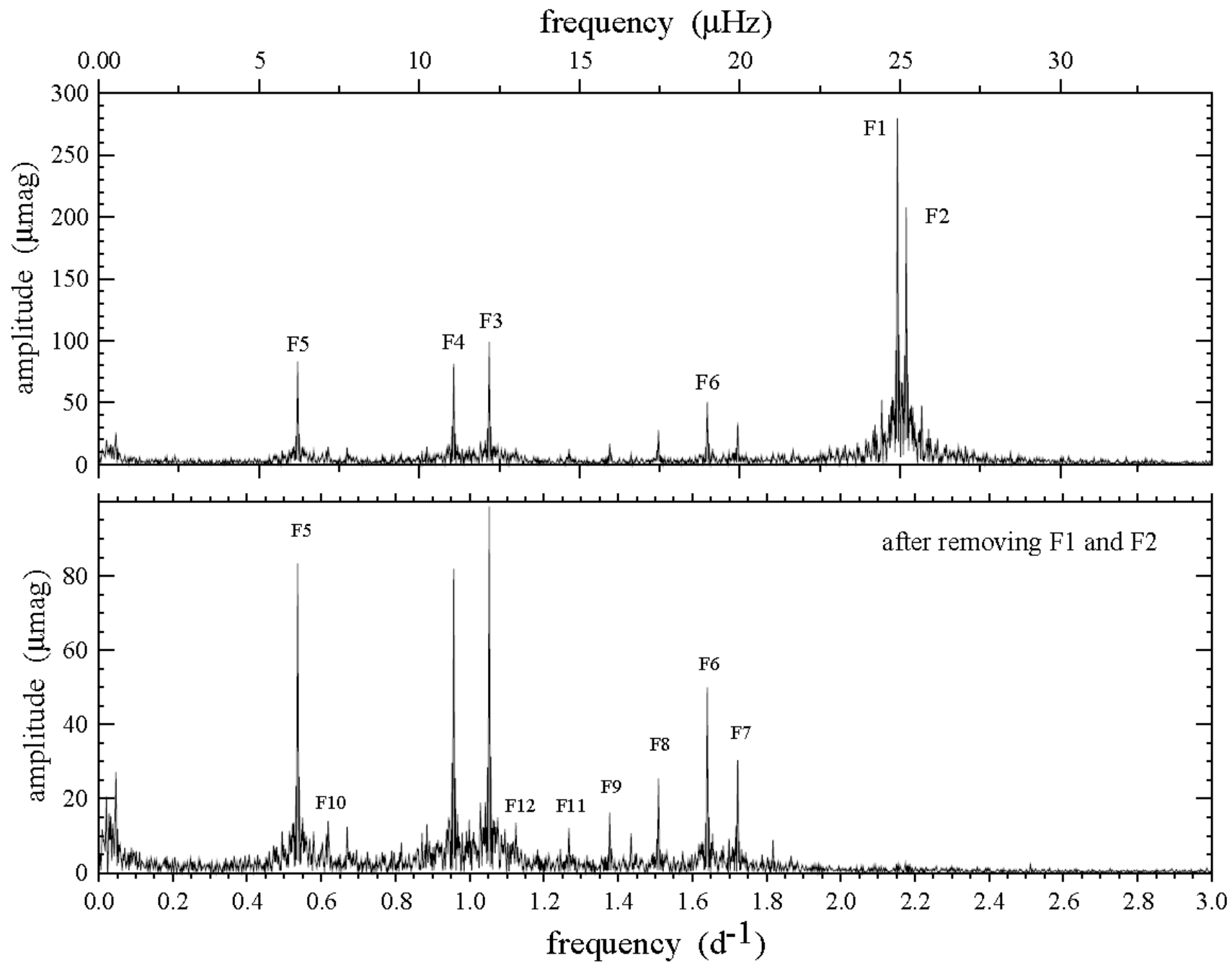


Collaborators: Kelly Hambleton, Susan Thompson, Don Kurtz, Avi Shporer, Ryan O'Leary



Heartbe





Exact Integer Harmonics!

# Tidally Induced Luminosity Fluctuations

- Tidal forcing can be written as sum over orbital harmonics

$$U_{lm} = -\frac{GM'W_{lm}r^l}{a^{l+1}}Y_{lm}(\theta, \phi_i) \sum_{N=-\infty}^{\infty} F_{Nm} e^{-iN\Omega t}$$

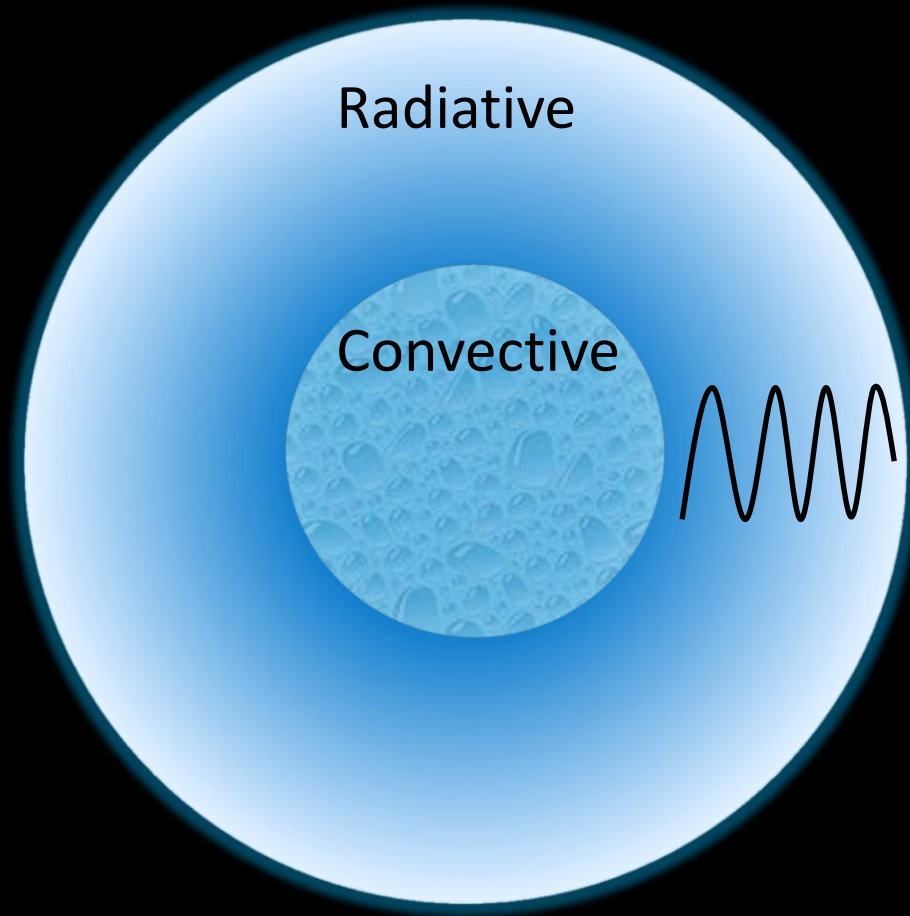
- Luminosity fluctuations occur at integer harmonics

$$\frac{\Delta L_{|N|}}{L} \simeq A_{\beta} \sin(N\Omega t + \Delta_{\beta})$$

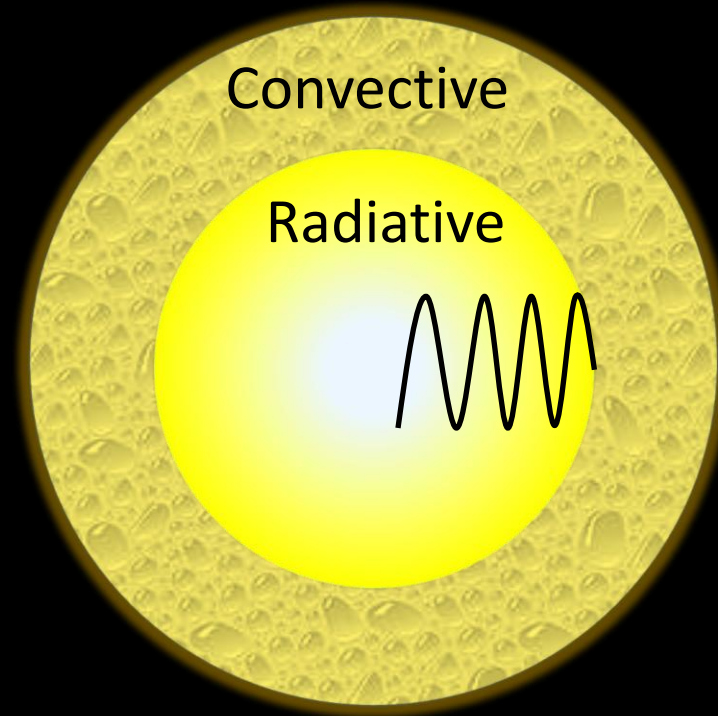
$$A_{\beta} = \epsilon_{lm} h_{klm} Q_{\beta} F_{Nm} P_{lm}(\theta_o) |L_{\beta}| \frac{\omega_{Nm}}{\sqrt{(\omega_{\beta} - \omega_{Nm})^2 + \gamma_{\beta}^2}}$$

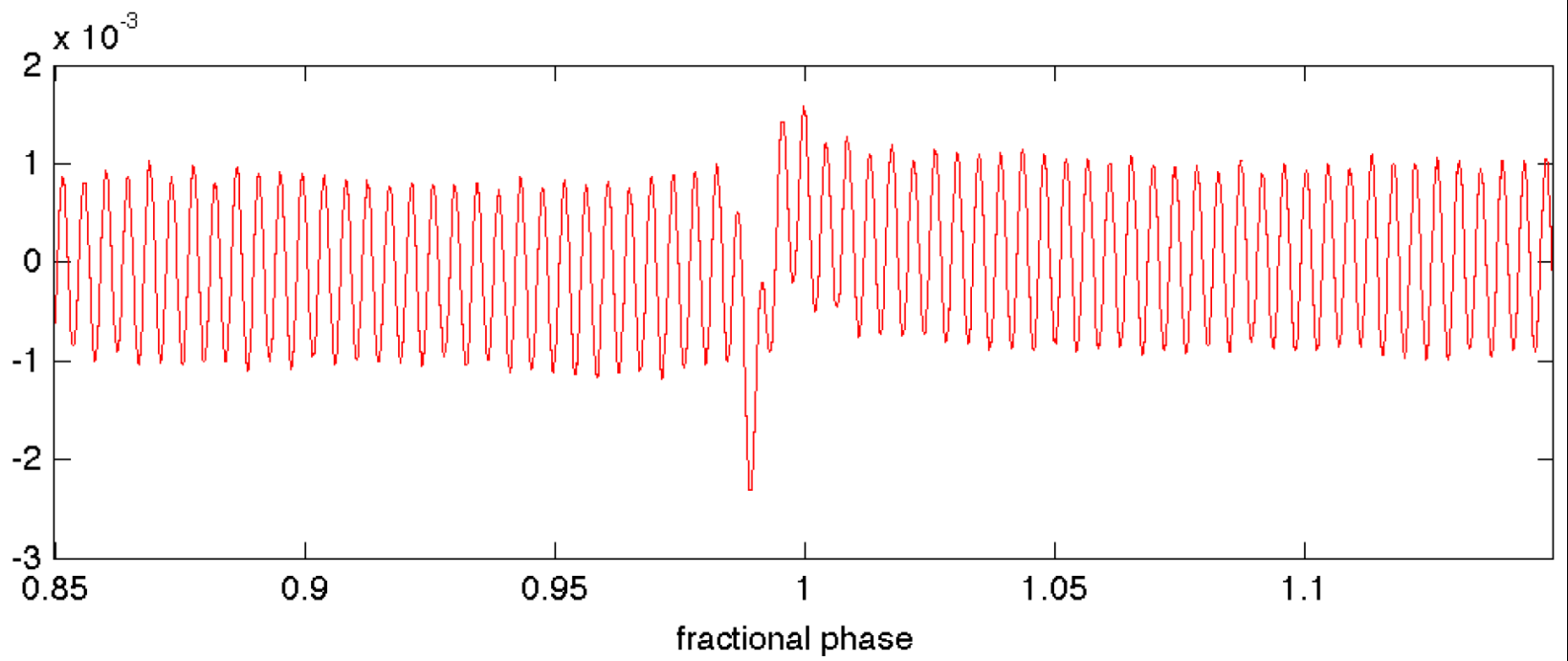
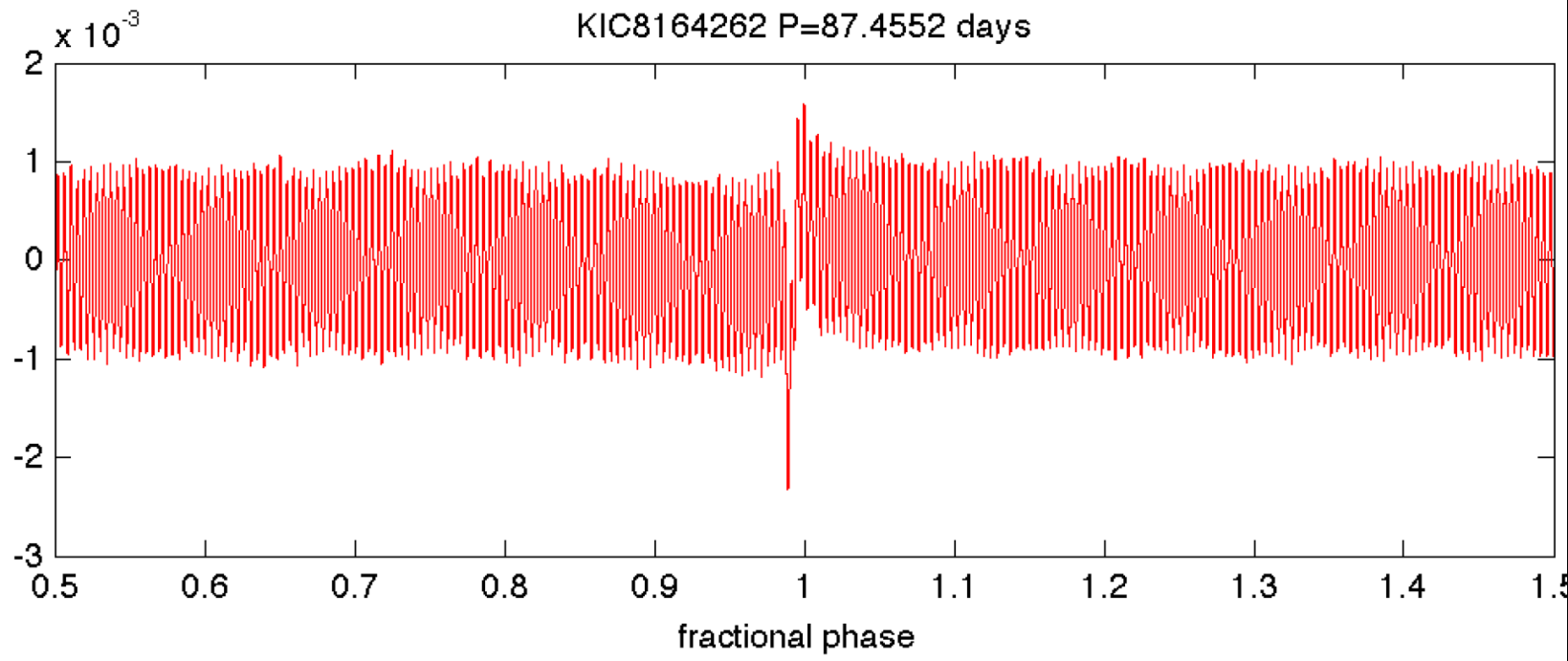
# Gravity waves in stars

High Mass Star



Low Mass Star





Preliminary:

$M=1.7 M_{\text{sun}}$

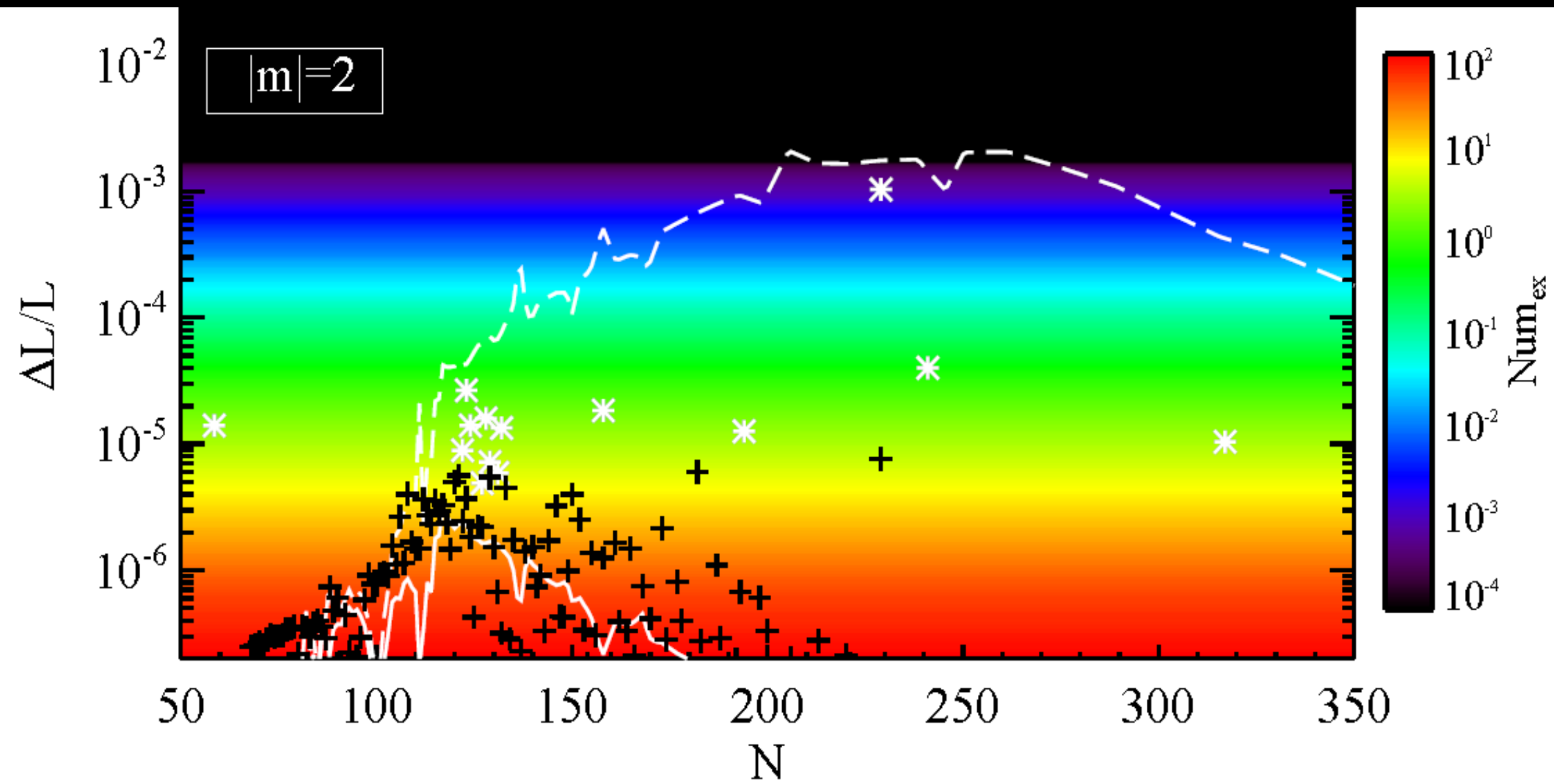
$R=2.05 R_{\text{sun}}$

$T=7000 \text{ K}$

$M'=0.3 M_{\text{sun}}$

$e=0.89$

Credits: Kelly Hambleton,  
Susan Thompson, Don Kurtz



# Resonance Locking

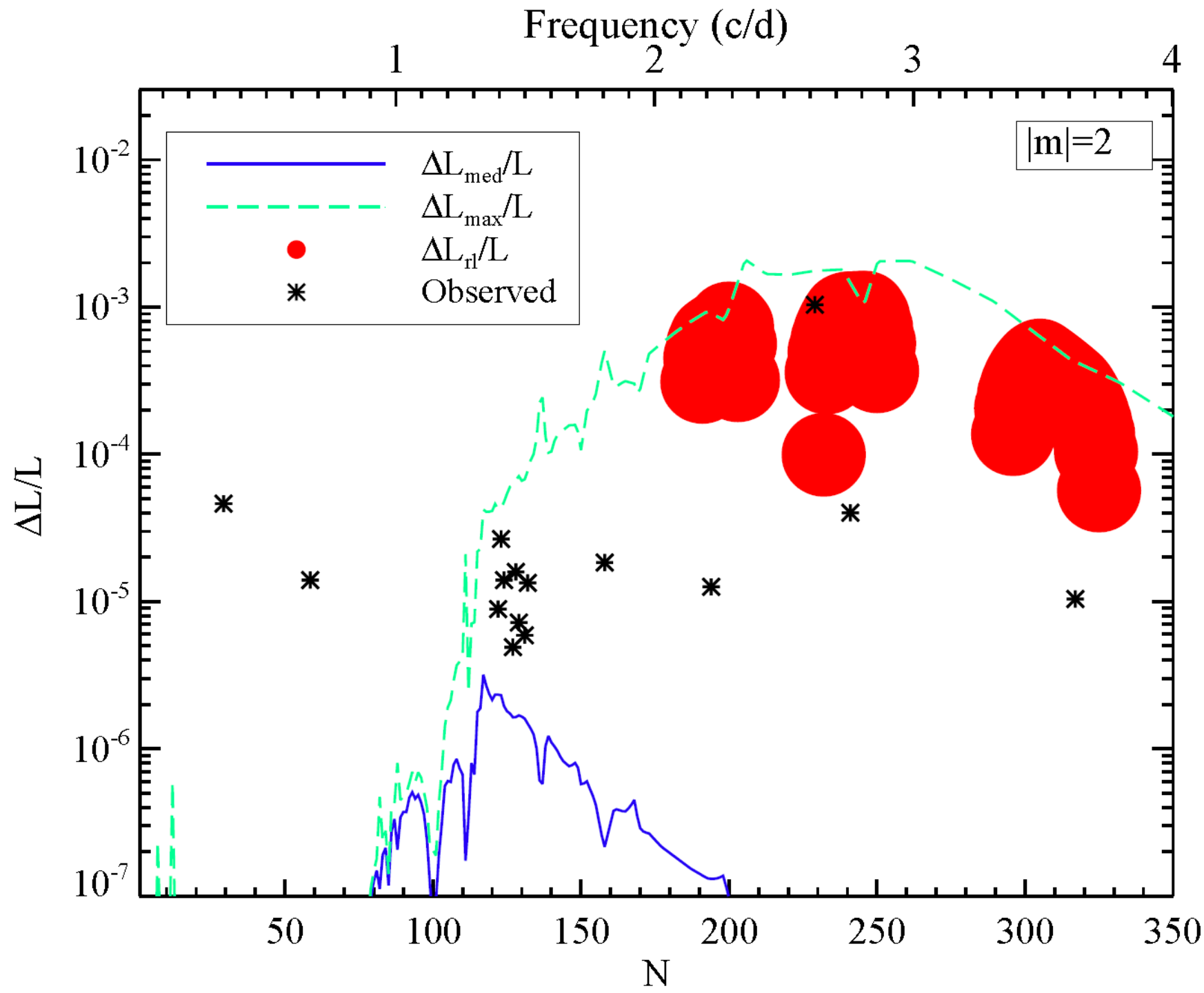
- Stellar mode frequencies evolve with time
- When they cross a resonance with tidal forcing frequency, mode can become locked in resonance such that

$$\dot{\sigma}_{\alpha} \simeq \dot{\sigma}_N = N\dot{\Omega}$$

- In this case, tidally excited mode amplitude can be calculated:

$$|a_{\alpha}|_{\text{lock}} = \frac{1}{2} \left[ \frac{N\Omega}{N\Omega - m\Omega_s} \frac{1}{\chi\gamma_{\alpha}t_{\alpha,\text{ev}}} \right]^{1/2}$$

- Resonantly locked amplitudes are large, generate enhanced tidal dissipation



Effective tidal  
quality factor:

$$Q \sim 2 \times 10^4$$



# Conclusions

- Coupled stellar/orbital evolution can lead to resonance locking, enhanced tidal dissipation
- Much more to be done!