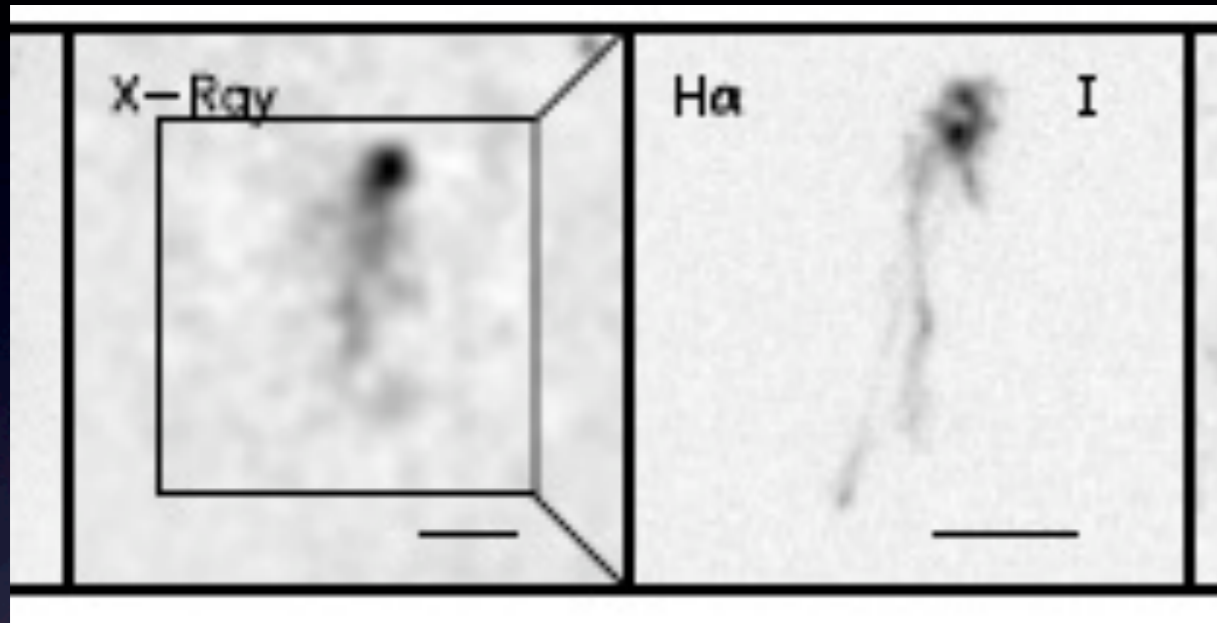


Thermal Instability & Feedback

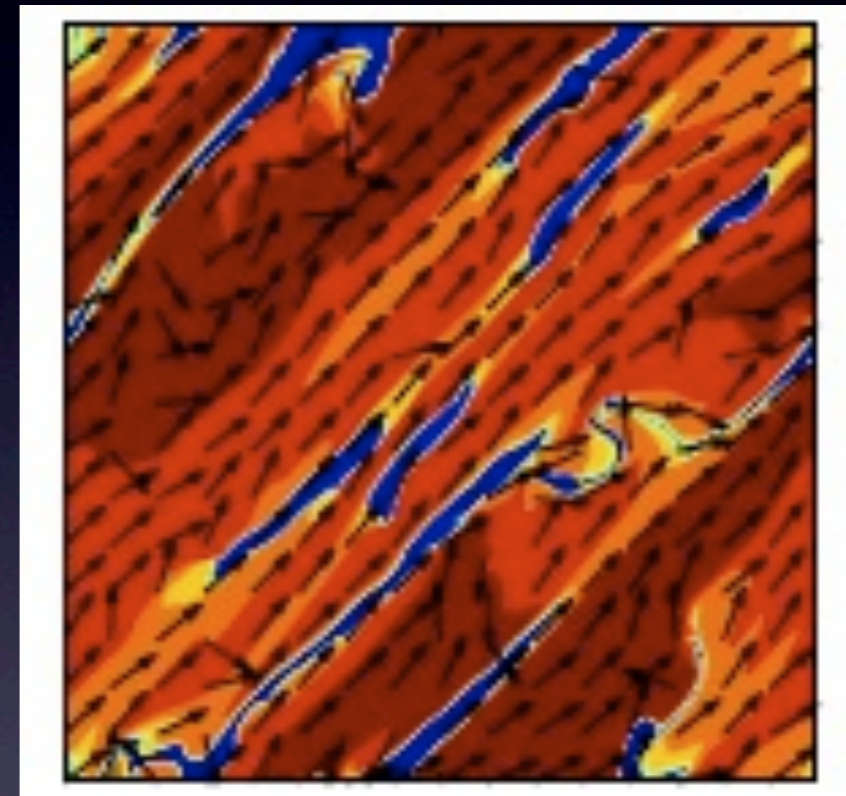
Prateek Sharma, Mike McCourt, Eliot Quataert, Ian
Parrish

Local TI in ICM

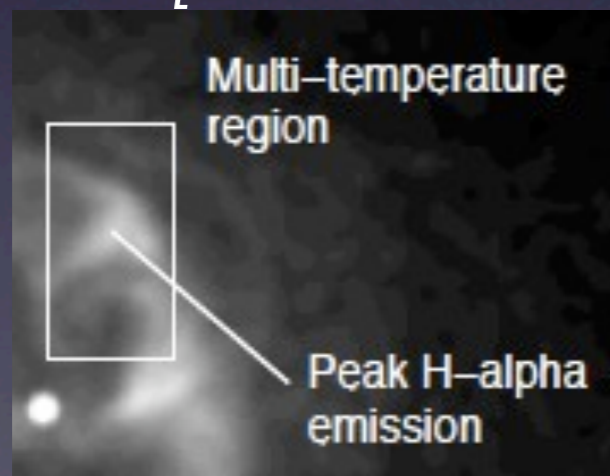
[McDonald et al. 2010]



[Sharma et al. 2010]



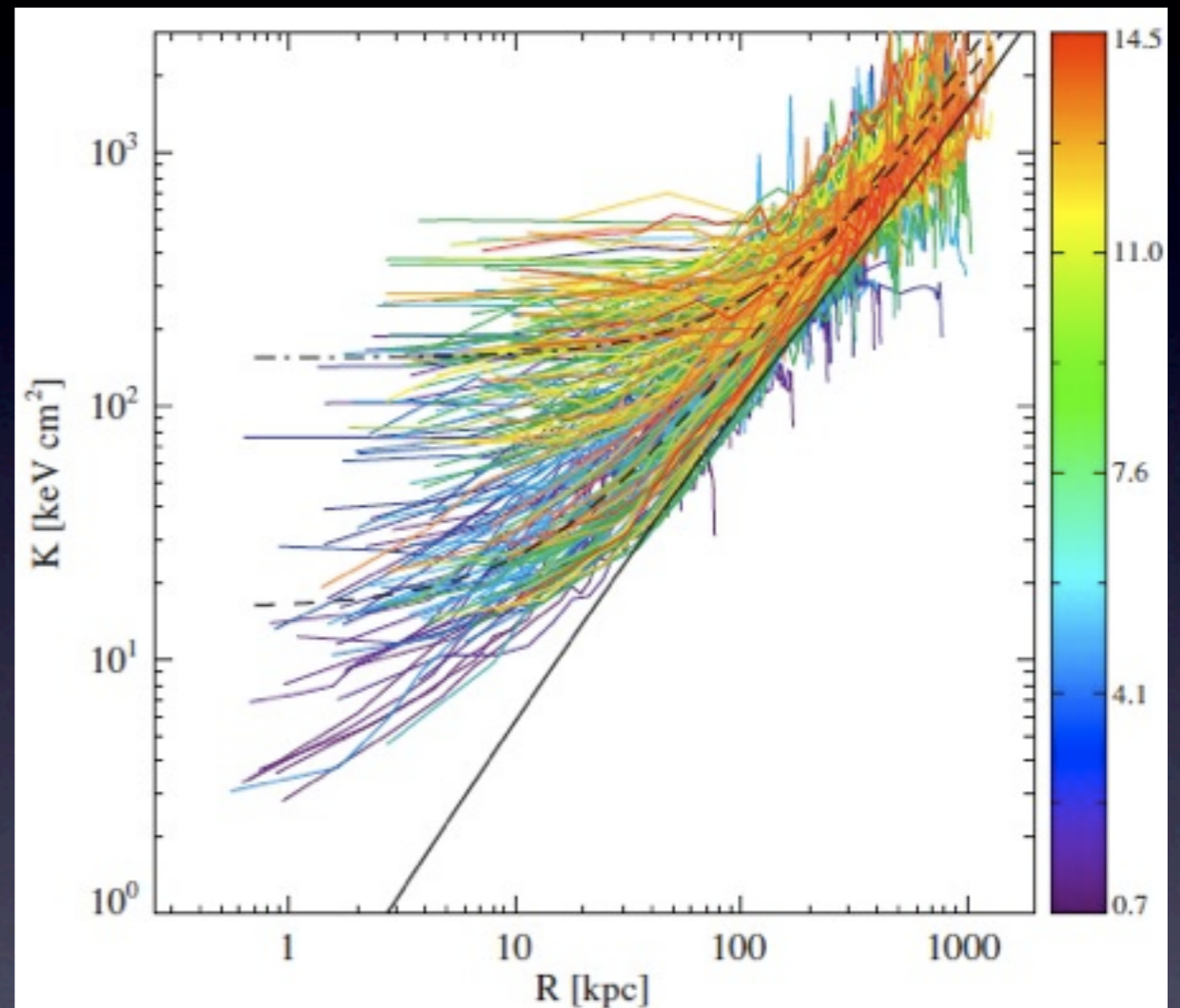
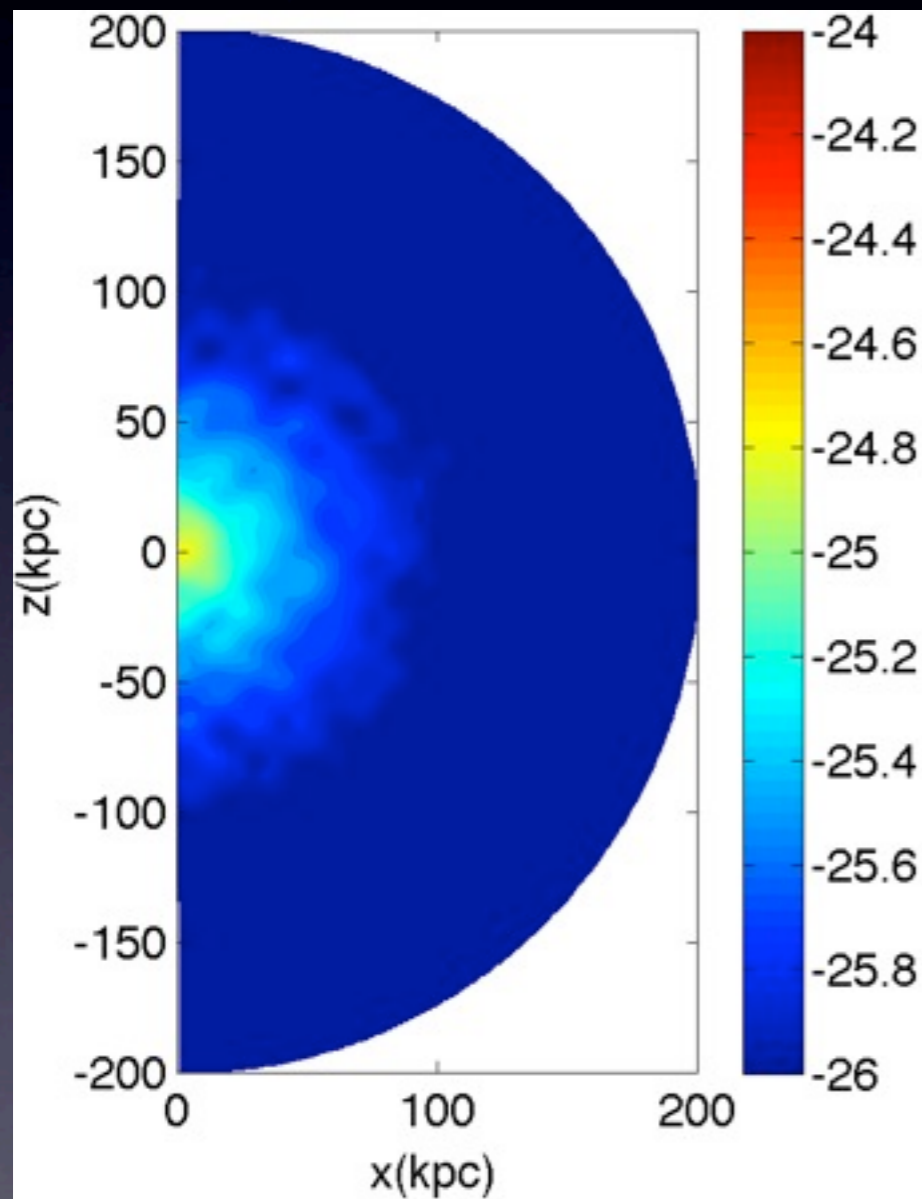
[de Plaa et al. 2010]



spatially resolved H α coincident w. soft X-rays

Spherical sims.

Log₁₀ density



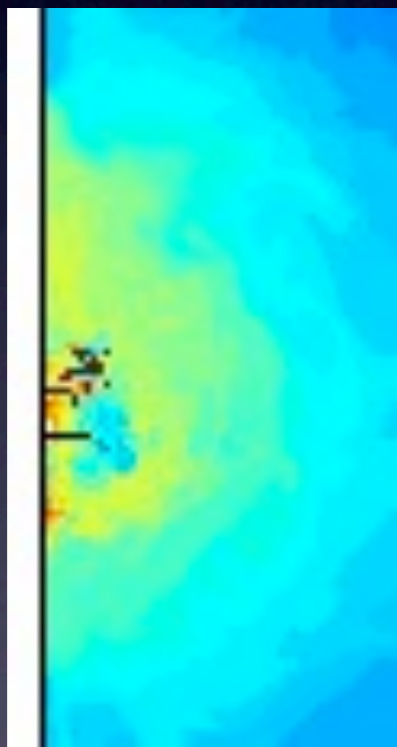
$$K(r) = K_0 + K_{100}(r/r_{100})^\alpha$$

$\delta\rho/\rho \sim 0.1$; hydro. eq. in NFW potl.

$H \sim L$

$\tau_{\text{TI}}/\tau_{\text{ff}} < 10 \Rightarrow$ filaments

multiphase
if local $\tau_{\text{cool}}/\tau_{\text{ff}} < 10$

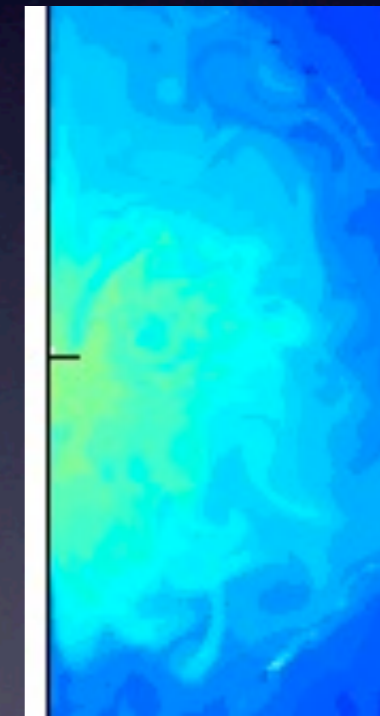


40 kpc

$$\tau_{\text{TI}} \approx \tau_{\text{cool}} = 3/2nkT/n^2\Lambda$$

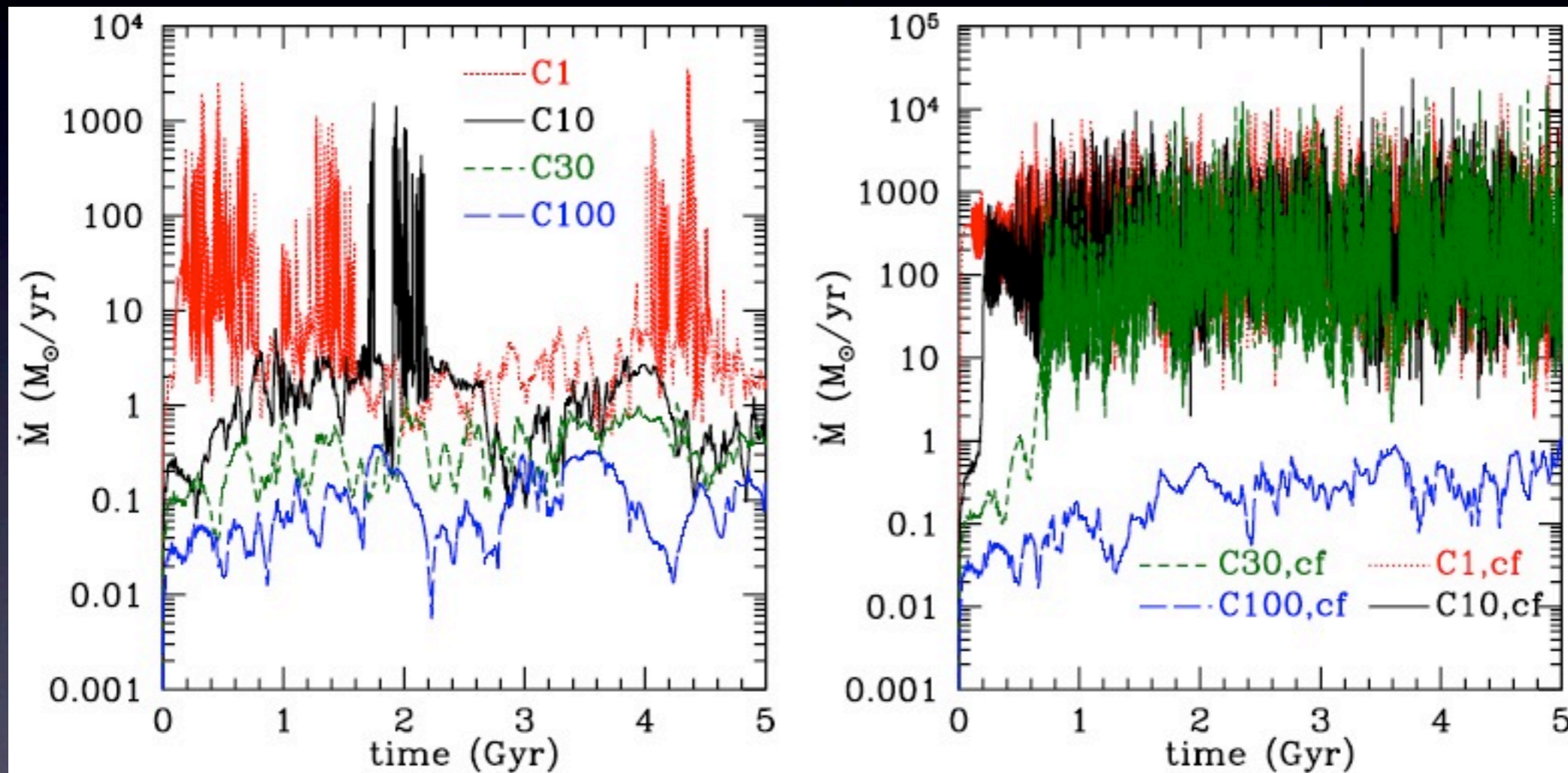
$$\tau_{\text{ff}} = (2r/g)^{1/2}; g = d\Phi/dr$$

only hot phase
if $\tau_{\text{cool}}/\tau_{\text{ff}} > 10$



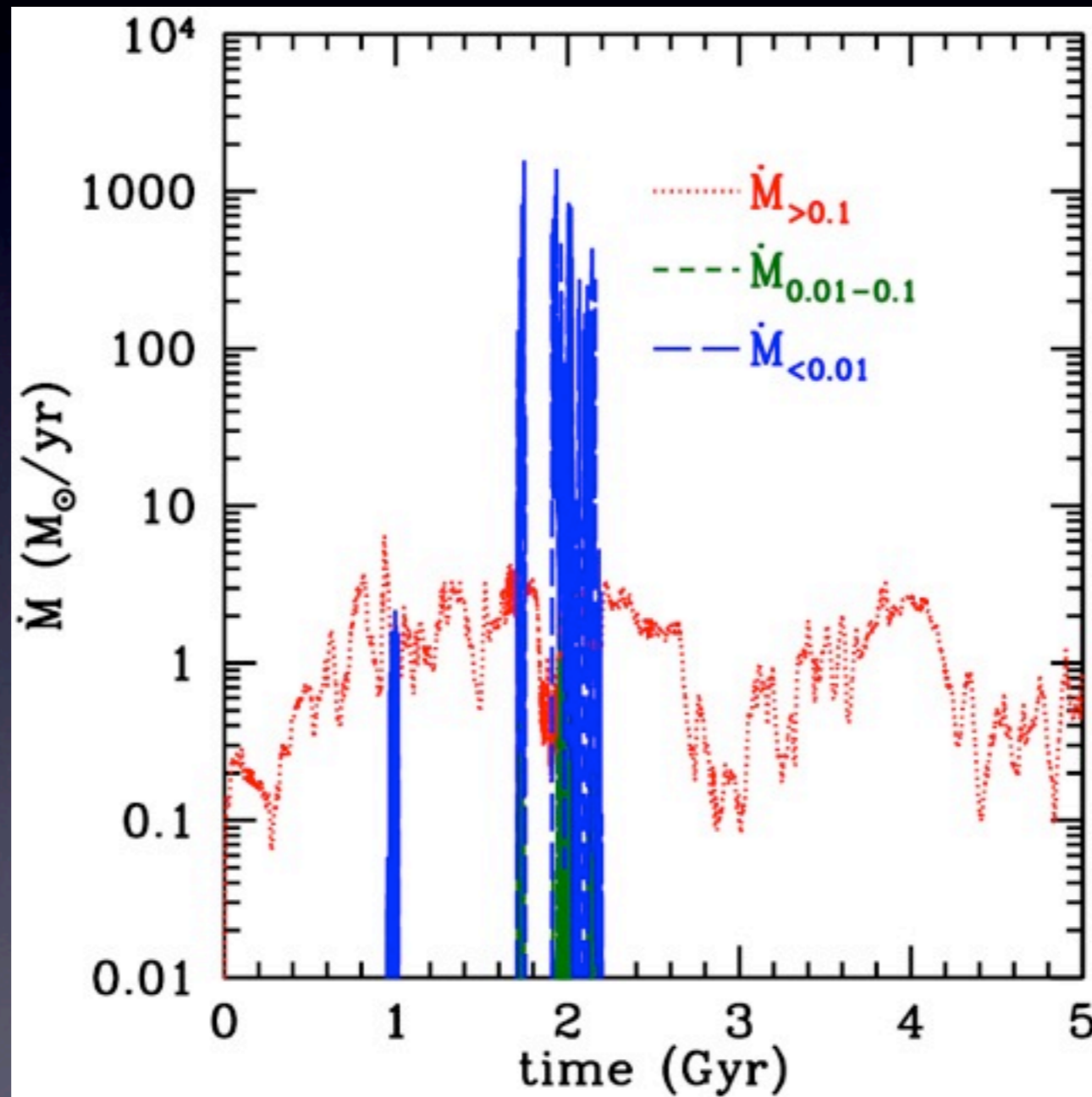
cool filaments when $\tau_{\text{TI}}/\tau_{\text{ff}} < 10$
spherical compression is quantitatively imp.

$H = \langle L \rangle \Rightarrow$ small dM/dt
as observed!



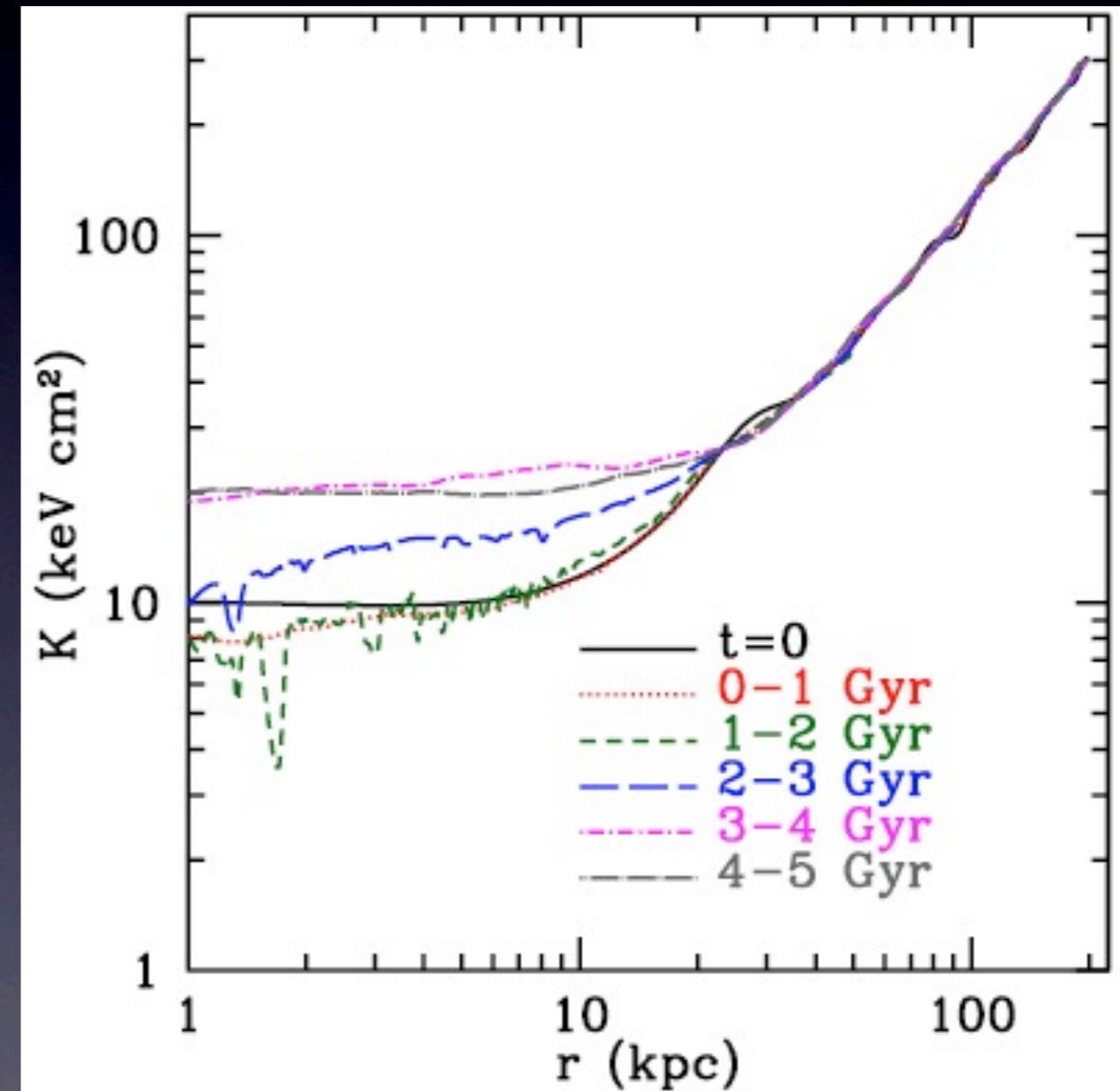
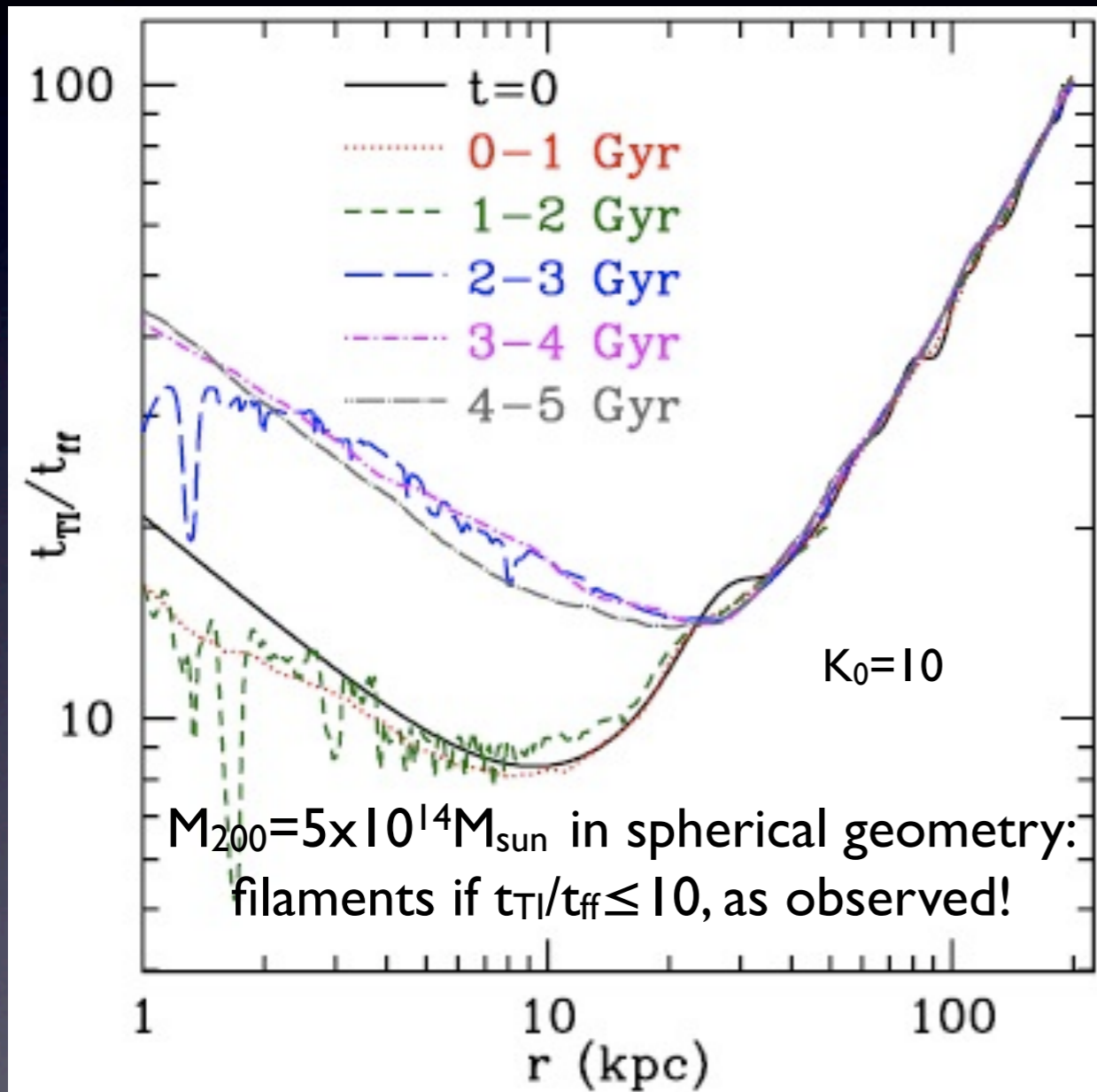
cold filaments drop out, hot phase maintained by H

Cold blobs => big dM/dt

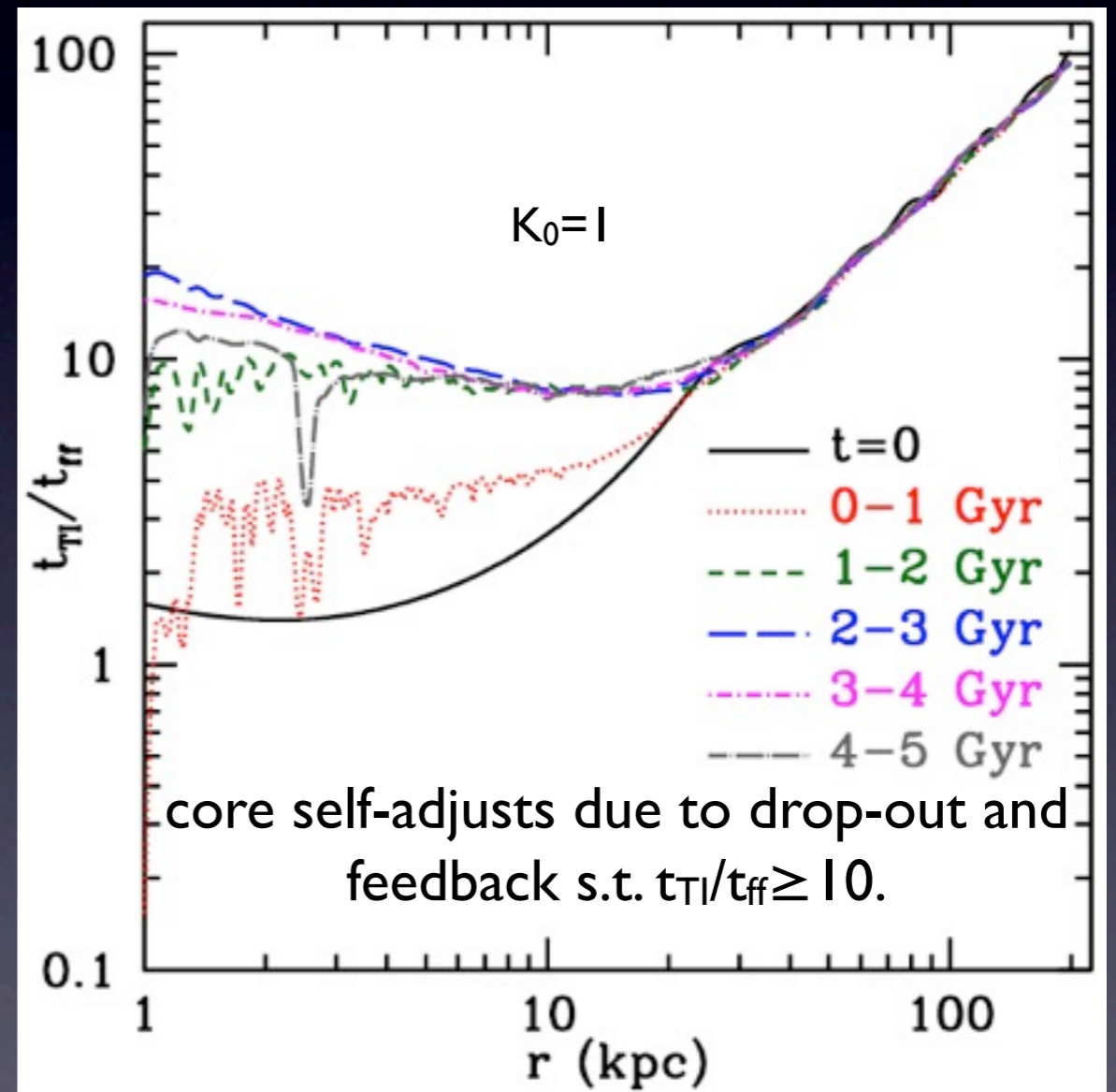
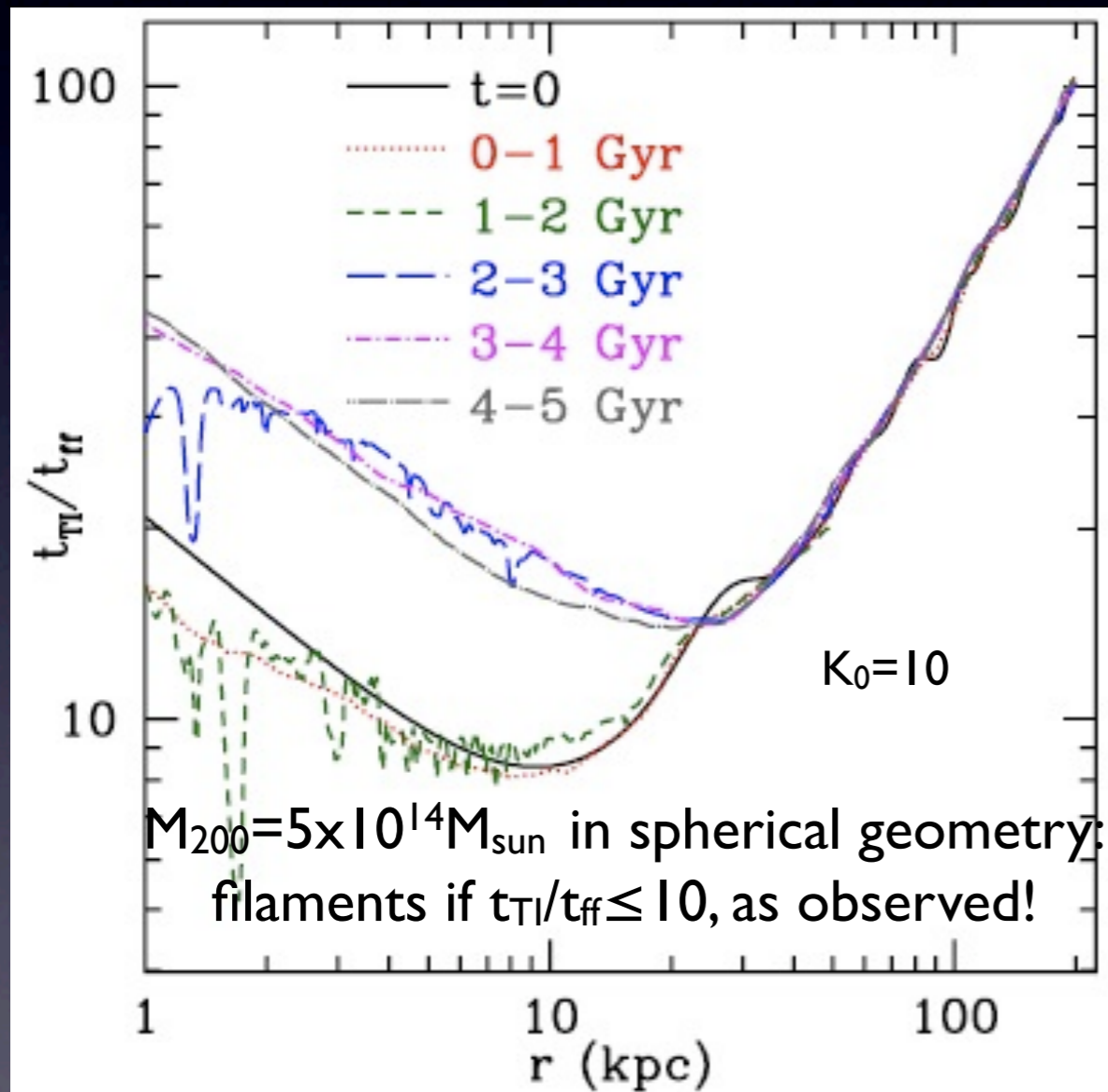


$K_0 = 10 \text{ keV cm}^2$

K & $t_{\text{TI}}/t_{\text{ff}}$



Self-regulation of halos

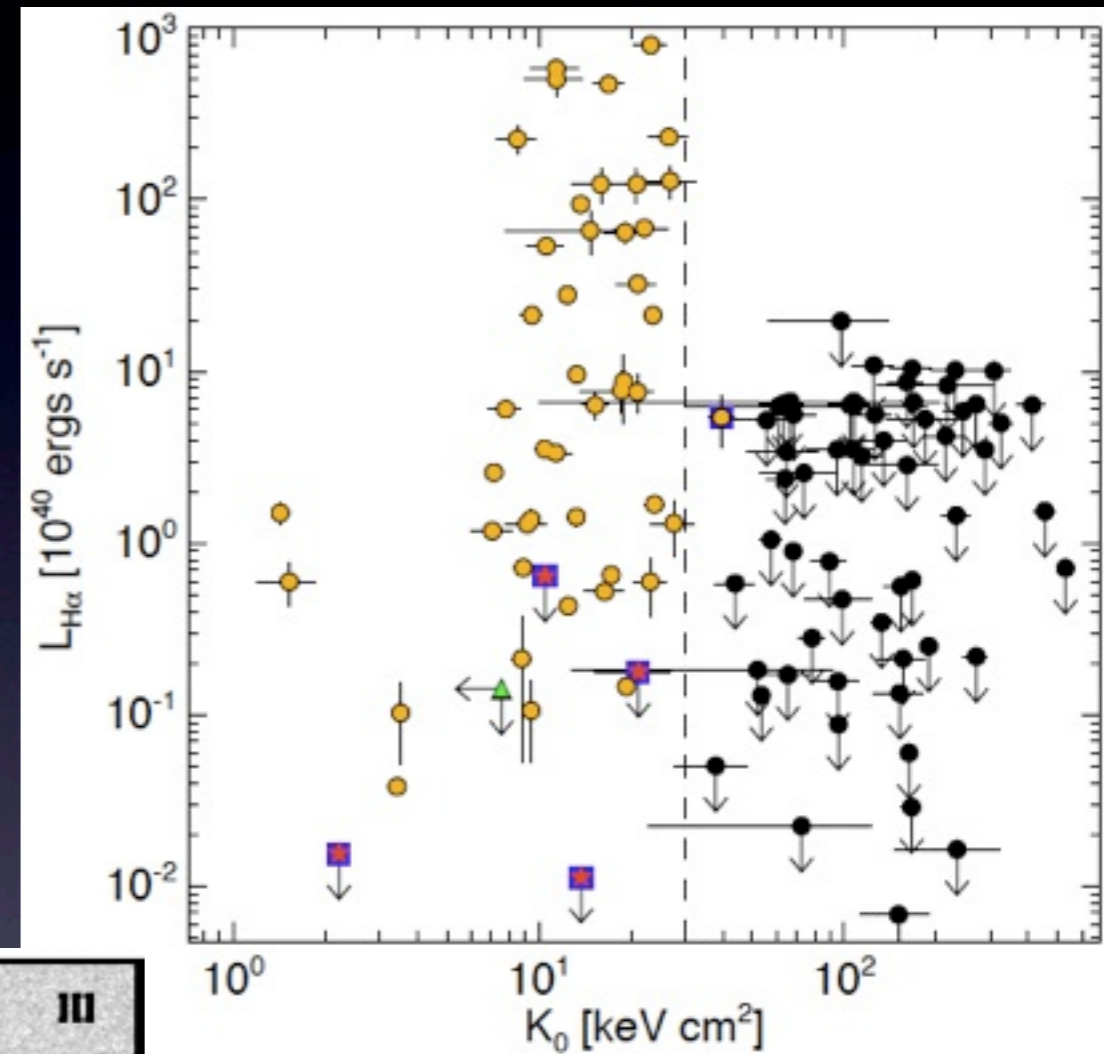
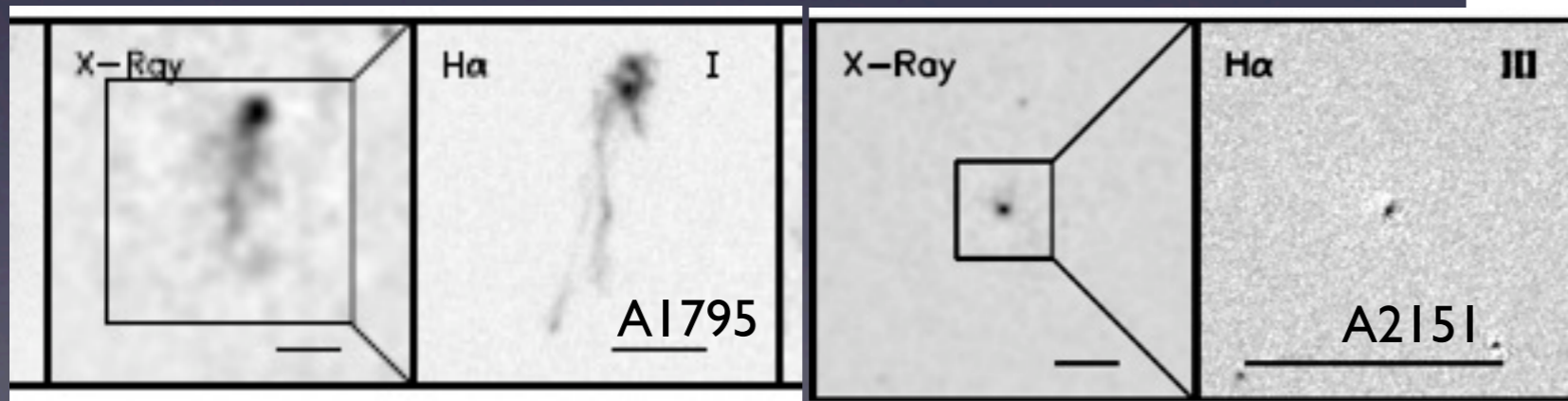


Cool Core vs non-CC

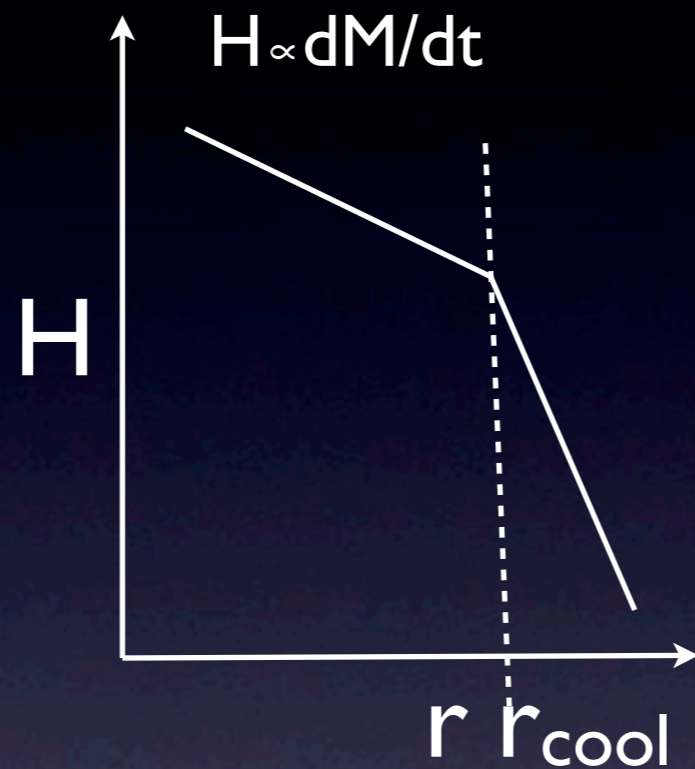
[Cavagnolo et al. 2008]

filaments (AGN fb) a fn. of entropy/
cooling time

[McDonald et al. 2010]

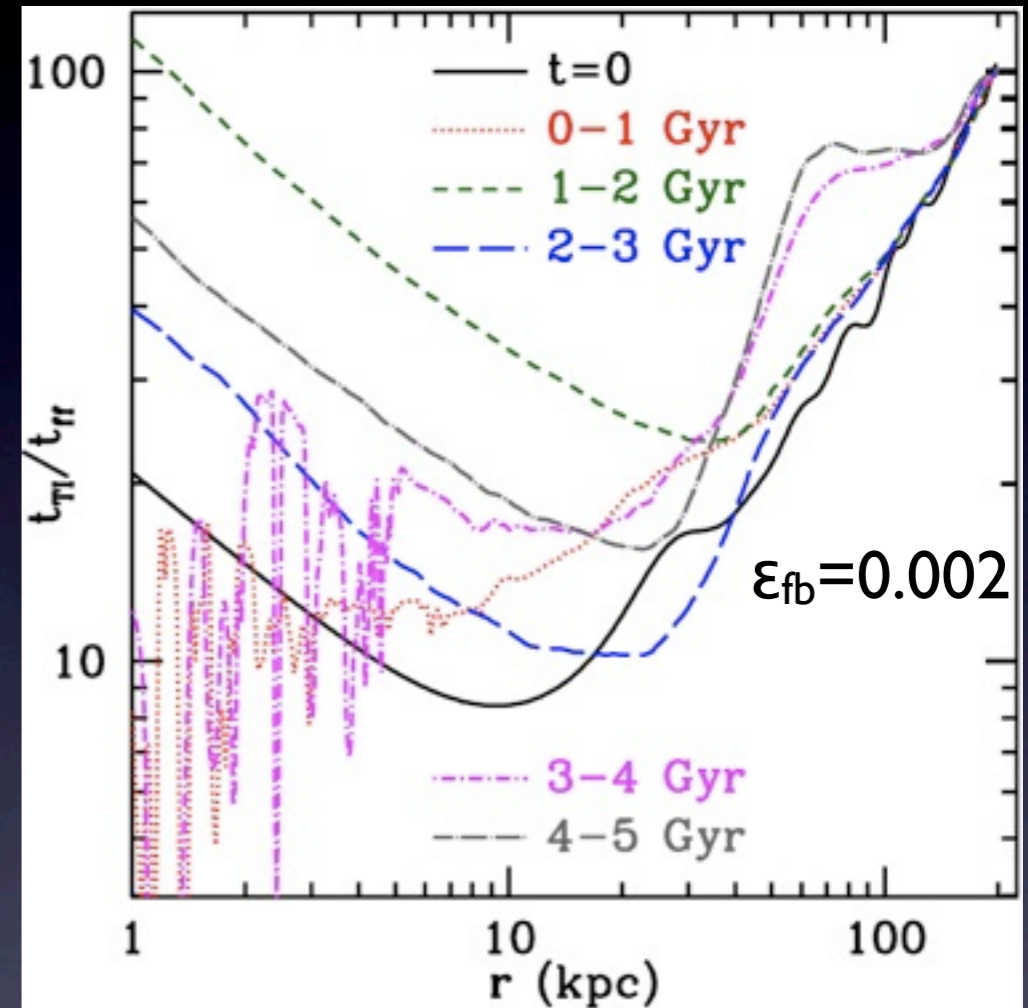
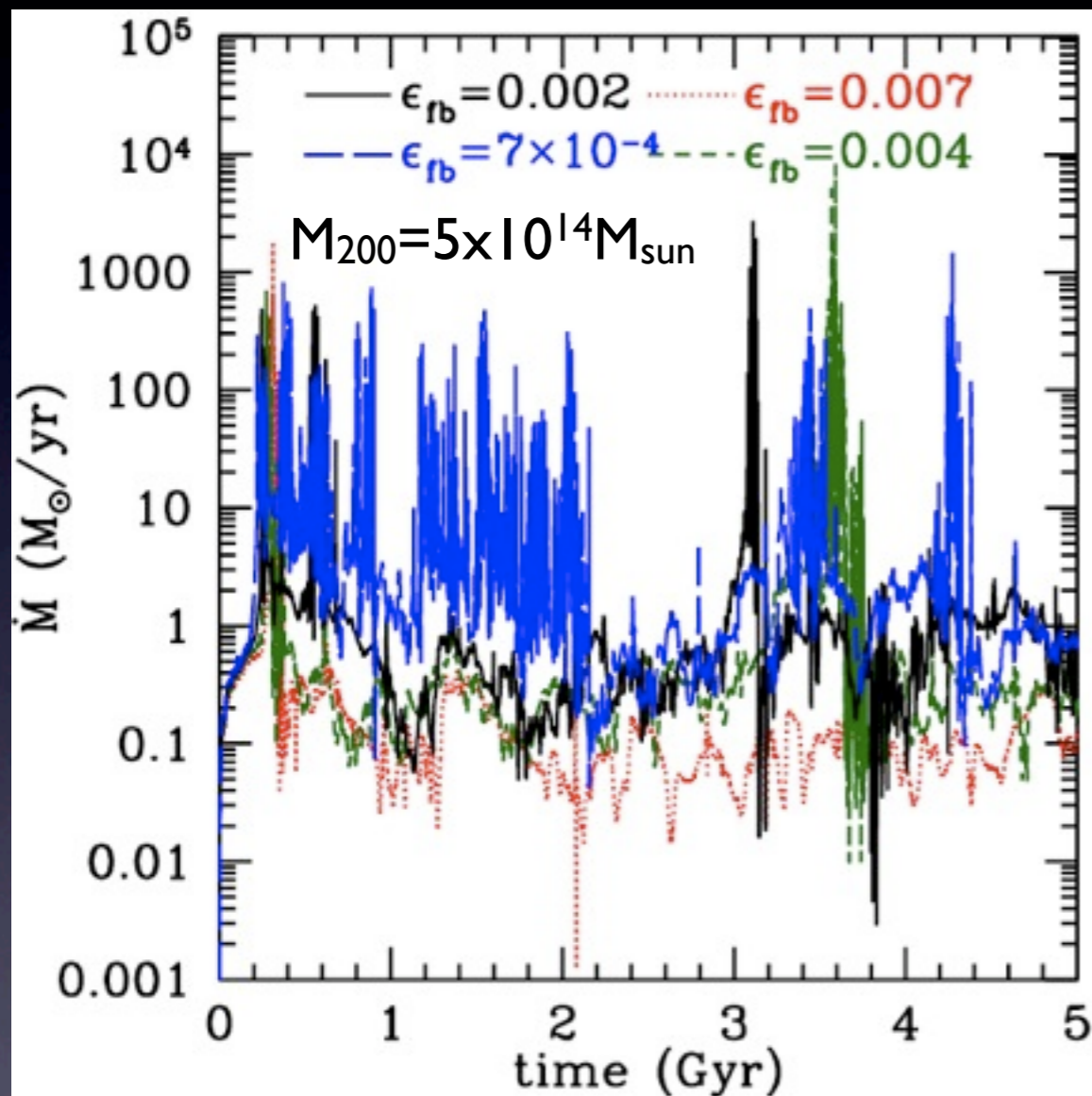


Feedback Sims.



isotropic, $H \propto dM/dt$ at r_{in}
most power deposited at r_{cool}
in reality anisotropic jets!

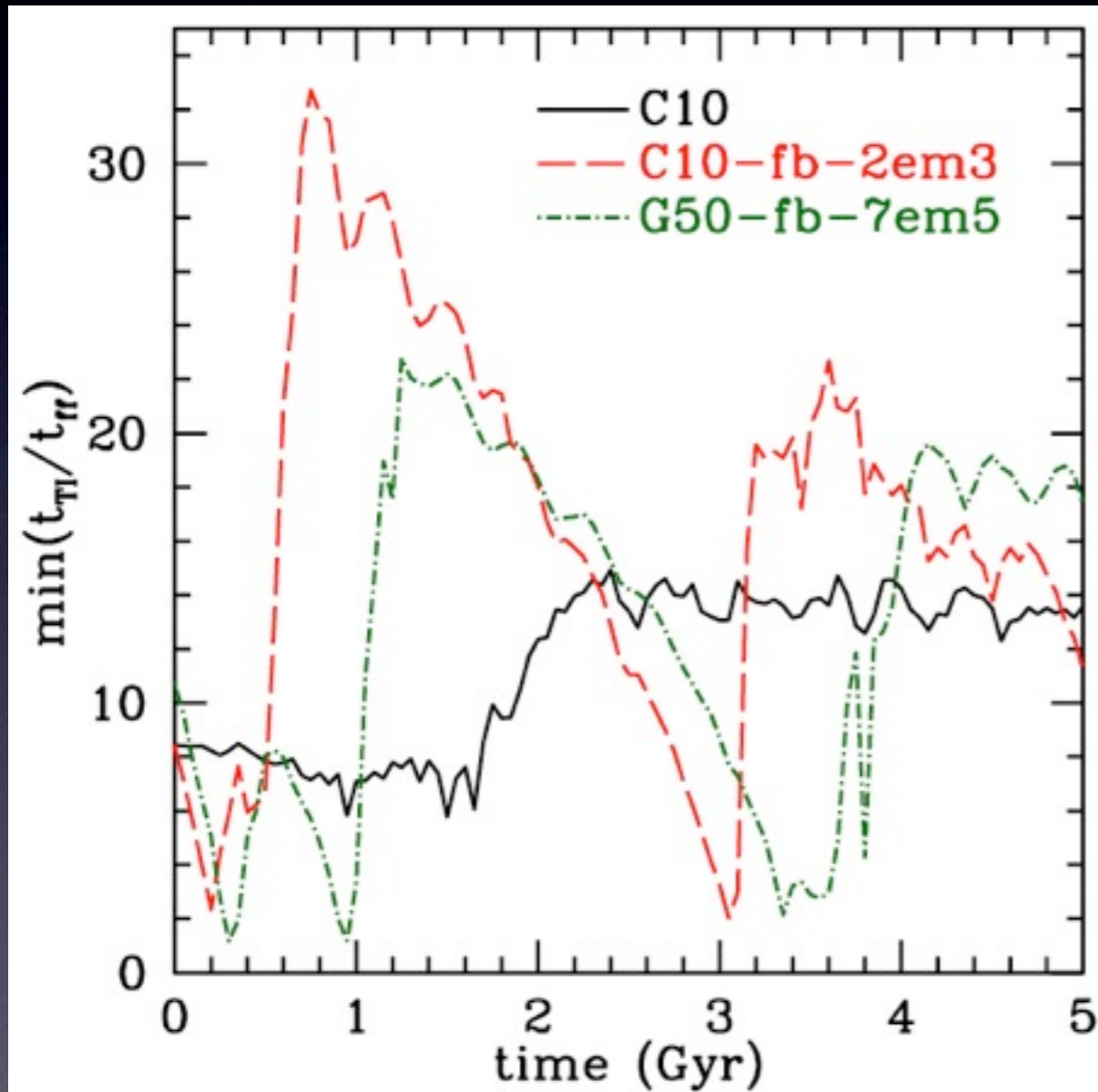
Feedback Sims.



reach a quasi-state where $H \approx L$
 TI => filaments whenever $t_{\text{TI}}/t_{\text{ff}} < 10$

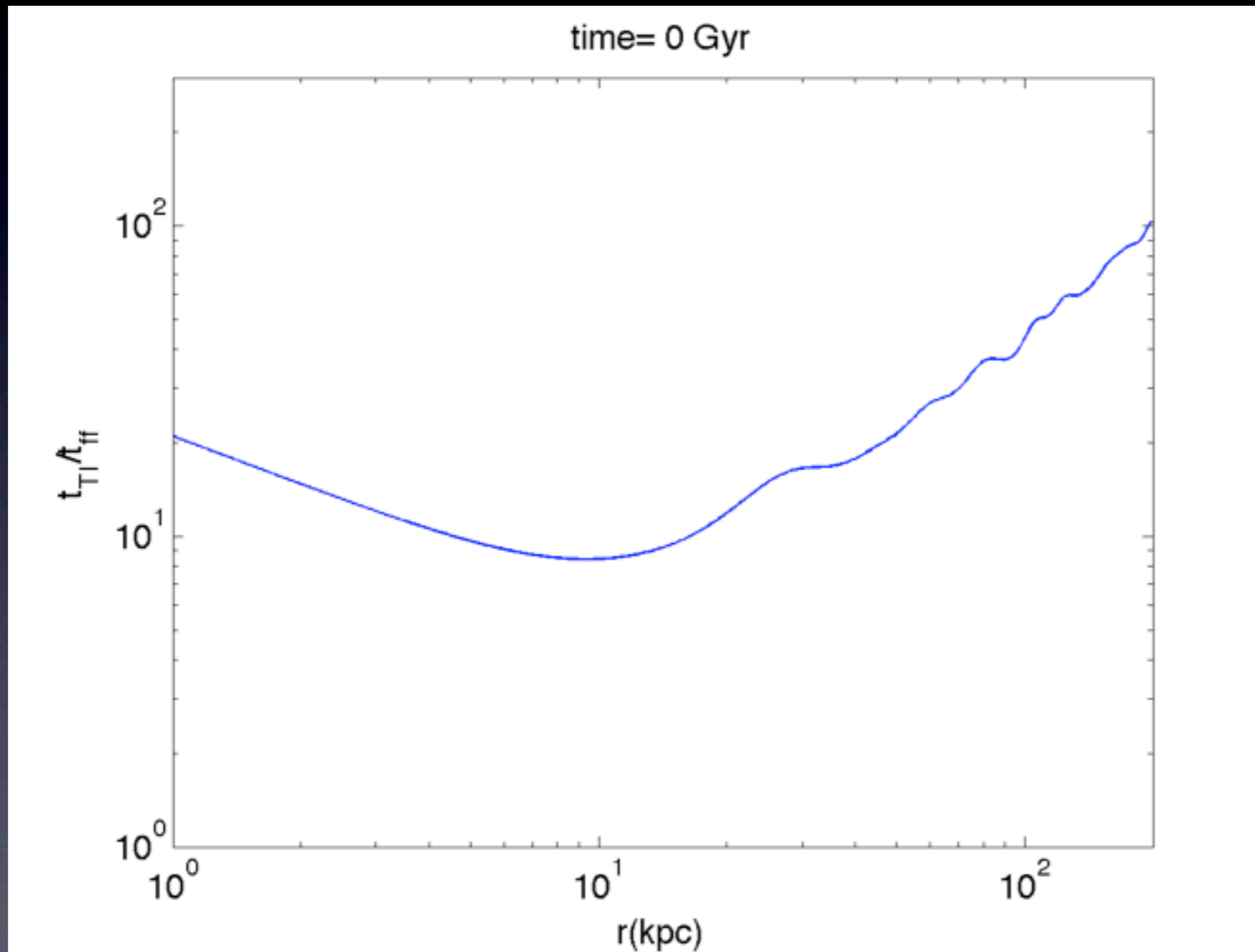
smaller ϵ required for lower mass halos bec. dM/dt similar in cold phase but
 M_g much smaller in lower mass halos

FB vs. idealized H



larger thermal imbalance w. FB

FB model self reg.



Conclusions

- $H=\langle L \rangle$ ansatz gives similar results as FB models, i.e., cool filaments when $t_{\text{TI}}/t_{\text{ff}} < 10$
- very different from a cooling flow; $dM/dt \sim$ few M_{sun}/yr
- peaks in dM/dt due to cool filaments
- self-adjustment of halos s.t. $t_{\text{TI}}/t_{\text{ff}} \geq 10$
- smaller FB eff. for groups than clusters