

Resonant absorption of dark matter in molecules

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[arXiv:1709.05354](https://arxiv.org/abs/1709.05354)

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KITP, 2018/03/29

“Problems” in particle physics

Standard Model of particle physics accurately describes every known experiment and observation to the measured and calculated precision*

theoretical frontiers:

mathematical structures | numerological curiosities | computational precision

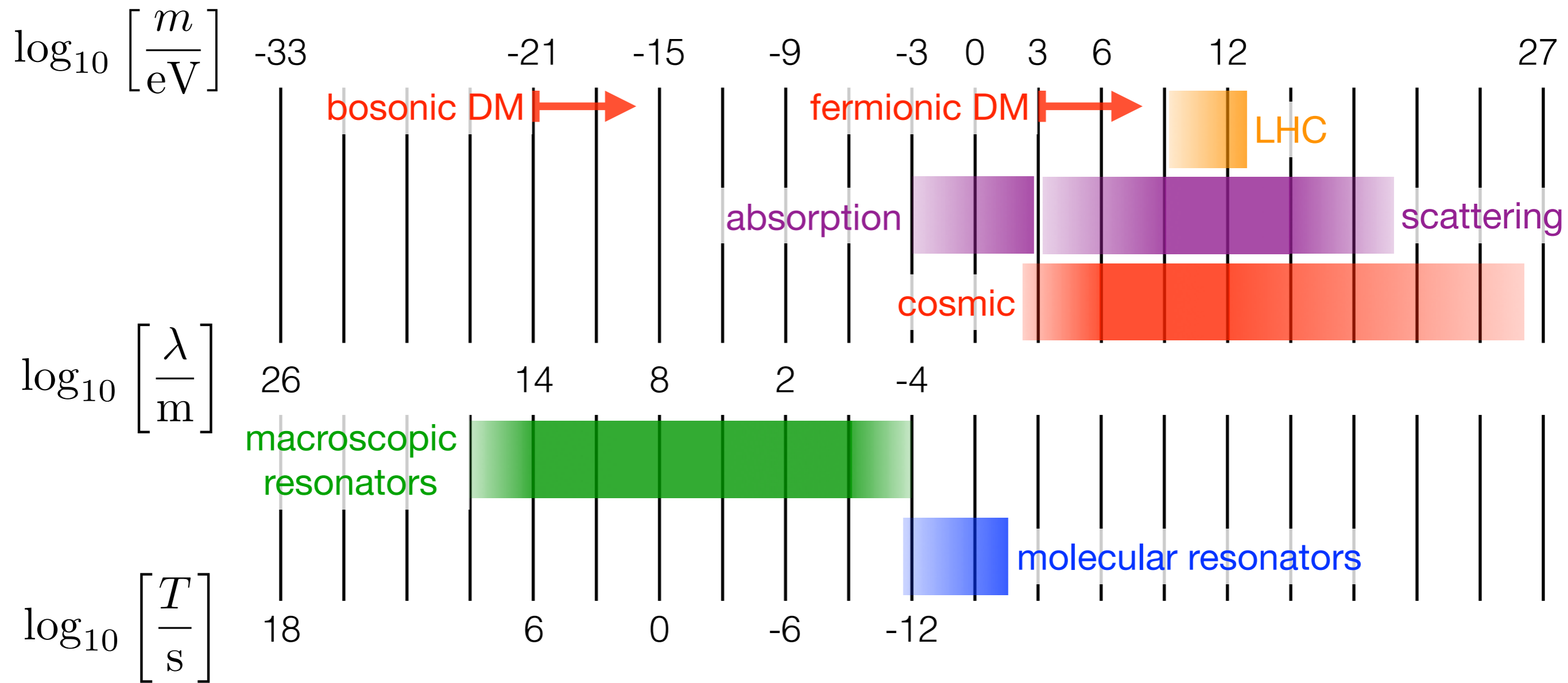
experimental frontiers:

high-energy | cosmic | intensity | precision

***parametrized unknowns:**

dark matter | baryon asymmetry | inflation

Scales of dark matter



absorption \rightarrow narrowband signal: $\omega = m \left(1 + \frac{v^2}{2} \right)$

Outline

(1) Dark matter fields can resonantly excite a molecular system

bosonic DM couplings | two-level system dynamics | molecular levels

(2) Experimental setup

configurations | photon detection | backgrounds | discrimination

(3) Dark matter sensitivity $0.2 \text{ eV} < m < 20 \text{ eV}$

hidden vectors | moduli | axions

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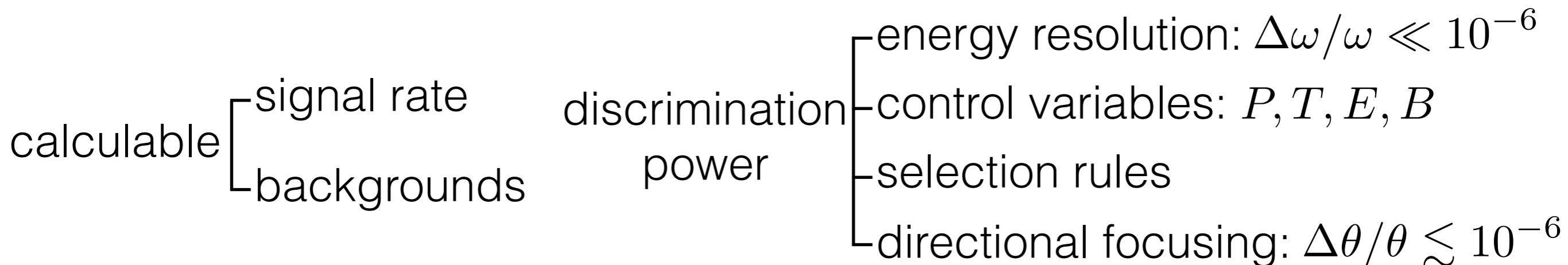
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Why is a molecular gas a good dark matter detector?



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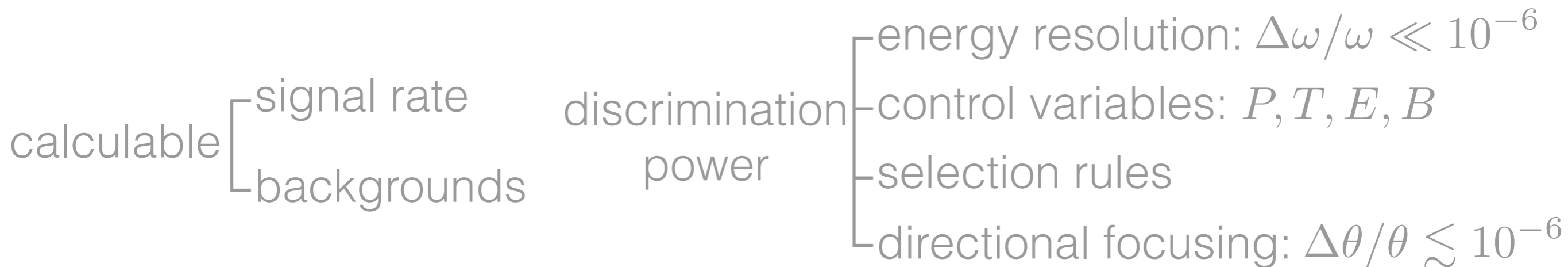
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Why is a molecular gas a good dark matter detector?



Bosonic dark matter couplings

EFT of DM bosons coupled to electrons, quarks, photons, gluons

spin-1

$$\epsilon A'_\mu J_{\text{EM}}^\mu$$

$$g A'_\mu J_{B-L}^\mu$$

spin-0

parity-even

$$\phi \bar{e}e$$

$$\phi \bar{q}q$$

$$\phi F^2$$

$$\phi G^2$$

parity-odd

$$(\partial_\mu a) \bar{e} \gamma^\mu \gamma_5 e$$

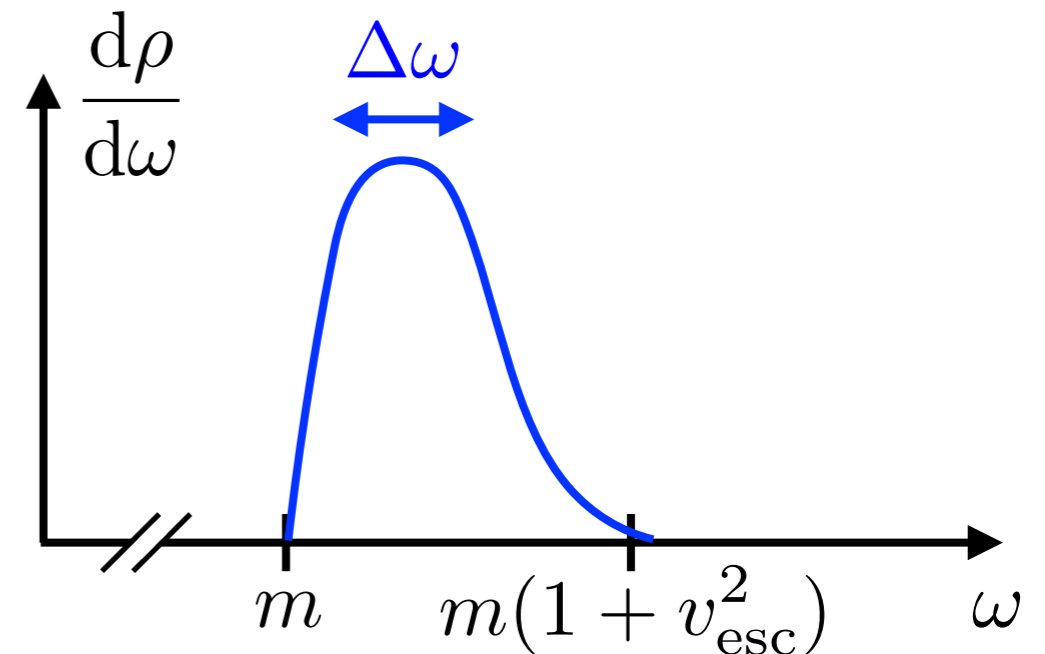
$$(\partial_\mu a) \bar{q} \gamma^\mu \gamma_5 q$$

$$a F \tilde{F}$$

$$a G \tilde{G}$$

Bosonic dark matter fields and interactions

$$\left. \begin{array}{l} \text{vector} \quad |\mathbf{A}'| \\ \text{scalar} \quad \phi \\ \text{pseudoscalar} \quad a \end{array} \right\} \simeq \frac{\sqrt{2\rho_{\text{DM}}}}{\omega} \cos(\omega t)$$



production: $\left\{ \begin{array}{l} \text{inflationary perturbations (spin-1)} \\ \text{misalignment mechanism (spin-0)} \end{array} \right.$

e.g. hidden photon: $\delta\mathcal{L} = \epsilon A'_\mu J_{\text{EM}}^\mu$

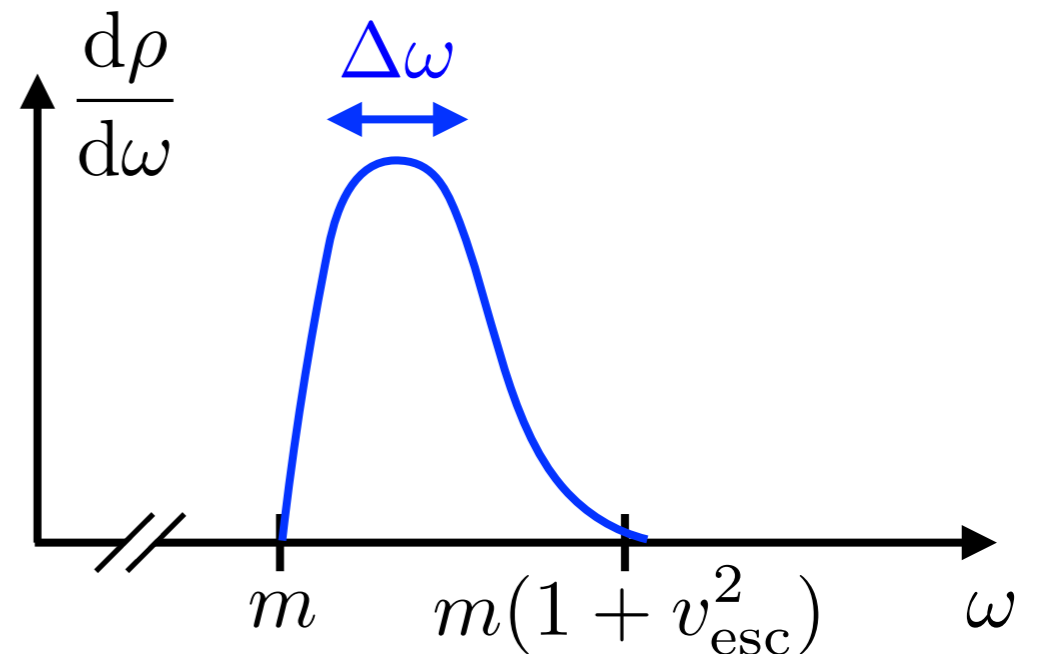
$$\delta H = \epsilon \mathbf{E}' \cdot \boldsymbol{\mu}_e$$

↓

3 kV/m

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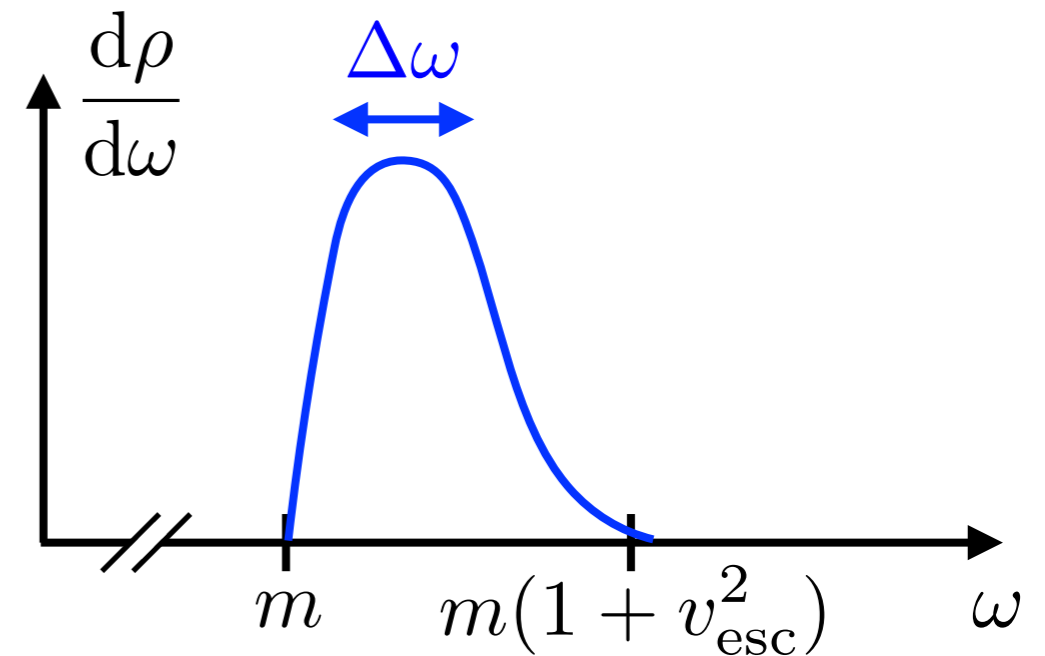
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$$\delta H = \epsilon \mathbf{E}' \cdot \boldsymbol{\mu}_e \quad \boldsymbol{\mu}_e = e \sum_{\psi} q_{\psi} \mathbf{r}_{\psi}$$

\downarrow
 3 kV/m

Bosonic dark matter fields and interactions

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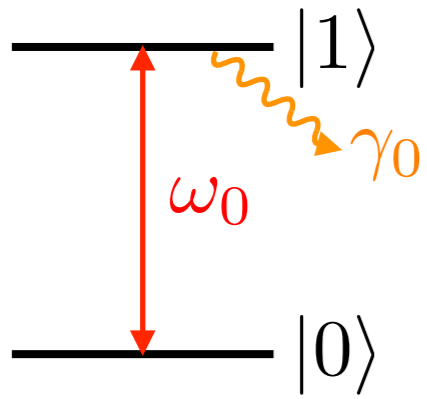
$$\delta H = \epsilon \mathbf{E}' \cdot \boldsymbol{\mu}_e \quad \boldsymbol{\mu}_e = e \sum_{\psi} q_{\psi} \mathbf{r}_{\psi}$$

3 kV/m

↳ equivalent to shining 20kW laser in a 1m² beam waist area →

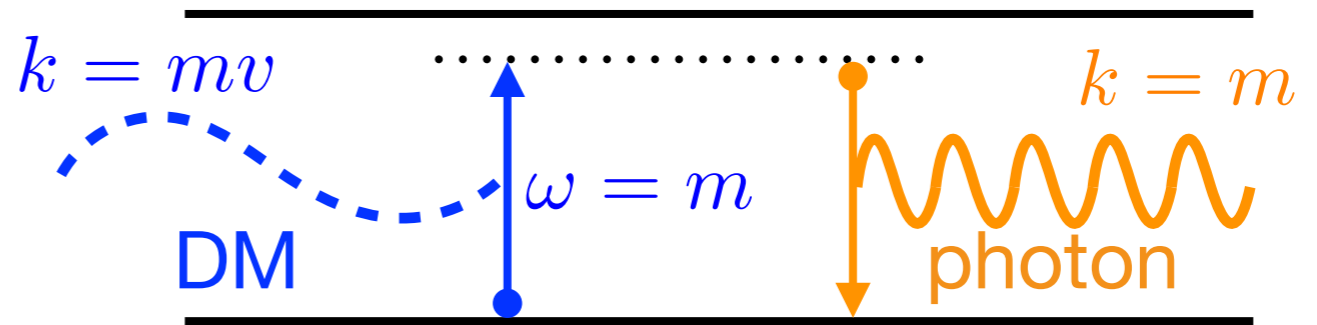
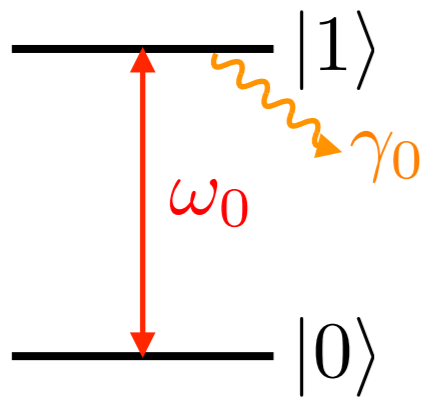


Resonant excitation of a two-level system



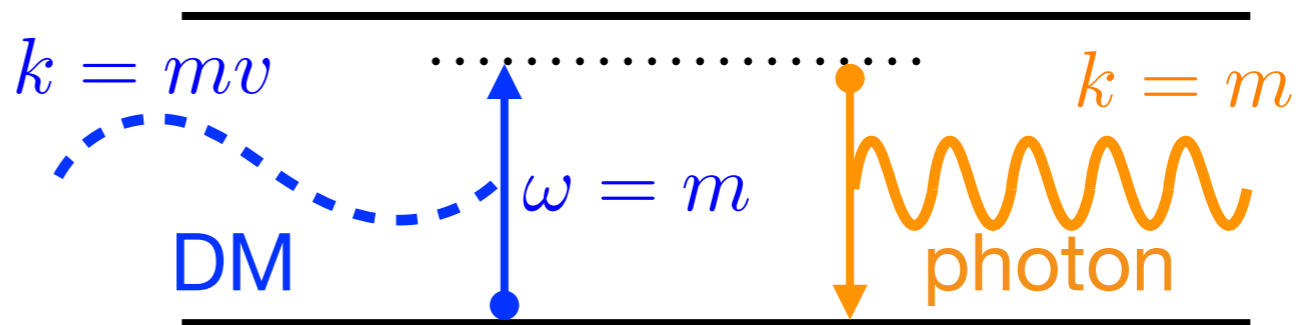
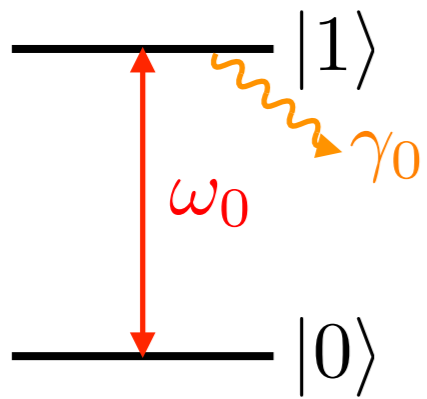
$$\langle 1 | \delta H | 0 \rangle \sim \Omega \cos(\omega t)$$

Resonant excitation of a two-level system



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Resonant excitation of a two-level system



$$\langle 1 | \delta H | 0 \rangle \sim \Omega \cos(\omega t)$$

$$\Gamma_{\text{abs}} = \underbrace{nV}_{\text{\# molecules}} \underbrace{\frac{|\Omega|^2}{\gamma}}_{\text{on-resonance absorption rate per molecule}} \underbrace{\frac{1}{1 + \frac{4(\omega_0 - \omega)^2}{\gamma^2}}}_{\text{lineshape}}$$

on-resonance
absorption rate per molecule

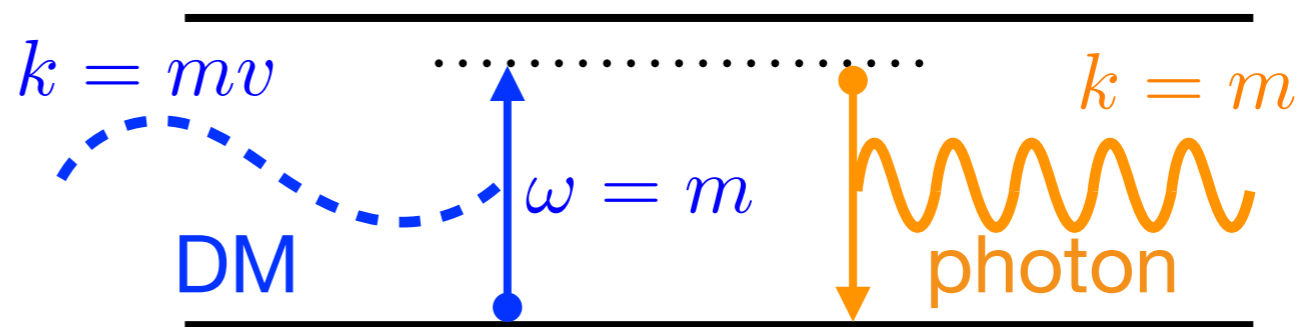
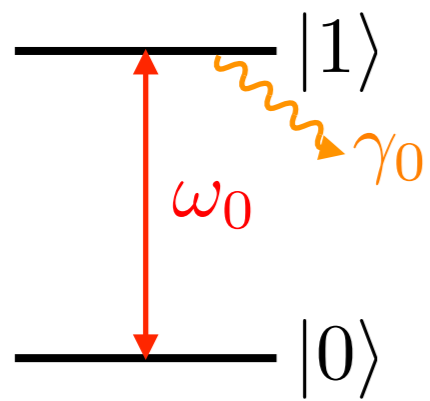
$$\gamma = \gamma_{\text{rad}} + 2\gamma_{\text{col}} + \frac{\Delta\omega}{\pi}$$

$$\gamma_{\text{rad}} \simeq \bar{r}\gamma_0 + \sum_i \gamma_i$$

$$\gamma_{\text{col}} = n\sigma_{\text{col}}v_{\text{mol}}$$

$\Delta\omega_{\text{Doppler}}$ negligible

Resonant excitation of a two-level system



$$\langle 1 | \delta H | 0 \rangle \sim \Omega \cos(\omega t) \quad |\Psi(t)'\rangle = e^{-i \int_0^t \delta H dt'} |0'\rangle \simeq |0'\rangle - \frac{i}{2} \Omega \frac{e^{i(\omega_0 - \omega)t} - 1}{i(\omega_0 - \omega)} |1'\rangle$$

$$\Gamma_{\text{abs}} = \underbrace{nV}_{\text{\# molecules}} \underbrace{\frac{|\Omega|^2}{\gamma}}_{\text{on-resonance}} \underbrace{\frac{1}{1 + \frac{4(\omega_0 - \omega)^2}{\gamma^2}}}_{\text{lineshape}}$$

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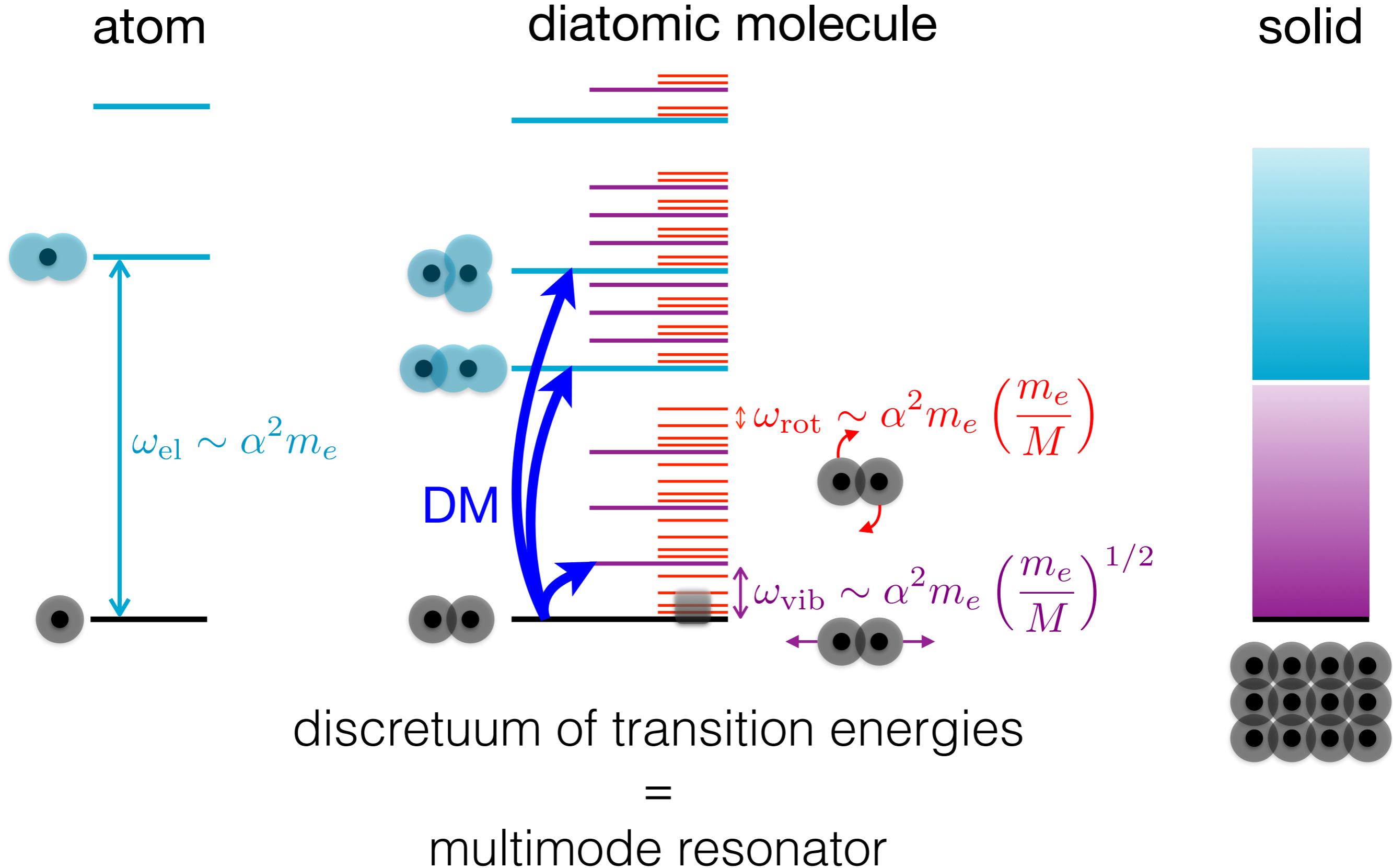
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Molecular levels and transitions



DM absorption rate

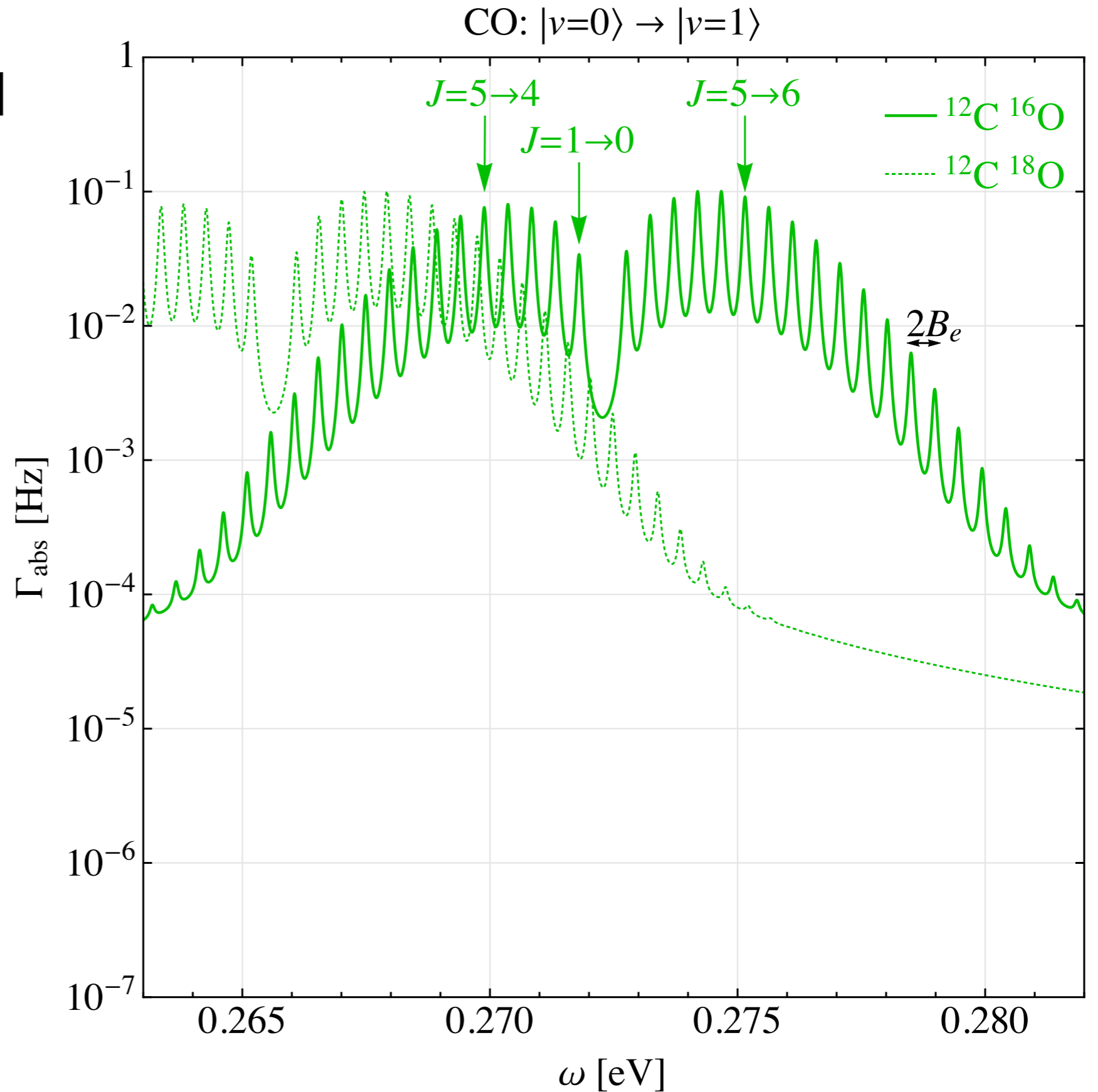
hidden photon DM

$$\epsilon = 10^{-12}$$

$$V = 300 \text{ cm}^3$$

$$T = 100 \text{ K}$$

$$P = 5 \text{ bar}$$



DM absorption rate

hidden photon DM

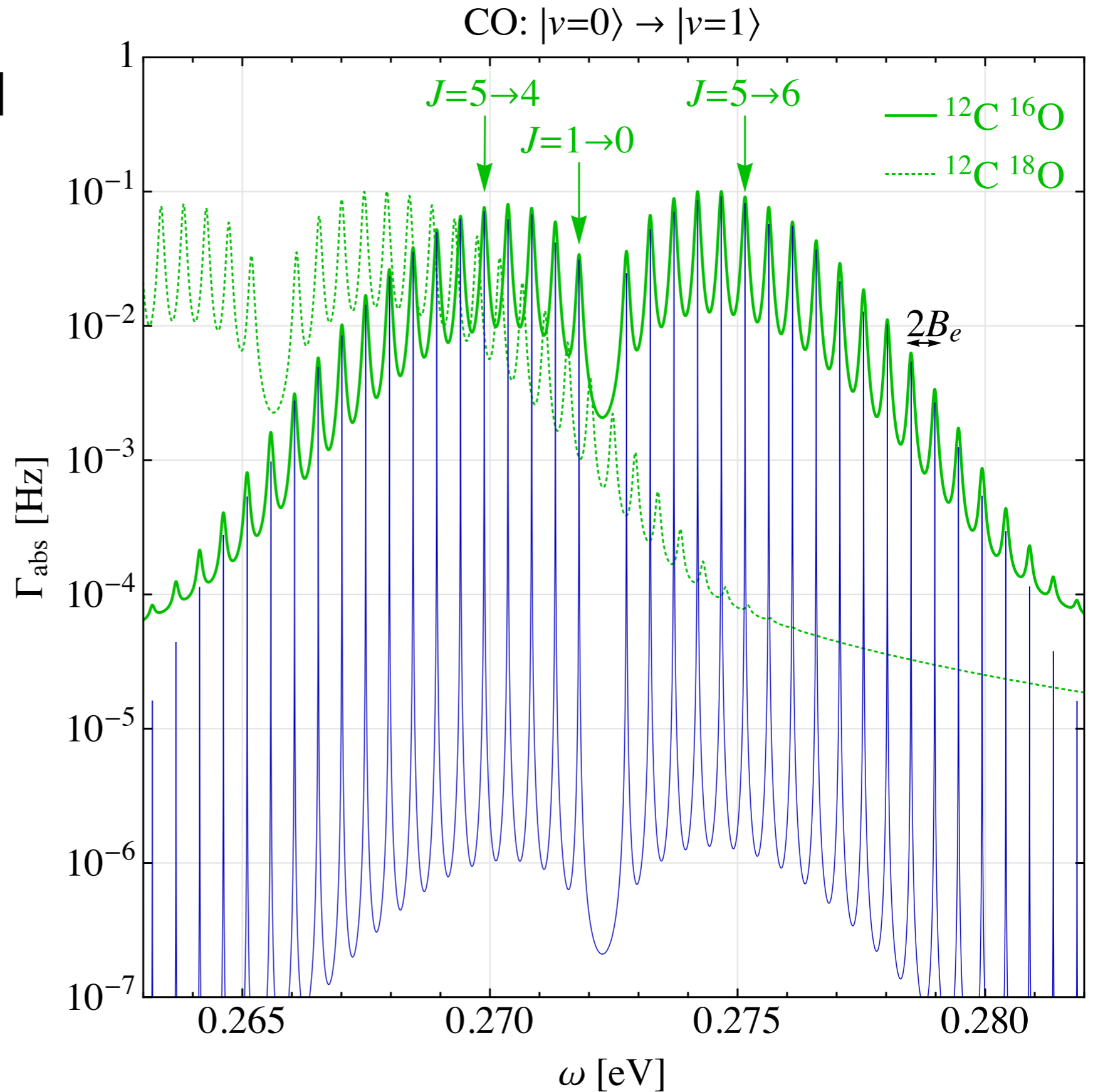
$$\epsilon = 10^{-12}$$

$$V = 300 \text{ cm}^3$$

$$T = 100 \text{ K}$$

$$P = 5 \text{ bar}$$

$$P = 0.05 \text{ bar}$$



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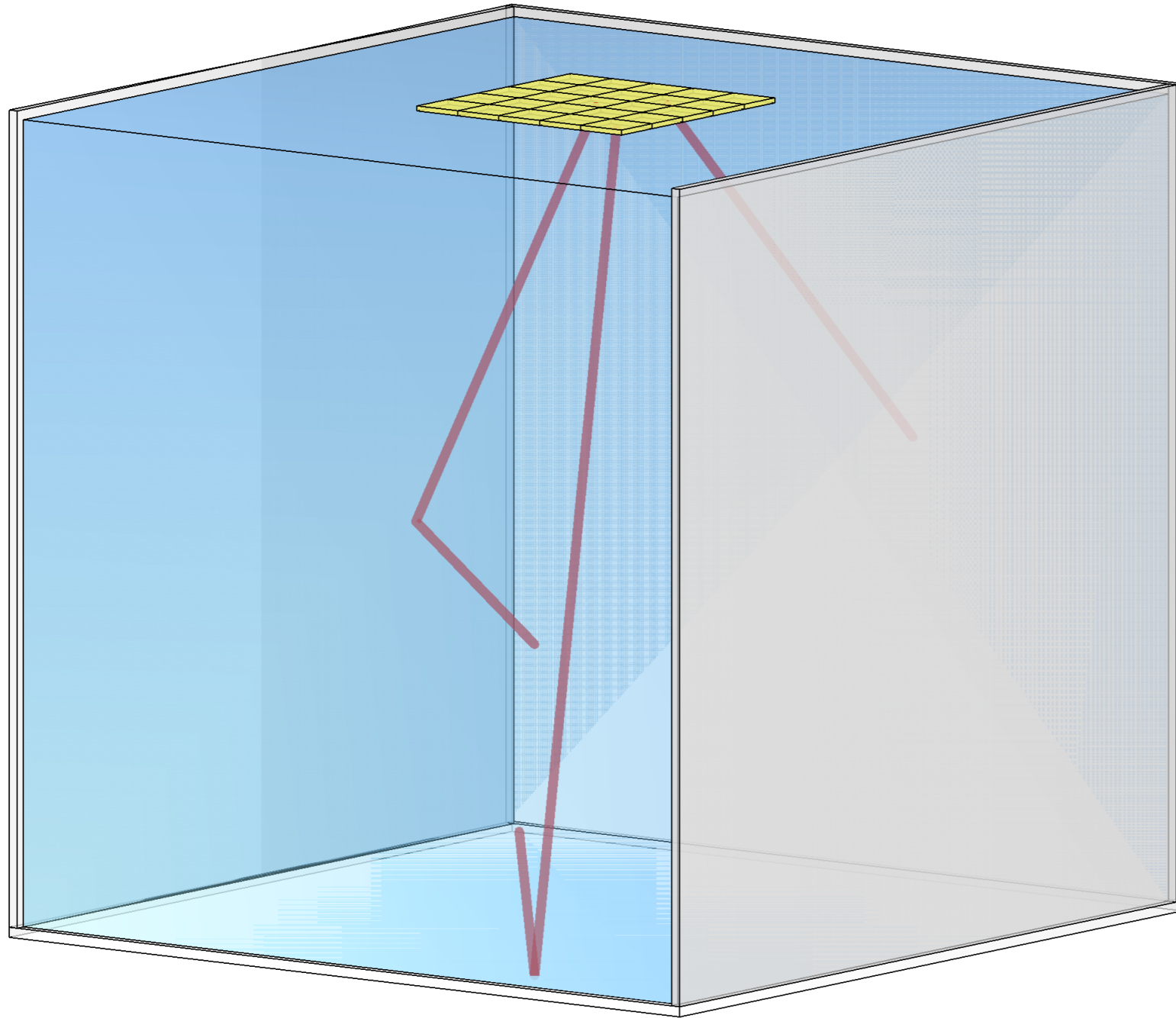
(3) Dark matter sensitivity $0.2 \text{ eV} < m < 20 \text{ eV}$

hidden vectors | moduli | axions

Why is a molecular gas a good dark matter detector?

calculable { signal rate
backgrounds } discrimination power { energy resolution: $\Delta\omega/\omega \ll 10^{-6}$
control variables: P, T, E, B
selection rules
directional focusing: $\Delta\theta/\theta \lesssim 10^{-6}$

Bulk configuration



Phase I

$$V = (0.3 \text{ m})^3$$

$$T = 300 \text{ K}$$

$$\text{DCR} = 1 \text{ Hz}$$

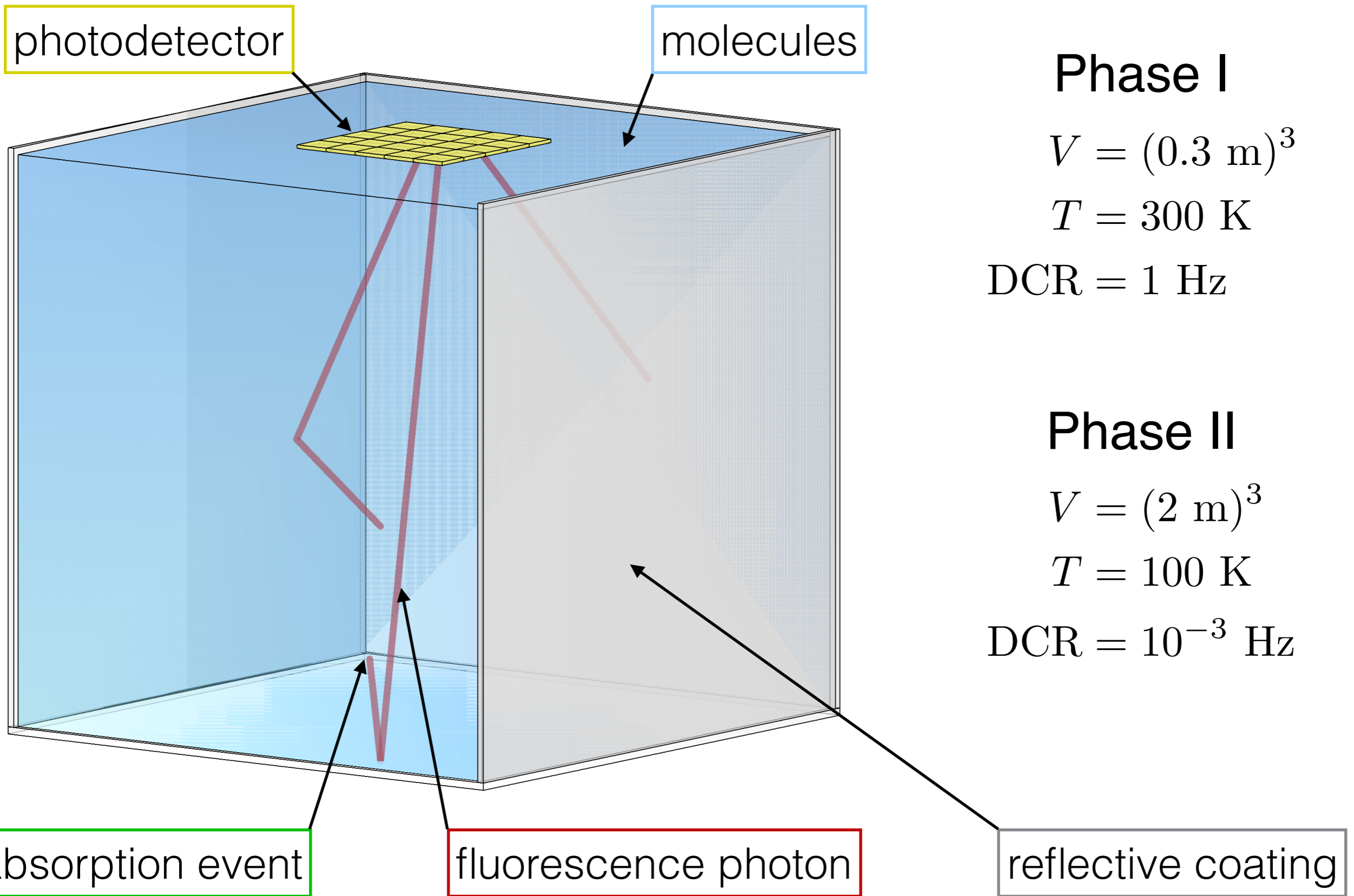
Phase II

$$V = (2 \text{ m})^3$$

$$T = 100 \text{ K}$$

$$\text{DCR} = 10^{-3} \text{ Hz}$$

Bulk configuration



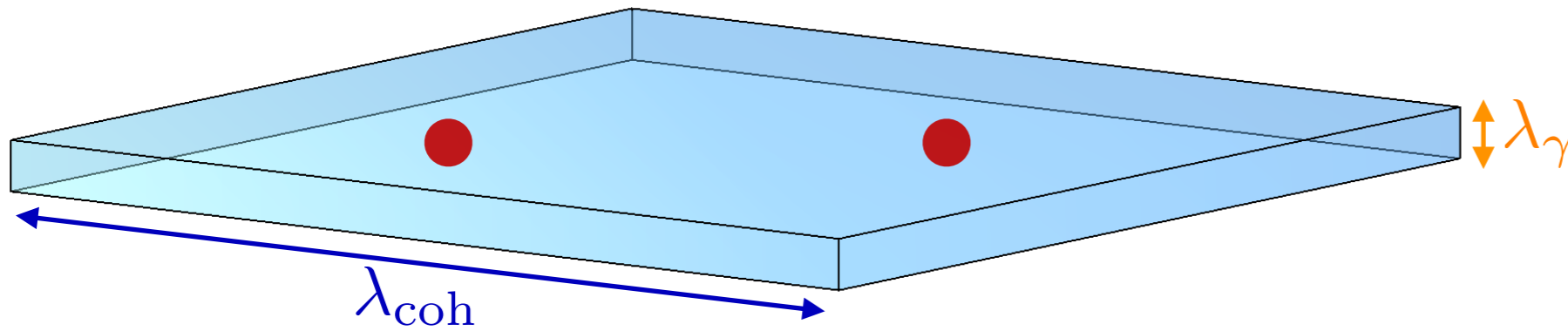
Cooperative radiation

$$f(\mathbf{v}) \propto \exp \left[-\frac{(\mathbf{v} - \mathbf{v}_{\text{lab}})^2}{v_0^2} \right]$$

coherence length: $\lambda_{\text{coh}} = \frac{2}{mv_0}$

typical deBroglie wavelength: $\lambda_{\text{dB}} \sim \frac{2\pi}{mv_{\text{lab}}}$

photon wavelength: $\lambda_\gamma = \frac{2\pi}{m}$



$$|0\rangle + \varepsilon e^{-i[mt + \varphi(\mathbf{x})]} |1\rangle$$

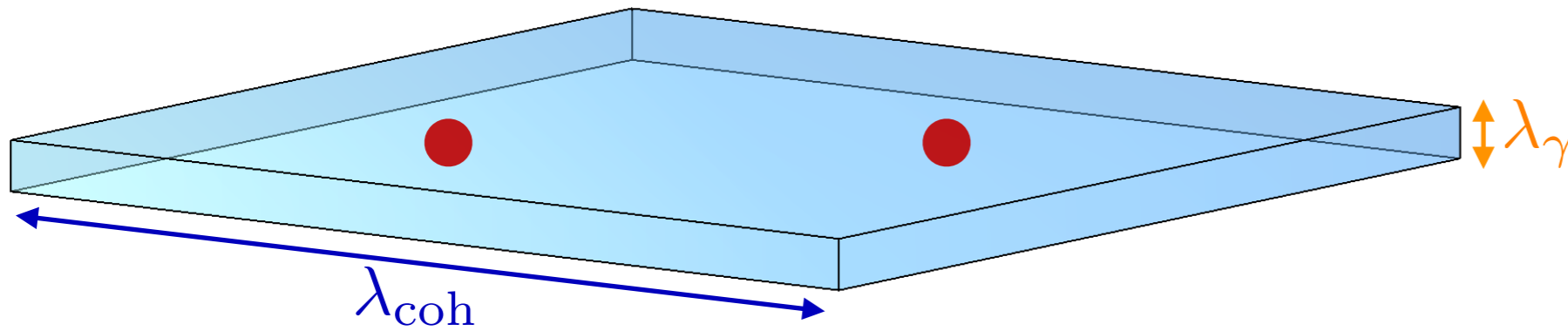
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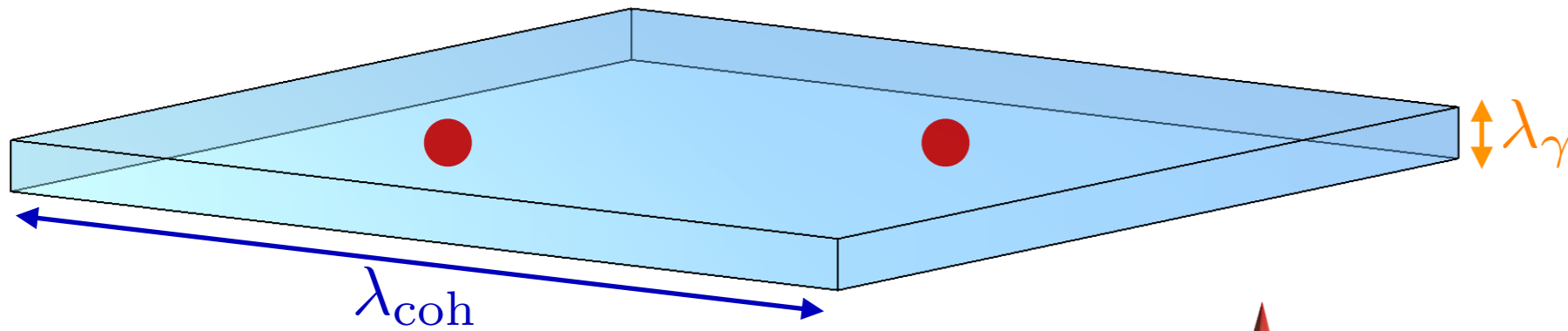
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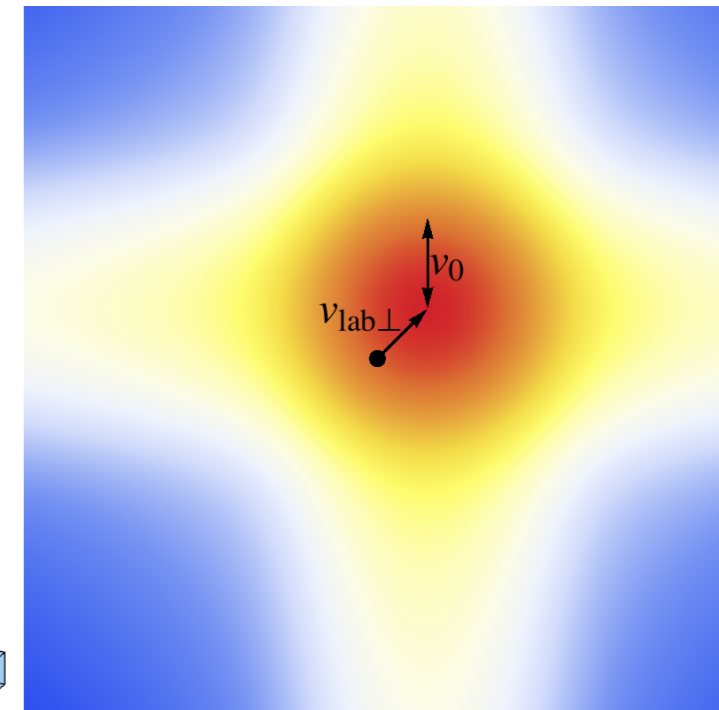
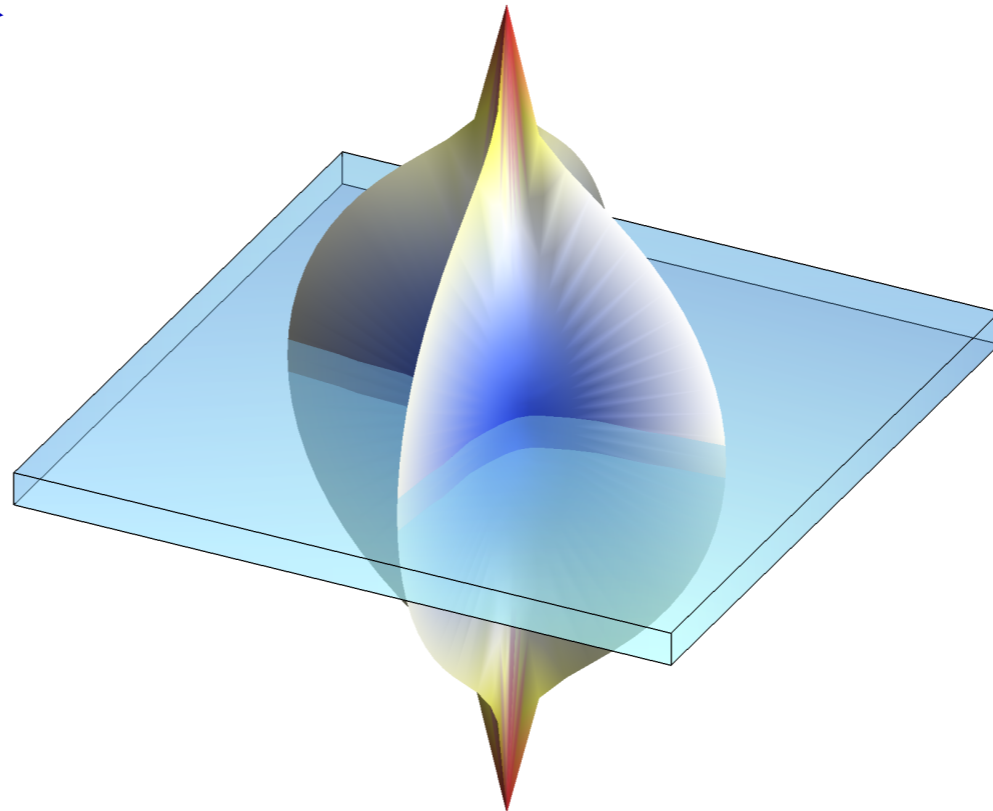
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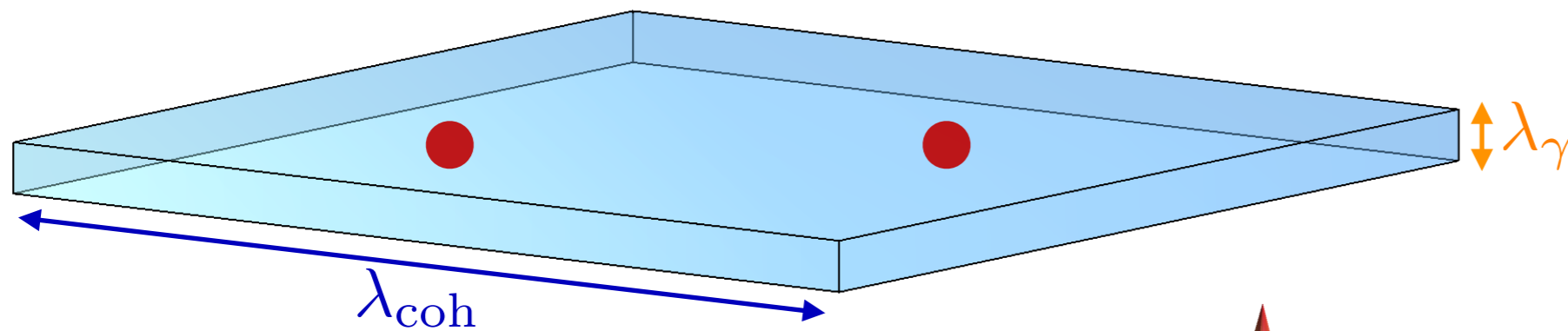
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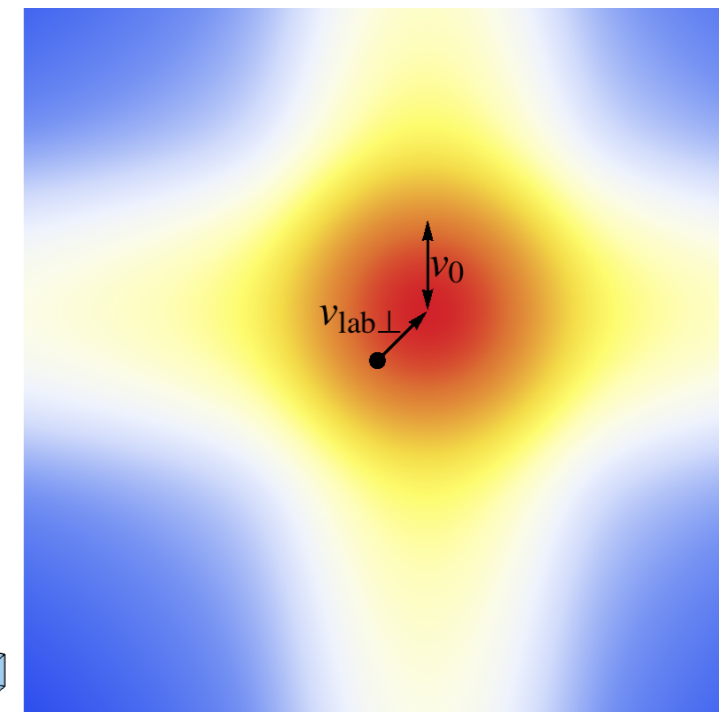
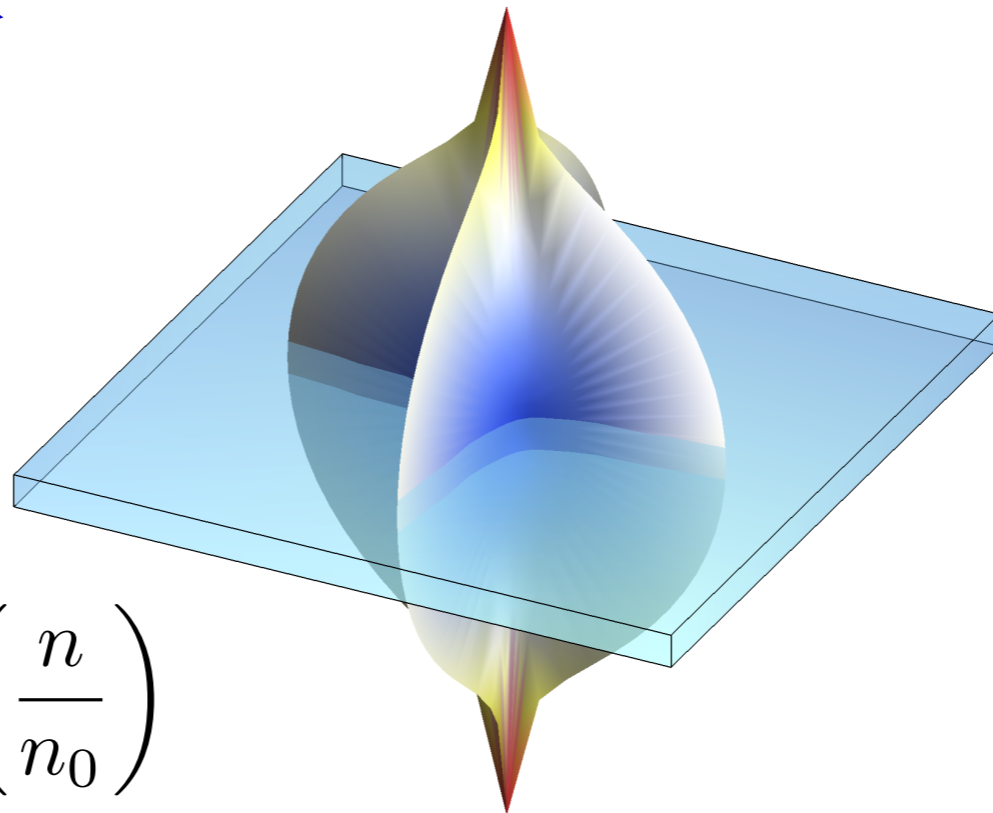
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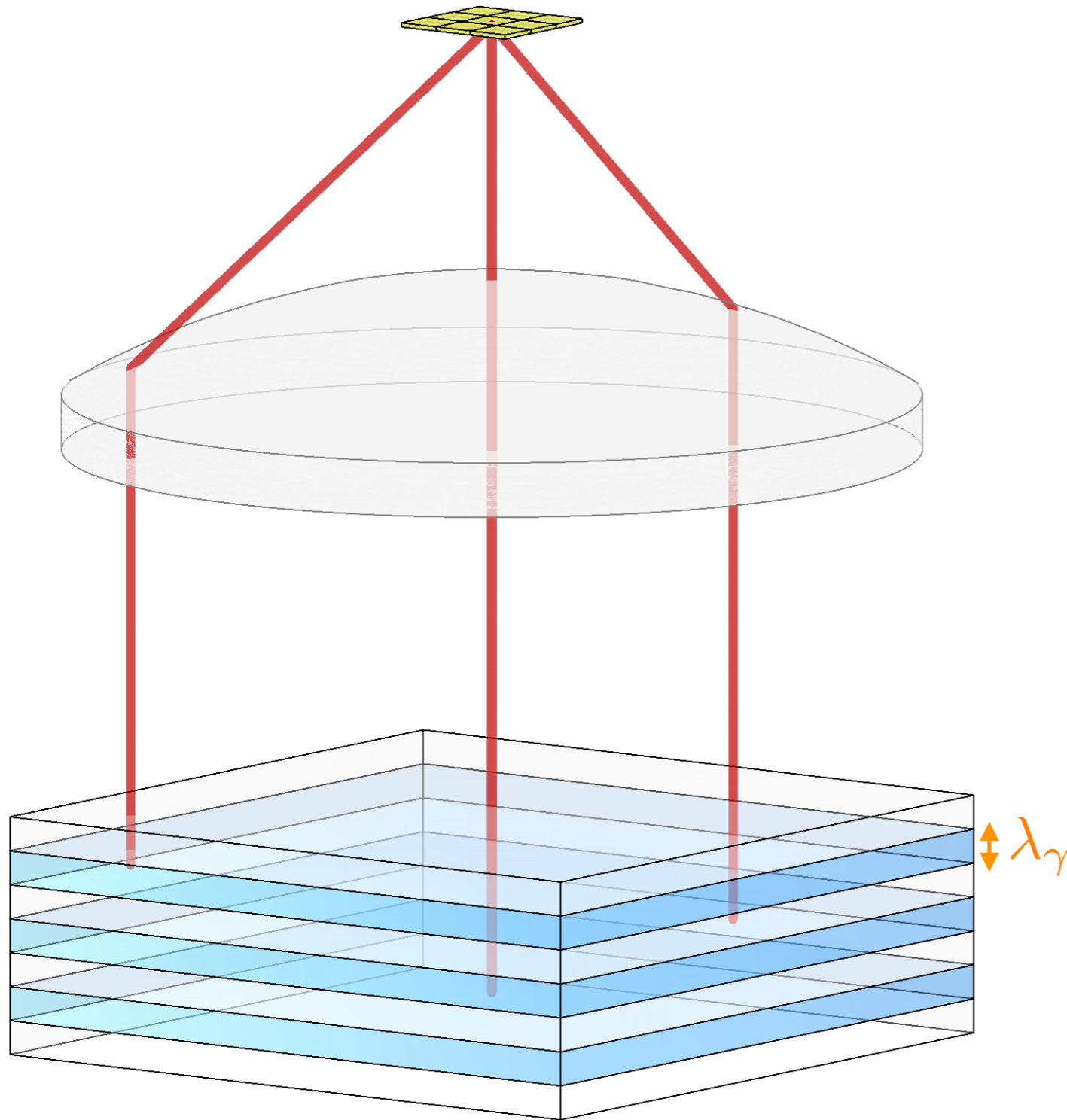
$$|0\rangle + \varepsilon e^{-i[mt + \varphi(\mathbf{x})]} |1\rangle$$



$$\bar{r} \simeq 1 + \frac{8\pi n}{m^4 R_z}$$

$$\simeq 1 + \frac{5 \times 10^6}{m R_z} \left(\frac{1 \text{ eV}}{m} \right)^3 \left(\frac{n}{n_0} \right)$$

Stack configuration



Phase I

$$A = \pi(0.3 \text{ m})^3$$

$$D = 1 \text{ mm}$$

$$\text{DCR} = 10^{-5} \text{ Hz}$$

Phase II

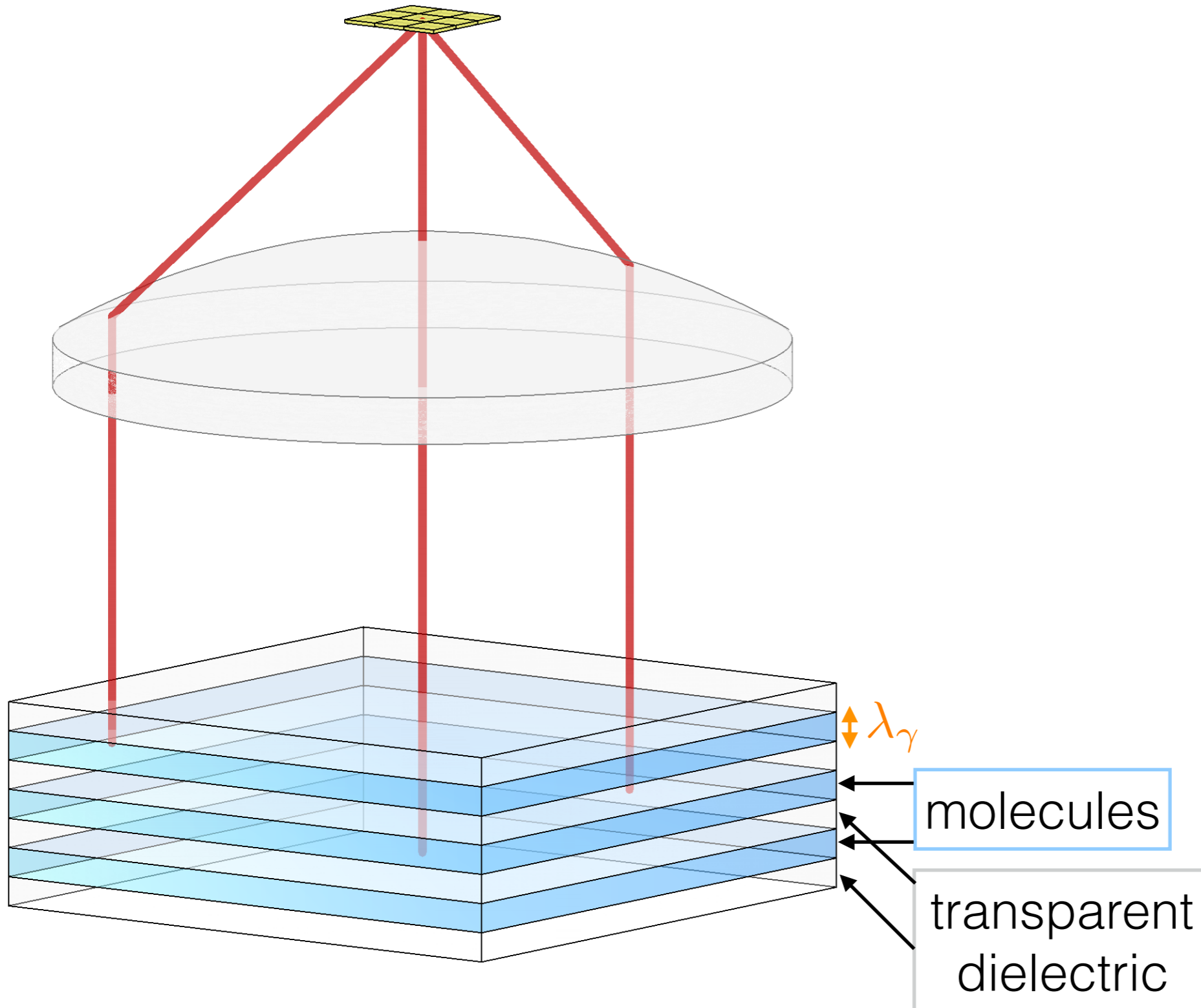
$$A = \pi(2 \text{ m})^3$$

$$D = 100 \text{ mm}$$

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$$T = 100 \text{ K}$$

Stack configuration



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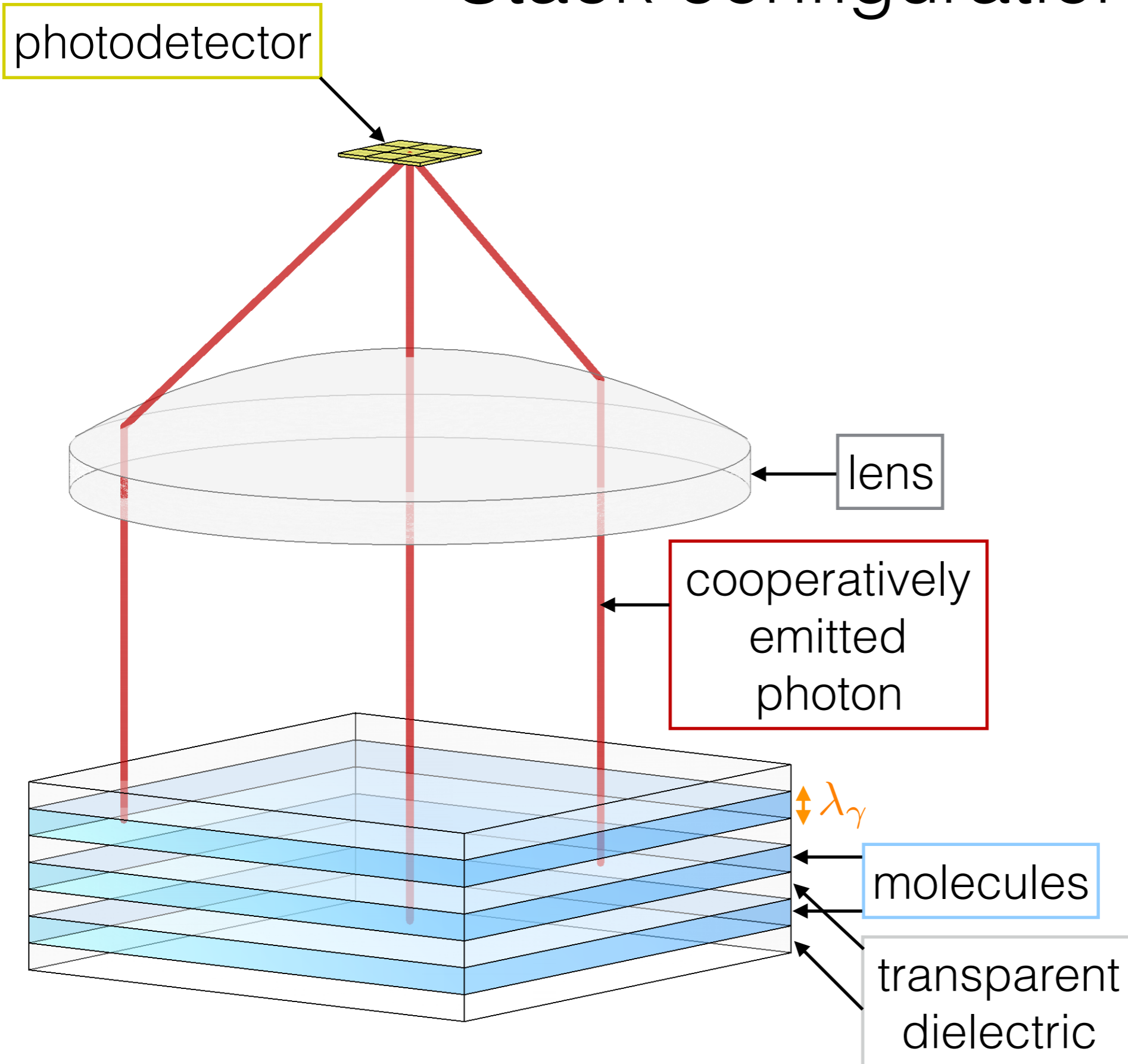
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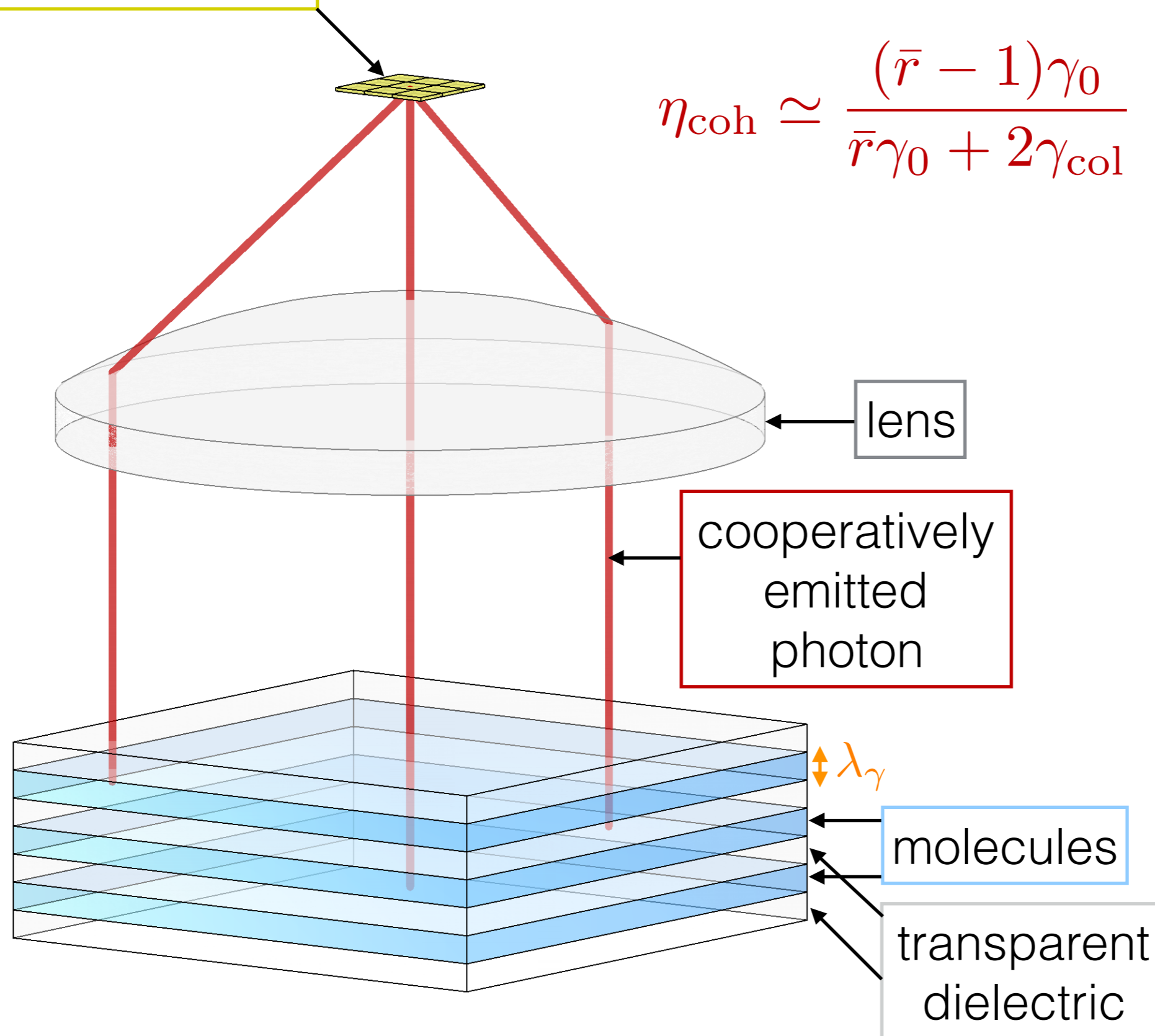
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Stack configuration

photodetector



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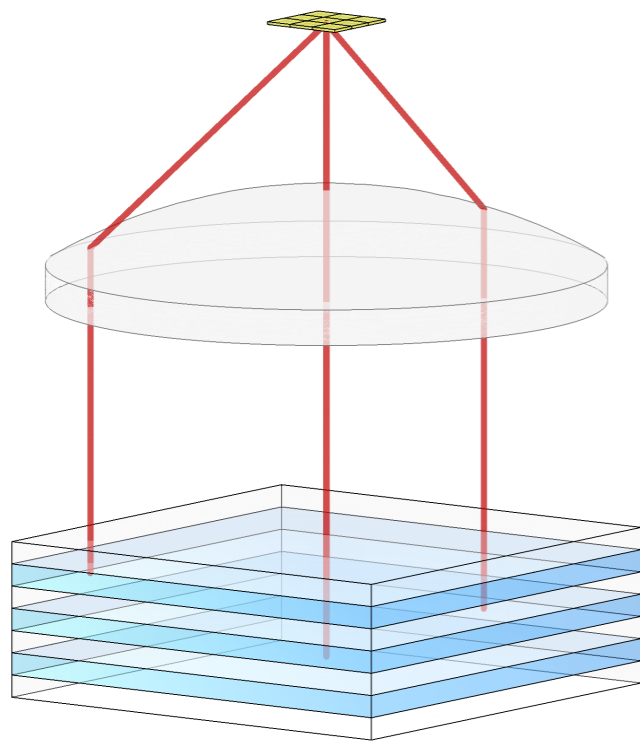
$$\text{DCR} = 10^{-7} \text{ Hz}$$

$$T = 100 \text{ K}$$

Absorption + Cooperative emission

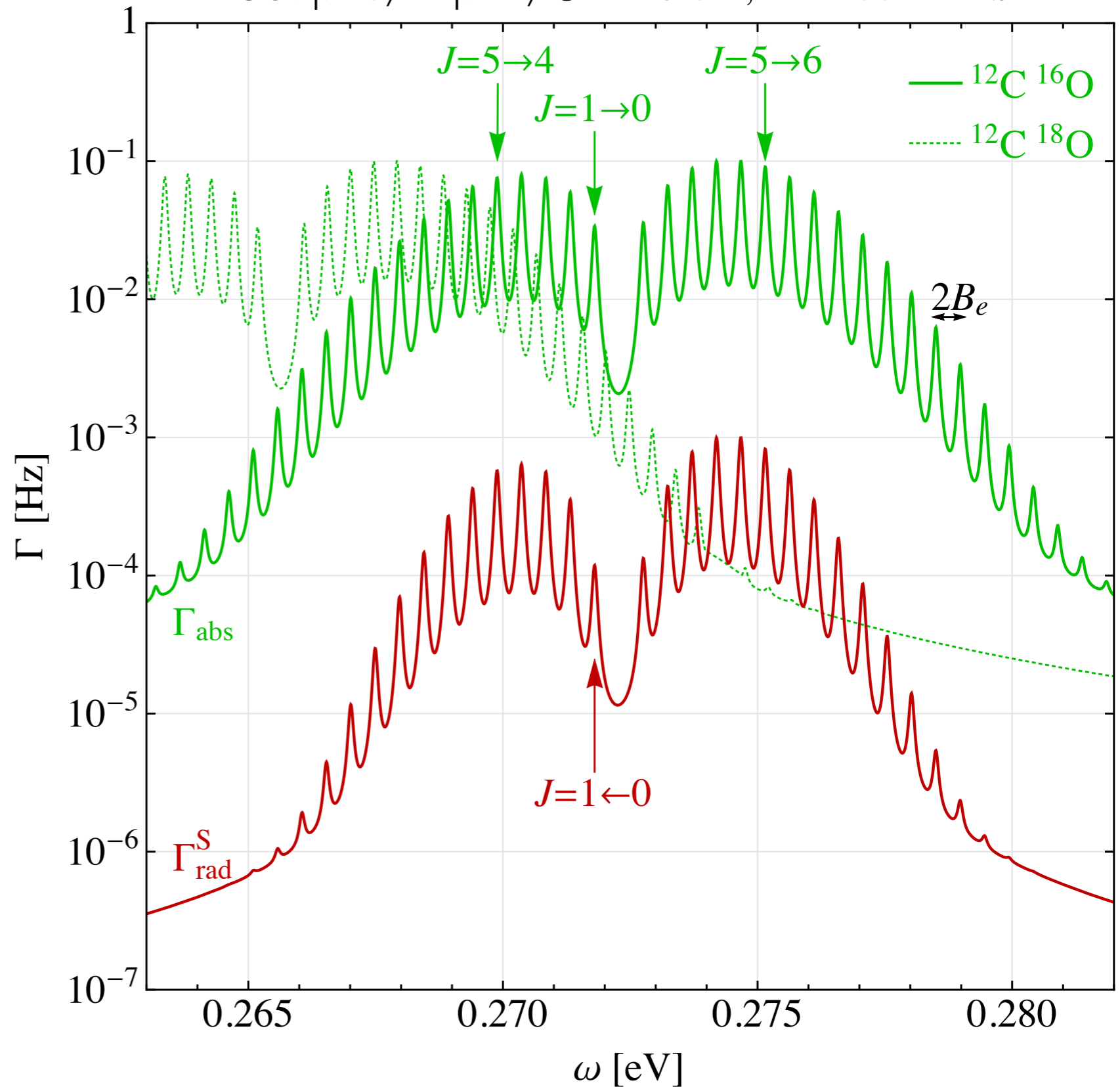
hidden photon DM

$$\epsilon = 10^{-12}$$



$$V = \pi(30 \text{ cm})^2(1 \text{ mm})$$

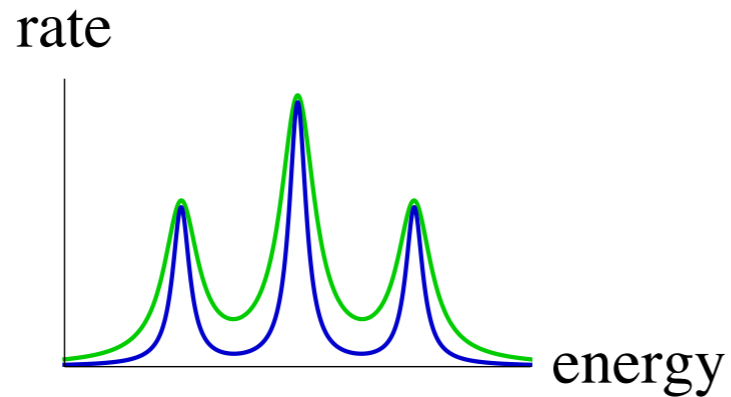
CO: $|v=0\rangle \rightarrow |v=1\rangle$ @ $P = 5 \text{ bar}$, $T = 100 \text{ K}$ in SI



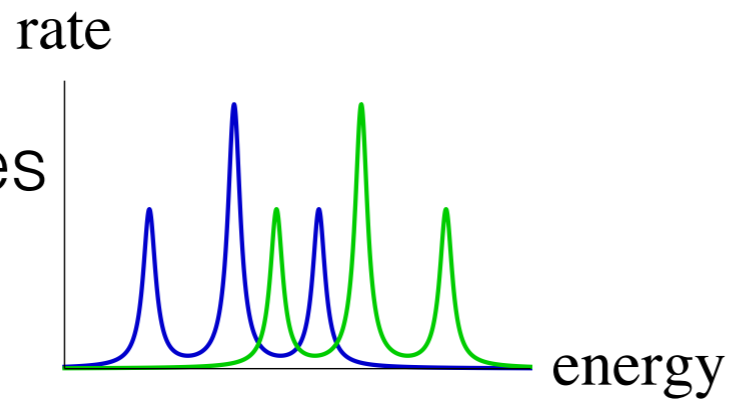
Key considerations

frequency coverage

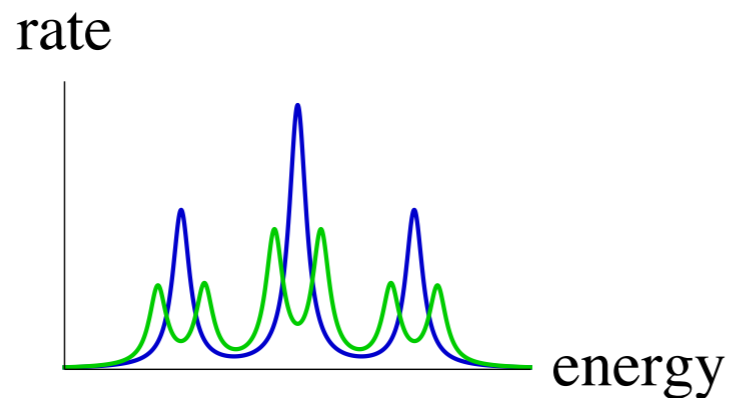
collisional broadening



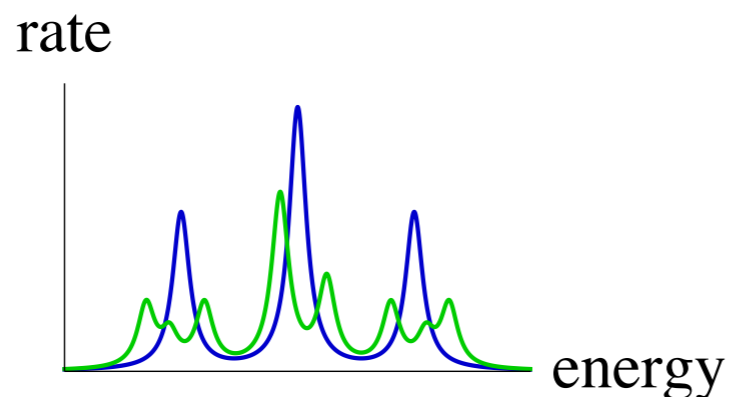
molecular species or isotope shift



Zeeman tuning (magnetic field)

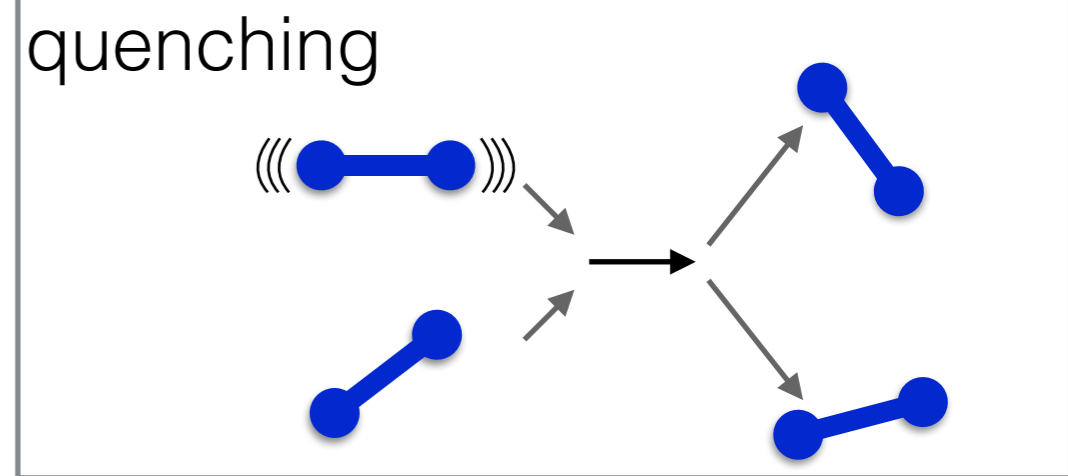
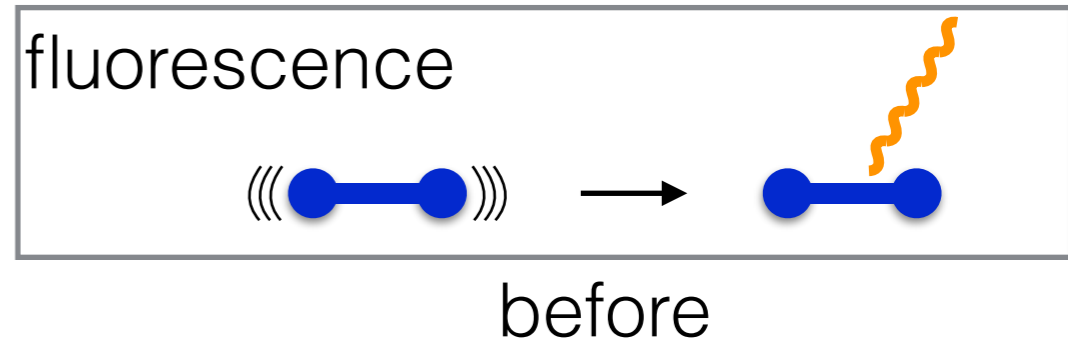


Stark tuning (electric field)

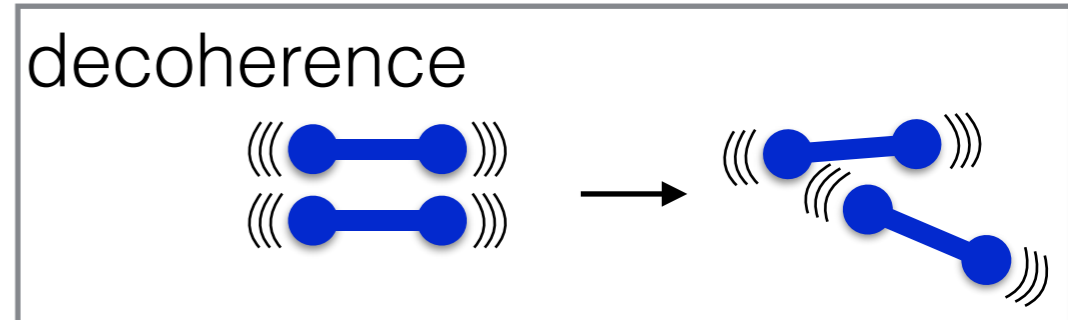
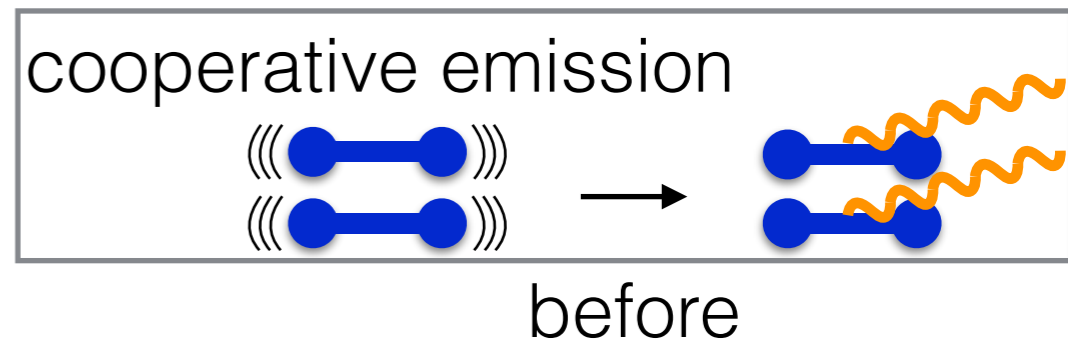


radiative efficiency

Bulk:



Stack:



Photon detection

$$\Gamma_{\text{det}} = \eta_{\text{det}} \Gamma_{\text{rad}} = \eta_{\text{det}} \eta_{\text{rad}} \Gamma_{\text{abs}} \quad \eta_{\text{det}} \sim \eta_{\text{opt-thin}} \eta_{\text{refl}} \eta_{\text{transm}} \eta_{\text{PD}}$$

Bulk Phase I: **PMT**

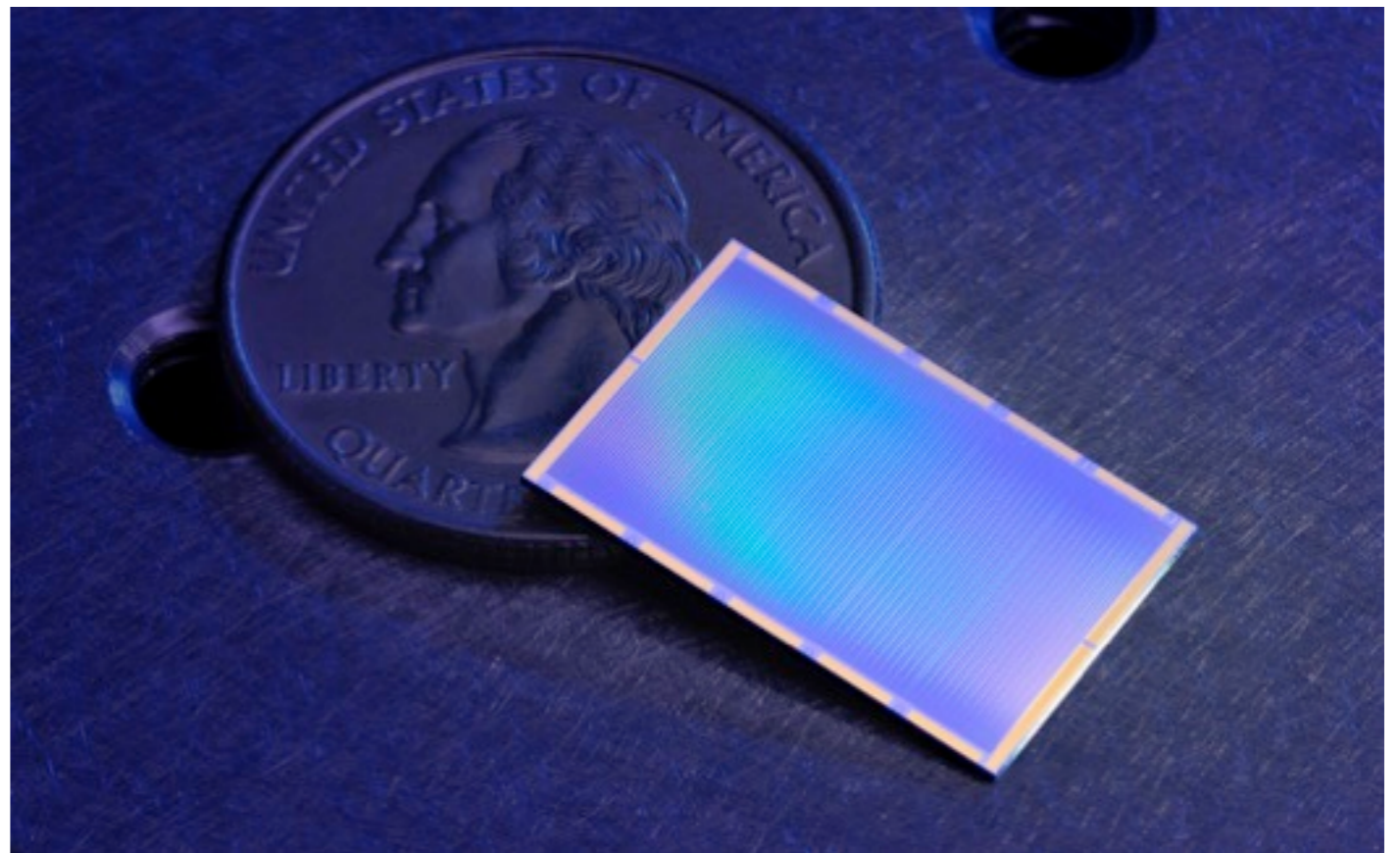
$$\eta_{\text{det}} \sim 0.3$$



Hamamatsu

Bulk Phase II:
Stack Phase I + II: **MKID**

$$\eta_{\text{det}} \sim 0.5 \quad \Delta E \sim 0.1 \text{ eV}$$



B. Mazin

Backgrounds

dark count rate (DCR): high-reflectivity coatings
cryogenic photodetectors: SNSPD, MKID, TES

thermal occupation / BBR: $nV e^{-\frac{\omega_0}{T}} \ll 1$

natural/cosmogenic radioactivity: high-purity shield + components

10^{-12} mass fraction $^{238}\text{U} \rightarrow \Gamma_{\text{RD}} \sim 10^{-2}$ Hz for meter-scale volume

veto trigger: $\left\{ \begin{array}{l} \text{many, high-}E \text{ particles} \\ \text{ionized electrons} \\ \text{timing + fast relaxation} \end{array} \right.$

Stack configuration: 84% of signal in 10^{-7} solid angle

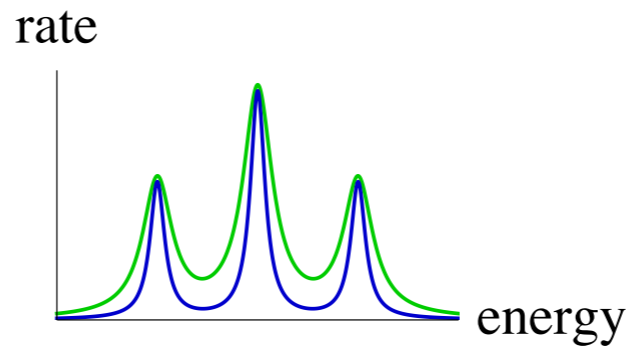
cosmic rays: underground and/or muon scintillator (99.9%)

Signal discrimination

DM

radioactivity
cosmic rays

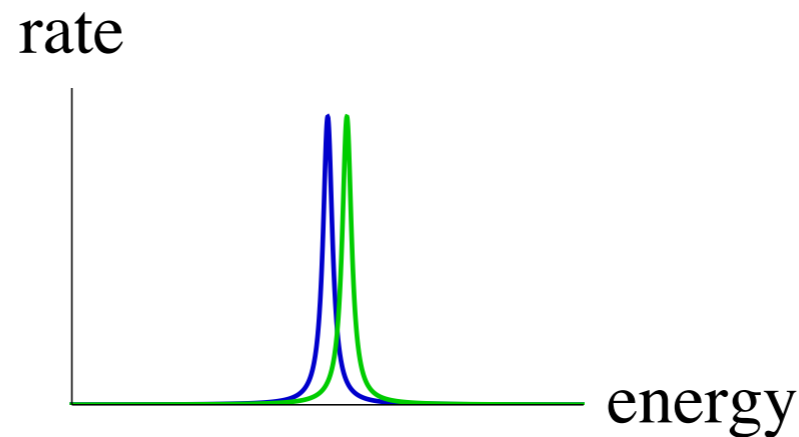
collisional broadening:



$$\propto n^1 \text{ or } \propto n^0$$

frequency shifts:

species
isotope
Stark
Zeeman



broadband

selection rules:

allowed
+
forbidden

allowed

directional emission:

solid angle $\sim 10^{-6}$

solid angle $\sim 4\pi$

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bosonic DM couplings | two-level system dynamics | molecular levels

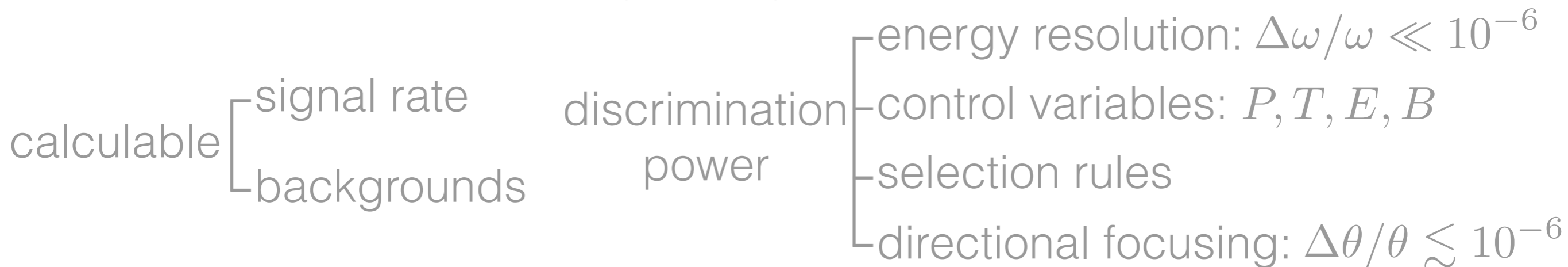
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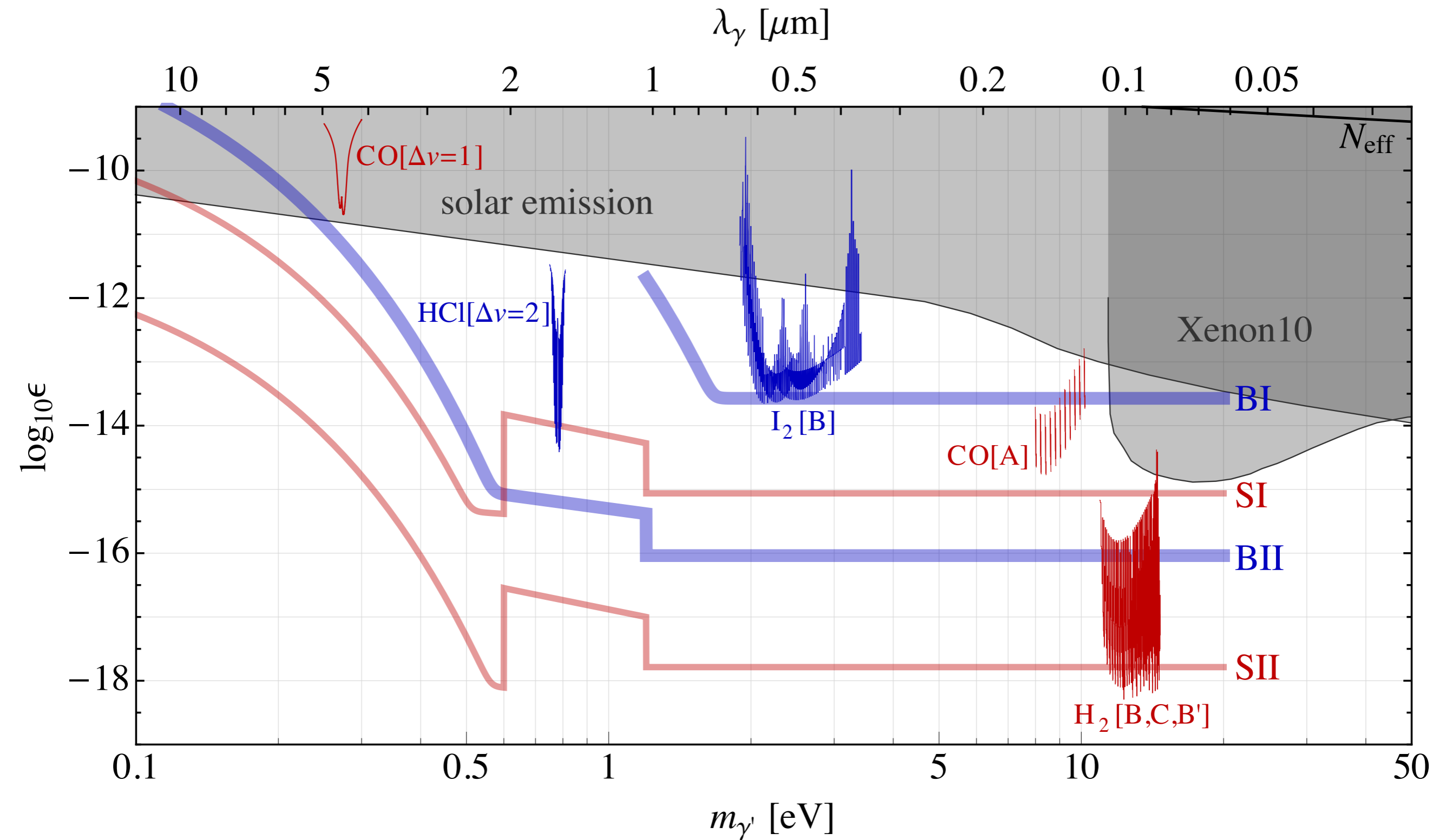
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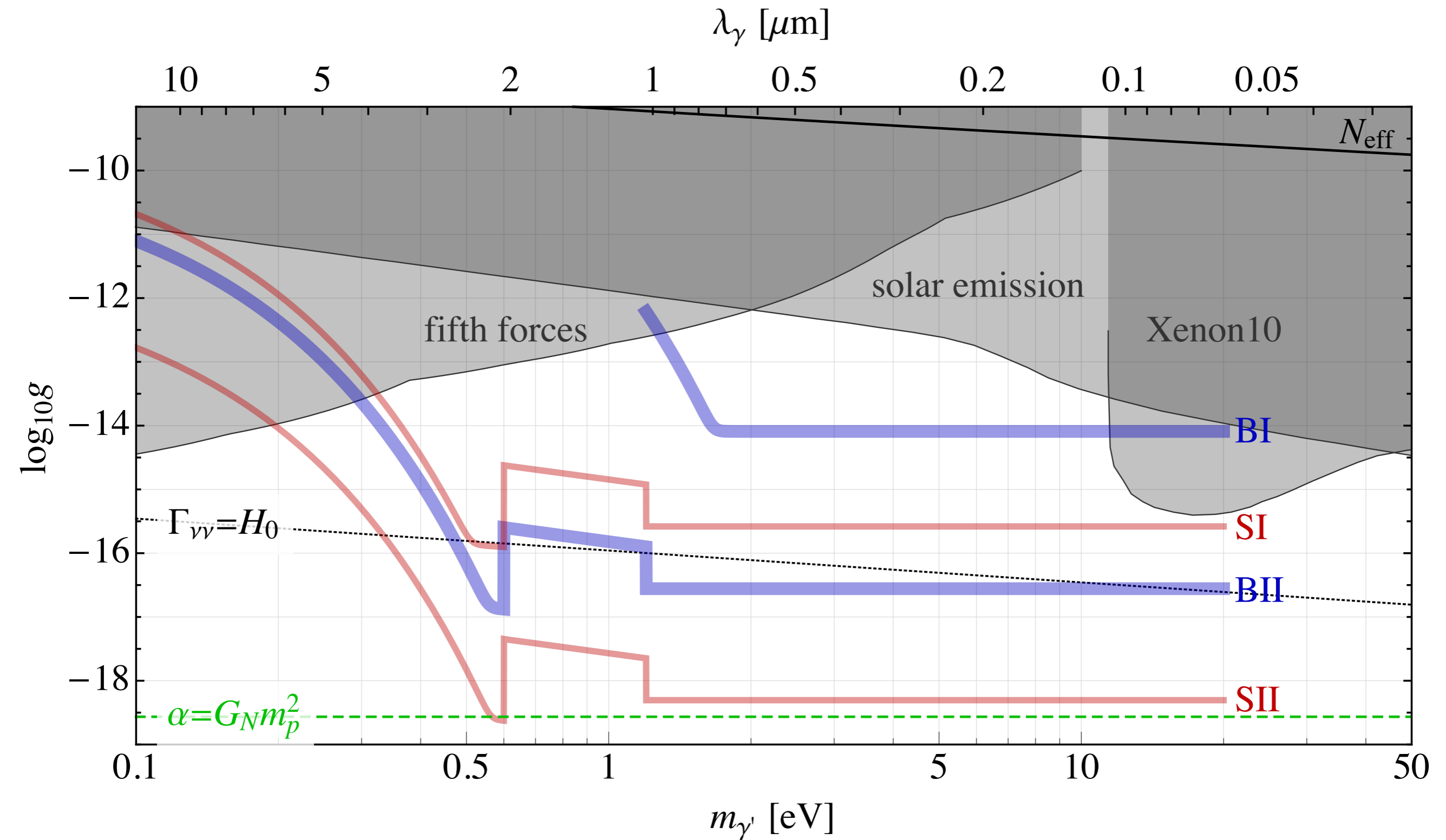
DM-induced transitions

DM type	Interaction Hamiltonian δH	Transition type and selection rules	Ω [rad s ⁻¹]
parity-even	$(d_{m_e} + d_e)\tilde{\phi}k_e R_e R$	vib $\Delta v = 1, \Delta J = 0$	$5.5 \times 10^{-9} \frac{d_{m_e}}{10^6}$
	$(3d_{m_e} + 4d_e)\tilde{\phi}\frac{k_e}{2}(R - R_e)^2$	vib $\Delta v = 2, \Delta J = 0$	$7.1 \times 10^{-10} \frac{d_{m_e}}{10^6}$
	$(d_g + Q_{\hat{m}_q} d_{\hat{m}_q})\tilde{\phi}\frac{\nabla_N^2}{2M}$	vib $\Delta v = 2, \Delta J = 0$	$2.4 \times 10^{-11} \frac{d_{\hat{m}_q}}{10^6} \frac{Q_{\hat{m}_q}}{0.1}$
	$(\Delta Q_i d_i)M(\nabla\tilde{\phi} \cdot \mathbf{R})$	vib $\Delta v = 1, \Delta J = \pm 1$	$3.0 \times 10^{-10} \frac{d_i}{10^6} \frac{\Delta Q_i}{10^{-2}}$
		rot $\Delta J = 1$	$4.1 \times 10^{-13} \frac{d_i}{10^2} \frac{\Delta Q_i}{10^{-2}}$
spin-0	$(d_{m_e} + d_e)\tilde{\phi}\frac{\nabla_e^2}{2m_e}$	el $\Delta\Lambda = 0, \Delta i = 0$	$9.5 \times 10^{-10} \frac{d_{m_e}}{10^6}$
	$d_{m_e} m_e \nabla\tilde{\phi} \cdot \mathbf{r}_e$	el $ \Delta\Lambda \leq 1, \Delta i = 1$	$7.5 \times 10^{-11} \frac{d_{m_e}}{10^6}$
	$G_{aNN} \partial_t a \boldsymbol{\sigma}_N \cdot \frac{-i\nabla_N}{M}$	vib $\Delta v = 1, \Delta J = \pm 1, \Delta S_N \leq 1$	$1.7 \times 10^{-10} \frac{G_{aNN}}{10^{-8}/\text{GeV}}$
		rot $\Delta J = 1, \Delta S_N = 1$	$2.5 \times 10^{-11} \frac{G_{aNN}}{10^{-8}/\text{GeV}}$
	parity-odd	$\frac{d_\theta}{f_a} a \boldsymbol{\sigma}_N \cdot \mathbf{E}$	vib $\Delta v = 1, \Delta J = \pm 1, \Delta S_N \leq 1$
rot $\Delta J = 1, \Delta S_N \leq 1$			$5.8 \times 10^{-13} \frac{10^8 \text{ GeV}}{f_a}$
	$G_{aee} \partial_t a \boldsymbol{\sigma}_e \cdot \frac{-i\nabla_e}{m_e}$	el $ \Delta\Lambda \leq 1, \Delta i = 1, \Delta S_e \leq 1$	$4.0 \times 10^{-10} \frac{G_{aee}}{10^{-10}/\text{GeV}}$
	$G_{a\gamma\gamma} \int d^3\mathbf{x} a \mathbf{E} \cdot \mathbf{B}$? ?	?
spin-1	kinetic mixing $\epsilon \boldsymbol{\mu}_e \cdot \mathbf{E}'$	el $ \Delta\Lambda \leq 1, \Delta i = 1$	$1.5 \times 10^{-6} \frac{\epsilon}{10^{-14}}$
		vib $\Delta v = 1, \Delta J = \pm 1$	$1.3 \times 10^{-5} \frac{\epsilon}{10^{-12}}$
		rot $\Delta J = 1$	$1.5 \times 10^{-2} \frac{\epsilon}{10^{-10}}$
	$B - L$ charge $\boldsymbol{\mu}_{B-L} \cdot \mathbf{E}_{B-L}$	el $ \Delta\Lambda \leq 1, \Delta i = 1$	$5.0 \times 10^{-6} \frac{g}{10^{-14}}$
		vib $\Delta v = 1, \Delta J = \pm 1$	$4.3 \times 10^{-7} \frac{g}{10^{-14}}$
	rot $\Delta J = 1$	$5.0 \times 10^{-10} \frac{g}{10^{-18}}$	

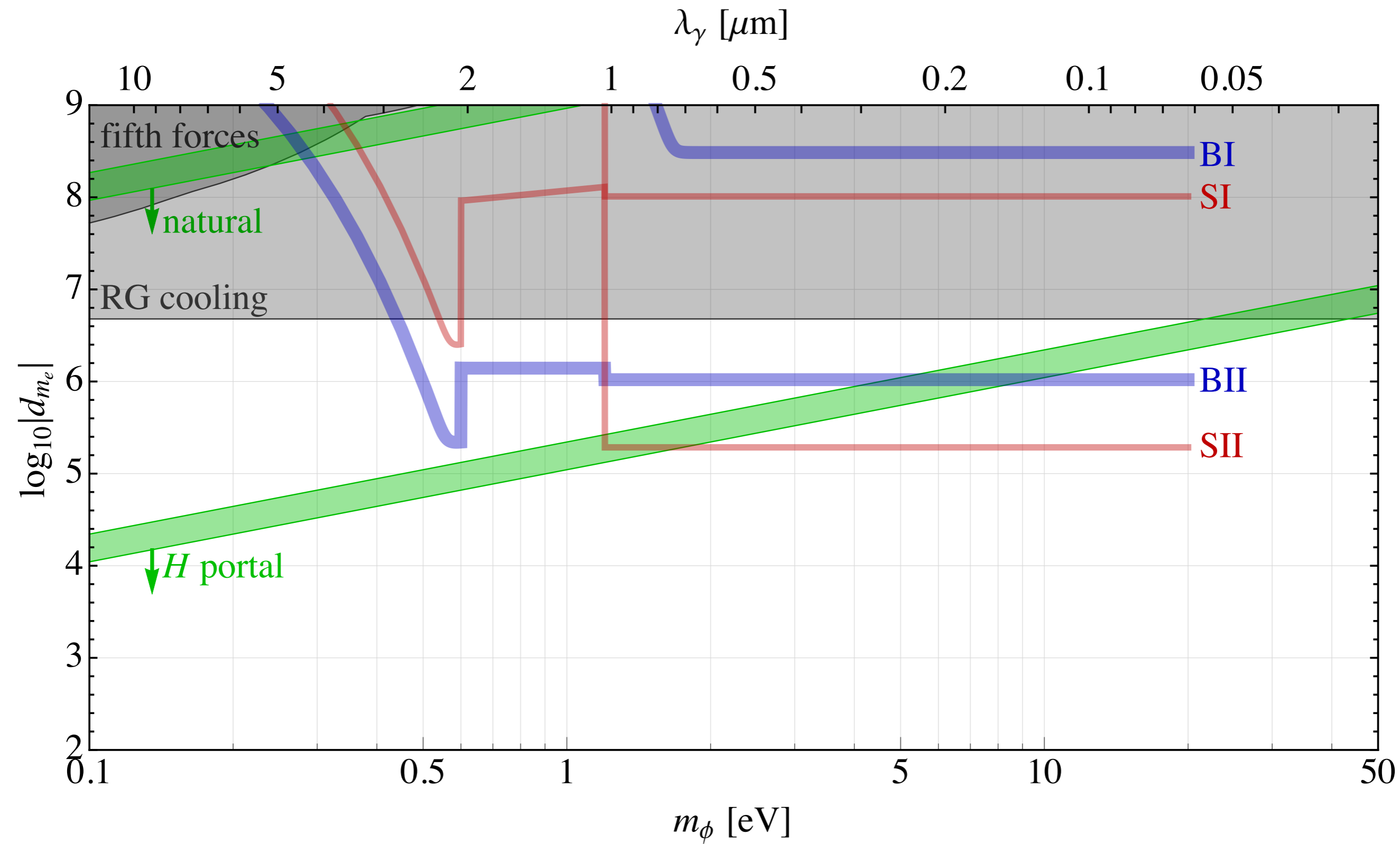
Kinematically mixed photon



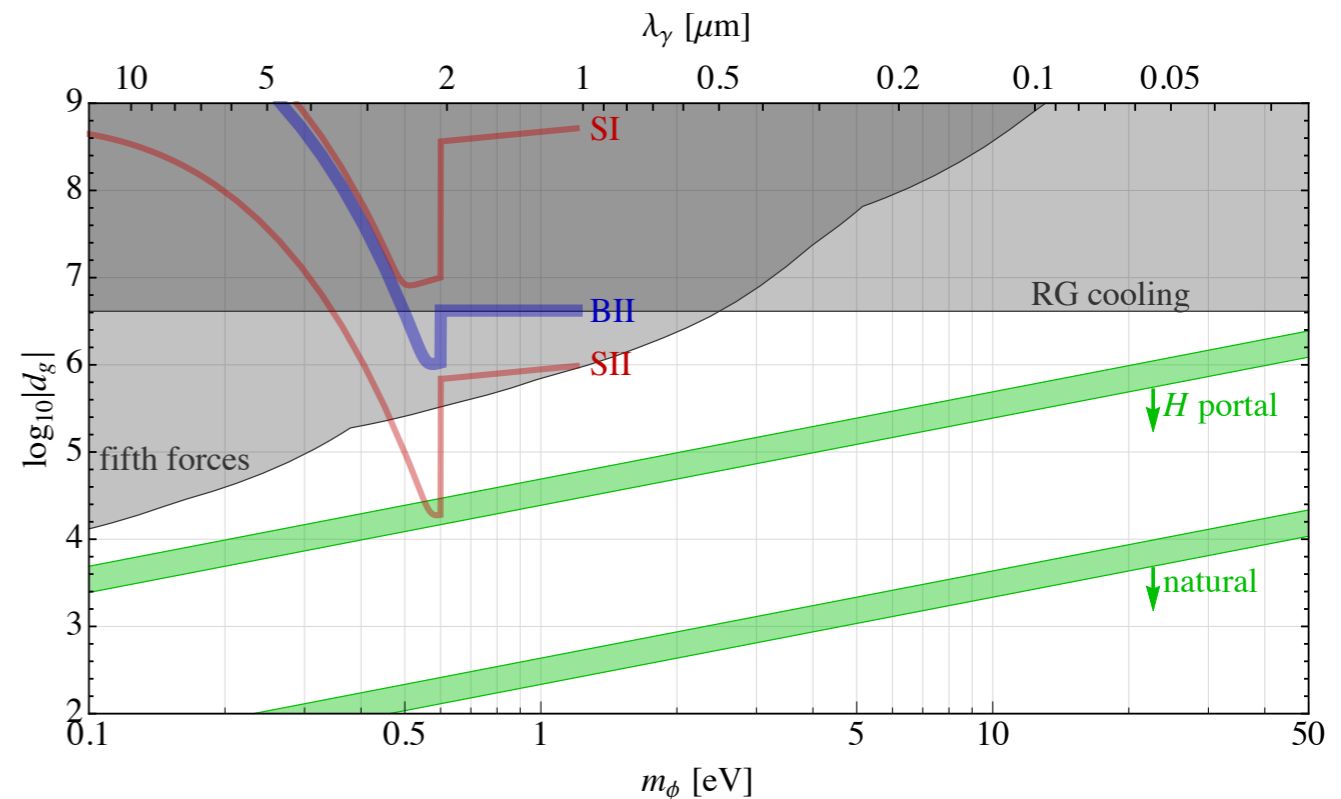
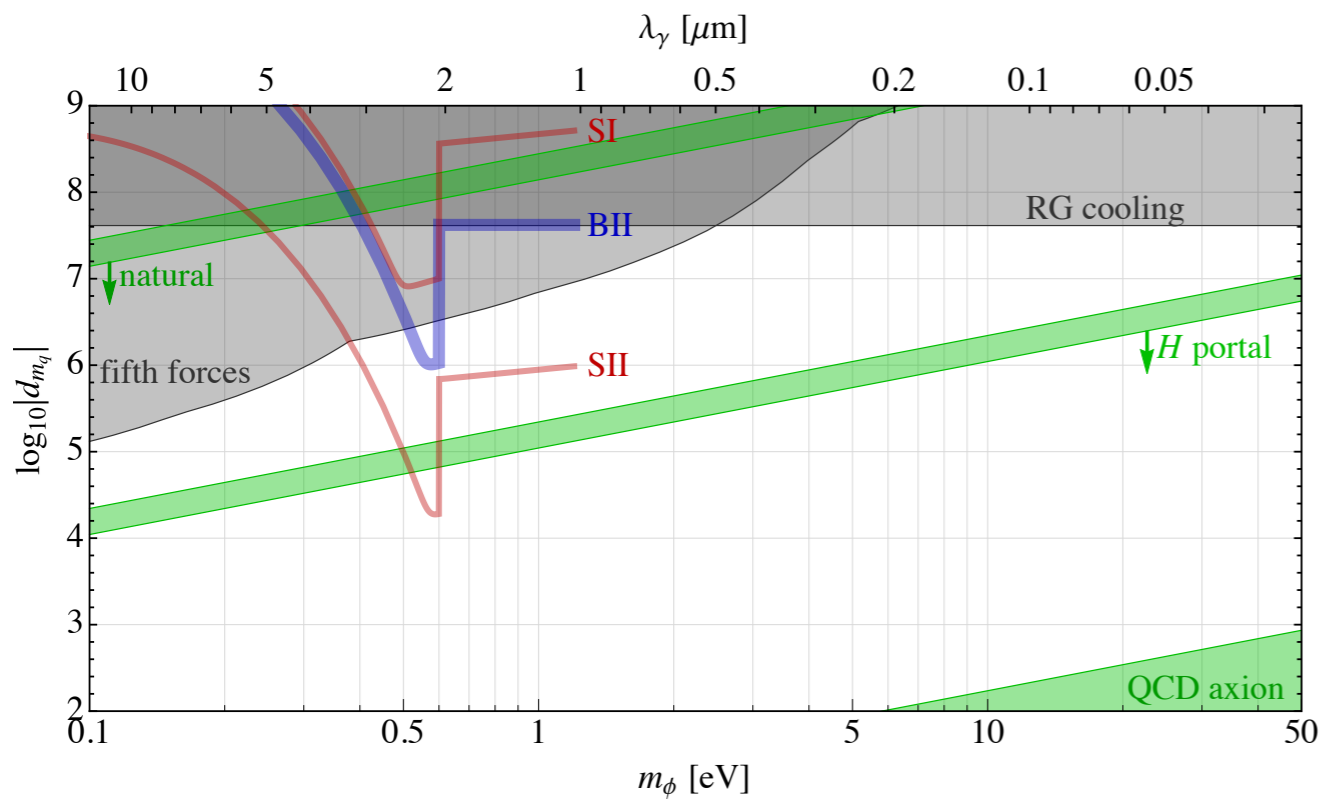
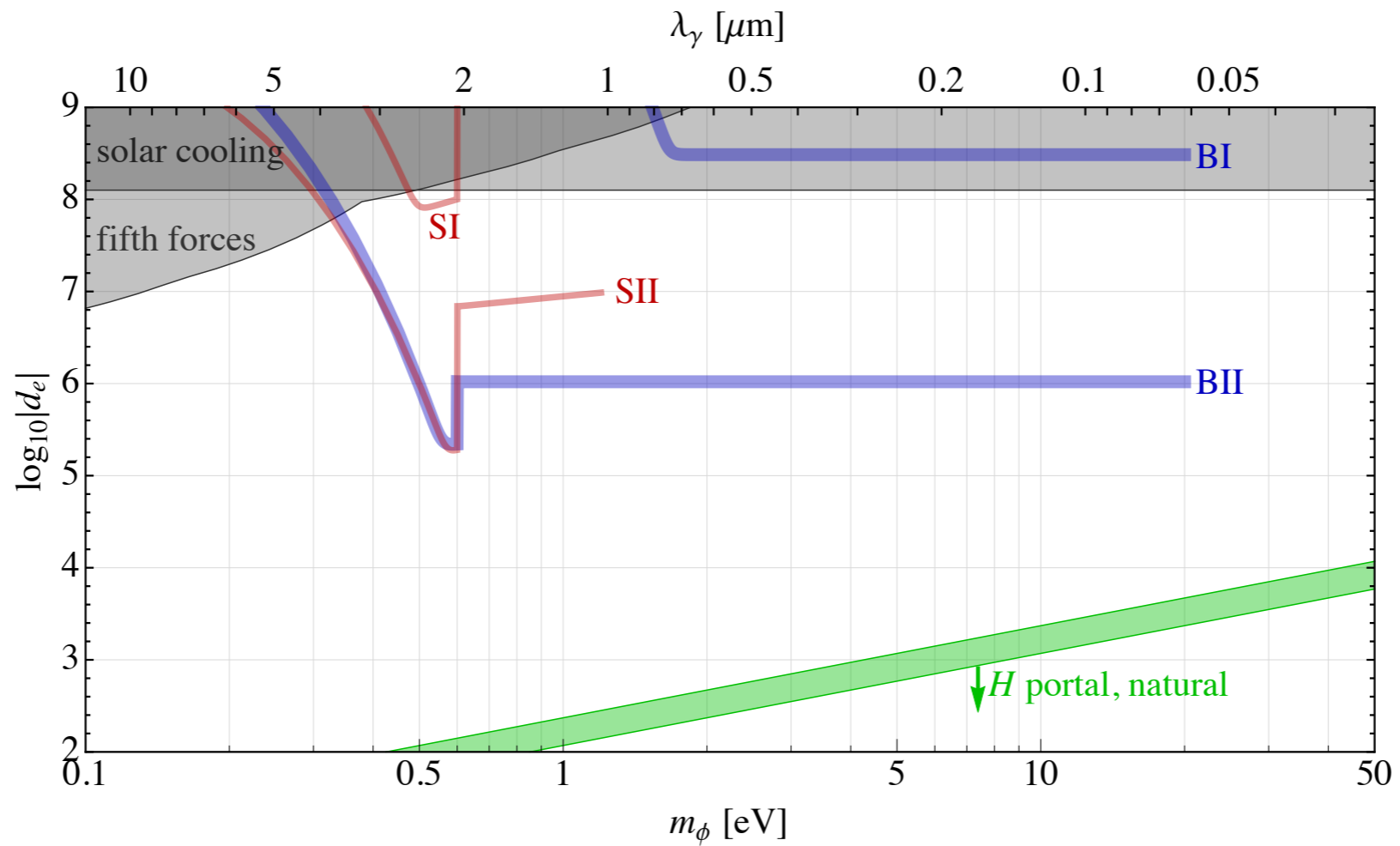
$B-L$ gauge boson



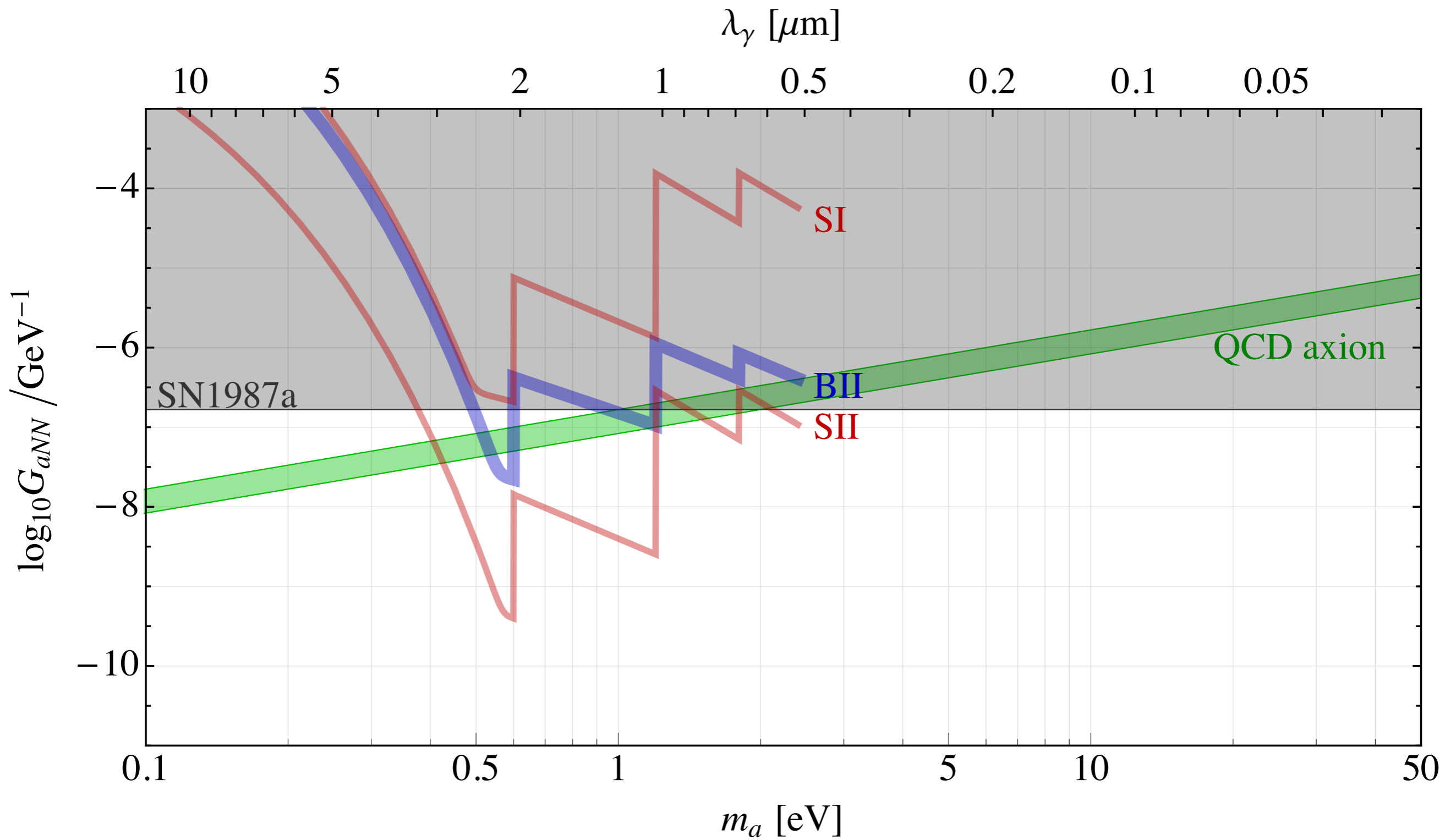
Modulus: electron coupling



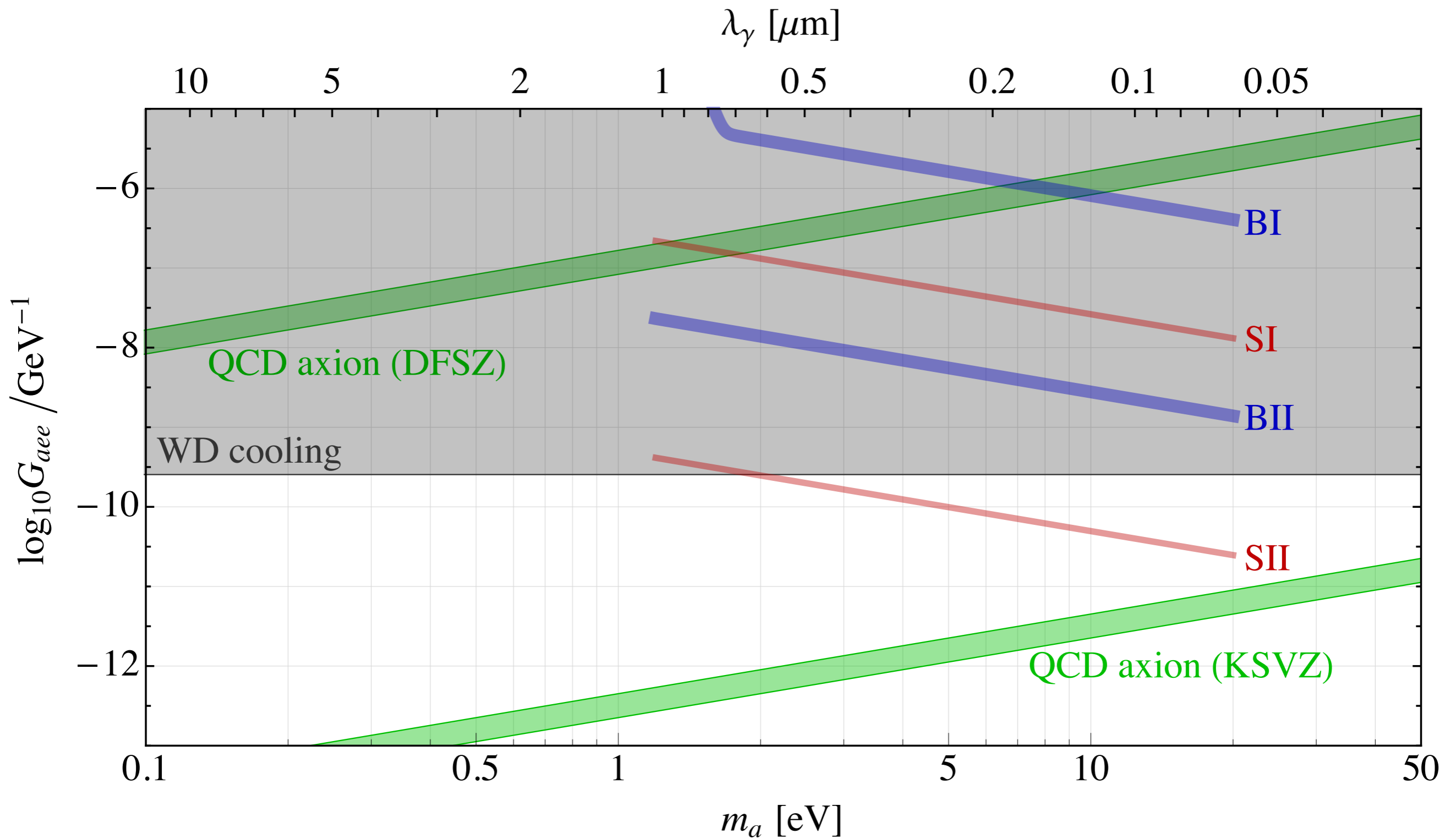
Modulus: photon, quark, gluon



Axion: nuclear coupling



Axion: electron coupling



Conclusions

(1) Dark matter fields can resonantly excite a molecular system

dark matter absorption = weak, monochromatic signal

gas of small molecules = multimode resonator with large signal rate

(2) Experimental setup

low background rates

tunable frequency response: superb discrimination power

established field with photodetector technology advancements

cooperative focusing effects: directional isolation and information

(3) Dark matter sensitivity $0.2 \text{ eV} < m < 20 \text{ eV}$

spin-1: mixed photon, $B-L$ gauge boson, ...

spin-0, parity-even: moduli fields for electron, quark, photon, gluon, ...

spin-0, parity-odd: axion coupling to electrons, nucleons, photons, ...

Future: $m < 0.2 \text{ eV}$ | $m > 20 \text{ eV}$ | other use? | map out molecular forest

Phase I prototypes

Phase II R&D: slab/stack manufacturing | optimize photodetectors

Bosonic dark matter couplings

EFT of DM bosons coupled to electrons, quarks, photons, gluons

spin-1

$$\epsilon A'_\mu J_{\text{EM}}^\mu$$

$$g A'_\mu J_{B-L}^\mu$$

spin-0

parity-even

$$\phi d_{m_e} \sqrt{4\pi G_N m_e} \bar{\psi}_e \psi_e$$

$$\phi d_{m_q} \sqrt{4\pi G_N m_q} \bar{\psi}_q \psi_q$$

$$\phi d_e \sqrt{4\pi G_N} \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$$\phi d_g \sqrt{4\pi G_N} \frac{\beta_3}{2g_3} G_{\mu\nu} G^{\mu\nu}$$

parity-odd

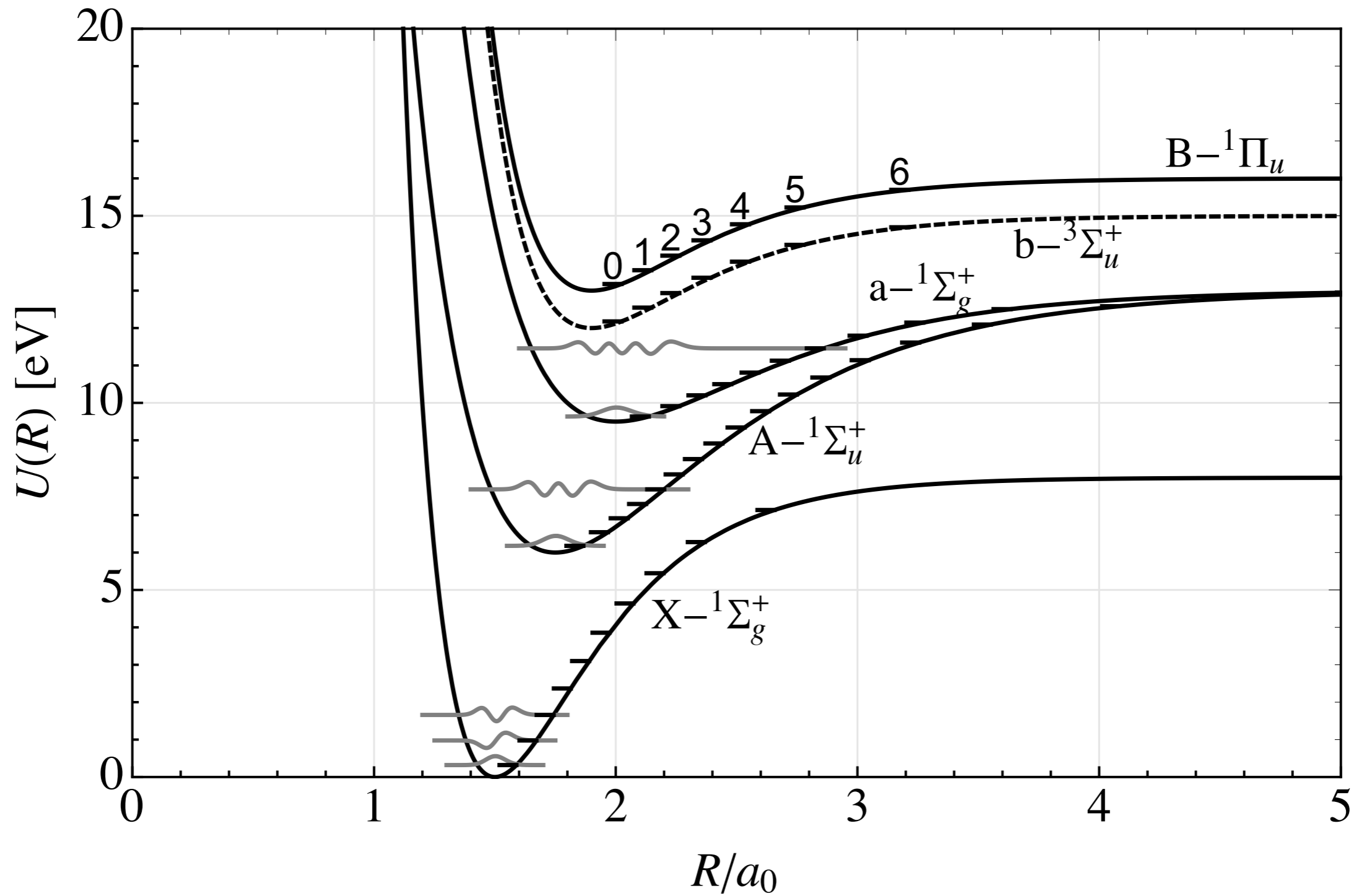
$$G_{aee} (\partial_\mu a) \bar{\psi}_e \gamma^\mu \gamma_5 \psi_e$$

$$G_{aqq} (\partial_\mu a) \bar{\psi}_q \gamma^\mu \gamma_5 \psi_q$$

$$G_{a\gamma\gamma} a \frac{1}{4} F_{\mu\nu} \tilde{F}^{\mu\nu}$$

$$\frac{\alpha_3}{8\pi} \frac{a}{f_a} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

Electronic states of a diatomic molecule



Molecular info

molecule	χ^{el}	T_e [cm $^{-1}$]	ω_e [cm $^{-1}$]	$\omega_e x_e$ [cm $^{-1}$]	B_e [cm $^{-1}$]	α_e [cm $^{-1}$]	R_e [Å]	T_b [K]
$^1\text{H}^1\text{H}$	c- $^3\Pi_u$	95938	2466.8	63.51	31.07	1.42	1.037	
	EF- $^1\Sigma_g^+$	100082.3	2588.9	130.5	32.68	1.818	1.011	
	C- $^1\Pi$	100089.8	2443.77	69.524	31.362	1.664	1.0327	
	B- $^1\Sigma_u^+$	91700	1358.09	20.888	20.015	1.184	1.2928	
	X- $^1\Sigma_g^+$	0	4401.21	121.33	60.853	3.062	0.74144	20.28
$^1\text{H}^2\text{H}$	C- $^1\Pi$	100092.9	2119.6	53.31	23.522	1.096	1.0329	
	B- $^1\Sigma_u^+$	91698.3	1177.16	15.59	15.071	0.820	1.2904	
	X- $^1\Sigma_g^+$	0	3813.1	91.65	45.655	1.986	0.74142	
$^2\text{H}^2\text{H}$	X- $^1\Sigma_g^+$	0	3115.50	61.82	30.443	1.0786	0.74152	
$^{16}\text{O}^{16}\text{O}$	X- $^3\Sigma_g^-$	0	1580.19	11.98	1.44563	0.0159	1.20752	90.19
$^{12}\text{C}^{16}\text{O}$	A- $^1\Pi$	65075.7	1518.2	19.40	1.6115	0.0232	1.2353	
	a'- $^3\Sigma^+$	55825.4	1228.60	10.468	1.3446	0.0189	1.3523	
	a- $^3\Pi$	48686.70	1743.4	14.3	1.69124	0.01904	1.20574	
	X- $^1\Sigma^+$	0	2169.81358	13.28831	1.93128	0.0175	1.12832	81.65
$^{12}\text{C}^{18}\text{O}$	X- $^1\Sigma^+$	0	2117.5	12.66	1.839	0.0163	1.128	
$^{14}\text{N}^{14}\text{N}$	A- $^3\Sigma_u^+$	50203.6	1460.64	13.87	1.4546	0.0180	1.2866	
	X- $^1\Sigma_g^+$	0	2358.57	14.324	1.99824	0.017318	1.09768	77.355
$^1\text{H}^{35}\text{Cl}$	X- $^1\Sigma^+$	0	2990.946	51.8	10.59341	0.30718	1.27455	188.10
$^{127}\text{I}^{127}\text{I}$	B- $^3\Pi_{0+u}$	15769.01	125.69	0.764	0.02903	0.000158	3.024	
	X- $^1\Sigma_g^+$	0	214.50	0.614	0.03737	0.000113	2.666	457.4

Configuration summary

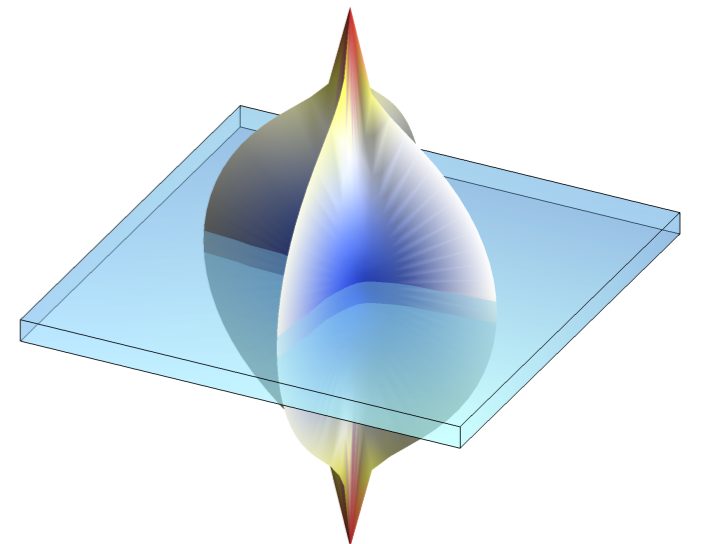
	Bulk	Stack
Phase I	$V = (0.3 \text{ m})^3$, $P \sim 0.1 \text{ bar}$, $T = 300 \text{ K}$ PMT, $\text{DCR} = 1 \text{ Hz}$, $A_{\text{det}} = (0.3 \text{ m})^2$, $\eta_{\gamma} = 0.3$ any electronic \rightarrow intermediate - Stark/Zeeman tuning, $t_{\text{shot}} = 10^2 \text{ s}$ $\delta\Omega \approx 2.9 \times 10^{-7} \text{ rad s}^{-1}$	$A = \pi(0.3 \text{ m})^2$, $D = 1 \text{ mm}$, $P \sim 10 \text{ bar}$, $T \sim 100 \text{ K}$ MKID/TES, $\text{DCR} \lesssim 10^{-5} \text{ Hz}$, $A_{\text{det}} = (0.3 \text{ mm})^2$, $\eta_{\gamma} = 0.5$ E1-allowed electronic E1-allowed vibrational collisional broadening, $t_{\text{shot}} = 10^6 \text{ s}$ $\delta\Omega \approx 9.4 \times 10^{-9} \text{ rad s}^{-1}$
Phase II	$V = (2 \text{ m})^3$, $P \sim 0.1 \text{ bar}$, $T \sim 100 \text{ K}$ MKID array, $\text{DCR} \lesssim 10^{-3} \text{ Hz}$, $A_{\text{det}} = (0.1 \text{ m})^2$, $\eta_{\gamma} = 0.5$ any electronic \rightarrow intermediate any vibrational with optically thin fluorescence Stark/Zeeman tuning, $t_{\text{shot}} = 10^3 \text{ s}$ $\delta\Omega \approx 9.9 \times 10^{-10} \text{ rad s}^{-1}$	$A = \pi(2 \text{ m})^2$, $D = 100 \text{ mm}$, $P \sim 10 \text{ bar}$, $T \sim 100 \text{ K}$ MKID/TES, $\text{DCR} \lesssim 10^{-7} \text{ Hz}$, $A_{\text{det}} = (2 \text{ mm})^2$, $\eta_{\gamma} = 1$ E1-allowed electronic E1-allowed vibrational collisional broadening, $t_{\text{shot}} = 10^7 \text{ s}$ $\delta\Omega \approx 1.8 \times 10^{-11} \text{ rad s}^{-1}$

Cooperation number

$$A(t, \mathbf{x}, \mathbf{x}') = \frac{q}{4\pi|\mathbf{x} - \mathbf{x}'|} \cos [\omega (t - |\mathbf{x} - \mathbf{x}'|) + \alpha_{\mathbf{v}} - m\mathbf{v} \cdot \mathbf{x}']$$

$$\begin{aligned} \langle A_{\text{tot}}(t, \mathbf{x})^2 \rangle_{\mathbf{v}, \alpha} &= \left\langle \left(\sum_{\mathbf{x}'} A(t, \mathbf{x}, \mathbf{x}') \right) \left(\sum_{\mathbf{y}'} A(t, \mathbf{x}, \mathbf{y}') \right) \right\rangle_{\mathbf{v}, \alpha} \\ &= \sum_{\mathbf{x}'} \left\langle A(t, \mathbf{x}, \mathbf{x}')^2 \right\rangle_{\mathbf{v}, \alpha} + \sum_{\mathbf{x}'} \sum_{\mathbf{y}' \neq \mathbf{x}'} \left\langle A(t, \mathbf{x}, \mathbf{x}') A(t, \mathbf{x}, \mathbf{y}') \right\rangle_{\mathbf{v}, \alpha} \\ &\simeq \int_V d^3 \mathbf{x}' n(\mathbf{x}') \frac{q^2}{2(4\pi)^2 L^2} \\ &\quad + \iint_V d^3 \mathbf{x}' d^3 \mathbf{y}' n(\mathbf{x}') n(\mathbf{y}') \frac{q^2}{2(4\pi)^2 L^2} \int d^3 \mathbf{v} f(\mathbf{v}) \cos [m(\hat{\mathbf{x}} - \mathbf{v}) \cdot (\mathbf{x}' - \mathbf{y}')] \\ &\equiv \frac{q^2}{2(4\pi)^2 L^2} nV r(\hat{\mathbf{x}}) \end{aligned}$$

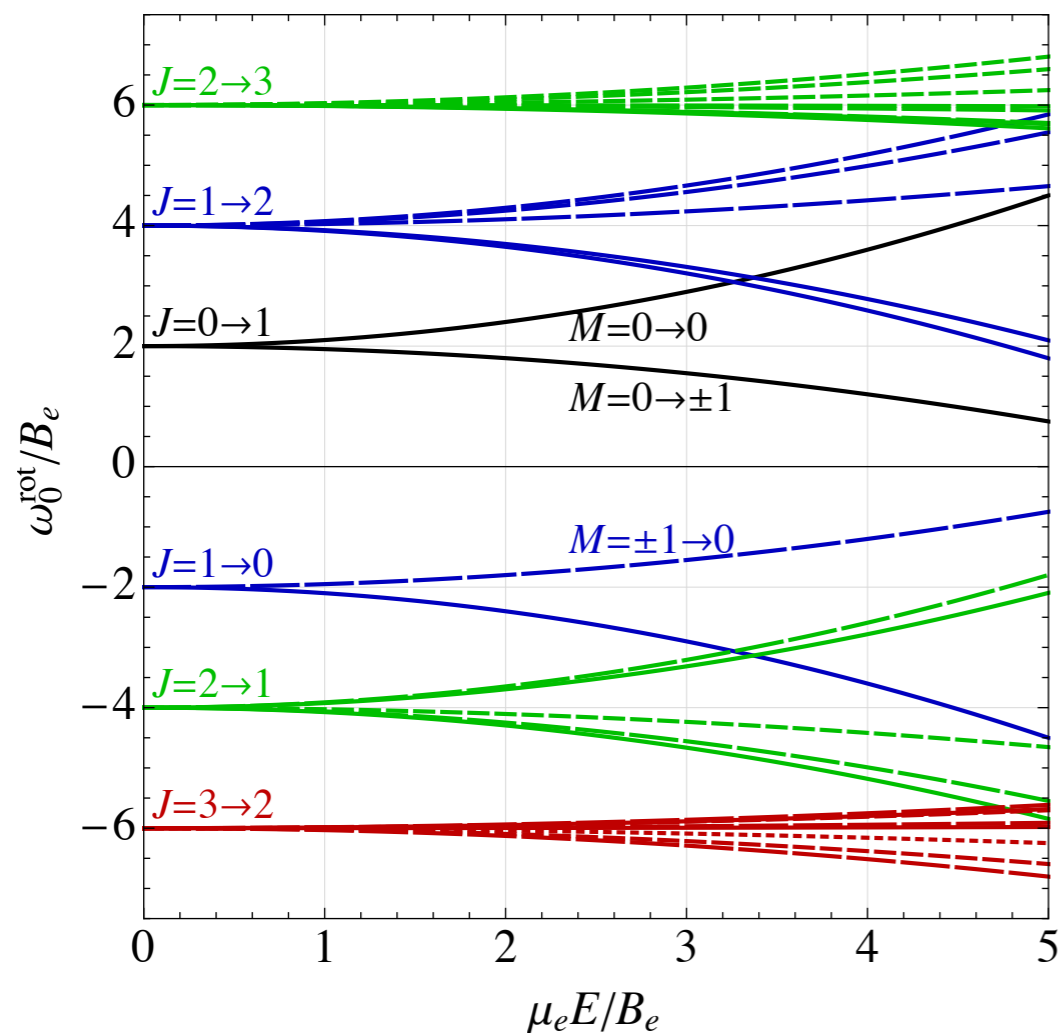
$$\bar{r} \simeq 1 + \frac{8\pi n}{m^4 R_z} \simeq 1 + \frac{5 \times 10^6}{m R_z} \left(\frac{1 \text{ eV}}{m} \right)^3 \left(\frac{n}{n_0} \right)$$



Key considerations

frequency coverage

- collisional broadening
- molecular species/isotope
- Zeeman tuning
- Stark tuning:



radiative efficiency

- Bulk: $\gamma_0 + \sum_i \gamma_i \gtrsim \gamma_{\text{quench}}$
- Stack: $(\bar{r} - 1)\gamma_0 \sim \gamma_{\text{col}}$

