

Theories of Neutrino Mass

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INFLATION 20, KITP, January 2020



The Standard Model Flavor Puzzle:

the origin of the fermion masses and mixing angles in the SM



traditional SM definition: 3 generations, massless neutrinos

13 out of 19 SM parameters

w/neutrino masses: at least 7 more (9 more if Majorana)



A long-standing question that predates and runs in parallel with the theoretical development of the SM itself (and its possible extensions). It has resulted in its own rich history, and multitude of ideas.



Key elements:✓V-A interactionsMarshak, Sudarshan (1957), Feynman, Gell-Mann (1958),...✓Cabibbo angle, weak interactions universalityCabibbo (1963),...✓GIM mechanism and FCNC suppressionGlashow, Iliopoulos, Maiani (1970),...✓quark sector CP violationKobayashi, Maskawa (1973),...

Much experimental input! In recent decades:

	\checkmark	discovery of top quark	CDF, D0 (1995)
-	\checkmark	precise determination of quark mixing matrix entries	many exps/analyses
	\checkmark	neutrino oscillations discovery and measurements	SuperK (1998) + many others
	\checkmark	Higgs discovery and coupling measurements to fermine	ons ATLAS, CMS (2012),

Discovery of neutrino oscillations (nonzero neutrino masses)

+ subsequent detailed measurements of lepton sector

a paradigm shift



(image credit: Sandbox Studio via symmetry magazine)

Goal here: overview of current status (not fully comprehensive) challenges and future outlook



(image credit: C. Wiens)

Some highlights:

2002: solar ν_e disappearance (SK) 2002: solar ν_e appear as ν_{μ}, ν_{τ} (SNO) 2004: reactor $\overline{\nu}_e$ oscillations (KamLAND) 2004: accelerator ν_{μ} disappearance (K2K) 2006: accelerator ν_{μ} disappearance (MINOS)

1998: atmospheric ν_{μ} disappearance (SK)

2011: accel. ν_{μ} appear as ν_{e} (T2K,MINOS) 2012: reactor $\overline{\nu}_e$ disappear (Daya Bay, reactor angle measured! **RENO,...)** CP violation hint? (T2K) 2014: 2015: normal hierarchy hint? (SK, T2K, NOvA) 2016: non-maximal atm hint? (NOvA) 2018: CP cons disfavored at 2σ (T2K) 2019: improved direct mass limit (KATRIN)

The emergent picture...

a (seemingly) robust 3-neutrino mixing scheme

$$\Delta m_{ij}^2 \equiv m_i^2 - m_j^2$$



✓ overall mass scale: limits from direct searches, cosmology

$$\sum m_{\nu} < 0.7 \text{ eV}$$

 $m_{\bar{\nu}_e} < 1.1 \text{ eV}$ KATRIN (2019)

Lepton mixing:

Global Fits:

"standard" PDG parametrization

Forero et al., '17

Capozzi et al.,'18

Gonzalez-Garcia et al., (www.nu-fit.org)

NuEIT 4 1 (2010)

$$\begin{split} |U|_{3\sigma}^{\text{w/o SK-atm}} &= \begin{pmatrix} 0.797 \rightarrow 0.842 & 0.518 \rightarrow 0.585 & 0.143 \rightarrow 0.156 \\ 0.244 \rightarrow 0.496 & 0.467 \rightarrow 0.678 & 0.646 \rightarrow 0.772 \\ 0.287 \rightarrow 0.525 & 0.488 \rightarrow 0.693 & 0.618 \rightarrow 0.749 \end{pmatrix} \\ |U|_{3\sigma}^{\text{with SK-atm}} &= \begin{pmatrix} 0.797 \rightarrow 0.842 & 0.518 \rightarrow 0.585 & 0.143 \rightarrow 0.156 \\ 0.243 \rightarrow 0.490 & 0.473 \rightarrow 0.674 & 0.651 \rightarrow 0.772 \\ 0.295 \rightarrow 0.525 & 0.493 \rightarrow 0.688 & 0.618 \rightarrow 0.744 \end{pmatrix} \end{split}$$

(assumptions: 3 active neutrinos only, unitarity)

here: "MNSP" more often: "PMNS"

Note

More details (NuFit 2019)

NuFIT 4.1 (2019)

		Normal Ordering (best fit)		Inverted Ordering $(\Delta \chi^2 = 6.2)$		
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range	
without SK atmospheric data	$\sin^2 \theta_{12}$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	$0.310\substack{+0.013\\-0.012}$	$0.275 \rightarrow 0.350$	
	$\theta_{12}/^{\circ}$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	
	$\sin^2 \theta_{23}$	$0.558\substack{+0.020\\-0.033}$	$0.427 \rightarrow 0.609$	$0.563\substack{+0.019\\-0.026}$	$0.430 \rightarrow 0.612$	
	$\theta_{23}/^{\circ}$	$48.3^{+1.1}_{-1.9}$	$40.8 \rightarrow 51.3$	$48.6^{+1.1}_{-1.5}$	$41.0 \rightarrow 51.5$	
	$\sin^2 \theta_{13}$	$0.02241^{+0.00066}_{-0.00065}$	$0.02046 \to 0.02440$	$0.02261\substack{+0.00067\\-0.00064}$	$0.02066 \rightarrow 0.02461$	
	$\theta_{13}/^{\circ}$	$8.61^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.99$	$8.65_{-0.12}^{+0.13}$	$8.26 \rightarrow 9.02$	
	$\delta_{ m CP}/^{\circ}$	222^{+38}_{-28}	$141 \to 370$	285^{+24}_{-26}	$205 \to 354$	
	$\frac{\Delta m^2_{21}}{10^{-5}~{\rm eV}^2}$	$7.39\substack{+0.21 \\ -0.20}$	$6.79 \rightarrow 8.01$	$7.39\substack{+0.21\\-0.20}$	$6.79 \rightarrow 8.01$	
	$\frac{\Delta m^2_{3\ell}}{10^{-3}~{\rm eV}^2}$	$+2.523\substack{+0.032\\-0.030}$	$+2.432 \rightarrow +2.618$	$-2.509\substack{+0.032\\-0.030}$	$-2.603 \rightarrow -2.416$	
		Normal Ordering (best fit)		Inverted Ordering $(\Delta \chi^2 = 10.4)$		
		Normal Ord	lering (best fit)	Inverted Orde	ring $(\Delta \chi^2 = 10.4)$	
		Normal Orc bfp $\pm 1\sigma$	dering (best fit) 3σ range	Inverted Orde bfp $\pm 1\sigma$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$	
	$\sin^2 \theta_{12}$	Normal Ord $bfp \pm 1\sigma$ $0.310^{+0.013}_{-0.012}$	$\frac{\text{dering (best fit)}}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$	
lata	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$	$\frac{\text{dering (best fit)}}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$	
eric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\sin^2 \theta_{23}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$	$\frac{\text{dering (best fit)}}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.433 \rightarrow 0.609$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.436 \rightarrow 0.610$	
spheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$ $48.6^{+1.0}_{-1.4}$	$\begin{array}{c} \text{lering (best fit)} \\ \hline 3\sigma \text{ range} \\ \hline 0.275 \rightarrow 0.350 \\ 31.61 \rightarrow 36.27 \\ \hline 0.433 \rightarrow 0.609 \\ 41.1 \rightarrow 51.3 \end{array}$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$ $48.8^{+1.0}_{-1.2}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \to 0.350$ $31.61 \to 36.27$ $0.436 \to 0.610$ $41.4 \to 51.3$	
atmospheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\sin^2 \theta_{13}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$ $48.6^{+1.0}_{-1.4}$ $0.02237^{+0.00066}_{-0.00065}$	dering (best fit) 3σ range $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.433 \rightarrow 0.609$ $41.1 \rightarrow 51.3$ $0.02044 \rightarrow 0.02435$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$ $48.8^{+1.0}_{-1.2}$ $0.02259^{+0.00065}_{-0.00065}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.436 \rightarrow 0.610$ $41.4 \rightarrow 51.3$ $0.02064 \rightarrow 0.02457$	
SK atmospheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/^{\circ}}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$ $48.6^{+1.0}_{-1.4}$ $0.02237^{+0.00066}_{-0.00065}$ $8.60^{+0.13}_{-0.13}$	lering (best fit) 3σ range $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.433 \rightarrow 0.609$ $41.1 \rightarrow 51.3$ $0.02044 \rightarrow 0.02435$ $8.22 \rightarrow 8.98$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$ $48.8^{+1.0}_{-1.2}$ $0.02259^{+0.00065}_{-0.00065}$ $8.64^{+0.12}_{-0.13}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.436 \rightarrow 0.610$ $41.4 \rightarrow 51.3$ $0.02064 \rightarrow 0.02457$ $8.26 \rightarrow 9.02$	
with SK atmospheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/^{\circ}}$ $\delta_{\rm CP}/^{\circ}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$ $48.6^{+1.0}_{-1.4}$ $0.02237^{+0.00066}_{-0.00065}$ $8.60^{+0.13}_{-0.13}$ 221^{+39}_{-28}	$\frac{\text{lering (best fit)}}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.433 \rightarrow 0.609$ $41.1 \rightarrow 51.3$ $0.02044 \rightarrow 0.02435$ $8.22 \rightarrow 8.98$ $144 \rightarrow 357$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$ $48.8^{+1.0}_{-1.2}$ $0.02259^{+0.00065}_{-0.00065}$ $8.64^{+0.12}_{-0.13}$ 282^{+23}_{-25}	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.436 \rightarrow 0.610$ $41.4 \rightarrow 51.3$ $0.02064 \rightarrow 0.02457$ $8.26 \rightarrow 9.02$ $205 \rightarrow 348$	
with SK atmospheric data	$\frac{\sin^2 \theta_{12}}{\theta_{12}/^{\circ}}$ $\frac{\sin^2 \theta_{23}}{\theta_{23}/^{\circ}}$ $\frac{\sin^2 \theta_{13}}{\theta_{13}/^{\circ}}$ $\frac{\delta_{\rm CP}/^{\circ}}{\frac{\Delta m_{21}^2}{10^{-5} \ {\rm eV}^2}}$	Normal Ord bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.76}$ $0.563^{+0.018}_{-0.024}$ $48.6^{+1.0}_{-1.4}$ $0.02237^{+0.00066}_{-0.00065}$ $8.60^{+0.13}_{-0.13}$ 221^{+39}_{-28} $7.39^{+0.21}_{-0.20}$	$\begin{array}{l} \label{eq:advector} \frac{\text{lering (best fit)}}{3\sigma \ \text{range}} \\ \hline & 3\sigma \ \text{range} \\ \hline & 0.275 \rightarrow 0.350 \\ 31.61 \rightarrow 36.27 \\ \hline & 0.433 \rightarrow 0.609 \\ 41.1 \rightarrow 51.3 \\ \hline & 0.02044 \rightarrow 0.02435 \\ 8.22 \rightarrow 8.98 \\ \hline & 144 \rightarrow 357 \\ \hline & 6.79 \rightarrow 8.01 \end{array}$	Inverted Orde bfp $\pm 1\sigma$ $0.310^{+0.013}_{-0.012}$ $33.82^{+0.78}_{-0.75}$ $0.565^{+0.017}_{-0.022}$ $48.8^{+1.0}_{-1.2}$ $0.02259^{+0.00065}_{-0.00065}$ $8.64^{+0.12}_{-0.13}$ 282^{+23}_{-25} $7.39^{+0.21}_{-0.20}$	$\frac{\text{ring } (\Delta \chi^2 = 10.4)}{3\sigma \text{ range}}$ $0.275 \rightarrow 0.350$ $31.61 \rightarrow 36.27$ $0.436 \rightarrow 0.610$ $41.4 \rightarrow 51.3$ $0.02064 \rightarrow 0.02457$ $8.26 \rightarrow 9.02$ $205 \rightarrow 348$ $6.79 \rightarrow 8.01$	

Caveat: sterile neutrino(s)?

Anomalies:

1995:	$\overline{\nu}_e$	appearance		(LSND)
2007:	$\overline{\nu}_e$	appearance	(Mi	niBooNE)
2012:	ν_e	appearance	(Miı	niBooNE)
1995:	ν_e	disappearanc	е	(Gallium)
2011:	ν_e	disappearanc	e	(Reactor)

[well-documented tension between appearance and disappearance data]



See: Huber's IPA 2017 talk for "scorecard" Maltoni's talk at Neutrino 2018

Restrict focus here to 3 active light families only

Compare to the quark sector:

(image credit: D0 Single Top group)





Quark mixing:

$$\begin{array}{ccc} u_{i} & & \mathcal{U}_{\mathrm{CKM}} \rangle_{ij} & & \mathcal{U}_{\mathrm{CKM}} = \mathcal{R}_{1}(\theta_{23}^{\mathrm{CKM}}) \mathcal{R}_{2}(\theta_{13}^{\mathrm{CKM}}, \delta_{\mathrm{CKM}}) \mathcal{R}_{3}(\theta_{12}^{\mathrm{CKM}}) & & \text{``standard'' PDG parametrization} \\ & & \mathcal{M}_{j} & & \mathcal{M}$$

Wolfenstein parametrization: $\lambda = \sin \theta_c$ (Cabibbo angle)

 $s_{12} = \lambda \quad s_{23} = A\lambda^2 \quad s_{13} = A\lambda^3(\rho + i\eta) = \frac{A\lambda^3(\bar{\rho} + i\bar{\eta})\sqrt{1 - A^2\lambda^4}}{\sqrt{1 - \lambda^2}(1 - A^2\lambda^4(\bar{\rho} + i\bar{\eta}))}$ (Cabibbo expansion) $\lambda \simeq 0.225 \quad A \simeq 0.83 \quad \bar{\rho} \simeq 0.1 \quad \bar{\eta} \simeq 0.35$ PDG (2018)

A plethora of interesting results to (try to) explain!

- Quark and lepton sectors look very different!
 - ✓ Quarks: hierarchical masses, small mixings, O(1) CP violation
 - Leptons: hierarchical charged lepton masses suppression of overall neutrino mass scale hierarchy apparently "milder" for neutrino masses two large mixing angles (or more**)





implications for quark-lepton unification and other BSM physics

Mass Generation

• Quarks and charged leptons:

Mass generation must proceed via Yukawa interactions w/Higgs:

$$\psi = \begin{pmatrix} \psi_L \\ \psi_R \end{pmatrix} \qquad Q_{Li} = \begin{pmatrix} u_{Li} \\ d_{Li} \end{pmatrix}, \ L_{Li} = \begin{pmatrix} \nu_{Li} \\ e_{Li} \end{pmatrix} \qquad u_{Ri}, \ d_{Ri}, \ e_{Ri}$$
$$\boxed{Y_{ij}H \cdot \bar{\psi}_{Li}\psi_{Rj}}$$

Dirac masses upon electroweak symmetry breaking (EWSB)

$$\mathcal{M}_{ij}^{\text{Dirac}} \equiv Y_{ij} \langle H \rangle$$

Yukawas arbitrary in SM: diagonalize

$$\mathcal{U}_{fL}^{\dagger}\mathcal{M}_{f}^{\mathrm{Dirac}}\mathcal{U}_{fR} = \mathcal{M}_{f}^{\mathrm{diag}}$$

$$\mathcal{U}_{\rm CKM} = \mathcal{U}_{uL}^{\dagger} \mathcal{U}_{dL}$$

Masses tied to EW scale: suppressions required (all but top)

Dirac



FB "The Same Oddly Asymmetric Picture of Paul Dirac Each Day"

• Neutrino Masses

main question: are neutrinos Dirac or Majorana?



Critically important question, to be settled by experiment! Many options for neutrino mass suppression in each case. Consider each in turn (Majorana first) Majorana neutrinos

lepton number violating

 $\Delta L = 2$

SM at NR level: Weinberg operator

Majorana neutrino masses upon EWSB

 $\frac{\lambda_{ij}}{\Lambda}L_iHL_jH$





Type I seesaw

✓ introduce right-handed neutrinos



(image credit: T. Ohlsson et al., Nat. Comm.)

$$Y_{ij}L_i\nu_{Rj}H + M_{R\,ij}\nu_{Ri}\nu_{Rj}^c \qquad \mathcal{M}_{\nu} \sim \langle H \rangle^2 Y M_R^{-1}Y^T$$

Advantages: economical, connection to grand unification, leptogenesis** Disadvantages: testability without model assumptions

Type II seesaw

Magg, Wetterich '80, Lazarides, Shafi, Wetterich '81, Mohapatra, Senjanovic '81, Cheng, Li '80,...

✓ introduce triplet Higgs scalar



(image credit: T. Ohlsson et al., Nat. Comm.)

$$\Delta \sim (\mathbf{3}, 2) \\ (SU(2)_L, U(1)_Y)$$
$$\Delta = \begin{pmatrix} \delta^+ / \sqrt{2} & \delta^{++} \\ \delta^0 & -\delta^+ / \sqrt{2} \end{pmatrix} \\ (Y_\Delta)_{ij} L_i L_j \Delta + \mu_\Delta H H \Delta$$
$$\mathcal{M}_{\nu} \sim \langle H \rangle^2 Y_\Delta \mu_\Delta / M_\Delta^2$$

can have clean LHC signatures of lepton # violation via decays of H^+, H^{++} if $M_{\Delta} \leq O(\text{TeV})$ Fileviez Perez et al. '08, Gavela et al. '09,... (also LFV) Advantages: testability (charged Higgs states probed at LHC)

Disadvantages: not as economical/minimal as Type I

✓ introduce electroweak triplet fermions

Type III seesaw



 $\Sigma \sim (\mathbf{3}, 0)$ $(SU(2)_L, U(1)_Y)$ $\Sigma_i = \left(\Sigma_i^0, \Sigma_i^{\pm}\right)$ $(Y_{\Sigma})_{ij} L_i \Sigma_j H + (M_{\Sigma})_{ij} \Sigma_i \Sigma_j$ $\mathcal{M}_{\nu} \sim \langle H \rangle^2 Y_{\Sigma} M_{\Sigma}^{-1} Y_{\Sigma}^T$

can have clean LHC signatures via mixing w/charged leptons $M_\Sigma \sim O({ m TeV})$ Franchesini, Hambye, Strumia '08,...

also highly predictive pattern of LFV signals

Abada et al. '07, Gavela et al. '09,...

Advantages: testability (new charged states probed at LHC, LFV) Disadvantages: not as economical/minimal as Type I



(image credit: T. Ohlsson et al., Nat. Comm.)

Radiative possibilities:

complete Weinberg operator via loops

Canonical example: "scotogenic" model

$$\mathcal{M}_{\nu} \sim \lambda \frac{\langle H \rangle^2}{16\pi^2} Y M_R^{-1} Y^T$$

"radiative Type I seesaw"

Ma '06

analogous construction with fermion triplet:

"radiative Type III seesaw" Ma, Suematsu '09

can also construct a "radiative Type II seesaw"

Fraser, Kownacki, Ma, Popov '15

Required: new symmetry (usually discrete) to forbid tree level term, new states

Zee '86, Babu '88, Ma '98,...

Can also consider alternatives to Weinberg operator

Many other ΔL	= 2 NR operators	(odd mas	ss dimension d>5)
Classification			
d=7	d=9		Babu and Leung '01 de Gouvea and Jenkins '07
$LLLe^{c}H$	$LLLe^{c}Le^{c}$	(Zee-Babu)	Bonnet et al. '12
$LLQd^{c}H$	$LLQd^{c}Qd^{c}$		
$LL\overline{Q}\overline{u}^{c}H$	+ many othe	ers	•••
$L\overline{e}^{c}\overline{u}^{c}d^{c}H$			

Generic advantage of radiative models:

New physics scale can be accessible at LHC (subject to LFV bounds)

Many explicit realizations!

See e.g. excellent review: Cai, Herrero-Garcia, Schmidt, Vicente, Volkas '17 Also potential connections to other, possibly accessible, new physics:



possible connections to flavor physics anomalies

Päs, Schumacher '15, Deppisch et al. '16,...

One way leptoquarks can manifest themselves:





Angel, Cai, Rodd, Schmidt, Volkas '13 Cai, Gargalones, Schmidt, Volkas '17 Many other ideas for Majorana masses

more complicated seesaws

e.g. double/inverse seesaw: $3\nu_R$ and 3 new singlet fermions

lower scales

Mohapatra, Valle '86

SUSY with R-parity violation /

. . .

Aulakh, Mohapatra '82, Hall, Suzuki '84,... Borzumati et al. '96, Grossman et al. '03

Warped extra dimensions

see e.g. Csaki, et al. '08,...



Dirac neutrinos

• Analogous to charged fermions, but much stronger suppression





Must forbid both types of "bare" mass terms Less intuitive, but mechanisms exist.

Example: radiative mass generation

Cheng, Li '78, Mohapatra '87, '88

Balakrishna, Mohaptra '89, Rajpoot '01

e.g. discrete symmetry: ν_R nontrivial one-loop: 2 topologies

Ma, Popov '16 Wang et al. '16, '17 Review: Cai et al. '17

...

variety of possibilities for new states

(many studies in GUT contexts)

other options: higher loops, loop-induced vev,...

✓ extended gauge sectors

non-singlet ν_R

higher-dimensional operators: e.g. U(1)' symmetry

forbids "bare" terms, simplest seesaws

see e.g. Langacker '11 for review

✓ SUSY breaking

. . .

symm+holomorphy forbids superpotential neutrino masses via Kahler potential terms

Arkani-Hamed et al. '00 Borzumati et al. '00,'01 Demir, LE, Langacker '08

✓ string constructions

exponentially suppressed interactions from stringy instanton effects...

Cvetic et al. '08,... Langacker review, '11

. . .

Neutrino sector much richer than charged fermions

but trade-off between simplicity and testability

Mass hierarchies and lepton mixing



(image credit: S. King)

(Dirac or Majorana neutrinos)

$$\mathcal{U}_{\nu}^{T}\mathcal{M}_{\nu}^{\mathrm{Maj}}\mathcal{U}_{\nu}=\mathcal{M}_{\nu}^{\mathrm{diag}}$$

$$\mathcal{U}_{fL}^{\dagger}\mathcal{M}_{f}^{\mathrm{Dirac}}\mathcal{U}_{fR} = \mathcal{M}_{f}^{\mathrm{diag}}$$

$$\mathcal{U}_{\mathrm{MNSP}} = \mathcal{U}_{eL}^{\dagger} \mathcal{U}_{\nu L}$$

$$\mathcal{U}_{\mathrm{MNSP}} = \mathcal{R}_1(\theta_{23})\mathcal{R}_2(\theta_{13},\delta)\mathcal{R}_3(\theta_{12})\mathcal{P}$$

Lepton sector properties:

charged lepton masses: strongly hierarchical hierarchy apparently "milder" for neutrino masses two large lepton mixing angles (or more**)

(can work within any framework for mass suppression. vast majority: Type I seesaw)

Lepton mixing structure:

Two large mixing angles: $\theta_{23}, \ \theta_{12}$

CP violation

Dirac phase: important goal of DUNE, HyperK Majorana phases: challenging

A basic question: is θ_{13} "large" or "small"?



Anarchy



(character: Watterson)

 \mathcal{U}_{ν} from a random draw of unbiased distribution of 3x3 unitary matrices

statistical tests: lower bound on $|\mathcal{U}_{e3}|^2$

basis independence:

distribution invariant upon unitary transformations

flat in Haar measure

Haba, Murayama '00

Post-reactor angle measurement: renewed focus

Some recent highlights: RG analysis

model-building + quark sector

de Gouvea, Murayama '12 Altarelli et al. '12 Bai and Torroba '12 ...

Brdar, Konig, Kopp '15 Babu et al. '16,... Fortin et al. '17

Note: anarchy hypothesis alone does not provide information on Δm^2



canonical three-family predictive examples: (ruled out: top quark mass + CKM)

- ✓ Fritzsch texture Fritzsch '78
- Georgi-Jarlskog texture Georgi, Jarlskog '79
- Y "Yukawa quilt"
 Ramond, Roberts, Ross '93

What persists: more intricate structures, lessons for high-scale embedding

GJ mass relations (high scale) $m_b = m_{\tau}$ $m_{\mu} = 3m_s$ $m_d = 3m_e$ (hanowitz et al. '77 Buras et al. '78 (factors of 3: # of colors (RG effects)) Achieve special structures via symmetries:

• Postulate family symmetry G_f

spontaneously broken by "flavon" fields $\{\varphi_a\}$

$$\epsilon = \left(rac{\langle \varphi
angle}{\Lambda}
ight) \qquad \Lambda > \langle \varphi
angle > \langle H
angle \qquad \epsilon \ll 1 \qquad$$
 Froggatt, Nielsen '79

Heavy sector >> TeV (avoid too-large FCNC)

natural identification:
$$\epsilon \sim O(\lambda)$$
 $\lambda = \sin \theta_c$

Cabibbo angle (or some power) as a flavor expansion parameter

Unique theoretical starting point: $U_{CKM} \sim 1 + O(\lambda)$

• Possibilities for G_f :

canonical example: U(1) family symmetries

Froggatt, Nielsen '79

Have guidance from enhanced symmetry for vanishing Yukawas:

✓ SM (no neutrino masses): $U(3)^5$

 $U(3)_Q \otimes U(3)_{u^c} \otimes U(3)_{d^c} \otimes U(3)_L \otimes U(3)_{e^c}$

 G_f subgroup of $U(3)^5$

Key input: O(1) top quark Yukawa coupling $U(3)_Q \otimes U(3)_{u^c} \xrightarrow{m_t} U(2)_Q \otimes U(2)_{u^c}$

Many examples! continuous Abelian, non-Abelian, discrete non-Abelian

Structure



Now, lepton sector:

Can still introduce family symmetry

 G_f spontaneously broken at scale M $\epsilon \sim \langle \varphi \rangle / M$

But quite different from the quark sector!

diagonalize \mathcal{M}_{ν} : 1 small, 2 large mixing angles (diagonal \mathcal{M}_{e} basis)

Arguably the most challenging* leading order pattern: (* for 3 families)





✓ No unique theoretical starting point for "Cabbibo-like" expansion

$$\mathcal{U}_{\text{MNSP}} \sim \mathcal{W} + O(\lambda') \qquad \lambda' \ll 1$$

"bare" mixing angles (diagonal charged lepton basis)

$$(\rho \nu \quad \rho \nu \quad \rho \nu)$$

 $(\theta_{12}^{\nu}, \theta_{23}^{\nu}, \theta_{13}^{\nu})$

First stage: symmetry breaking to generate nontrivial ${\cal W}$ different unbroken subgroups for neutrinos, charged leptons

large mixing angles

Next stage: corrections as expansion in λ'

"Bare" mixing angles generically shift due to $O(\lambda')$ corrections

✓ *A priori,* expansions in quark and lepton sectors unrelated unification paradigm (broad sense): $\lambda' = \lambda$

deas of quark-lepton complementarity and "Cabibbo haze"
Raidal '04, Minakata+Smirnov '04,...

$$\theta_{23} = \theta_{12} + \theta_c$$
 ("haze" terminology from
Datta, L.E., Ramond '05)

not an unreasonable approach given the data

$$\theta_{13} \sim O(\lambda)$$

pre-measurement, idea that θ_{13} might be a Cabibbo effect:

 $\theta_{13}^{
u}=0 \qquad heta_{13}=\lambda/\sqrt{2}$ Vissiani '98, '01

Ramond '04

Possible theoretical starting points:



All can be obtained via SSB of discrete non-Abelian family symmetries

Family symmetry models

usual choices: SU(3), SO(3) subgroups: \checkmark

 $\mathcal{A}_4 \quad \mathcal{S}_4 \quad \mathcal{A}_5 \quad \Delta(3n^2) \quad \Delta(6n^2) \quad \mathcal{D}_n \quad \mathcal{T}' \quad \mathcal{I}' \quad \dots \quad \text{``Platonic solid'' groups''}$

+ double covers

example (Majorana):



Flavons:

$$\phi^l, \phi^\nu$$

Residual symmetries:

 $T\langle \phi^l \rangle \approx \langle \phi^l \rangle$ $S, U\langle \phi^{\nu} \rangle \approx \langle \phi^{\nu} \rangle$

(or broken further, e.g. only S or U unbroken)

also needed: "driving fields" (singlets)

Many papers and authors! Some authors (not comprehensive):

Babu, Chen, Ding, L.E., Feruglio, Grimus, Hagedorn, King, Lam, Luhn, Ma, Merle, Ohlsson, Rodejohann, Stuart,...

Residual Symmetries

model-independent approach:

Lam '08, '09, Grimus et al. '09, Ge et al. '11, Toroop et al. '11, He et al. '12, Hernandez et al. '12,'13 Holthausen et al. '12, King et al. '13, Hagedorn et al. '14, Lavoura, Ludl '14, Fonseca, Grimus '14

. . .

determine rows and columns in \mathcal{U}_{MNSP} as pure numbers, independent of masses, depending on preserved subgroups of finite group G_f $T^{\dagger}\mathcal{M}_{e}\mathcal{M}_{e}^{\dagger}T = \mathcal{M}_{e}\mathcal{M}_{e}^{\dagger} \qquad S^{\dagger}\mathcal{M}_{\nu}S = \mathcal{M}_{\nu}$ $G_f \to G_e, \ T \in G_e \qquad G_f \to G_\nu, \ S \in G_\nu$ $\mathcal{U}_{eL}^{\dagger} T \mathcal{U}_{eL} = T^{\text{diag}} \qquad \mathcal{U}_{\nu}^{\dagger} S \mathcal{U}_{\nu} = S^{\text{diag}}$ Majorana: $G_{\nu} \supseteq Z_2 \times Z_2$ (Klein group) systematic classification of possible mixing matrices Fonseca, Grimus '14

Very different from texture zeros (mixing angles as ratios of masses)!

Corrections



✓	RG effects	more significant for IO,	Antusch et al. '03
		heavy neutrino masses	
		(can be significant for sum rule analysis)	Gehrlein et al. '16

Example: tri-bimaximal mixing (TBM/HPS)

(Type I seesaw)

"minimal" flavor group (contains *S*,*T*,*U* generators)

Lam '11

✓ Residual symmetries: $Z_3 \sim T$ $Z_2 \times Z_2 \sim S, U, SU$

Can further break down Klein symmetry:

1 column only of HPS matrix preserved

see e.g. King '17 for review



Notable (often-found) feature: phase input needed

phase required in $U_{\nu} \sim U^{(\rm HPS)}$ for consistency with mixing angle data numerical example: $\delta \simeq \pm 1.3\pi, J \simeq \mp 0.03$



many recent papers! see King '17 for review

Connection (or not) to quark sector

not particularly straightforward (subjective!)

many examples, authors recent notable case: Hagedorn et al. '18**

✓ discrete non-Abelian symmetry models:

quarks can require alternate embeddings

e.g. often $L_i \sim \mathbf{3}$ but $Q_i \sim \mathbf{2} \oplus \mathbf{1}$

want groups with both doublets and triplets

(larger groups, double covers $\mathcal{T}', \mathcal{I}'$)

✓ work explicitly in SUSY GUT framework (Type I seesaw)

 $SO(10) \text{ with family symmetry:} Dermisek et al. '05, '06 Poh, Raby '15 D_3 \times U(1) \times Z_3 \times Z_3 2 \oplus 1$ 14 fermion sector inputs 6 fermion sector predictions
Pati-Salam version (lighter superpartners) Poh, Raby, Wang '17 (17 fermion sector inputs)



$$Q_i, L_i \sim \mathbf{3}$$



requires intricate sector of flavons+ driving fields (characteristic)

• String constructions:

- ✓ variety of possibilities for mass suppression higher-dimensional operators (field theoretic) geometric suppression (braneworlds) $Y \sim e^{-A}$ worldsheet instantons (nonperturbative)
- ✓ Yukawa unification often not retained even in GUT scenarios
- ✓ not necessarily just minimal Type I seesaw

 ν_R candidates often not pure gauge singlets

explorations in heterotic orbifolds Giedt et al. '05, Buchmuller et al.'07...

"Mixed" scenarios (e.g. seesaw + R-parity violation)

e.g. G2 models F theory GUTs Acharya et al. '08... Beasley, Heckman, Vafa '08,...

Concluding remarks

• Neutrino masses and lepton mixing angles: paradigm shift

many ways to suppress overall neutrino mass scale lepton mixings: anarchical or structural masses/mass ordering "secondary" not easy to unify/wed with quark sector needed physics often at high scales

A great amount of theoretical effort, but still seeking

compelling, complete, testable theories

More observational handles needed!

New ideas welcomed!