KITP, Santa Barbara, August 17, 2011

LFV, DM and LHC: how's SUSY health these days?

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Progress on SUSY



Within the constrained SSM models we are crossing the border of excluding gluinos and squarks up to 1TeV and beyond. The air is getting thin for constrained SUSY. More conclusive results after summer.

SUSY in 0-lepton channel



MSUGRA/CMSSM: tanB=10, A₀=0, μ>0 Equal mass case: m_q=m_g > 980 GeV

D. CHARLTON EPS-HEP 2011

Simplified model with two q generations, m($\tilde{\chi}_{_{1}}^{\ 0}$)~0 m_s>800 GeV m_q>850 GeV Equal mass case: m_g=m_q>1.075 TeV

SUSY in 1-lepton channel D. CHARLTON EPS-HEP 2011



	ATLAS Searches* - 95% CL Lower Limits (EPS-HEP 2011)				
MSUGRA/CMSSM : 0-lep + E _{Loise}	CALCE (5 ⁻¹ COTS (and industra)	swew ĝ = ĝ mass	I		
Simplified model (light $\tilde{\chi}_i^*$) : 0-lep + E_{trains}	canne is ^a party premiusly	tataraa 🧃 = ĝimass	ATLAS		
Simplified model (light $\tilde{\chi}_i^{*}$) : 0-lep + E_{total}	CALON IN ¹ (2017) (protiminary)	ទទេ tey 🛛 ថ្មី mass	Preliminary		
Simplified model (light $\overline{\chi}^*$) : 0-lep + E_{trains}	Coll. 66 In ¹¹ (2017) (protinizary)	alease ğ mess			
Simplified model : 0-lep + b-jets + E _{Loise}	LYCERS TO ¹ (2017) (AT LASSCONF 2017-2033)	220 See] j mass (lor $m(b)$ < 600 GeV)	$I dl = (0.031 - 1.21) \text{ (b}^{-1}$		
Phono-MSSM (light $\overline{\chi}_{i}^{0}$) : 2-lap SS + E_{Laiss}	an manyakang panang pananan na manang	an av 🦷 🏹 maas	Jean (sees the type		
Phene-MSSM (light $\tilde{\chi}_{i}^{0}$) : 2-lep OS _{pr} + E _{train}	L+35 pt/ ² (2910) [#99 x 1100.3248]	ssi cevi 🛱 masa	$\sqrt{s} = 7 \text{ TeV}$		
GMSB (GGM) + Simpl. model : 77 + E	E.r.(85, pitc ⁴⁾ (28-40) [p-01 + 41 07, 054 4]	an an Ö mass			
GMSB : stable τ	2-07 p2 ⁴ (2010)[a05x1103.4410] 108 9ay 🕄 🕅	355			
Stable massive particles : R-hadrons	2.434 pb ⁻¹ (2010) (a/04 x 1100, 1944)	IN KeV 👌 mass			
Stable messive particles : R-hadrons	2.×04 µ0 ⁻⁰ (2010) (a-01 × 1100.1014)	asen di mass			
Stable massive particles : R-hadrons	2.208 pix ² (2010) (a:00x 1100, 1844)	Die Gew I Mittels S			
RPV (λ_{11} =0.01, λ_{112} =0.01) : high-mass e.u	L-0.17-6 ⁻¹ (2015) (collisione))	AND CALV V _e TRANS			
Large ED (ADD) : monojet	LET BE B ¹ (2015) [ATLAS-CONF-3011-086]	aatav M _d	, (ð=2)		
UED : $\gamma \gamma + E_{\text{trains}}$	Ex105 ato ¹⁰ (2010) (arXiv:1107.3544)	encev Compact scale I/R			
RS with $h/M_{\rm Pl} = 0.1 \pm m_{co}$	C430 (#0 ⁴ (2010) (#11.43 (GOMP 2011-044)	set see Greviton meso			
RS with $k/M_{\mu_1} = 0.1 : m_{\mu_1 \mu_2}$	L+1.64-1.31 Bs ¹ (2013) [assiminary]	isstavi Graniton mass			
RS with top couplings $g_i = 1.0, g_i = 4.0 : m_i$	E-800 µb ⁻¹ (2010) [ATLA3-GONF-3011-087]	eaveer IKK gluon mass			
Quantum black hole (QBH) : m _{aija} , F(χ)	LASS ato ¹² (2010) (arXiv:1100.3894)	3 61 Tay	М ₀ (б=С)		
OBH : High-mass σ_{i+x}	er an fer _d benoù feurren roeen e er rari	LINDA MO			
ADD BH ($M_0/M_0=3$) : multijot $\Sigma p_{+}, N_{jete}$	LYES (\$1) (ATLAS CONFERENCE)	ranne M _☉ (5–6)			
ADD BH $(M_0/M_0=3)$: SS dimuon $N_{ob,part}$	2.401 pt; ⁴ (2040) [ATLAS (CONT 3.511-005]	1.20 TW N (8=6)			
qqqq contact interaction : $F_{\chi}(m_{ m eqc})$	Cx35 pc1 ² (2010) (a56 x 1100,3664 (857 calent imit))		LT HY A		
qqμμ contact interaction : m	End2 at ² (2010) (a00 × 1104 4094)	49	tev A		
SSM : m _{eviu}	C/1.06-1.01 fb [*] (2011) [pistiminory]	taunar Z' mass			
SSM : m _{ter}	Coll 64 (6 ¹) (2015) (and industry)	skina Wimass			
Scalar LQ pairs (∂=1) : kin. vars. in eejj, evjj –	Cr05 pb ² (2910) (a00 x 1104,444 t)	aweev 1 st gen. LQ mass			
Scalar LQ pairs (β=1) : kin. vars. in μμjj, μνjj	E-485 ato* (2010) (arXiv:1104.4481)	azaa 2 rd gen Lü mass			
4 ²¹ family : coll. mass in Q ₄ Q ₄ → WqWq	even ve _{la} terio) (autres cores evel-ceri	an and O _d make			
4^m family : $d_a \overline{d}_a \rightarrow WtWt$ (SS dilepton)	C+34 pt/1 (2010) [preniminary]	encew d, mass			
Major. neutr. (V _{4-tern} , A=1 TeV) : S5 dilepton	Z.: 28 pix ¹ (2010) (positivity arg)	ALBERT NITASS			
Excited quarks : m _{diet}	CROBE IN ¹ (2015) (WILAS-CONF-OR 1-085)	g"m" g"m	835		
Axigluons : m _{dijer}	C/O RE 15 ¹ OR TO (ATLAS-CONFORT-DRS)	seinev Ax	gluon mass		
Color octot scalar : m _{dijet}	CARE IN DRIVEN CONTRACTORY	Lense Scalar report	lance mass		
	10"	1	10		

Impressive bounds on squarks and gluinos, into TeV range...

What do we learn? \rightarrow Papucci talk

 Plain vanilla SUSY models (like MSSM with flavor-universal soft masses) are being pushed into a corner

but

Rychkov EPS-HEP 2011

Several other, theoretically motivated, scenarios remain very poorly constrained by existing searches

<u>"Flavor-Split" spectra</u> (heavy 1st-2nd gen squarks, gluino below 1-1.5 TeV, light 3rd gen)

<u>"Squashed" spectra</u> (everything below ~500GeV but splittings are small, O(10GeV)) Low MET scenarios (not necessarily RPV)

WHY TO GO BEYOND THE SM

"OBSERVATIONAL" REASONS

•HIGH ENERGY PHYSICS NO (but A_{FB}, bb) •FCNC, CP \neq NO (but CPV in Bs, sin2 β tension...)

•HIGH PRECISION LOW-EN.

NO (but $(g-2)_{\mu}$...)

•NEUTRINO PHYSICS

 $(\mathbf{YES})_{\nu} \neq 0, \ \theta_{\nu} \neq 0$

•COSMO - PARTICLE PHYSICS YESDM, ΔB_{COSm} , INFLAT., DE)

THEORETICAL REASONS

•INTRINSIC INCONSISTENCY OF SM AS QFT

(spont. broken gauge theory without anomalies)

•NO ANSWER TO QUESTIONS THAT "WE" CONSIDER "FUNDAMENTAL" QUESTIONS TO BE ANSWERED BY "FUNDAMENTAL" THEORY

(hierarchy, unification, flavor)

EVIDENCE OF NP ALONG THE HIGH INTENSITY ROAD?

• "FLAVOR COLDS for the SM:



But *tension* in the UT fit even neglecting CPV in the B_s mixing

Lenz, Nierste + CKMfitter (2010)



Theoretical analyses without CPV in B_s mixing



+++ adding hadronic uncertainty δΔSΨκ=0.021



- discrepancies in the determinations of V_{ub} from inclusive semileptonic decays B→X_uIv, exclusive semileptonic decays B→πIv, and leptonic decay B→τv ("V_{ub} crisis")
- large difference of (14.4±2.9)% in the direct CP asymmetries measured in B⁰→K⁺π⁻ vs. B⁺→K⁺π⁰ decays, which is in conflict with the prediction of (2.2±2.4)% from QCD factorization ("B→Kπ puzzle")
- enhanced B_s→µ⁺µ⁻ branching ratio observed by CDF (but not by LHCb and CMS ☺) NEUBERT, EPS11

For many years, there has been a persistent discrepancy between determinations of $|V_{ub}|$ from **inclusive and exclusive semileptonic decays** of B mesons (B \rightarrow X_uIv vs. B \rightarrow πIv). HFAG quotes:

$$|V_{ub}|_{\text{incl}} = (4.32 \pm 0.16 \pm 0.22) \cdot 10^{-3}$$
$$|V_{ub}|_{\text{excl}} = (3.51 \pm 0.10 \pm 0.46) \cdot 10^{-3}$$

Measurement of the purely leptonic decay $B \rightarrow \tau v$ sharpen the discrepancy further:



 interesting rare decays, which can be much enhanced in models with a warped extra dimension or SUSY models with large tanβ

Excess in B_s mode reported by CDF:



$\mathcal{B}(B_s \to \mu^+ \mu^-) = (1.8^{+1.1}_{-0.9}) \cdot 10^{-8}$	SM: $(3.2 \pm 0.2) \cdot 10^{-9}$
$\mathcal{B}(B_d \to \mu^+ \mu^-) < 6.0 \cdot 10^{-9}$	SM: $(1.0 \pm 0.1) \cdot 10^{-10}$

Unfortunately no excess seen at LHCb (CMS):

$$\mathcal{B}(B_s \to \mu^+ \mu^-) < 1.5 \ (1.9) \cdot 10^{-8}$$

$$\mathcal{B}(B_d \to \mu^+ \mu^-) < 5.2 \ (4.6) \cdot 10^{-9}$$
 (at 95% CL)

NEUBERT EPS11 These bounds to not rule out the CDF result, but without refined LHC measurements the situation is inconclusive!

Relevant Parameter Space for 2 fb⁻¹

Jones at the EPS-HEP 2011 on the work in progress by Calibbi, Hodgkinson, Jones, A.M. and Vives



Flavour 3σ Constraints



The Role of B $\rightarrow \mu \mu$ **IMPACT ON THE SUSY** • LHCb with 2 fb⁻¹ PARAMETER SPACE - Exclusion of BR(B $\rightarrow \mu \mu$) down to 4x10⁻⁹, 95% C.L. - 3σ evidence of BR(B -> $\mu \mu$) down to $5x10^{-9}$. - 5σ discovery of BR(B -> $\mu \mu$) down to $9x10^{-9}$.

R. Lambert @ Moriond

• CDF with 7 fb⁻¹

 $-BR(B_{s} \rightarrow \mu \mu) = (1.8 \pm 1) \times 10^{-8}$

CDF Collaboration (1107.2304 [hep-ex])

Exclusion due to $B_s \rightarrow \mu \mu$

 $BR(B_s \to \mu\mu) < 4 \times 10^{-9}$



CHJMV work in progress

A Large $B_s \rightarrow \mu \mu$

 $\square BR(B_s \to \mu\mu) > 5 \times 10^{-9} \qquad \square BR(B_s \to \mu\mu) > 9 \times 10^{-9}$



A Large $B_s \rightarrow \mu \mu$

 $\square BR(B_s \to \mu\mu) > 5 \times 10^{-9} \qquad \square BR(B_s \to \mu\mu) > 9 \times 10^{-9}$



The muon g-2: the experimental result



Today: aµ^{EXP} = (116592089 ± 54_{stat} ± 33_{sys})x10⁻¹¹ [0.5ppm].

Future: new muon g-2 experiments proposed at:

- Fermilab (P989), aiming at 0.14ppm STAGE-1 APPROVAL!!
- J-PARC aiming at 0.1 ppm

[D. Hetzog & N. Saito, U.Paris, Feb 2010; B. Lee Roberts & T. Mibe, Tau2010]

Are theorists ready for this (amazing) precision? [not yet]

The muon g-2: Standard Model vs. Experiment

Adding up all contributions, we get the following SM predictions and comparisons with the measured value:

 a_{μ}^{EXP} = 116592089 (63) × 10⁻¹¹

E821 – Final Report: PRD73 (2006) 072 with latest value of $\lambda = \mu_{\mu}/\mu_{p}$ (CODATA'06)

$a_{\mu}^{\text{SM}} imes 10^{11}$	$(\Delta a_{\mu} = a_{\mu}^{\text{EXP}} - a_{\mu}^{\text{SM}}) \times 10^{11}$	σ
[1] 116591782(59)	307 (86)	3.6
[2] 116591802(49)	287 (80)	3.6
$[3] \ 116591830(52)$	259 (82)	3.2
[4] 116591894(54)	195 (83)	2.4

with $a_{\mu}^{HHO}(IbI) = 105 (26) \times 10^{-11}$

- [1] F. Jegerlehner, A. Nyffeler, Phys. Rept. 477 (2009) 1
- [2] Davier et al, arXiv:1010.4180, Oct 2010 (includes BaBar and KLOE10 2π)
- [3] HLMNT10: Hagiwara et al, Tau 2010, Sep. 2010 (incl BaBar and KLOE10 2π)
- [4] Davier et al, arXiv:1010.4180, Oct 2010, τ data.

Note that the th. error is now about the same as the exp. one



✓ axigluons, diquarks, new weak bosons, EDs etc..

✓ Or gluon radiations modeling at NLO?

Is it possible that there is "only" a light higgs boson and no NP?

- This is acceptable if one argues that no ultraviolet completion of the SM is needed at the TeV scale simply because there is no actual fine-tuning related to the higgs mass stabilization (the correct value of the higgs mass is "environmentally" selected). This explanation is similar to the one adopted for the cosmological constant
- Barring such wayout, one is lead to have TeV NP to ensure the unitarity of the elw. theory at the TeV scale

% FINE-TUNING FOR THE NEW PHYSICS AT THE ELW. SCALE

- Elementary Higgs →In the MSSM % fine-tuning among the SUSY param. to avoid light SUSY particles which would have been already seen at LEP and Tevatron and now also at LHC
- Elementary Higgs → PSEUDO-GOLDSTONE boson in the LITTLE HIGGS model → Λ² div. cancelled by new colored fermions, new W,Z, γ, 2Higgs doublets... → % fine-tuning to avoid too large elw. Corrections
- COMPOSITE HIGGS in a 5-dim. holographic theory (Higgs is a PSEUDO-GOLDSTONE boson and the elw. symmetry breaking is triggered by bulk effects (in 5 dim. the theory is WEAKLY coupled, but in 4 dim. the bulk looks like a STRONGLY coupled sector) → also here % fine-tuning needed to survive the elw. precision tests

The Energy Scale from the "Observational" New Physics



AT THE ELW. SCALE

THE DM ROAD TO NEW **PHYSICS BEYOND THE SM**: IS DM A PARTICLE OF THE NEW PHYSICS AT THE ELECTROWEAK ENERGY SCALE ?

CONNECTION DM – ELW. SCALE THE WIMP MIRACLE :STABLE ELW. SCALE WIMPs

1) ENLARGEMENT OF THE SM	SUSY (χ ^μ , θ)	EXTRA DIM . (X ^{μ,} j ⁱ⁾	LITTLE HIGGS. SM part + new part
	Anticomm.	New bosonic	to cancel Λ^2
	Coord.	Coord.	at 1-Loop
2) SELECTION RULE	R-PARITY LSP	KK-PARITY LK	P T-PARITY LTP
→DISCRETE SYMM.	Neutralino spin 1/2	spin1	spin0
→STABLE NEW PART.			
3) FIND REGION (S)	m↓ LSP	m ↓ LKP	↓ m _{LTP}
WHERE THE "I " NEW	~100 - 200	~600 - 800	~400 - 800
PART. IS NEUTRAL + $\Omega_1 h^2 OK$	GeV *	GeV	GeV

Bottino, Donato, Fornengo, Scopel

IS THE "WIMP MIRACLE" AN ACTUAL MIRACLE?

USUAL STATEMENT

Many possibilities for DM candidates, but WIMPs are really special: peculiar coincidence between particle physics and cosmology parameters to provide a VIABLE DM CANDIDATE AT THE ELW. SCALE

HOWEVER

when it comes to quantitatively reproduce the precisely determined DM density \rightarrow once again the fine-tuning threat...

LHC reach in the SUSY parameter space (example CMSSM – A, M, m, $tan\beta$, μ)



(see e.g., Ellis, Ferstl, Olive)

Cerdeno '09

Recent Status



11/07/26



DM and NON-STANDARD COSMOLOGIES BEFORE NUCLEOSYNTHESIS

- NEUTRALINO RELIC DENSITY MAY DIFFER FROM ITS STANDARD VALUE, i.e. the value it gets when the expansion rate of the Universe is what is expected in Standard Cosmology (EX.: SCALAR-TENSOR THEORIES OF GRAVITY, KINATION, EXTRA-DIM. RANDALL-SUNDRUM TYPE II MODEL, ETC.)
- WIMPS MAY BE "COLDER", i.e. they may have smaller typical velocities and, hence, they may lead to smaller masses for the first structures which form **GELMINI, GONDOLO**

WHY H
$$\neq$$
 H_{GR}
 $H_{GR}^2 = \frac{1}{3M_p^2} \rho_{tot} \simeq 2.76 g_* \frac{T^4}{M_p^2}$

Change the number of relativistic d.o.f.'s, g_* ;

R. Catena

- - Kination
 P. Salati, Phys. Lett. B 571 (2003) 121
- Consider theories where the effective Planck mass is different from the constant M_p:
 - Scalar-Tensor theories
 R. C., N. Fornengo, A. Masiero, M. Pietroni and F. Rosati, Phys. Rev. D 70 (2004) 063519
 - Extradimensions
 L. Randall and R. Sundrum, Phys. Rev. Lett. 83 (1999) 4690

DIRECT AND INDIRECT SEARCHES FOR WIMPs

• PROBING NEW PHYSICS AT THE ELW. SCALE

 INFORMATION ON THE EVOLUTION OF THE EARLY UNIVERSE BEFORE THE NUCLEOSYNTHESIS TIME, i.e. at times < 1 sec.

4. ELW. SYMM. BREAKING STABILIZATION **VS. FLAVOR PROTECTION: THE SCALE TENSION**

$$M(B_{d}-\bar{B}_{d}) \sim c_{SM} \frac{(v_{t} V_{tb} * V_{td})^{2}}{16 \pi^{2} M_{W}^{2}} + c_{new} \frac{1}{\Lambda^{2}}$$

If $c_{new} \sim c_{SM} \sim 1$ Isidori
 $\Lambda > 10^{4} \text{ TeV for } O^{(6)} \sim (\bar{s} d)^{2}$
 $[K^{0}-\bar{K^{0}} \text{ mixing }]$ $\Lambda > 10^{3} \text{ TeV for } O^{(6)} \sim (\bar{b} d)^{2}$
 $[B^{0}-\bar{B^{0}} \text{ mixing }]$

 $\begin{bmatrix} B^0 - B^0 \text{ mixing} \end{bmatrix}$

UV SM COMPLETION TO STABILIZE THE ELW. SYMM. BREAKING: $\Lambda_{UV} \sim O(1 \text{ TeV})$

Flavor Structure in the SM and Beyond NEUBERT EPS-HEP 2011



Generic bounds without a flavor symmetry
$K - \overline{K}$	8×10^{-7}	6×10^{-9}
$D - \overline{D}$	$5 imes 10^{-7}$	$1 imes 10^{-7}$
$B - \overline{B}$	$5 imes 10^{-6}$	1×10^{-6}
$B_s - \overline{B_s}$	$2 imes 10^{-4}$	$2 imes 10^{-4}$

SMALLNESS OF THE NP COUPLINGS IF THE NP SCALE IS 1 TEV

THE FLAVOUR PROBLEMS

FERMION MASSES

What is the rationale hiding behind the spectrum of fermion masses and mixing angles (our "**Balmer lines**" problem)

LACK OF A FLAVOUR "THEORY"

(new flavour – horizontal symmetry, radiatively induced lighter fermion masses, dynamical or geometrical determination of the Yukawa couplings, ...?)

FCNC

Flavour changing neutral current (FCNC) processes are suppressed.

In the SM two nice mechanisms are at work: the **GIM mechanism** and the structure of the **CKM mixing matrix.**

How to cope with such delicate suppression if the there is new physics at the electroweak scale?

<u>MSSM 🗴 FAMILY SYMM.</u>

- AMBITION: simultaneously accounting for the "correct" SM fermion masses and mixings (SM Flavor Puzzle) and a structure of the SUSY soft breaking masses allowing for adequate FCNC suppression + possible "explanation" of the alleged SM FCNC difficulties (SUSY Flavor Puzzle)
- Mechanism a la Frogatt Nielsen with abelian or non-abelian family symmetry

Froggatt-Nielsen mechanism and flavour symmetry to understand

small Yukawa elements. Example: $U(1)_{fl}$





ROBERTS, ROMANINO, ROSS, VELASCO-SEVILLA; ROSS, VELASCO-SEVILLA, VIVES

- $Q, L \sim 3$ and $d^c, u^c, e^c \sim 3$; flavon fields: $\theta_3, \theta_{23} \sim \overline{3}, \overline{\theta}_3, \overline{\theta}_{23} \sim 3$
- Family Symmetry breaking: $SU(3) \xrightarrow{\langle \theta_3 \rangle} SU(2) \xrightarrow{\langle \theta_{23} \rangle} \emptyset$

$$\theta_3, \overline{\theta}_3 = \begin{pmatrix} 0\\0\\a_3 \end{pmatrix}, \ \theta_{23}, \overline{\theta}_{23} = \begin{pmatrix} 0\\b\\b \end{pmatrix}$$
 with $\begin{pmatrix} \underline{a_3}\\M \end{pmatrix} \sim \mathcal{O}(1), \ \begin{pmatrix} \underline{b}\\M_u \end{pmatrix} \simeq \begin{pmatrix} \underline{b}\\M_d \end{pmatrix}^2 = \varepsilon \sim 0.05.$

• Yukawa superpotential: $W_Y = H\psi_i\psi_j^{\circ} \left[\theta_3^i\theta_3^j + \theta_{23}^i\theta_{23}^j \left(\theta_3\theta_3\right) + \epsilon^{ikl}\theta_{23,k}\theta_{3,l}\theta_{23}^j \left(\theta_{23}\theta_3\right)\right]$

O. VIVES

$$Y^{f} = \begin{pmatrix} 0 & a \varepsilon^{3} & b \varepsilon^{3} \\ a \varepsilon^{3} & \varepsilon^{2} & c \varepsilon^{2} \\ b \varepsilon^{3} & c \varepsilon^{2} & 1 \end{pmatrix} \frac{|a_{3}|^{2}}{M^{2}},$$

LFV CONSTRAINTS IN THE $M_0 - M_{1/2}$ **SUSY PLANE**



FLAVOR BLINDNESS OF THE NP AT THE ELW. SCALE?

- THREE DECADES OF FLAVOR TESTS (Redundant determination of the UT triangle → verification of the SM, theoretically and experimentally "high precision"
 FCNC tests, ex. b → s + γ, CP violating flavor conserving and flavor changing tests, lepton flavor violating (LFV) processes, …) clearly state that:
- A) in the HADRONIC SECTOR the CKM flavor pattern of the SM represents the main bulk of the flavor structure and of (flavor violating) CP violation;
- B) in the LEPTONIC SECTOR: although neutrino flavors exhibit large admixtures, LFV, i.e. non – conservation of individual lepton flavor numbers in FCNC transitions among charged leptons, is extremely small: once again the SM is right (to first approximation) predicting negligibly small LFV

What to make of this triumph of the CKM pattern in hadronic flavor tests?

New Physics at the Elw. Scale is Flavor Blind CKM exhausts the flavor changing pattern at the elw. Scale

MINIMAL FLAVOR VIOLATION

MFV : Flavor originates only from the SM Yukawa coupl.

New Physics introduces

NEW FLAVOR SOURCES in addition to the CKM pattern. They give rise to contributions which are <10% in the "flavor observables" which have already been observed!

SuperB vs. LHC Sensitivity Reach in testing Λ_{SUSY}

	superB	general MSSM	high-scale MFV
$ \left(\delta^d_{13}\right)_{LL} ~(LL\gg RR)$	$1.8 \cdot 10^{-2} \frac{m_q}{(350 {\rm GeV})}$	1	$\sim 10^{-3} rac{(350 { m GeV})^2}{m_{ ilde{q}}^2}$
$ \left(\delta^d_{13}\right)_{LL} ~(LL\sim RR)$	$1.3 \cdot 10^{-3} \frac{m_{\tilde{q}}}{(350 \text{ GeV})}$	1	_
$ \left(\delta^{d}_{13}\right)_{LR} $	$3.3 \cdot 10^{-3} \frac{m_{\tilde{q}}}{(350 \text{GeV})}$	$\sim 10^{-1} aneta rac{(350 { m GeV})}{m_{\tilde{q}}}$	$\sim 10^{-4} {\rm tan} \beta \frac{(350 {\rm GeV})^3}{m_{\rm q}^3}$
$ \left(\delta^{d}_{23}\right)_{LR} $	$1.0 \cdot 10^{-3} \frac{m_{\tilde{q}}}{(350 \mathrm{GeV})}$	$\sim 10^{-1} aneta rac{(350 { m GeV})}{m_{ m Q}}$	$\sim 10^{-3} \tan\beta \frac{(350 {\rm GeV})^3}{m_{\rm q}^3}$

SuperB can probe MFV (with small-moderate tan β) for TeV squarks; for a generic non-MFV MSSM \longrightarrow sensitivity to squark masses > 100 TeV ! Ciuchini, Isidori, Silvestrini SLOW-DECOUPLING OF NP IN FCNC

Estimates of error for 2015					
Hadronic matrix element	Current lattice error	6 TFlop Year	60 TFlop Year [2011 LHCb]	1-10 PFlop Year [2015 SuperB	
$f_{+}^{K\pi}(0)$	0.9% (22% on 1-f ₊)	0.7% (17% on 1-f ₊)	0.4% (10% on 1-f ₊)	< 0.1% (2.4% on 1-f ₊)	
Â _K	11%	5%	3%	1%	
f _B	14%	3.5 - 4.5%	2.5 - 4.0%	1 – 1.5%	
$\mathbf{f}_{\mathbf{B}s}\mathbf{B}_{\mathbf{B}s}^{1/2}$	13%	4 - 5%	3 - 4%	1 – 1.5%	
ξ	5% (26% on ξ-1)	3% (18% on ξ-1)	1.5 - 2 % (9-12% on ξ-1)	0.5 – 0.8 % (3-4% on ξ-1)	
$\mathcal{F}_{B \rightarrow D/D^* l \nu}$	4% (40% on 1- <i>F</i>)	2% (21% on 1- <i>F</i>)	1.2% (13% on 1-F)	0.5% (5% on 1-F)	
$f_{+}^{B\pi},$	11%	5 .5 - 6.5%	4 - 5%	2-3%	
$T_1^{B \rightarrow K^*/\rho}$	13%			3-4%	

SUSY SEE-SAW

• UV COMPLETION OF THE SM TO STABILIZE THE ELW. SCALE:

LOW-ENERGY SUSY

 COMPLETION OF THE SM FERMIONIC SPECTRUM **TO ALLOW FOR NEUTRINO MASSES:** NATURALLY SMALL PHYSICAL NEUTRINO **MASSES WITH RIGHT-**HANDED NEUTRINO WITH A LARGE MAJORANA MASS

SEE-SAW

LFV and NEW PHYSICS

- Flavor in the HADRONIC SECTOR: CKM paradigm
- Flavor in the LEPTONIC SECTOR:
 - Neutrino masses and (large) mixings
 - Extreme smallness of LFV in the charged lepton sector of the SM with massive neutrinos:

$$I_k$$
 suppressed by $(m_v_i^2 - m_v_k^2) / M_W^2$

NEW BOUND OF MEG AT THE EPS 2011

The MEG Experiment $\mu^+ ightarrow { m e}^+ \gamma$

SHOWN AT ICHEP 2010



Blue lines are 1(39.3 % included inside the region w.r.t. analysis window), 1.64(74.2%) and 2(86.5%) sigma regions.

For each plot, cut on other variables for roughly 90% window is applied.

Numbers in figures are ranking by Leg/(Lnub+Leg). Same numbered dots in the right and the left figure are an identical event.



MEG summary

2009+2010 data consistent w/ no signal

 New physics is now constrained by 5× tighter upper limit: BR < 2.4×10⁻¹² @90% C.L. (Preprint will be posted at arXiv today)

 MEG is accumulating more data this and next year to reach O(10⁻¹³) sensitivity; So stay tuned!

Detector improvements/upgrades
 T. MORI AT THE EPS-HEP 2011

SUSY SEESAW: Flavor universal SUSY breaking and yet large lepton flavor violation Borzumati, A. M. 1986 (after discussions with W. Marciano and A. Sanda)

$$L = f_l \ \overline{e}_R Lh_1 + f_v \ \overline{v}_R Lh_2 + M \ v_R v_R$$

$$\stackrel{\tilde{L}}{\longrightarrow} \stackrel{\tilde{L}}{\longrightarrow} (m_{\tilde{L}}^2)_{ij} \square \underbrace{\frac{1}{8\pi^2}}_{3\pi^2} (3m_0^2 + A_0^2) (f_v^{\dagger} f_v)_{ij} \log \frac{M}{M_G}$$

Non-diagonality of the slepton mass matrix in the basis of diagonal lepton mass matrix depends on the unitary matrix U which diagonalizes $(f_v + f_v)$

How Large LFV in SUSY SEESAW?

- 1) Size of the **Dirac neutrino couplings** f_v
- 2) Size of the diagonalizing matrix U

In **MSSM seesaw** or in **SUSY SU(5)** (Moroi): not possible to correlate the neutrino Yukawa couplings to know Yukawas;

In SUSY SO(10) (A.M., Vempati, Vives) at least one neutrino Dirac Yukawa coupling has to be of the order of the top Yukawa coupling \longrightarrow one large of O(1) f_v

U **—** two "extreme" cases:

a) U with "small" entries
b) U with "large" entries with the exception of the 13 entry
U = PMNS matrix responsible for the diagonalization of the neutrino mass matrix

THE STRONG ENHANCEMENT **OF LFV IN SUSY SEESAW MODELS** CAN OCCUR EVEN IF THE MECHANISM **RESPONSIBLE FOR SUSY BREAKING IS** ABSOLUTELY **FLAVOR BLIND**

LFV in SUSYGUTs with SEESAW

Scale of appearance of the SUSY soft breaking terms resulting from the spontaneous breaking of supergravity

M_{GUT} M_R

Mpi

Low-energy SUSY has "memory" of all the multi-step RG occurring from such superlarge scale down to M_W

potentially large LFV

M

Barbieri, Hall; Barbieri, Hall, Strumia; Hisano, Nomura,
Yanagida; Hisano, Moroi, Tobe Yamaguchi; Moroi;A.M.,, Vempati, Vives;
Carvalho, Ellis, Gomez, Lola; Calibbi, Faccia, A.M, Vempati
LFV in MSSMseesaw: μ eγ Borzumati, A.M.
τ μγ Blazek, King;
General analysis: Casas Ibarra; Lavignac, Masina,Savoy; Hisano, Moroi, Tobe, Yamaguchi; Ellis,
Hisano, Raidal, Shimizu; Fukuyama, Kikuchi, Okada; Petcov, Rodejohann, Shindou, Takanishi;
Arganda, Herrero; Deppish, Pas, Redelbach, Rueckl; Petcov, Shindou

$\mu \rightarrow e + \gamma$ in SUSYGUT: past and future

$\mu ightarrow e \, \gamma \,$ in the ${\it U}_{e3}$ = 0 PMNS case



CFMV

Comparing CKM and PMNS



$\mu ightarrow e$ in Ti and **PRISM/PRIME** conversion experiment



LFV from SUSY GUTs

Lorenzo Calibbi

Antusch, Arganda, Herrero, Teixeira



LFV, g – 2, EDM: a promising correlation in SUSY SEESAW



DEVIATION from μ - e UNIVERSALITY A.M., Paradisi, Petronzio

• Denoting by $\Delta r_{NP}^{e-\mu}$ the deviation from $\mu - e$ universality in $R_{K,\pi}$ due to new physics, i.e.:

$$R_{K,\pi} = R_{K,\pi}^{SM} \left(1 + \Delta r_{K,\pi NP}^{e-\mu} \right),$$

• we get at the 2σ level:

 $-0.063 \le \Delta r_{KNP}^{e-\mu} \le 0.017 \text{ NA48/2}$

$-0.0107 \le \Delta r_{\pi NP}^{e-\mu} \le 0.0022 \text{ PDG}$

Presently: error on R_{κ} down to the **1% level** (KLOE (09) and NA48 (07 data);using 40% of the data collected in 08, NA62 is now decreasing the uncertainty at the **0.7% level Prospects**: Summer conf. we'll have the result concerning the 40% data analysis by NA62 and when the analysis of the whole sample of data is accomplished **the stat. uncertainty will be < 0.3%**

HIGGS-MEDIATED LFV COUPLINGS

- When non-holomorphic terms are generated by loop effects (HRS corrections)
- And a source of LFV among the sleptons is present
- Higgs-mediated (radiatively induced) H-lepton-lepton LFV couplings arise
 Babu, Kolda; Sher; Kitano,Koike,Komine, Okada; Dedes, Ellis, Raidal; Brignole,Rossi; Arganda,Curiel,Herrero,Temes; Paradisi; Brignole,Rossi

H mediated LFV SUSY contributions to R_K

$$R_{K}^{LFV} = \frac{\sum_{i} K \to e\nu_{i}}{\sum_{i} K \to \mu\nu_{i}} \simeq \frac{\Gamma_{SM}(K \to e\nu_{e}) + \Gamma(K \to e\nu_{\tau})}{\Gamma_{SM}(K \to \mu\nu_{\mu})} , \quad i = e, \mu, \tau$$



LFU breaking occurs in a **LF conserving** case because of the splitting in slepton masses



Figure 2: Left: $\Delta r_K^{e/\mu}$ as a function of the mass splitting between the second and the first (left-handed) slepton generations. Red dots can saturate the $(g-2)_{\mu}$ discrepancy at the 95% C.L., i.e. $1 \times 10^{-9} < (g-2)_{\mu} < 5 \times 10^{-9}$. Right: $\Delta r_K^{e/\mu}$ as a function of M_{H^+} .

A.M., PARADISI. PETRONZIO

LFU breaking occurs with LFV

SUSY GUTs

• UV COMPLETION OF THE SM TO STABILIZE THE ELW. SCALE:

LOW-ENERGY SUSY

TREND OF UNIFICATION OF THE SM GAUGE COUPLINGS AT HIGH SCALE:

GUTs

Large v mixing - large b-s transitions in SUSY GUTs

In SU(5) $d_R \longrightarrow I_L$ connection in the 5-plet Large $(\Delta^{I}_{23})_{LL}$ induced by large f_v of O(f_{top}) is accompanied by large $(\Delta^{d}_{23})_{RR}$

In SU(5) assume large f_v (Moroi) In SO(10) f_v large because of an underlying Pati-Salam symmetry (Darwin Chang, A.M., Murayama)

See also: Akama, Kiyo, Komine, Moroi; Hisano, Moroi, Tobe, Yamaguchi, Yanagida; Hisano, Nomura; Kitano,Koike, Komine, Okada

FCNC HADRON-LEPTON CONNECTION IN SUSYGUT



GUT -RELATED SUSY SOFT BREAKING TERMS

$$\begin{split} m_Q^2 &= m_{\tilde{e}c}^2 = m_{\tilde{u}c}^2 = m_{10}^2 \\ m_{\tilde{d}c}^2 &= m_L^2 = m_{\overline{5}}^2 \\ A_{ij}^e &= A_{ji}^d \,. \end{split}$$

SU(5) RELATIONS

	Relations at weak-scale	Relationss at $M_{\rm GUT}$
(1)	$(\delta^u_{ij})_{\mathrm{RR}}~pprox~(m^2_{e^c}/m^2_{u^c})~(\delta^l_{ij})_{\mathrm{RR}}$	$m^2_{u^c_0} = m^2_{e^c_0}$
(2)	$(\delta^q_{ij})_{\mathrm{LL}} \approx (m_{e^c}^2/m_Q^2) \ (\delta^l_{ij})_{\mathrm{RR}}$	$m_{Q_0}^2 = m_{e^c_0}^2$
(3)	$(\delta^d_{ij})_{\mathrm{RR}}~pprox~(m_L^2/m_{d^c}^2)~(\delta^l_{ij})_{\mathrm{LL}}$	$m^2_{d^c_0} = m^2_{L_0}$
(4)	$(\delta^d_{ij})_{\mathrm{LR}} \approx (m^2_{L_{avg}}/m^2_{Q_{avg}}) (m_b/m_{\tau}) (\delta^l_{ij})^{\star}_{\mathrm{LR}}$	$A^e_{ij_0} = A^d_{ji_0}$

Bounds on the hadronic $(\delta_{12})_{RR}$ as modified by the inclusion of the LFV correlated bound



3 QUESTIONS

- Are we sure that there is new physics (NP) at the TeV scale? YES (barring an antropic approach)
- If yes, are we sure that LHC will see something "new", i.e. beyond the SM with its "standard higgs boson"? YES
- If there is new physics at the TeV scale, what can flavor and DM physics tell to LHC and viceversa? (or, putting it in a less politically correct fashion: if LHC starts seeing some new physics signals, are flavor and DM physics still a valuable road to NP, or are they definitely missing that train? NO, actually to catch the "right train" it is highly desirable, though maybe strictly not necessary, to make use of all the three roads at the same time

A FUTURE FOR FLAVOR PHYSICS IN OUR SEARCH BEYOND THE SM?

- The traditional competition between direct and indirect (FCNC, CPV) searches to establish who is going to see the new physics first is no longer the priority, rather
- COMPLEMENTARITY between direct and indirect searches for New Physics is the key-word
- Twofold meaning of such complementarity:
- i) synergy in "reconstructing" the "fundamental theory" staying behind the signatures of NP;
- ii) coverage of complementary areas of the NP parameter space (ex.: multi-TeV SUSY physics)