modeling molecular gas and associated star formation in cosmological simulations

Andrey Kravtsov

with



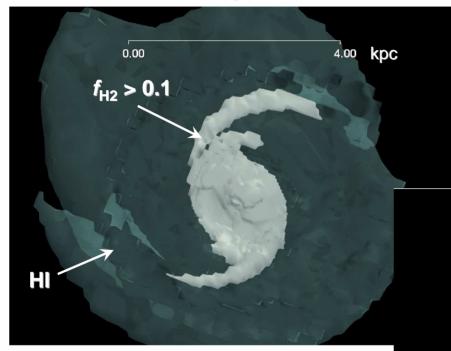
Nick Gnedin (Fermilab/U.Chicago)



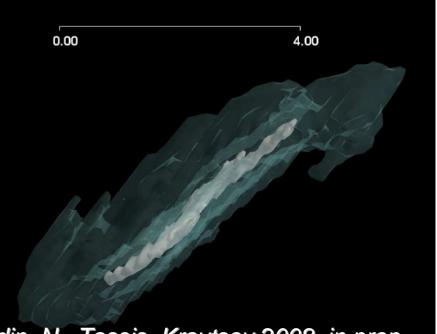
Kostas Tassis (U.Chicago->JPL)

Approximate model of H2 formation and radiative transfer

dense, molecular gas traces densest, high-pressure regions of the ISM



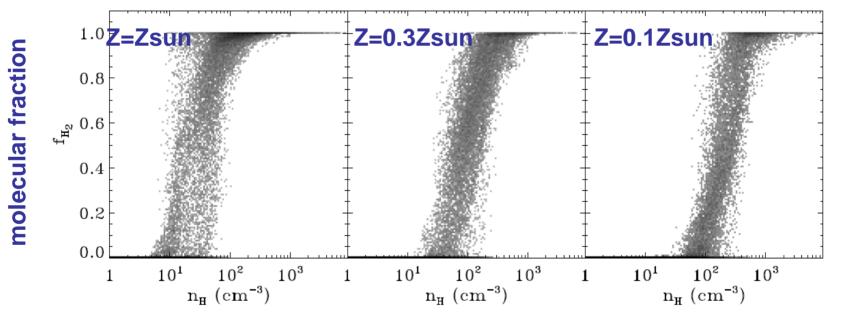
Galaxy formation simulation (ART code) with approximate 3D radiative transfer and a model for H2 formation on dust with approximate self-shielding using Sobolev approximation face-on and edge-on views of HI and H2 distribution in a z~4 gas disk



Gnedin, N., Tassis, Kravtsov 2008, in prep

Molecular fraction as a function of gas surface and local 3D density

Strong trends with metallicity (dust content) and local UV flux

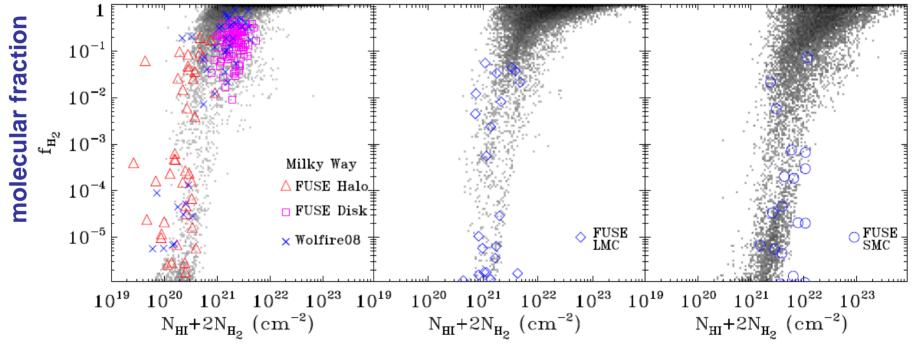


3D gas number density

Gnedin, N., Tassis, Kravtsov 2008, in preparation

Molecular fraction as a function of gas surface and local 3D density

Strong trends with metallicity (dust content) and local UV flux



HI+H2 column density

Gnedin, N., Tassis, Kravtsov 2008, in preparation