# Low-energy nu cross sections, and implications (with Ulrich Mosel)



## Japan & Santa Barbara (Wiki)

During World War II, Santa Barbara was home to Marine Corps Air Station Santa Barbara; Naval Reserve Center Santa Barbara at the harbor; was near to the Army's Camp Cook, present-day Vandenberg Air Force Base; and contained a hospital for treating servicemen wounded in the Pacific Theatre. On February 23, 1942, not long after the outbreak of war in the Pacific, the Japanese submarine *I-17* surfaced offshore and lobbed 16 shells at the Ellwood Oil Field, about 10 miles (16 km) west of Santa Barbara, in the first wartime attack by an enemy power on the U.S. mainland since the War of 1812. Although the shelling was inaccurate and only caused about \$500 damage to a catwalk, panic was immediate. Many Santa Barbara residents fled, and land values plummeted to historic lows.



## **CP** violation: warning

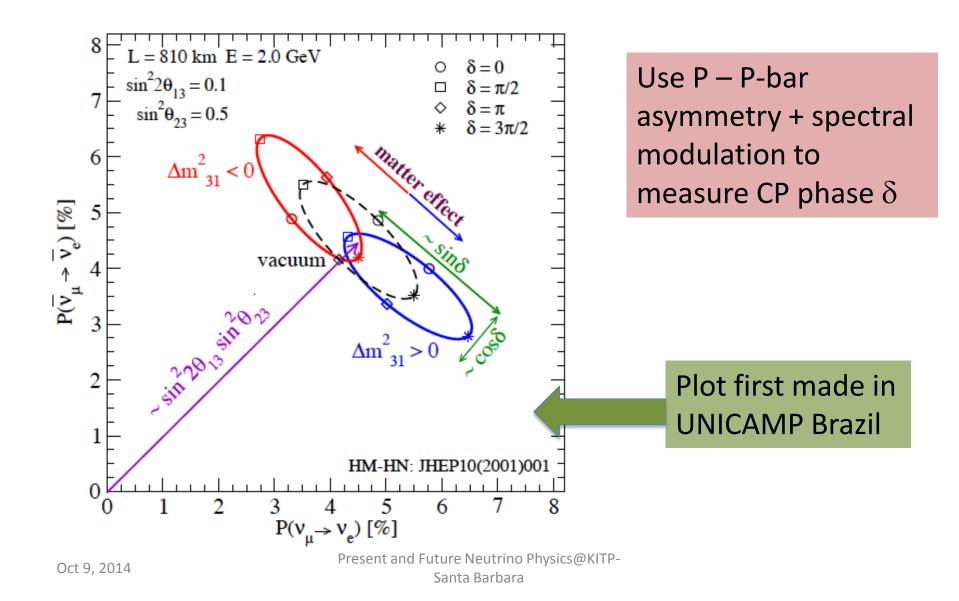
- Not possible to carry out unambiguous CPV detection: " $K_{|} \rightarrow 2 \pi \square$  does not appear to be possible (T-Violation would do, but ..)
- One cannot avoid CPV mimicking by other cause, matter effect, cross section errors, etc.
- May be we (theorists) should spend more time to think about possible mechanism which can mimic "CP phase effect"



## Motivating for v cross sections: CP

Present and Future Neutrino Physics@KITP-Santa Barbara

#### CP, matter effect, and all that



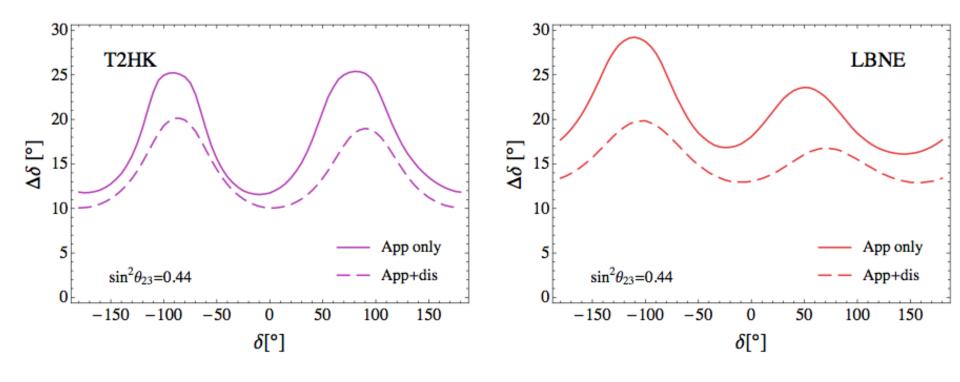
## CP measurement: basics 1

- For narrow band beam (T2K, NOVA) a set of measurement of (P, bar-P) determines  $\delta$
- However, effect of  $\delta$  (cosine and sine) has to be tiny
- Because it is suppressed by two small numbers:  $\Delta m_{21}^2 / \Delta m_{31}^2 \sim 0.031$ ,  $J_r = c_{12} s_{12} c_{23} s_{23} s_{13} \sim 0.035$
- Altogether it is  $P_{\mu e} \sim 10^{-3} \rightarrow$  we need big detector/intense neutrino beam

## CP measurement: basics 2

- For narrow band beam (T2K, NOVA) a set of measurement of (P, bar-P) determines  $\delta$
- There might be issues like, degeneracy etc. but let's ignore them
- Probably the most crucial issues are:
- To know nu and bar-nu flux accurately
- To know nu and nu-bar cross sections accurately

## Expected CP sensitivity: $\Delta\delta \sim 10-20^{\circ}$



P.Coloma, HM, S.J.Parke 1406.2551 T2HK: 7.5 MW year LBNE: 20 MW year (year =  $10^7$  sec)

Present and Future Neutrino Physics@KITP-

Santa Barbara

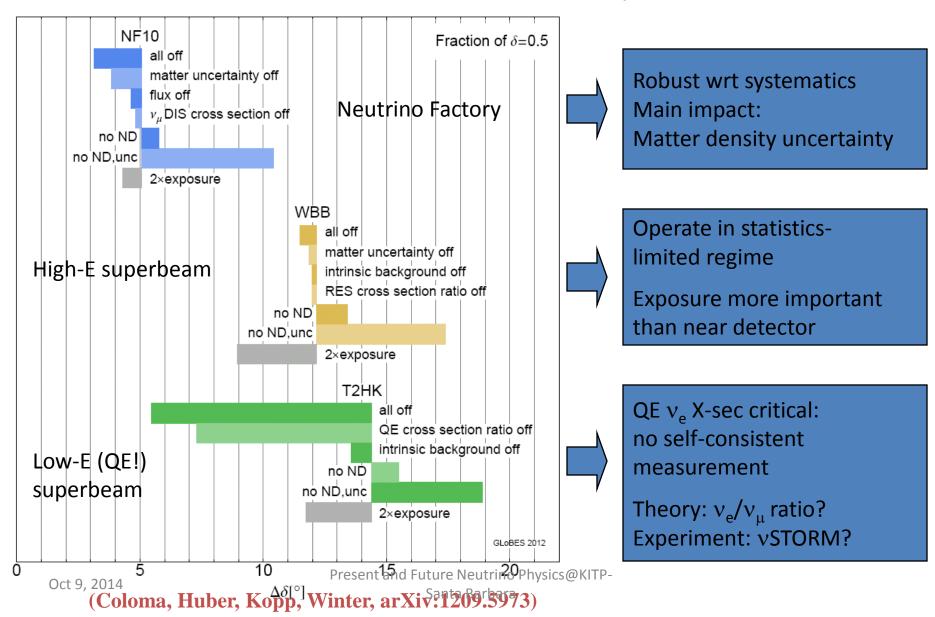
## Let us examine some of them

	L (km)	Detector (kton)	Beam Power	$E_p$ (GeV)	Flux peak	$(t_{\nu}, t_{\overline{\nu}})^{\dagger} \times 10^7 s$
LBNE	1300	LAr - 34	1.2 MW	120	$3~{ m GeV}$	(8.25, 8.25)
T2HK	295	WC - 560	$0.75 \ \mathrm{MW}$	30	$0.6~{\rm GeV}$	(3, 7)
$\mathrm{ESS}\nu\mathrm{SB}$	540	WC - 500	5 MW	2.5	$0.3~{ m GeV}$	(3.4, 13.6)
IDS-NF	2000	MIND - 100	$10^{21}\;\mu^{\pm}/10^{7}\;{\rm sec}$	NA	6 (9) GeV	$(10, 10)^{\ddagger}$

	Energy range	$\nu$ app.	$\bar{\nu}$ app.	$\nu$ dis.	$\bar{\nu}$ dis.
LBNE	$0.5$ - $8.0~{\rm GeV}$	1095/314	324/208	7340/82	3873/27
T2HK	$0.4$ - $1.2~{\rm GeV}$	3984/1705	2161/1928	26237/716	19232/735
$ESS\nu SB$	0.1 - 1.0  GeV	270/85	244/82	6198/113	4128/79
IDS-NF	$0.1$ - $9.0~{\rm GeV}$	20241/476	5257/269	171133/7370	106077/3279

P. Coloma, HM, & S.Parke, Note it's a very balanced collaboration! arXiv:1406.2551

# Main cause of error for CP (slide borrowed from Walter, modified)

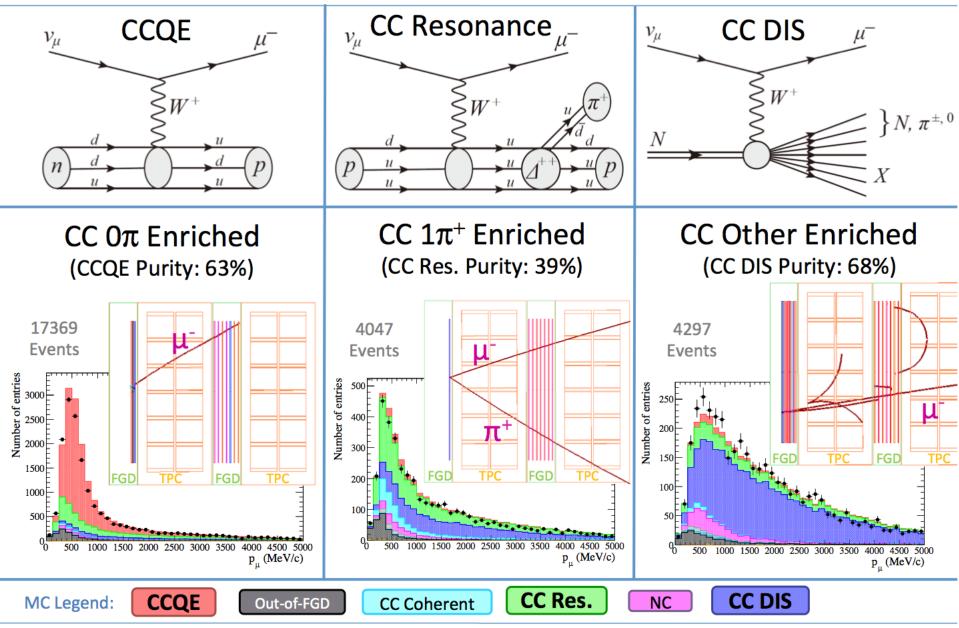




Measuring cross section must be easy



## ND280 Tracker Event Selection



Rencontres de Moriond, March 17, 2014

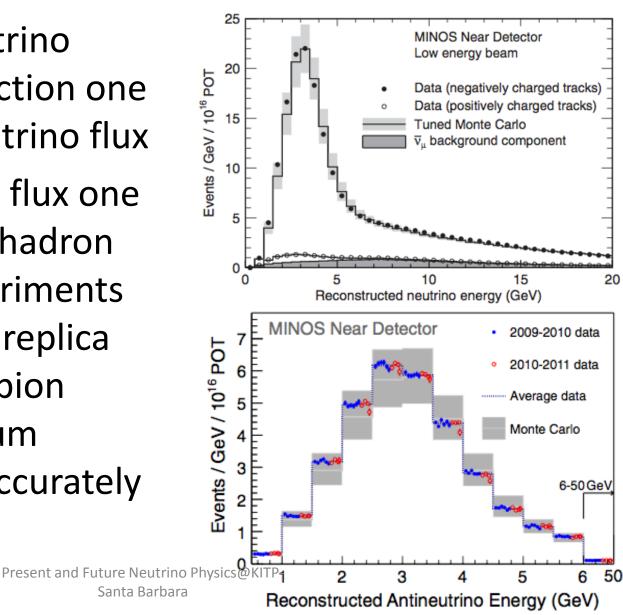
Patrick de Perio: T2K Neutrino Oscillation Results

# Do you know v flux? No!

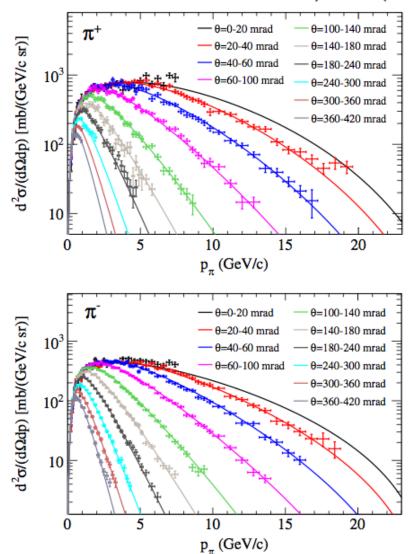


#### To measure neutrino nucleus cross section

- To measure neutrino nucleus cross section one has to know neutrino flux
- To determine nu flux one has to carry out hadron production experiments (preferably with replica target) to know pion (kaon) momentum distribution as accurately as possible



## For T2K: NA61/SHINE @ CERN



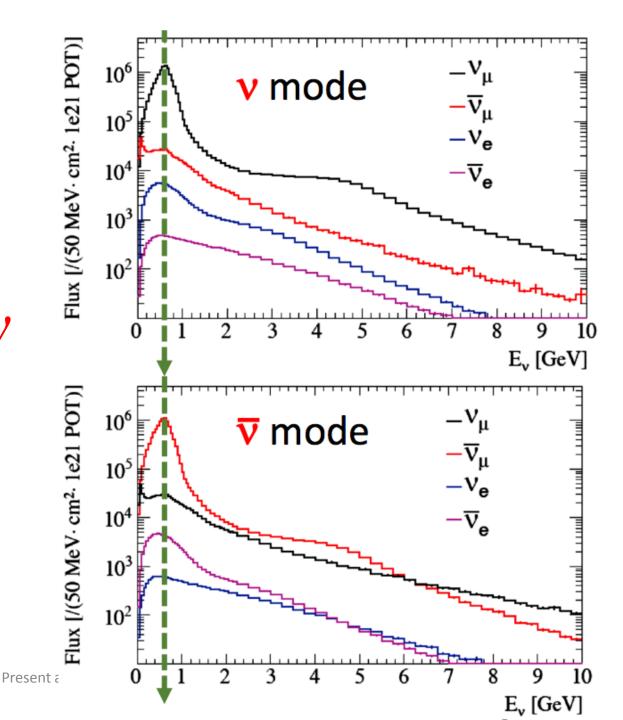
PHYSICAL REVIEW D 87, 012001 (2013)

Thin target

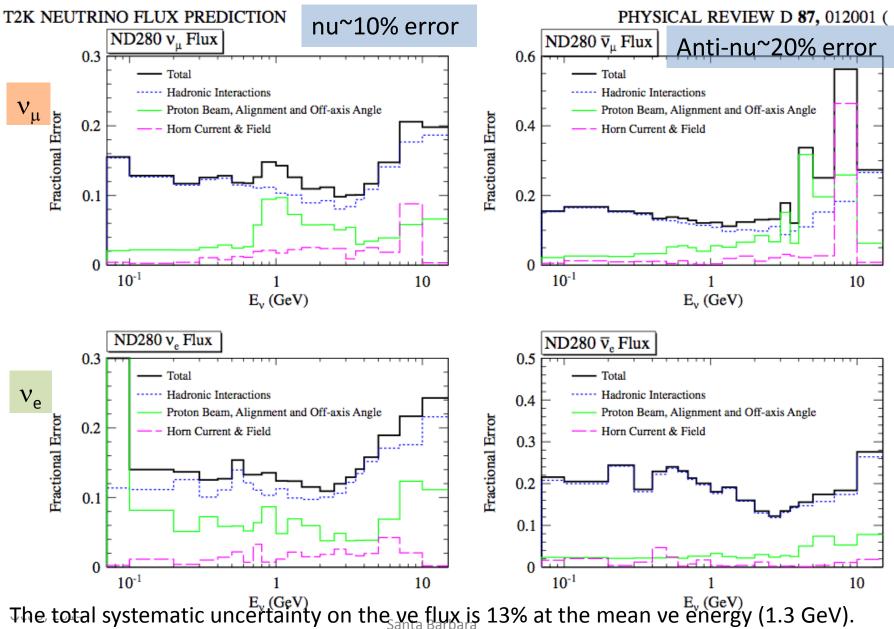


FIG. 28 (color online). The BMPT fits to the NA61/SHINE pion production data.

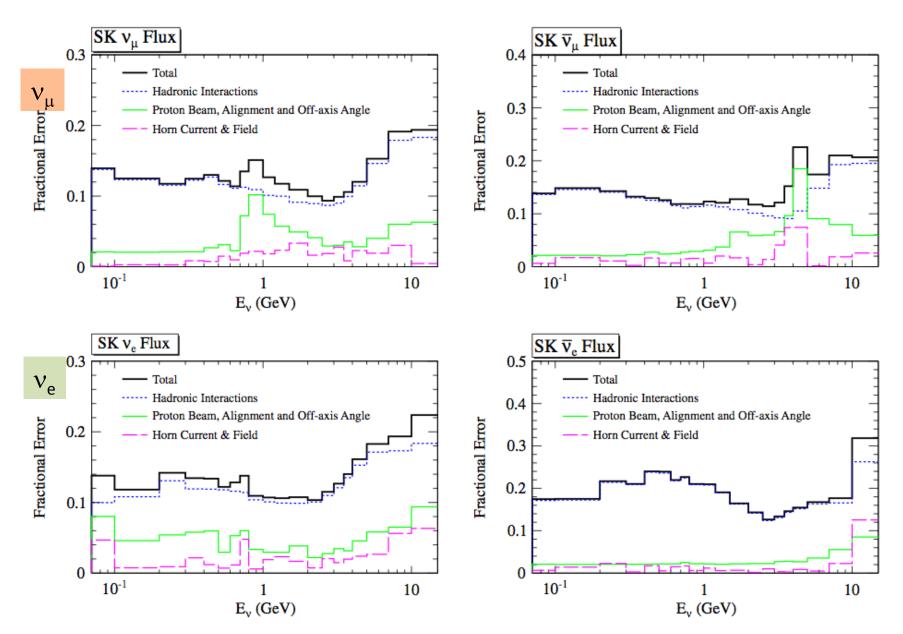
## T2K flux prediction: v and anti-v



## nu flux error at ND280



#### nu flux error at SK



#### Uncertainty: Hadron production dominates

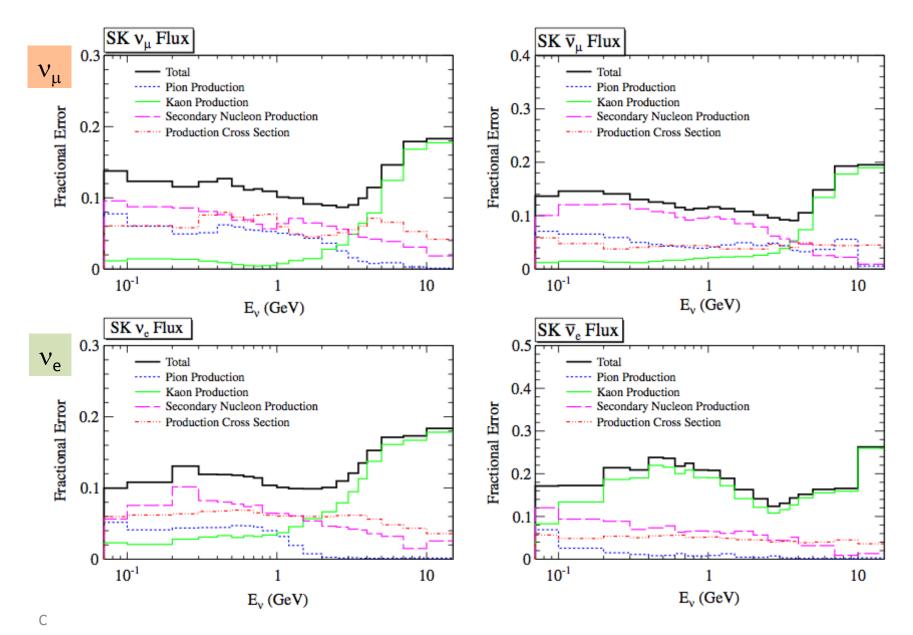
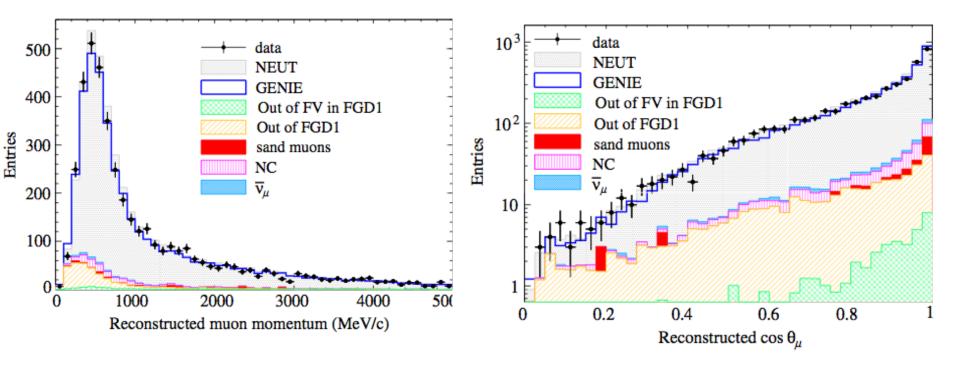


FIG. 38 (color online). Fractional flux error due to hadron production uncertainties.

## T2K $\nu_{\mu}$ CC inclusive



#### PRD D 87, 092003 (2013)

FIG. 9 (color online). The muon momentum and angle distribution for the selected events in the data and MC. The GENIE prediction is shown by the solid blue line, while the backgrounds and signal as derived from NEUT are shown in filled colors.

## T2K fluxaveraged $v_{\mu}$ CC cross section

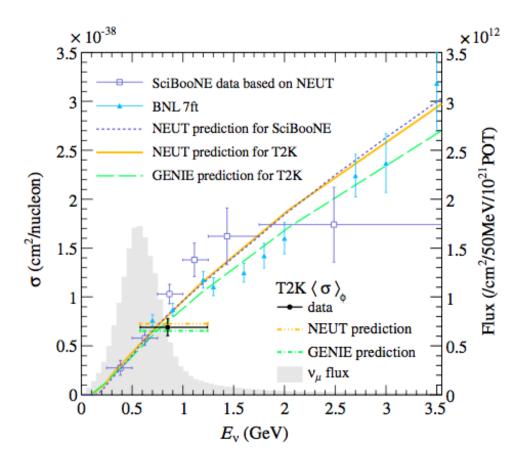
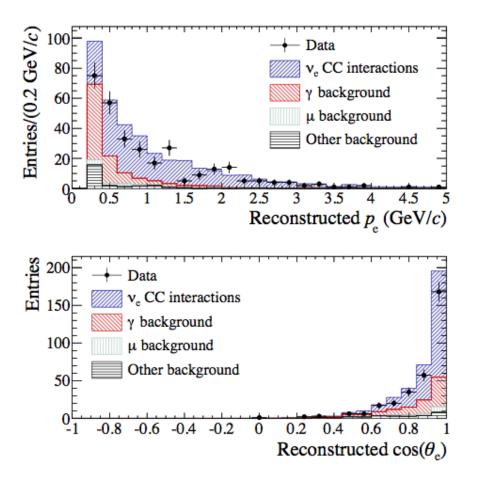


FIG. 13 (color online). The T2K total flux-averaged cross section with the NEUT and the GENIE prediction for T2K and SciBooNE. The T2K data point is placed at the flux mean energy. The vertical error represents the total (statistical and systematic) uncertainty, and the horizontal bar represent 68% of the flux at each side of the mean energy. The T2K flux distribution is shown in grey. The predictions for SciBooNE have been done for a C<sub>8</sub>H<sub>8</sub> target [52] which is comparable to the mixed T2K target. BNL data <sup>Present</sup> new prediction of deuterium [53].

## T2K $\nu_e$ CC



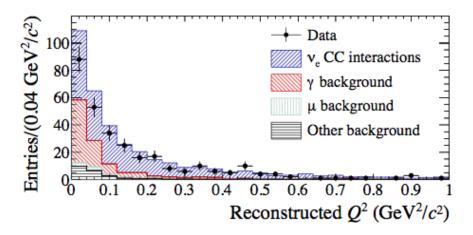


FIG. 1. Reconstructed  $p_e$  (top),  $\cos(\theta_e)$  (middle) and  $Q^2$  (bottom) distributions of  $\nu_e$  event candidates. The NEUT Monte Carlo prediction is separated into the  $\nu_e$  CC interaction signal, background from  $\gamma \rightarrow e^+e^-$  conversions, background from  $\mu^-$  tracks and all other backgrounds. The last bins in the top and bottom plots do not include the overflow of events.

## T2K $\nu_e$ CC

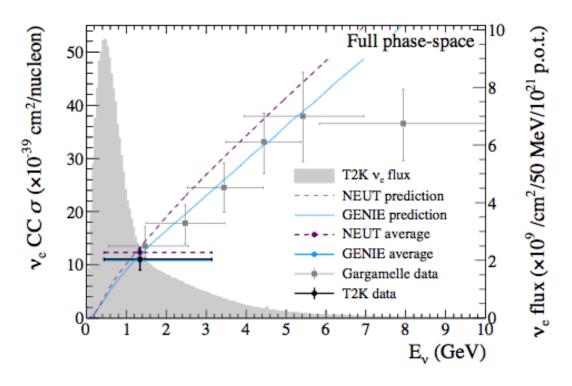


FIG. 3. Total  $\nu_e$  CC inclusive cross-section when unfolding through  $Q^2$ . The T2K data point is placed at the  $\nu_e$  flux mean energy. The vertical error represents the total uncertainty, and the horizontal bar represents 68% of the flux each side of the mean. The T2K flux distribution is shown in grey. The NEUT and GENIE predictions are the total  $\nu_e$  CC inclusive predictions as a function of neutrino energy. The NEUT Pread a GENIE averages are the flux-averaged predictions. The Gargamellet dataais taken from Ref. [9].

1407.7389 315 ve CC interaction candidates are selected, with an expected purity of 65%.

How to deal with ve cross section?



## T2K ne flux is mostly from muon decay

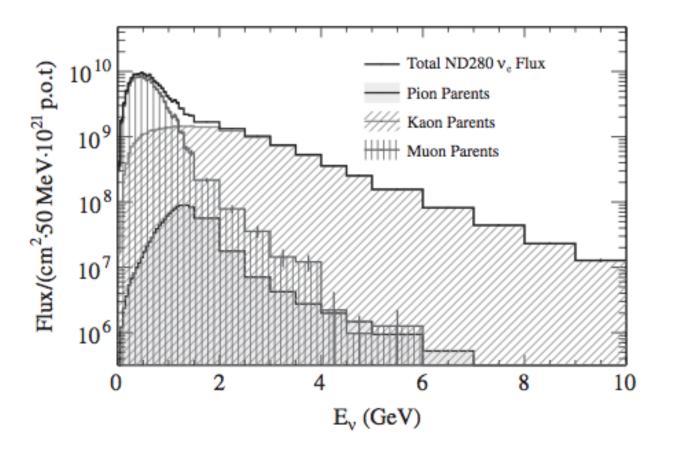


FIG. 1. The prediction of the  $\nu_e$  flux at ND280 broken down by the neutrino parent particle type [4].

TABLE VI: The parameters used to vary the NEUT cross section model along with the values used in the ND280 fit (input value) and uncertainties prior to the ND280 and SK data fits.

Parameter	Input Value	Uncertainty	
$M_A^{QE}$ (GeV)	1.21	0.43	
$x_1^{QE}$	1.00	0.11	
$x_2^{QE}$	1.00	0.30	
$x_3^{ar{Q}E}$	1.00	0.30	
$x_{SF}$	0.0	1.0	
$p_F(^{12}{ m C})~({ m MeV}/c)$	217	30	
$p_F(^{16}{ m O}) ~({ m MeV}/c)$	225	30	
$M_A^{RES}$ (GeV)	1.16	0.11	
$x_1^{CC1\pi}$	1.63	0.43	
$x_2^{CC1\pi}$	1.00	0.40	
$x^{NC1\pi^0}$	1.19	0.43	
$x_{1\pi E_ u}$	$\operatorname{off}$	on	
$W_{ m eff}$	1.0	0.51	
$x_{\pi-less}$	0.2	0.2	
$x^{CCcoh.}$	1.0	1.0	
$x^{NCcoh.}$	1.0	0.3	
$x^{NCother}$	1.0	0.3	
$x_{CCother}$ (GeV)	0.0	0.4	
$x_{ u_e/ u_\mu}$	1.0	0.03	

## $v_e/v_\mu$ determined at 3% error

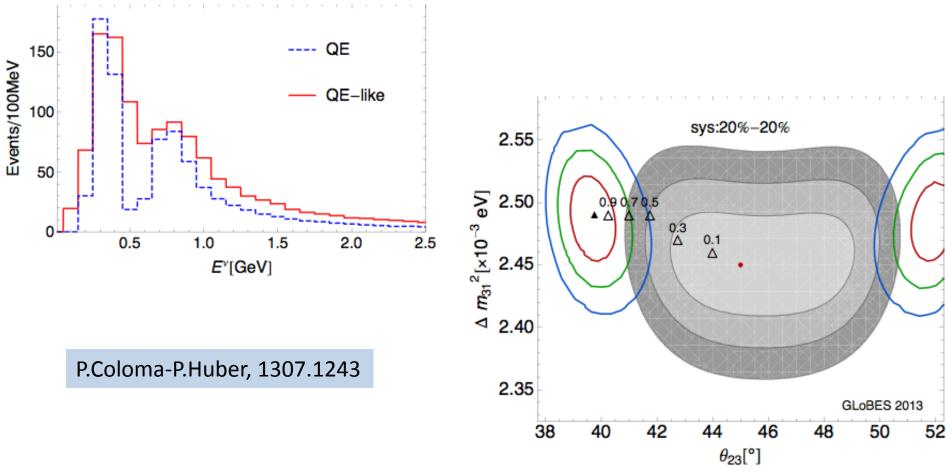
arXiv:1304.0841

# Effect on $\theta_{23}$



Present and Future Neutrino Physics@KITP-Santa Barbara

# If QE and non-QE are not separated appropriately ...



Present and Future Neutrino Physics@KITP-

Santa Barbara



## ArgoNeuT sees nucleon correlations

Present and Future Neutrino Physics@KITP-Santa Barbara

## ArgoNeut saw SRC in nuclei

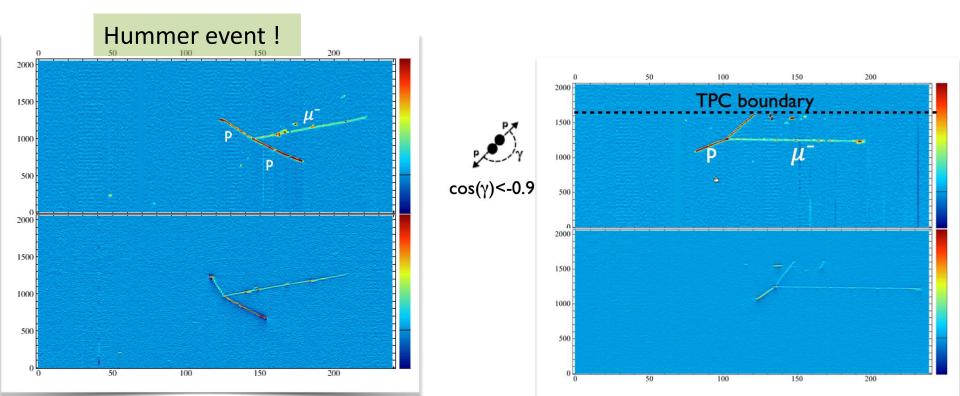
#### 240 Kg active volume

 $47 \times 40 \times 90$  cm<sup>3</sup>, wire spacing 4 mm

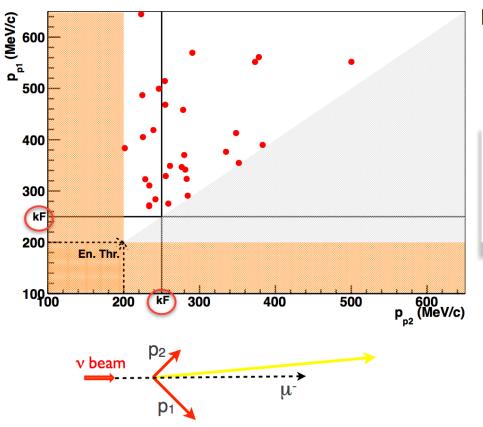
LAr TPC

~7000 CC events

 $(\mu$ -+2p) triple coincidence topology events 30 (19 collected in the anti-neutrino mode run and 11 in the neutrino mode run) fully reconstructed events



## (µ-+2p) data sample

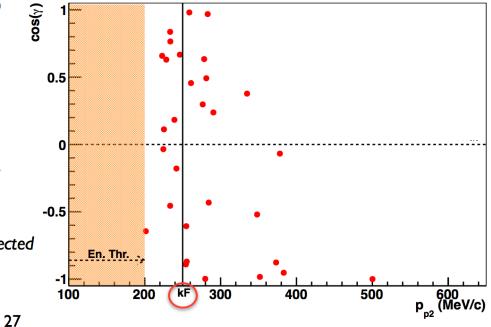


 $\mbox{cos}(\gamma)$  vs momentum of the least energetic proton  $p_{P^2}$  in the pair

 $\gamma$ =angle in space between the two detected proton tracks in the Lab reference frame

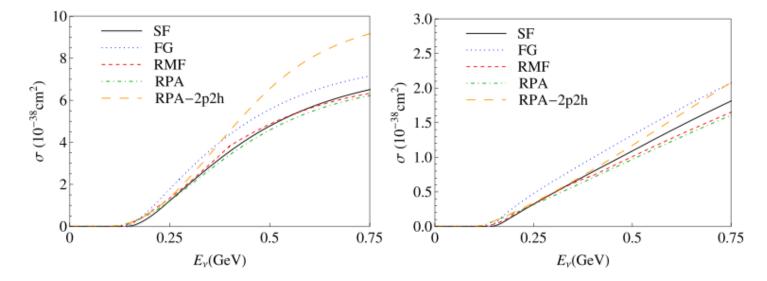
Momentum of the more energetic proton  $\mathbf{p}_{P1}$  in the pair vs. momentum of the other (less energetic) proton  $\mathbf{p}_{P2}$ 

Most of the events (19 out of 30) have both protons above Fermi momentum of the Ar nucleus (k<sub>F</sub>~250 MeV/c)



# Fermi gas model tends to overestimate cross sections

E. Fernandez Martinez, D. Meloni / Physics Letters B 697 (2011) 477-481

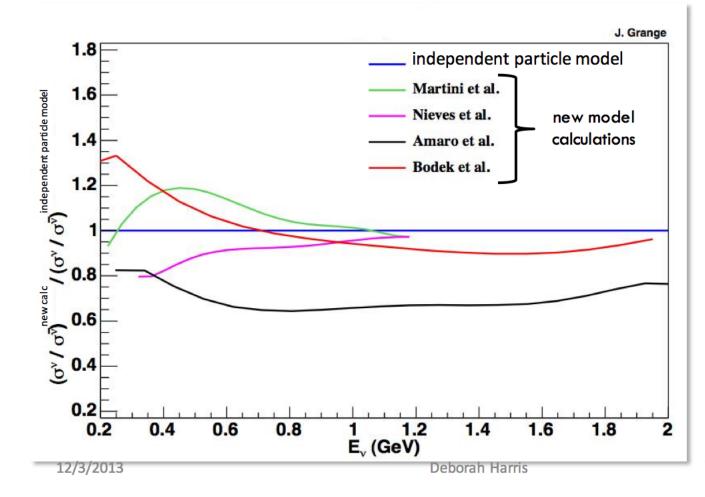


**Fig. 1.** Total charged current QE cross sections for the  $\nu_{\mu}{}^{16}O \rightarrow \mu^{-}X$  (left panel) and  $\bar{\nu}_{\mu}{}^{16}O \rightarrow \mu^{+}X$  (right panel) processes in the energy range  $E_{\nu} \sim [0, 0.75]$  GeV.

Present and Future Neutrino Physics@KITP-Santa Barbara

# What about neutrino/antineutrino cross section ratios?

Different models predict different ratios



And these are  $v_{\mu}$  cross section ratios, Also need this for  $v_{e}$ 

And the theorists are assuming we'll know these ratios (times acceptances) to 1-3% each...

Slide courtesy G. Zeller

## Conclusion

- CP phase effect is a small effect, and so flux and cross section errors must be controlled
- For flux, dedicated hadron production measurement indispensable
- For cross section, near detector can do the job, but the current error is ~10% for  $\nu_{\mu}$  and ~20% for  $\nu_{e}$  at T2K
- Serious effects ongoing to reduce the errors of v flux and v cross sections

replica target soon!