# Thermalization of boost-invariant plasma, AdS/CFT and numerical relativity

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## Outline

Key question

- 2 AdS/CFT, hydrodynamics and nonequilibrium processes
- Boost-invariant flow
- The AdS/CFT approach to evolving plasma

#### 5 Numerical relativity setup

- Initial conditions
- The metric ansatz and numerical formalism

# 🜀 Main results

- Nonequilibrium vs. hydrodynamic behaviour
- Entropy
- Characteristics of thermalization

# Conclusions

Point of reference: heavy-ion collision at RHIC/LHC:



Collision Fireball

isotropization thermalization

expansion

freezout hadronization

#### Key question:

Understand the features of (early) thermalization for an evolving (*boost-invariant*) plasma system

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- At weak coupling the obvious definition would be to require thermal momentum distributions for quarks and gluons...
- At strong coupling, the picture of a gas of gluons is not really valid

   alternatively require that observables such as 2-point functions/spatial
   Wilson loops/ entanglement entropy are the same as for a thermal system...
- This is very good for studying relaxation processes where the final state is some uniform static plasma system this is not so for the plasma undergoing expansion
- For an expanding plasma fireball we need *local* equilibrium bilocal probes get contaminated by collective flow
- We adopt an *operational* definition of thermalization the point when plasma starts being describable by (viscous) hydrodynamics.

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- Full nonlinear hydrodynamic equations follow now from  $\partial_{\mu}T^{\mu
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#### Linearized hydrodynamics

- Look at small disturbances of the uniform static plasma...
- If  $T_{\mu\nu}$  is described by (1<sup>st</sup> order viscous) hydrodynamics then one can derive dispersion relation of long wavelength modes from hydrodynamic equations: shear modes:

$$\omega_{shear} = -i\frac{\eta}{E+p}k^2$$

sound modes:

$$\omega_{sound} = \frac{1}{\sqrt{3}}k - i\frac{2}{3}\frac{\eta}{E+p}k^2$$

- If we were to include terms in  $T_{\mu\nu}$  with more derivatives (higher order viscous hydrodynamics), we would get terms with higher powers of k in the dispersion relations...
- Hypothetical resummed *all-order* hydrodynamics would predict the full dispersion relation for these modes  $\omega_{shear}(k)$ ,  $\omega_{sound}(k)$

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- The uniform static plasma system is described as a static planar black hole
- Small disturbances of the uniform static plasma ≡ small perturbations of the black hole metric (≡ quasinormal modes (QNM))

$$g^{5D}_{lphaeta} = g^{5D, black\ hole}_{lphaeta} + \delta g^{5D}_{lphaeta}(z) e^{-i\omega t + ikx}$$

• Dispersion relation fixed by linearized Einstein's equations. Results for the sound channel

#### from Kovtun, Starinets hep-th/0506184

- This is equivalent to summing contributions from *all-order* viscous hydrodynamics
- But, **in addition**, there is an infinite set of higher QNM effective degrees of freedom not contained in the hydrodynamic description at all!

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- But, **in addition**, there is an infinite set of higher QNM effective degrees of freedom not contained in the hydrodynamic description at all!
- contain all-order viscous hydrodynamic modes (with specific values of all transport coefficients)
- in addition contain the dynamics of genuine nonhydrodynamical modes
- incorporate their interactions in a fully nonlinear (and unique) way

## **Consequence:**

Einstein's equations can serve to study nonequilibrium processes in strongly coupled  $\mathcal{N}=4$  SYM and are an effective tool for exploring physics beyond hydrodynamics

# Question:

In the case of boost-invariant plasma expansion can we unambigously determine i) whether these nonhydrodynamical modes are really important or

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Assume a flow that is invariant under longitudinal boosts and does not depend on the transverse coordinates.



- In a conformal theory,  $T^{\mu}_{\mu} = 0$  and  $\partial_{\mu}T^{\mu\nu} = 0$  determine, under the above assumptions, the energy-momentum tensor completely in terms of a single function  $\varepsilon(\tau)$ , the energy density at mid-rapidity.
- The longitudinal and transverse pressures are then given by

$$p_L = -\varepsilon - \tau \frac{d}{d\tau} \varepsilon$$
 and  $p_T = \varepsilon + \frac{1}{2} \tau \frac{d}{d\tau} \varepsilon$ .

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**Method:** Describe the time dependent evolving strongly coupled plasma system through a dual 5D geometry — given e.g. by

$$ds^{2} = \frac{g_{\mu\nu}(x^{\rho},z)dx^{\mu}dx^{\nu} + dz^{2}}{z^{2}} \equiv g_{\alpha\beta}^{5D}dx^{\alpha}dx^{\beta}$$

i) use Einstein's equations for the time evolution

$$R_{lphaeta}-rac{1}{2}g^{5D}_{lphaeta}R-6\,g^{5D}_{lphaeta}=0$$

ii) read off  $\langle T_{\mu
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Chesler and Yaffe adopted a different way of preparing the initial state:

- () Start from the vacuum of  $\mathcal{N} = 4$  SYM (no plasma)
- Change the physical 4D metric of gauge theory spacetime in a time-dependent manner
- This will produce some nonequilibrium state
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- We want to study the evolution right from  $\tau = 0$  with energy-momentum conservation satisified throughout the evolution
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- Note that the initial hypersurface au=0 is partly light-like...
- The initial conditions are determined in terms of a *single* function, say  $c_0(z)$ .  $a_0(z) = b_0(z)$  are determined through a constraint equation.
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#### The metric ansatz and numerical formalism



black line - dynamical horizon, arrows - null geodesics, colors represent curvature

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• We set the lapse to always vanish at the boundary in the bulk

• Consequently, we set the (nondynamical) function a(u) to

$$a(u) = \cos\left(\frac{\pi}{2}\frac{u}{u_0}\right)$$

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- We use Chebyshev spectral methods for the spatial derivatives (hence very strong sensitivity to boundary conditions)
- We need very accurate spatial derivatives at the boundary in order to reliably extract the physical energy density from the numerical geometry
- For the time evolution we use an adaptive 8<sup>th</sup>/9<sup>th</sup>-order Runge-Kutta method (gnu scientific library)

- We monitor ADM constraints during evolution
- <sup>(2)</sup> The energy density  $\varepsilon(\tau)$  extracted from simulations made with different lapses/cut-offs for the same initial condition should coincide
- (a) We compare the numerical  $\varepsilon(\tau)$  with the power series solution in its region of convergence

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#### Results

- We have considered 20+9 initial conditions, each given by a choice of the metric coefficient c(τ = 0, u).
- We have chosen quite different looking profiles e.g.

$$c_{1}(u) = \cosh u$$

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- When and how does the transition to hydrodynamics (≡ thermalization/ isotropization) occur?
- To what extent would higher order (even all-order) viscous hydrodynamics explain plasma dynamics or do we need to incorporate genuine nonhydrodynamic degrees of freedom in the far from equilibrium regime
- Does there exist some physical characterization of the initial state which determines the main features of thermalization and subsequent evolution?
- What is the produced entropy from  $\tau = 0$  to  $\tau = \infty$  (asymptotically perfect fluid regime)

$$\langle T_{\tau\tau} \rangle \equiv \varepsilon(\tau) \equiv N_c^2 \cdot \frac{3}{8} \pi^2 \cdot T_{eff}^4$$

# Key physical questions

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$$\frac{F_{hydro}(w)}{w} = \frac{2}{3} + \frac{1}{9\pi w} + \frac{1 - \log 2}{27\pi^2 w^2} + \frac{15 - 2\pi^2 - 45\log 2 + 24\log^2 2}{972\pi^3 w^3} + \dots$$

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- Genuine nonequilibrium dynamics would, in contrast, lead to several curves...

A plot of F(w)/w versus w for various initial data

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• An observable sensitive to the details of the dissipative dynamics (e.g. hydrodynamics) is the pressure anisotropy

$$\Delta p_L \equiv 1 - rac{p_L}{arepsilon/3} = 12F(w) - 8$$

• For a perfect fluid  $\Delta p_L \equiv 0$ . For a sample initial profile we get

- For  $w = T_{eff} \cdot \tau > 0.63$  we get a very good agreement with viscous hydrodynamics
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- We want to study systematically the properties of the plasma at the point when the dynamics becomes describable by viscous hydrodynamics...
- We adopted a numerical criterion for thermalization

$$\left|\frac{\tau \frac{d}{d\tau}w}{F_{hydro}^{3^{rd} \text{ order}}(w)} - 1\right| < 0.005$$

- We looked at the following features of thermalization:
  - ]) the dimensionless quantity  $w=T_{eff}\cdot au$
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- N.B. sample initial conditions for hydrodynamics at RHIC ( $\tau_0 = 0.25 \text{ fm}$ ,  $T_0 = 500 \text{ MeV}$ ) assumed in [Broniowski, Chojnacki, Florkowski, Kisiel] correspond to w = 0.63
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- It is interesting to consider the ratio of the temperature at thermalization to the initial effective temperature
- This gives information on which part of the cooling process occurs in the far from equilibrium regime and which part occurs during the hydrodynamic evolution

- Note: for initial profiles with large  $s_{initial}$ , the energy density initially rises and only then falls  $\longrightarrow$  even for  $T_{th}/T_{eff}(0) \sim 1$  there is still sizable nonequilibrium evolution
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- For  $w = T_{th} \cdot \tau_{th} > 0.7$  we observe hydrodynamic behaviour but with sizeable pressure anisotropy (described wholly by viscous hydrodynamics)
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