

Partons in Phase Space

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Nonequilibrium Phenomena
in Cosmology and Particle Physics
Kavli Institute for Theoretical Physics
Santa Barbara, CA
Feb 25 - 29, 2008



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2 γ in Phase Space

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- γ Phase-Space Distributions

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- Mass & Medium
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Introduction

Large systems such as Au+Au at RHIC: final state a result of history where different processes are localized in phase space.

?Can such a perspective be developed for elementary processes?

?Elementary processes embedded in a large system?

- Perturbative processes
- Partons: folding in cross sections
- Generalized Wigner representation:

$$A(x, p) = \int d^4 q A\left(p + \frac{q}{2}\right) A^*\left(p - \frac{q}{2}\right) e^{-iqx}$$

where $A(p)$ is an amplitude



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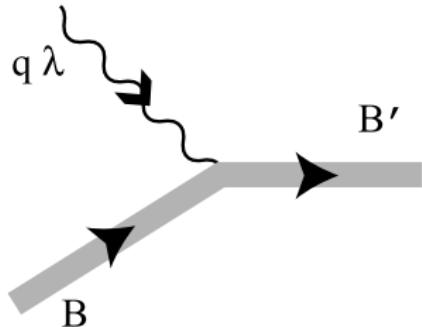
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γ Absorption



Transition Amplitude

$$S_{\gamma B \rightarrow B'}$$

$$= \int d^4x \langle 0 | A^\mu(x) | \vec{q}, \lambda \rangle \langle B' | j_\mu(x) | B \rangle$$

$$\begin{aligned} \text{Transition No} &= |S_{\gamma B \rightarrow B'}|^2 = \int d^4x d^4k \delta^{(4)}(q - k) \epsilon_\mu \epsilon_\nu^* J_{BB'}^{\mu\nu}(x k) \\ &\equiv \int d^4x \mathcal{W}_{\gamma B \rightarrow B'}(x q) \end{aligned}$$

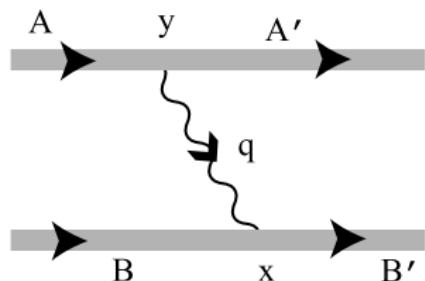
where

$$J_{BB'}^{\mu\nu}(x k) \equiv \int d^4\tilde{k} e^{-i\tilde{k}x} \langle B | j^{\dagger\nu}(k - \tilde{k}/2) | B' \rangle \langle B' | j^\mu(k + \tilde{k}/2) | B \rangle$$

⇒ Wigner transform of transition current
 \equiv Phase-Space density for γ absorption/emission



γ -Exchange



Amplitude

$$S_{AB \rightarrow A'B'} = \int d^4x d^4y \langle B'|j^{B\nu}(x)|B\rangle \\ \times D_{\nu\mu}(x-y) \langle A'|j^{A\mu}(y)|A\rangle$$

Transition No = $|S_{AB \rightarrow A'B'}|^2$

$$= \int d^4y d^4x d^4q J_{AA'}^{\mu\nu}(y|q) D_{\mu\nu\mu'\nu'}(y-x, q) J_{BB'}^{\mu'\nu'}(x|q)$$

where the Wigner transform of a **propagator** is

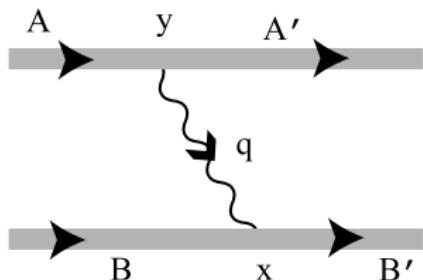
$$D_{\mu\nu\mu'\nu'}(x, q) = \int d^4\tilde{x} e^{iq\tilde{x}} D_{\mu\nu}(x + \tilde{x}/2) D_{\mu'\nu'}(x - \tilde{x}/2)$$

In Phase Space:

- A emits γ of momentum q at y
- γ propagates from y to x
- B absorbs γ at x



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Towards Weizsäcker-Williams Interpretation

?Vector potential of object A?

$$A^\mu(x) = \int d^4y D_{\mu\nu}(x-y) J_A^\nu(y)$$

Wigner transform of the vector potential:

$$\begin{aligned} A_{\mu\nu}(x q) &= \int d^4\tilde{x} e^{i\tilde{x}q} A_\mu(x + \tilde{x}/2) A_\nu(x - \tilde{x}/2) \\ &= \int d^4y D_{\mu\nu\mu'\nu'}(x-y, q) J_A^{\mu'\nu'}(y, q) \end{aligned}$$

Source-Propagator Form: the current of A creates γ with momentum q at position y and the γ propagates to x

For γ -exchange, the WW interpretation applied in phase space:

$$\begin{aligned} \text{Transition No} &= |S_{AB \rightarrow A'B'}|^2 = \int d^4x d^4q A_{\mu\nu}(xq) J_B^{\mu\nu}(xq) \\ &\equiv \int d^4x d^4q \frac{d n_\gamma}{d^3x d^4q} W_{\gamma B \rightarrow B'}(xq) \end{aligned}$$

In standard WW: γ distribution in energy *only*



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Issue of Current in WW

In standard WW approximation, A -body current classical:

$$\begin{aligned} j_\mu(x) &= Z e v_\mu \delta^3(\vec{x} - x^0 \vec{v}) \\ \text{Then } J_{\text{classical}}^{\mu\nu}(xq) &= \int d^4 \tilde{x} e^{i\tilde{x}q} j_\mu(x + \tilde{x}/2) j_\nu^\dagger(x - \tilde{x}/2) \\ &= Z^2 e^2 v_\mu v_\nu \delta(q v) \delta^{(3)}(\vec{x} - x^0 \vec{v}) \end{aligned}$$

$\delta(q v)$ stems from current conservation, photons purely spacelike in emitter frame; $q_0 = q_L$, $v_L \simeq q_L$ in lab frame

Us: good approximation for $q \ll p$, when one can construct an A -packet such that $q \ll (\Delta x_A)^{-1} \ll p_A$

The current of body A in terms of Wigner functions:

$$\begin{aligned} J_{AA'}^{\mu\nu}(xq) &= \int d^4 \tilde{q} e^{-i\tilde{q}x} \langle A | j^\dagger{}^\nu(q - \tilde{q}/2) | A' \rangle \langle A' | j^\mu(q + \tilde{q}/2) | A \rangle \\ &= \int d^4 p_i d^4 p_f \delta^{(4)}(p_i - p_f - q) \Gamma^{\mu\nu}(q, p_i, p_f) f_A(x p_i) f_{A'}(x p_f) \end{aligned}$$

For $q \ll (\Delta x_A)^{-1} \ll p_A$: $\Gamma_{\mu\nu}(q, p_i, p_f) \simeq Z^2 e^2 (p_i + p_f)_\mu (p_i + p_f)_\nu$
 \Rightarrow If packet localized on q^{-1} scale, we could take $J^{\mu\nu} \simeq J_{\text{classical}}^{\mu\nu}$



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Propagator in γ -Exchange

$$A^\mu(x) = \int d^4y D_{\mu\nu}(x-y) J_A^\nu(y)$$

Wigner transform of what propagator ?

→ Feynman doable analytically but acausal!

→ Retarded – the same physics results for the same asymptotic states but simpler to interpret

Scalar propagators:

$$G^+(p) = \frac{1}{p^2 - m^2 + i\epsilon p_0} \quad G^c(p) = \frac{1}{p^2 - m^2 + i\epsilon}$$

Propagator in phase space:

$$G(x, q) = \int d^4\tilde{x} e^{iq\tilde{x}} G(x + \tilde{x}/2) G(x - \tilde{x}/2)$$

Massless scalar

$$G^+(x p) = \theta(x_0) \theta(x^2) \theta(\lambda^2) \frac{\sin(2\sqrt{\lambda^2})}{\sqrt{\lambda^2}}$$

$$\text{where } \lambda^2 = (x p)^2 - x^2 p^2. \rightarrow D_{\mu\nu\mu'\nu'}^+(x q) = g_{\mu\nu} g_{\mu'\nu'} G^+(x q)$$



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Features of Retarded Phase-Space Propagator

- causal, propagation inside light cone only
- ptcles do not propagate beyond $0 \leq \lambda^2 \lesssim 1$

q^2 -Case	Special Frame	Moving Frame $(q^0, q^L, \vec{0}^T)$
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0		$x^0 - x^L \lesssim \frac{1}{ q^0 }, \vec{x}^T \lesssim \sqrt{\frac{x^0}{ q^0 }}$

Off-shell propagation limited, from emission point,
by the inverse of energy, momentum or mass

On-shell propagation limited, from the classical trajectory,
by the inverse of energy

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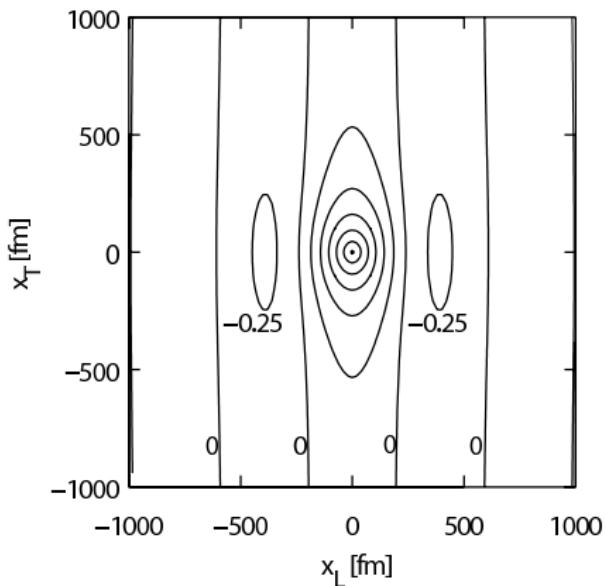
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γ Phase-Space Distribution for Static Charge

Putting current + propagation together yields γ distribution:

$$A^{\mu\nu}(x q) = Z^2 e^2 v^\mu v^\nu \frac{\delta(qv)}{\sqrt{-q^2}} \mathcal{A}$$

where \mathcal{A} is dimensionless, semi-analytic and gives shape:

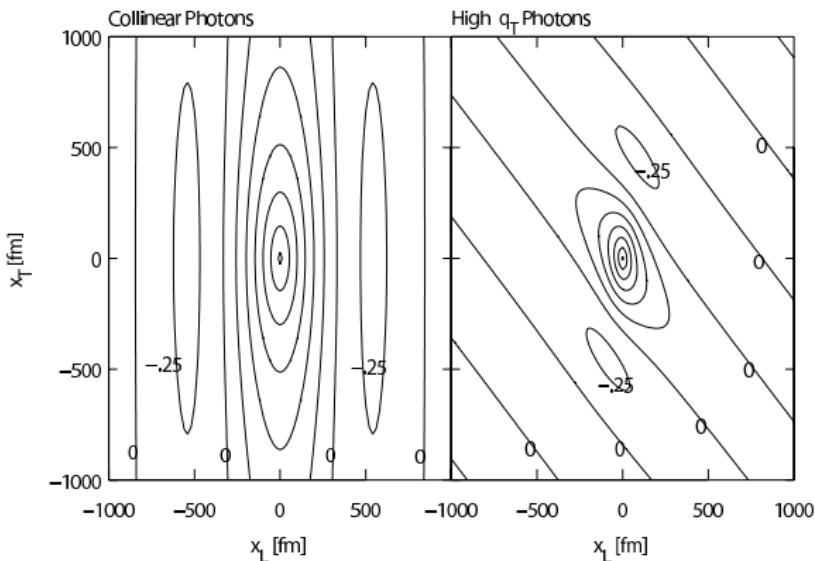


photon 4-momentum
 $q_\mu = (0, 0.788, \vec{0}_T)$ MeV/c

Lorentz contraction
in the direction of
photon momentum

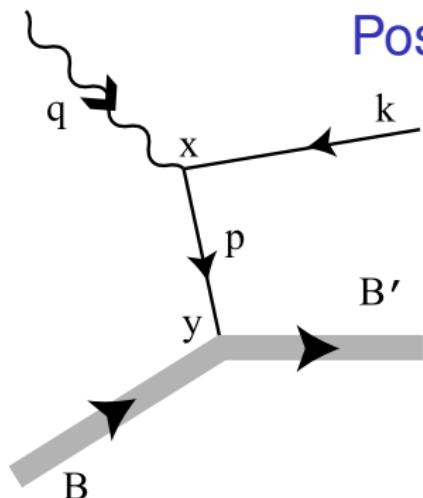
γ Phase-Space Distribution for Moving Charge

Charge moving at $\vec{v} = (0.9c, \vec{0}_T)$



Photon $q^\mu = (m_e, m_e/v_L, \vec{0}_T)$ & $q^\mu = (m_e, m_e/v_L, 0.56 \text{ MeV}/c, 0)$
 Again contraction in the *photon direction*





Positron Production

$$\gamma + B \rightarrow \bar{e} + B'$$

Process amplitude

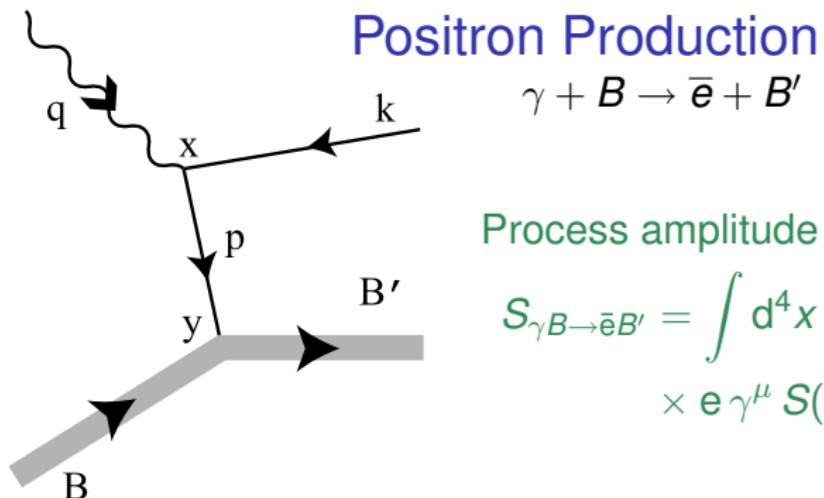
$$S_{\gamma B \rightarrow \bar{e} B'} = \int d^4x d^4y A_\mu(x) \psi_{\bar{e}}(x s) \times e \gamma^\mu S(x - y) \mathcal{V}_{B \bar{e} \rightarrow B'}(y)$$

$$\text{Transition No} = |S_{\gamma B \rightarrow \bar{e} B'}|^2 = e^2 \int d^4y d^4p d^4x d^4q \frac{d^3 k}{|k^0|} A_{\mu\nu}(x q) \times G(y - x, p) \delta^{(4)}(k + q - p) \text{Tr}\{\mathcal{V}_{B \bar{e} \rightarrow B'}(y, p) (\not{p} - \dots)\}$$

Distribution of virtual electrons for a photon:

$$\frac{dn_{\bar{e}}}{d^3y d^3p dp^2} \propto e^2 \int d^4x \frac{d^3 k}{|k^0|} G(y - x, p) A_{\mu\nu}(x, p - k)$$





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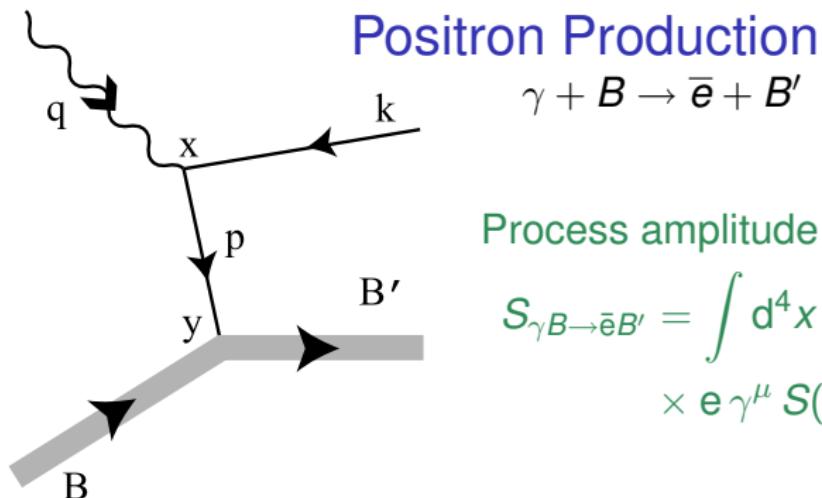
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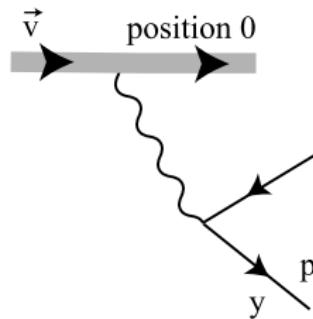
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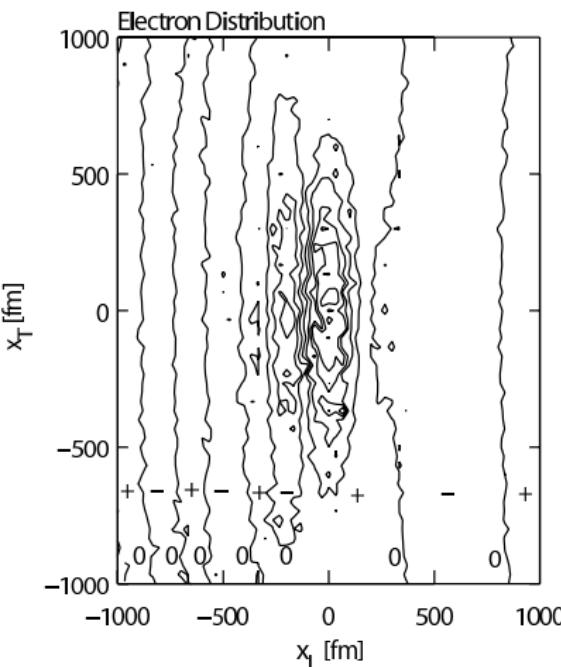
Virtual Electron Distribution

Point charge, moving at $0.9c$ to the right, is the source of photons that are the source of virtual electrons

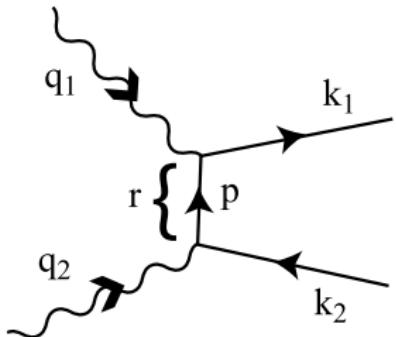


Electron momentum
 $p^\mu = (2.0, 2.05, \vec{0}) \text{ MeV/c}$

Forward-backward
asymmetry due to \bar{e} recoil



Interference in 2- γ Interaction



$$\text{Transition No} = |S_{12 \rightarrow 1'2'e\bar{e}}|^2 = \int d^4R d^4r \dots$$

$$\begin{aligned}
 & \times G(r p) \delta^{(4)}(q_1 + q_2 - k_1 + k_2) \\
 & \times \{ A_1^{\mu\mu'}(R - r/2, q_1) A_2^{\nu\nu'}(R + r/2, q_2) \\
 & \times \delta^{(4)}(q_1 + p - k_1) + A_1^{\nu\nu'}(R + r/2, q_1) \\
 & \times A_2^{\mu\mu'}(R - r/2, q_2) \delta^{(4)}(q_1 + k_2 - p) \\
 & + \int d^4\tilde{r} 2 \cos [\tilde{r} \cdot (-p + \frac{k_1 + k_2}{2}) - r \cdot (q_1 - q_2)] \\
 & \times A_1^{\nu\mu'}(R - \tilde{r}/4, q_1) A_2^{\mu\nu'}(R + \tilde{r}/4, q_2) \}
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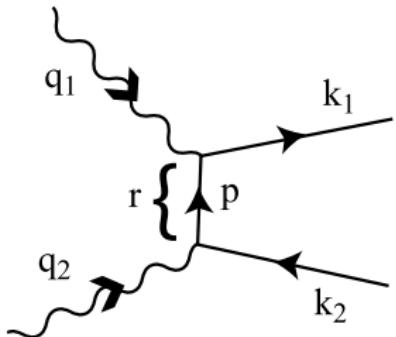
2 direct terms + HBT-like interference term

The direct terms factorize into
e-density \times absorption.

The interference term does not.

⇒ Scale separation needed to achieve a factorization

Interference in 2- γ Interaction



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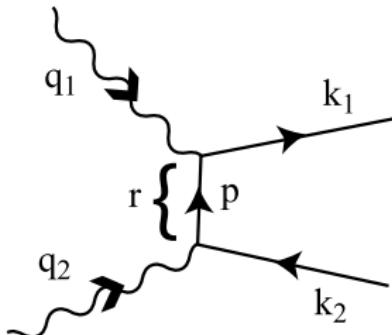
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Interference in 2- γ Interaction



$$\text{Transition No} = |S_{12 \rightarrow 1'2'e\bar{e}}|^2 = \int d^4R d^4r \dots$$

$$\begin{aligned} & \times G(r p) \delta^{(4)}(q_1 + q_2 - k_1 + k_2) \\ & \times \{ A_1^{\mu\mu'}(R - r/2, q_1) A_2^{\nu\nu'}(R + r/2, q_2) \\ & \times \delta^{(4)}(q_1 + p - k_1) + A_1^{\nu\nu'}(R + r/2, q_1) \\ & \times A_2^{\mu\mu'}(R - r/2, q_2) \delta^{(4)}(q_1 + k_2 - p) \\ & + \int d^4\tilde{r} 2 \cos [\tilde{r} \cdot (-p + \frac{k_1 + k_2}{2}) - r \cdot (q_1 - q_2)] \\ & \times A_1^{\nu\mu'}(R - \tilde{r}/4, q_1) A_2^{\mu\nu'}(R + \tilde{r}/4, q_2) \} \end{aligned}$$

2 direct terms + HBT-like interference term

The direct terms factorize into
e-density \times absorption.

The interference term **does not**.

⇒ Scale separation needed to achieve a factorization.



Propagation with Mass and in Medium

?Massive propagator?

$$G^+(x p) = G_0^+(x p) - \frac{2m}{\pi} \theta(x^0) \theta(x^2)$$

$$\times \int_0^{\sqrt{x^2}} d\xi J_1(2m\xi) \frac{\sin(2\sqrt{\lambda^2 + \xi^2 p^2})}{\sqrt{\lambda^2 + \xi^2 p^2}}$$

G_0^+ – massless propagator; the correction removes pieces close to the light cone...

For an on-shell ptcle, as before $|x_0 - \vec{v} \cdot \vec{x}| \lesssim \frac{1}{|p_0|}$

In medium, fluctuation-dissipation theorem: $G^{\gtrless} = G^+ \Sigma^{\gtrless} G^-$

In Wigner representation:

$$\begin{aligned} G^{\gtrless}(x p) &= \int d^4y d^4p' G^+(x p; y p') \Sigma^{\gtrless}(y p') \\ &\simeq \int d^4y G^+(x - y, p) \Sigma^{\gtrless}(y p') \end{aligned}$$

G^{\gtrless} – density, Σ^{\gtrless} – source

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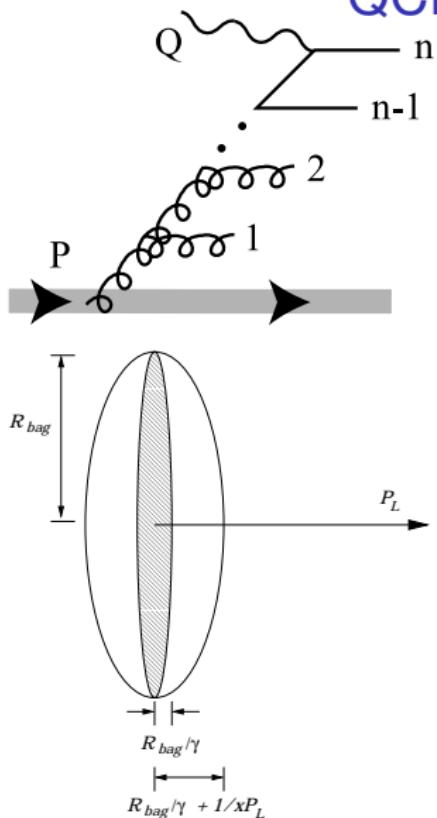
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QCD DGLAP Cloud



Large Q^2 & $-q_1^2 \ll -q_2^2 \dots \ll -q_n^2$
increasing virtuality

decreasing portion of beam momentum

$$1 \leq x_1 \leq \dots \leq x_{n-1} \leq x_n$$

Each parton a source for the subsequent one. Propagation determines size of virtual cloud.

First source localized to $(\frac{R_{bag}}{\gamma}, R_{\perp bag})$.

Propagation increases R_{\perp} by $\sim \frac{1}{\sqrt{-q_n^2}}$.

Because of large virtuality $R_n^{\perp} \sim R_{bag}^{\perp}$.

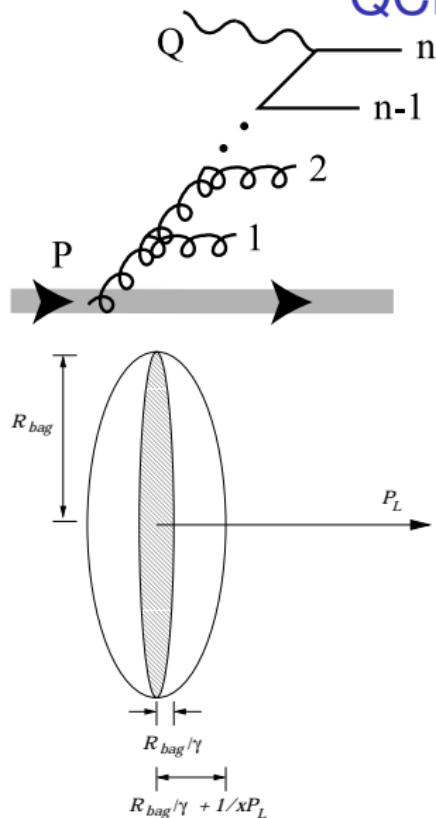
In \parallel direction size increases by $\frac{1}{|q_{n0}|}$,

$$\text{i.e. } R_n^{\parallel} \sim \frac{R_{bag}}{\gamma} + \frac{1}{x_n P}$$

If $\frac{1}{x_n P} > \frac{R_{nuc}}{\gamma}$ partons delocalized beyond the nucleus...



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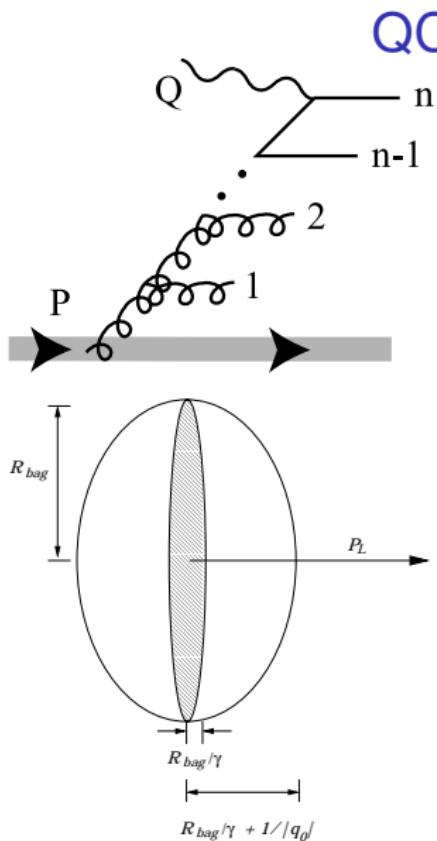
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QCD BFKL Cloud

$1 \gg x_1 \gg \dots \gg x_{n-1} \gg x_n$
strongly ordered low- x

No constraints on q_{\perp}^2 ;
in practice $q_{\perp 1}^2, \dots, q_{\perp n}^2 \gg 1/R_{\text{bag}}$.

If $\frac{1}{x_n P} > \frac{R_{\text{nuc}}}{\gamma}$ partons delocalized beyond the nucleus.

Small- x partons can travel beyond the longitudinal nuclear size and then see color charge of any other nucleon in a longitudinal tube centered on the parent nucleon.

⇒ Support for McLerran-Venugopalan



Summary

- Without interference, probabilities from tree-type diagrams can be expressed in terms of transparent phase-space convolutions of source-propagator form.
- Massless phase-space propagators in analytic form; massive involve 1D integral.
- Interferences produce HBT-type terms in probability.
- Phase-space propagation is causal for the retarded propagators. Deviations from classical motion are limited by energy, momentum or mass.
- Interference terms will get suppressed when there is space-momentum scale separation and that is the case in different limits where the partonic model is employed.
- Wee partons extend farthest from a nucleus and should be first to interact.



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