

Comparison of Boltzmann Kinetics with Quantum Dynamics for Relativistic Quantum Fields

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Nonequilibrium Phenomena in Cosmology and Particle Physics

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Boltzmann
Kinetics
vs. Quantum
Dynamics

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Motivations for going into the subject

The situation

- Many interesting phenomena in particle physics and cosmology require the description of **systems out of thermal equilibrium**.
- Very often, such nonequilibrium situations are treated by means of (approximations to) **Boltzmann equations**.
- However, Boltzmann equations are only a **classical approximation to** the quantum thermalization process described by **Kadanoff-Baym equations**.

An obvious question

How reliable are Boltzmann equations as **compared to** Kadanoff-Baym equations?

Boltzmann Equation

for a **spatially homogeneous** system in the framework of a **real scalar Φ^4** quantum field theory:

$$\partial_t n(t, \mathbf{k}) = \frac{\lambda^2 \pi}{48} \int \frac{d^3 p}{(2\pi)^3} \int \frac{d^3 q}{(2\pi)^3} \int d^3 r \left[\frac{1}{E_k E_p E_q E_r} \right. \\ \times \delta(\mathbf{k} + \mathbf{p} - \mathbf{q} - \mathbf{r}) \delta(E_k + E_p - E_q - E_r) \\ \times \left. \left(\underbrace{(1 + n_k)(1 + n_p) n_q n_r}_{\text{gain term}} - \underbrace{n_k n_p (1 + n_q)(1 + n_r)}_{\text{loss term}} \right) \right]$$

Momentum conservation Energy conservation

Isotropy: 9 dimensional integral \implies 2 dimensional integral.
Important for numerics! [Dolgov, Hansen, Semikoz (1997)]

Complete Schwinger-Keldysh Propagator

Definition

$$G(x, y) = \langle T_{\mathcal{C}} \{ \Phi(x) \Phi(y) \} \rangle$$

The index \mathcal{C} denotes time ordering along the closed Schwinger-Keldysh real-time contour.

Decomposition [Aarts, Berges (2001)]

$$G(x, y) = G_F(x, y) - \frac{i}{2} \text{sign}_{\mathcal{C}}(x^0 - y^0) G_{\rho}(x, y)$$

- **Statistical propagator** \implies effective particle number
- **Spectral function** \implies thermal mass, decay width

Effective Energy and Particle Number Densities

Free-field ansatz [Berges (2002)]

Effective kinetic energy density:

$$\omega^2(t, k) = \left(\frac{\partial_{x^0} \partial_{y^0} G_F(x^0, y^0, k)}{G_F(x^0, y^0, k)} \right)_{x^0=y^0=t}$$

Effective particle number density:

$$n(t, k) = \omega(t, k) G_F(t, t, k) - \frac{1}{2}$$

Advantages of these definitions

- They furnish a particle number density which thermalizes.
- They do not rely on any quasi-particle assumption.
- They comprise conserved charges, if present in the theory.

Kadanoff-Baym Equations

for a **spatially homogeneous** and **isotropic** system in the framework of a **real scalar Φ^4** quantum field theory:

$$\begin{aligned} & \left[\partial_{x^0}^2 + k^2 + M^2(x^0) \right] G_F(x^0, y^0, k) \\ &= \int_0^{y^0} dz^0 \Pi_F(x^0, z^0, k) G_\varrho(z^0, y^0, k) \\ & \quad - \int_0^{x^0} dz^0 \Pi_\varrho(x^0, z^0, k) G_F(z^0, y^0, k) \end{aligned}$$

Effective mass: $M^2(x^0) = m^2 + \text{---} \bullet \text{---} \text{---} \bigcirc \text{---} \text{---} \bullet \text{---} \text{---}$

Nonlocal self-energy: $\Pi(x^0, z^0, k) = \text{---} \bullet \text{---} \bigcirc \text{---} \bullet \text{---}$

Internal lines represent the complete Schwinger-Keldysh propagator!

Initial Conditions

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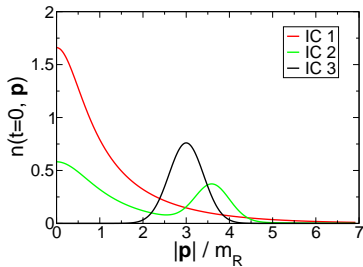
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- All initial conditions correspond to the **same (conserved) average energy density**.
- The initial conditions **IC1** and **IC2** correspond to the **same initial total particle number**.

Universality

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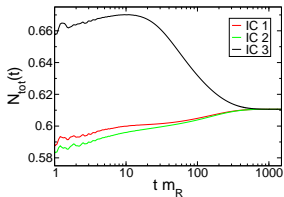
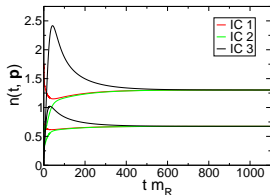
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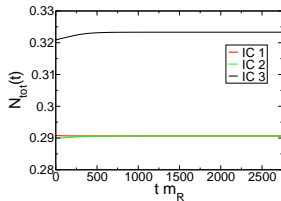
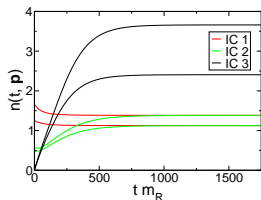


Full Universality

Evolution of
the particle
number
densities

Evolution of
the total
particle
numbers

Boltzmann



Restricted Universality

Chemical Equilibration

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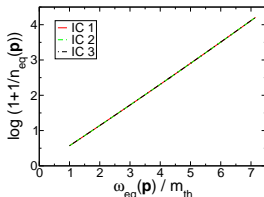
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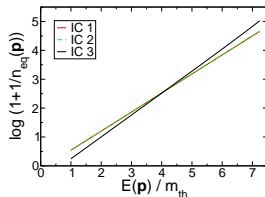
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- Full Universality
- Chemical Equilibration

Boltzmann

Equilibrium
particle
number
densities



- Restricted Universality
- No Chemical Equilibration

[Manfred Lindner, MMM (2006)]

Separation of Time Scales

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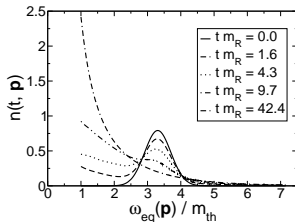
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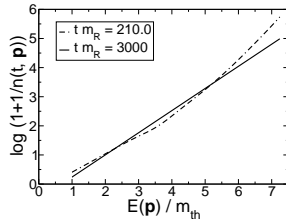
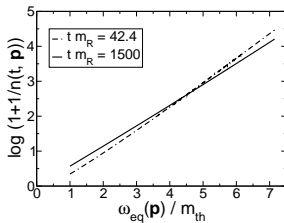
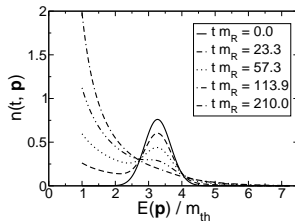
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Boltzmann




Generalization to fermionic theories

$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$ symmetric Yukawa model


$$\lambda (\Phi_a \Phi_a)^2 + i\eta \bar{\Psi} \Phi_a (\sigma_a P_R - \sigma_a^\dagger P_L) \Psi$$

Effective scalar mass:


$$M^2(x^0) = m^2 + \text{---} \bullet \text{---} \text{---} \text{---}$$


Nonlocal Self Energies:

scalars: $\Pi(x^0, z^0, k) = \text{---} \bullet \text{---} \text{---} \text{---}$



fermions: $\Sigma(x^0, z^0, k) = \text{---} \bullet \text{---} \text{---} \text{---}$



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Kadanoff-Baym Equations [Manfred Lindner, MMM (2008)]

- Full universality [Berges et al. (2003)]
- Quantum-chemical equilibration [Berges et al. (2003)]
- Prethermalization [Berges et al. (2004)]

Boltzmann equations [Manfred Lindner, MMM (2008); MMM (2006)]

- Restricted universality
- Classical, but **no quantum**-chemical equilibration
- No separation of time scales

Conclusions

Quantum Dynamics (Kadanoff-Baym equations)

- take **memory** and **off-shell** effects into account.
- respect **full** universality.
- include **chemical** equilibration.
- **separate** time scales between **kinetic** and **chemical** equilibration.

Classical Kinetics (Standard Boltzmann equations)

- **do not** take **memory** and **off-shell** effects into account (**molecular chaos** for **quasi-particles**).
- comprise **fake constants of motion**.
- respect **only a restricted** universality.
- **do not** include **quantum chemical** equilibration, and therefore
- **cannot separate** time scales between **kinetic** and **chemical** equilibration.

Outlook

Renormalization of the 2PI effective action for a real scalar $\lambda\Phi^4/4!$ theory at three-loop order

Standard approximate perturbative renormalization

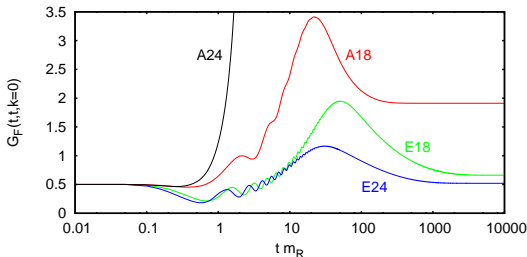
A18: $\lambda = 18, m_B^2 = -6.87 m_R^2$

A24: $\lambda = 24, m_B^2 = -9.49 m_R^2$

Exact nonperturbative renormalization at zero temperature

E18: $\lambda_R = 18, \lambda_B = 37.18, m_B^2 = -14.39 m_R^2$

E24: $\lambda_R = 24, \lambda_B = 63.43, m_B^2 = -25.14 m_R^2$



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