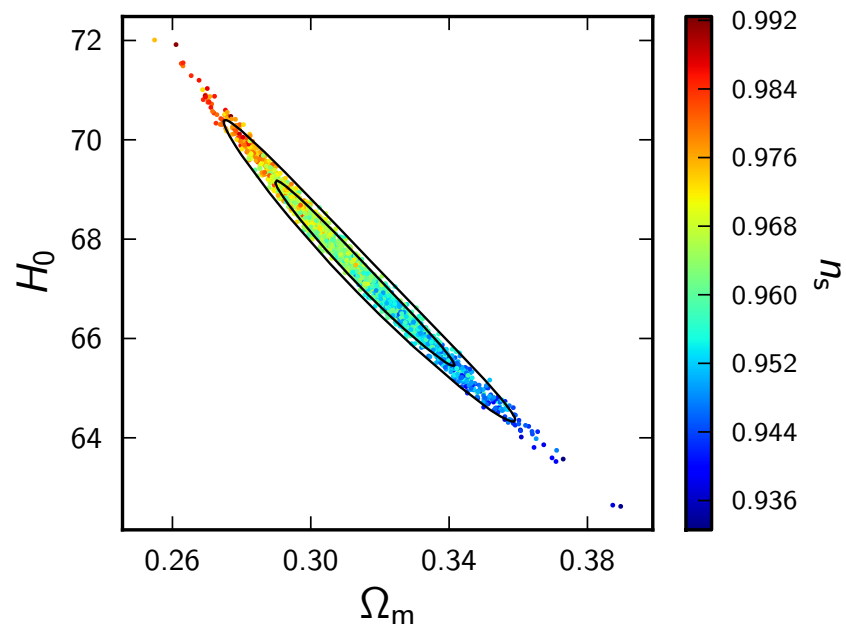


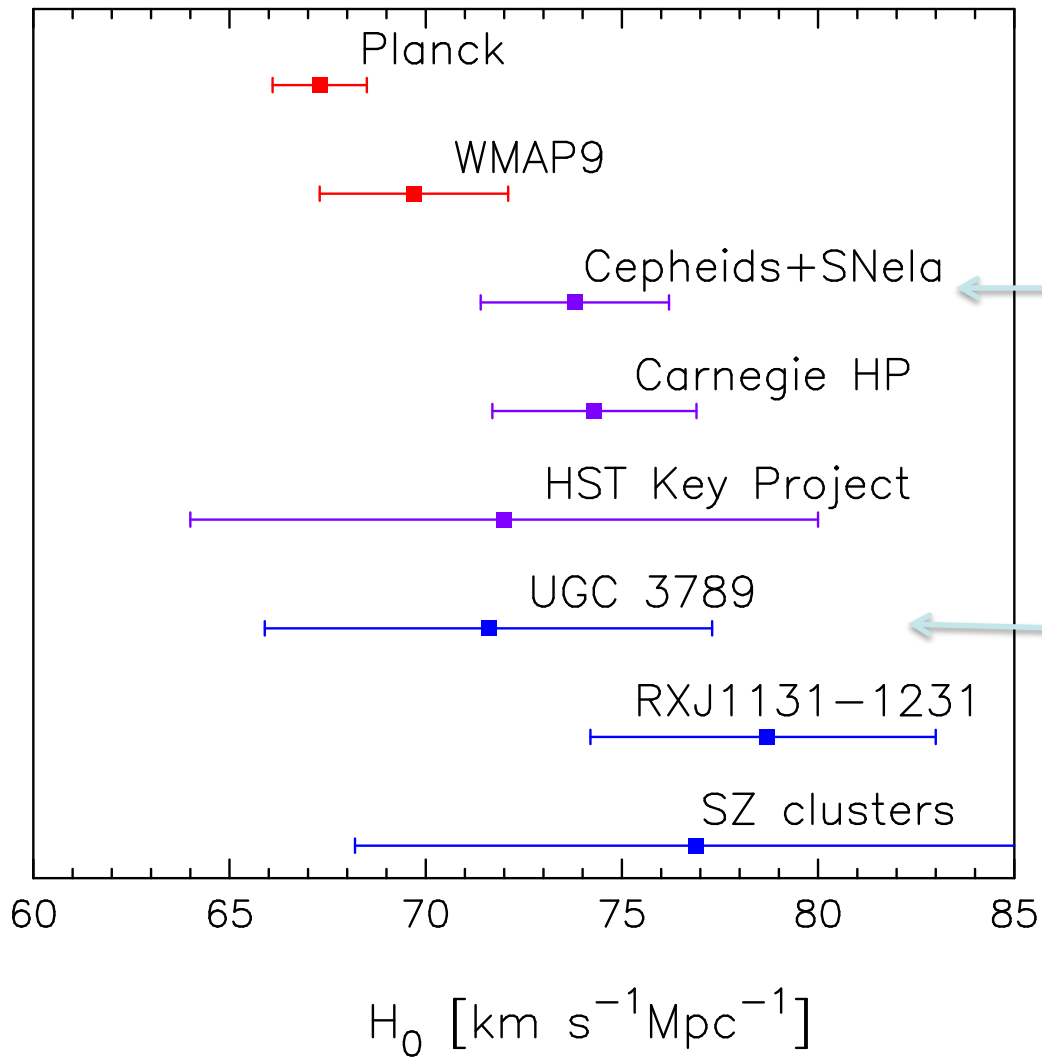
The Hubble constant from Planck



Jo Dunkley (followed by Eiichiro)

Questions

- Is the Planck-derived Hubble constant now discrepant with direct measurements?
- Is it consistent with WMAP?
- Why is it lower than the WMAP value?
- What 'new physics' could reconcile Planck and local measurements?



Uses three 1st-rung
calibrators: 71.3 – 75.7

UGC, at 50 Mpc: now
 $H_0=68.9\pm 7.1$

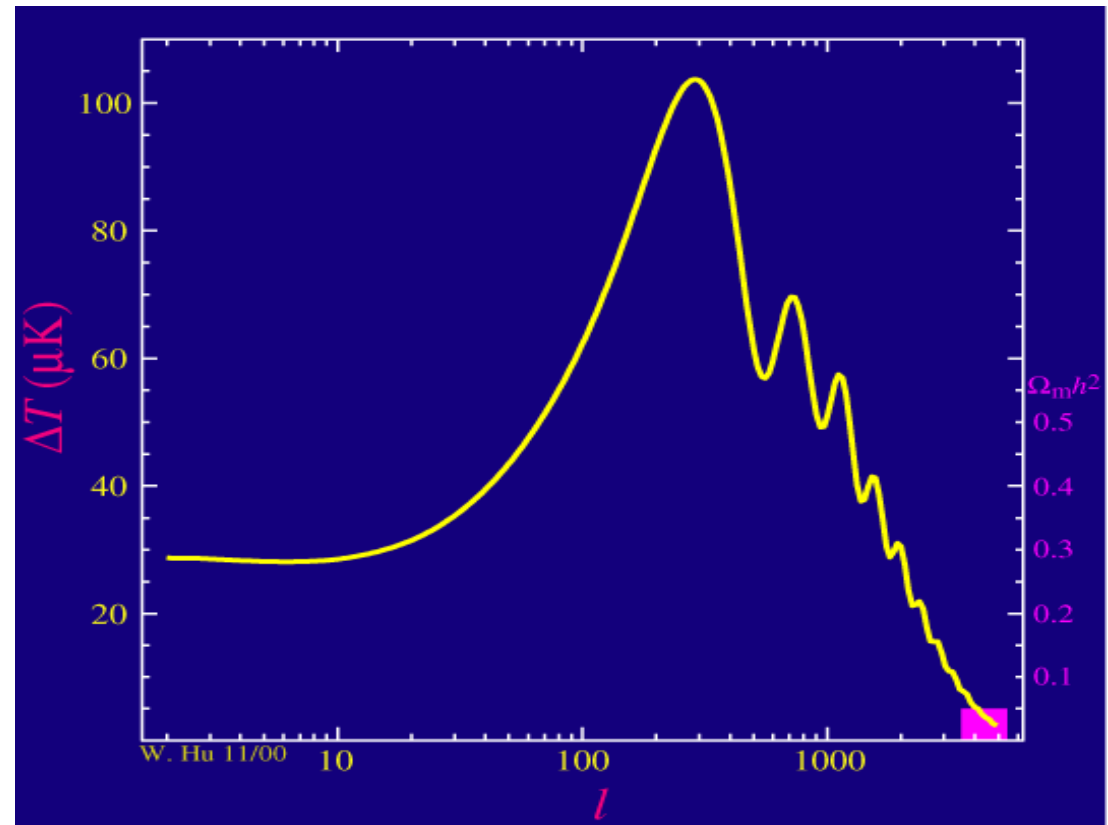
LCDM parameters $\rightarrow H_0$

For flat universe:

peak positions $\sim \Omega_m h^3$

2nd peak height $\sim \Omega_b h^2$

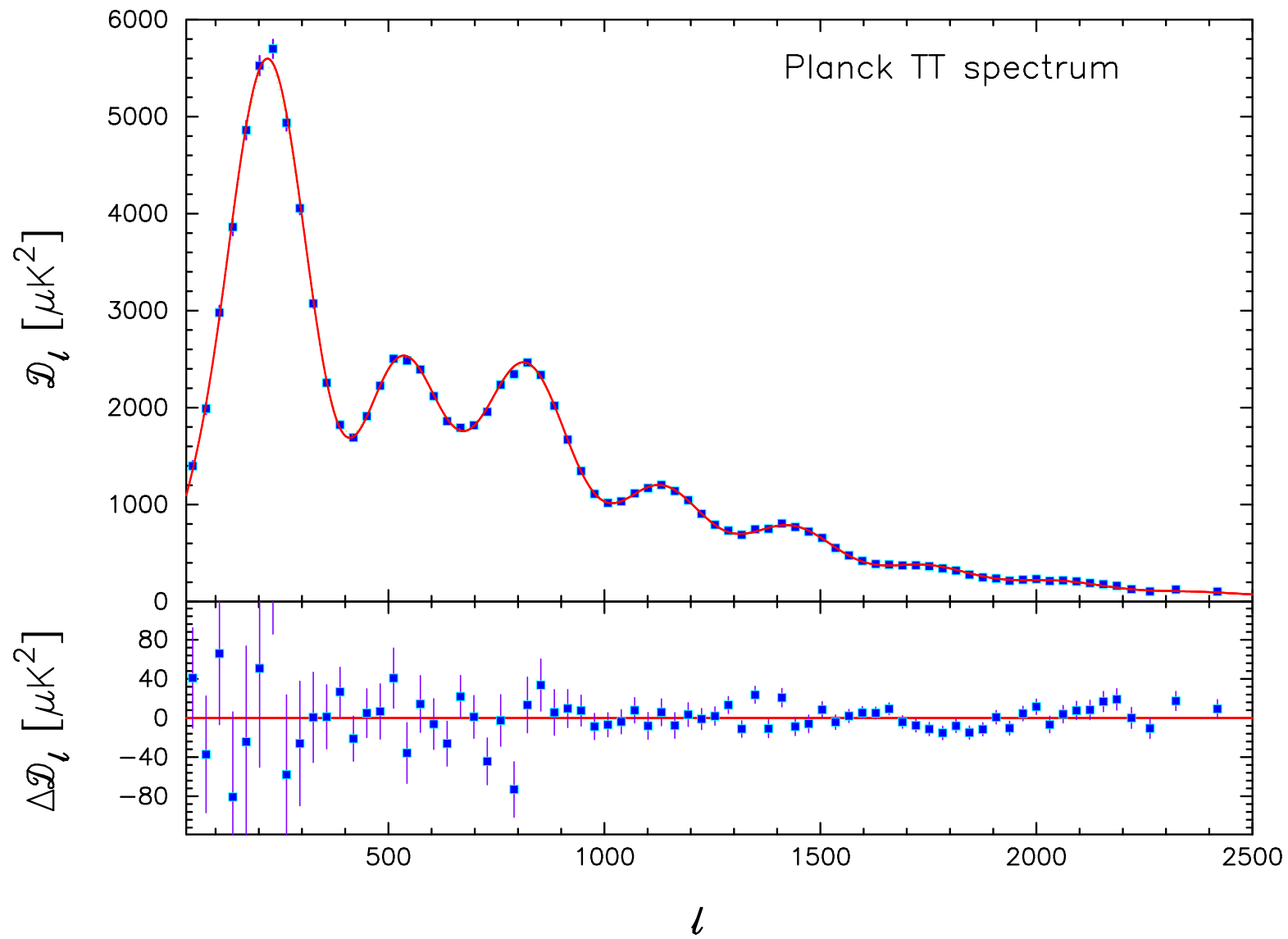
1st and 3rd/2nd peak height $\sim \Omega_m h^2$



Measures matter-radiation equality

But also: more matter density \rightarrow more damping. Also more lensing.

From Wayne Hu



WMAP v Planck

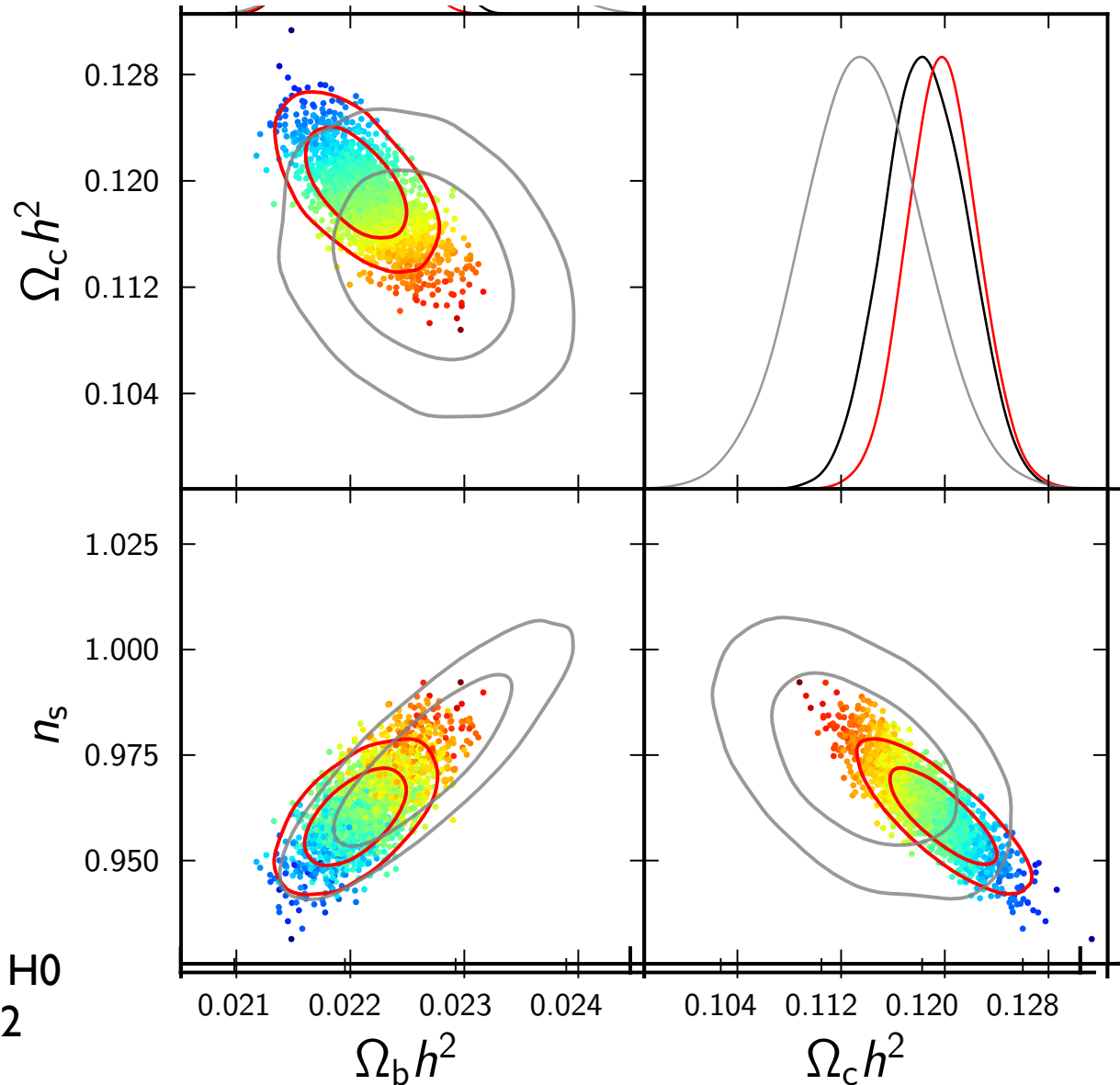
$$H_0 = 70.2 \pm 2.2$$

(WMAP9)

$$H_0 = 67.3 \pm 1.2$$

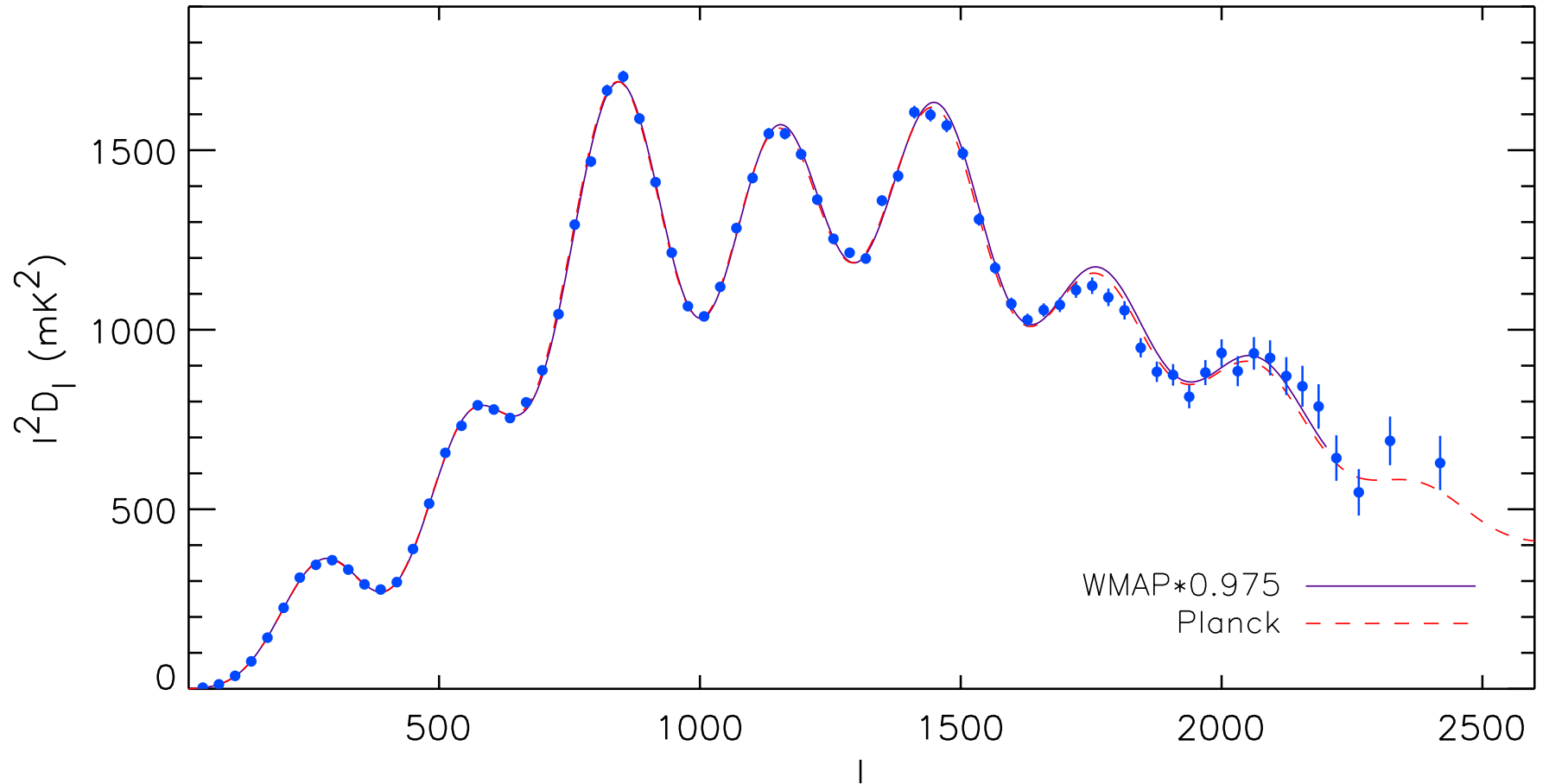
(Planck+WP)

Planck: higher $\Omega_c h^2 \rightarrow$ lower H_0
Also has lower n_s , lower $\Omega_b h^2$



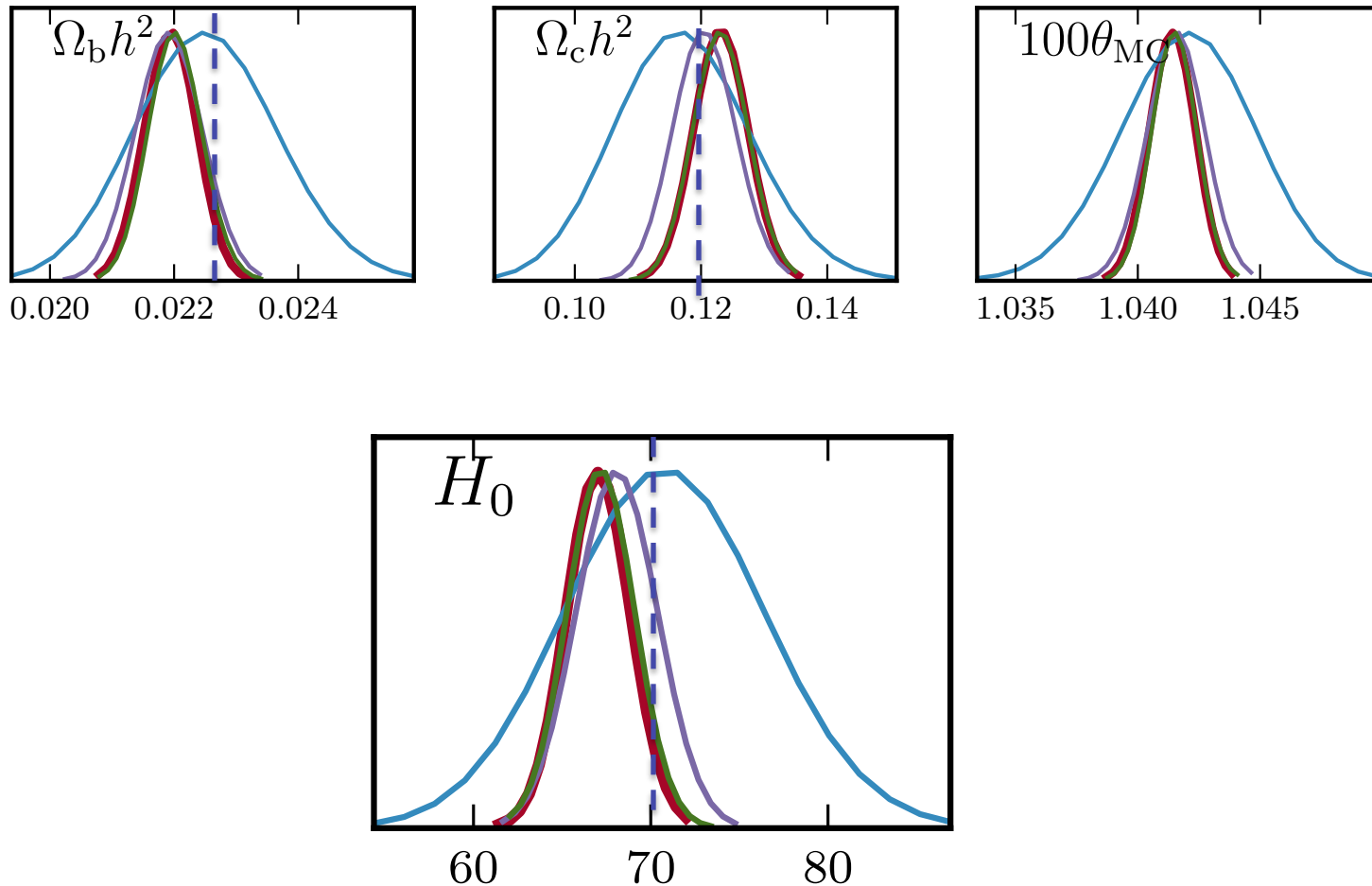
[NB: if we had used $m\nu=0.06$ for WMAP, H_0 would be 69.7 ± 2.2]

Where is difference?



Is $ell < 1000$ data consistent?

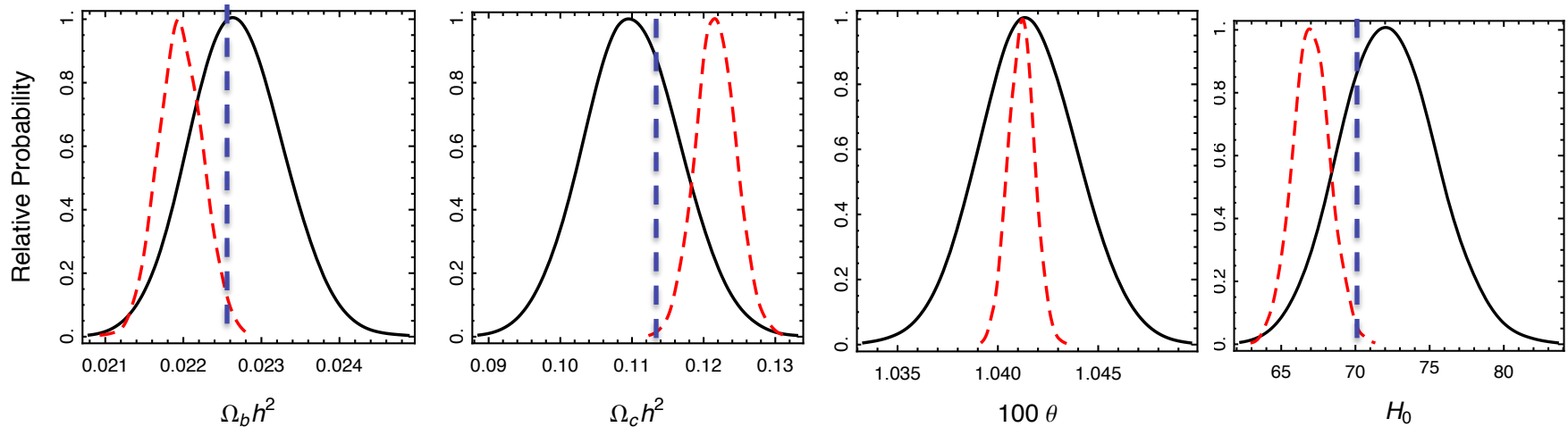
Varying Planck ell range



Lower ell range moves back similar to WMAP ones

Is $ell < 1000$ data consistent?

Using just LFI data

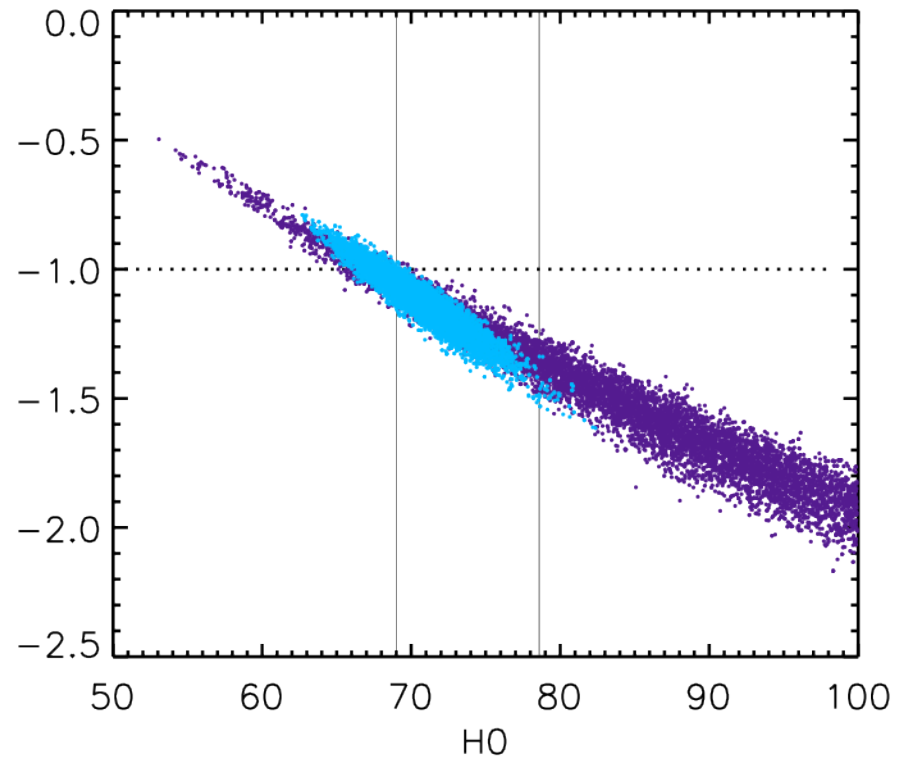
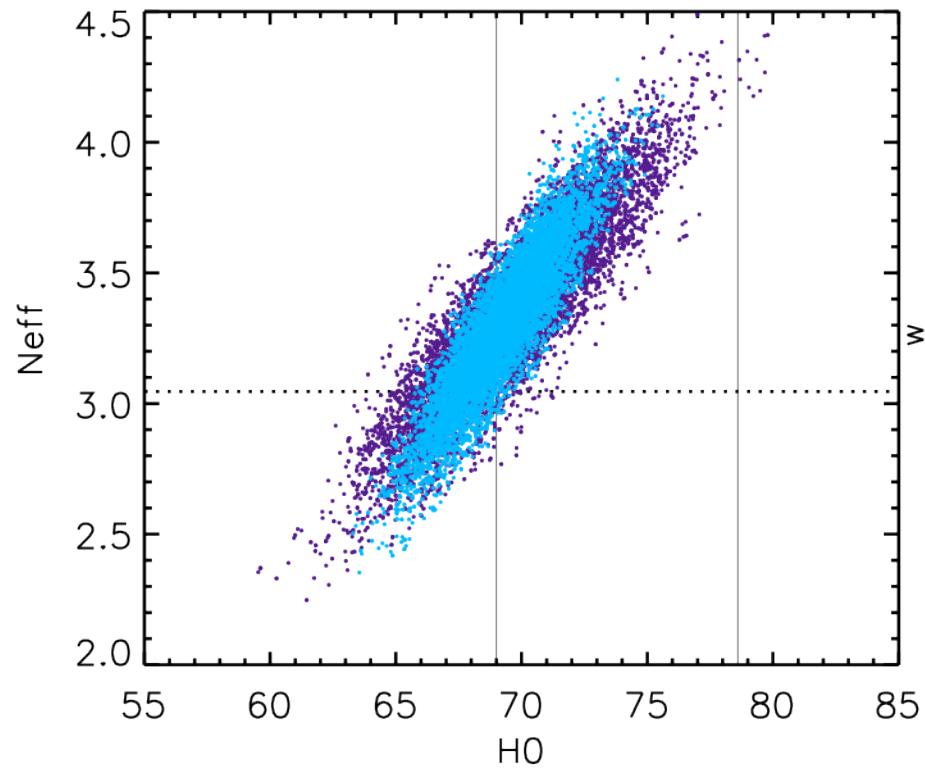


LFI parameters – from 70 GHz - move back similar to WMAP ones
(slightly higher H_0 and lower $\Omega_c h^2$)

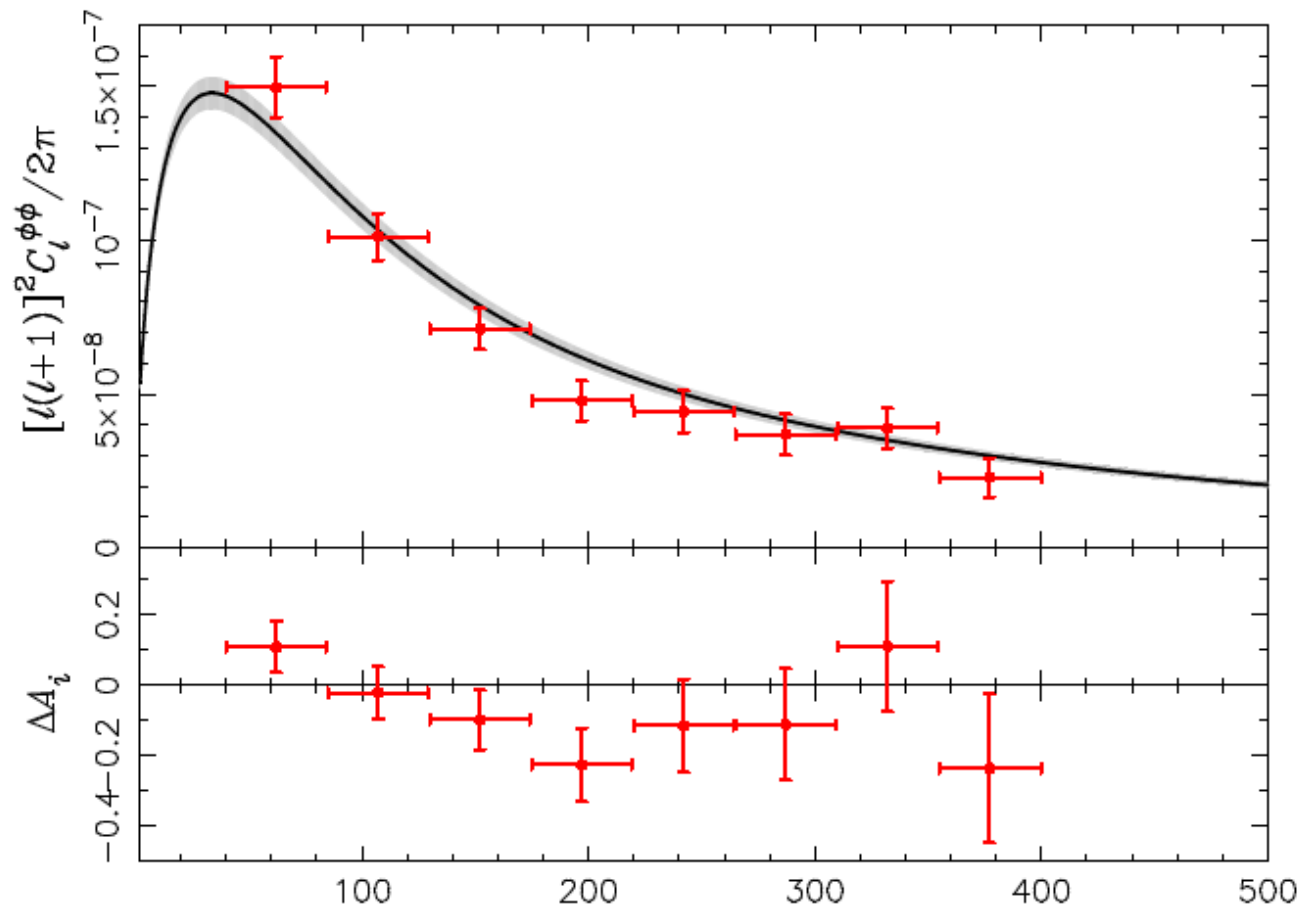
What if difference is real?

Not n_{eff} : anti-correlates with H_0

If error is \sim halved and mean doesn't change, would need to go beyond LCDM



What does lensing tell us?



Not much yet for LCDM ($H_0 = 67.9 \pm 1.0$ with lensing), but beautifully breaks geometric degeneracy ($H_0 = 64.6 \pm 3.3$ in curved universe)

Questions

- Is the Planck-derived Hubble constant now discrepant with direct measurements?

Different at 2.5 sigma from HST measurements.

Consistent with new water maser measurement

- Is it consistent with WMAP?

Best-fit differs by ~ 1 -sigma, but using same ell range see consistent parameters (not driven by amplitude diff)

- Why is it lower than the WMAP value?

$4^{\text{th}}-6^{\text{th}}$ peaks are lower than WMAP best-fit model. Need $n_s < 1$ and more damping \rightarrow more $\Omega_{\text{ch}2}$ and less $\Omega_{\text{bh}2} \rightarrow$ lower H_0 .

- What 'new physics' could reconcile Planck and local measurements?

$3 < N_{\text{eff}} < 3.5$ or $-1.3 < w < -1.1$ could fit

