Explosions from Double Neutron Star Mergers

Chris Fryer Los Alamos National Laboratory/ University of Arizona

- Population Synthesis
 Studies: Rates, Delays,
 Locations and Problems
- Merger: Comparison between DNS and stellar collapse
- Nucleosynthesis









Theoretical estimates of merger rates are plagued by uncertainties: we are slowly making progress on these uncertainties.

UNCLASSIFIED

Progenitors: Merger Rates

 Observational predictions are plagued by small number statistics.



2/11

MS

Standard, FBB Kick

-High Wind, FBB Kick

Standard, Maxw. Kick

10

100

Merger Time (Myr)

1000

merger times

300

Bin

Time

log per 500

of Binaries

Number (

0.25

0.2

Large Radius, FBB Kick

Fryer et al.

1999

104

New Formation Scenario Changes Distribution of Merger Times

- Lots of population studies: Belcyznski, Kalogera and collaborators; Lee, Park, Brown and collaborators; Tutukov, Yungelson and collaborators; Bloom et al., Tout et al., Portegies Zwart, van den Heuvel...
- Belczynski & Kalogera (2001) argue for Hegiant formation scenario. By 2004, they argue that this scenario contributes half of their DNS rate.





Progenitors: Outstanding Problems

- Stellar Radii
- Stellar Structure
- Winds
- Binary Mass Transfer
- Common Envelope Evolution
- Neutron Star Kicks
- NS and BH distribution functions (see stellar structure)



Diehl et al. 2009



EST. 1943



DNS as GRB engines: First, let's review the collapsar engine







Merger provides angular momentum

• Accretion rates higher, shorter, and more disruptive. The energy requirements for short bursts are less extreme. Hence, mergers can produce GRBs through neutrino annihilation.

• Merger outbursts should be less beamed than Can magnetic fields produce a narrow jet?

• Fallback can produce late-time accretion (Rosswog 2007 – although see Metzger et al. 2009)







Lots of nice work here, see also Arnould et al. 2007; Metzger et al. 2008



DNS and r-Process Nucleosynthesis

Although DNS mergers may produce r-process elements, they have a distribution problem:

- r-process yields appear along with iron and other explosive yields.
- If deposited at the star formation site, the DNS rate would need to be nearly as fast as the supernova rate.



Argast et al. 2004



os Alamos







- Many DNS mergers occur in the halo. Hence, the yields can be mixed more simply.
- A 1-zone model including these yields through infall can be made to match the observations.

LA-UR 09-05528

10/11

Summary

• Population studies plagued by uncertainties in stellar evolution and binary effects. The rates are uncertain to at least 1 order of magnitude. Although these mergers should be more distributed than stellar collapse, the fraction of mergers far from the disk depends upon population

synthesis calculations.

- Advanced LIGO should place strong constraints both on the rates and on the neutron star equation of state.
- DNS mergers may still produce r-process yields.





