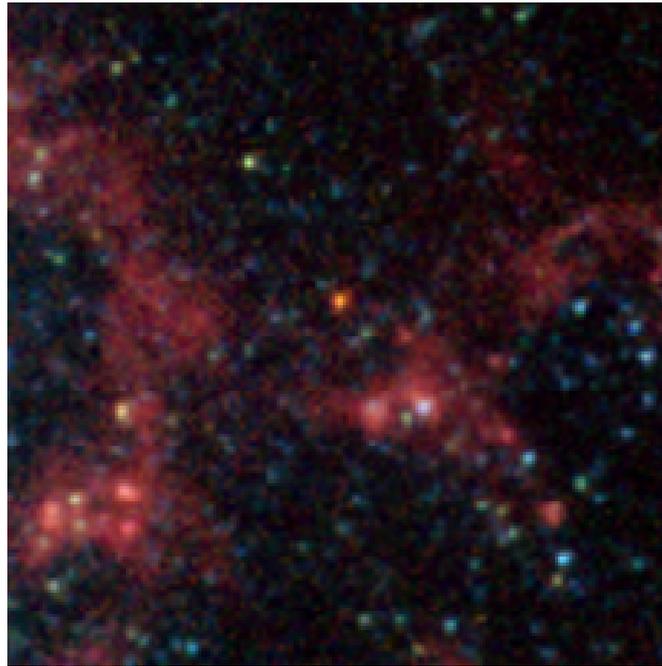


# A Spitzer Spectrum of the 2008 Luminous Transient in NGC 300: Connection to Proto-Planetary Nebulae



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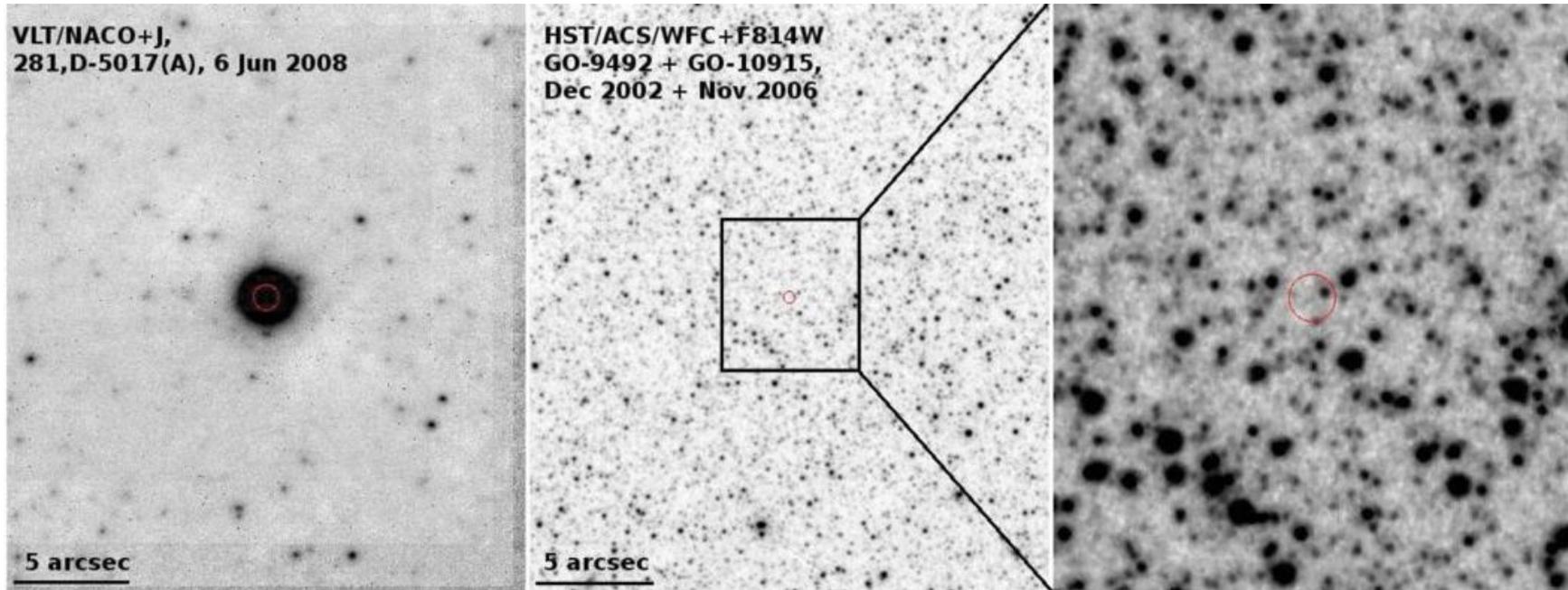
# Luminous Transient in NGC 300



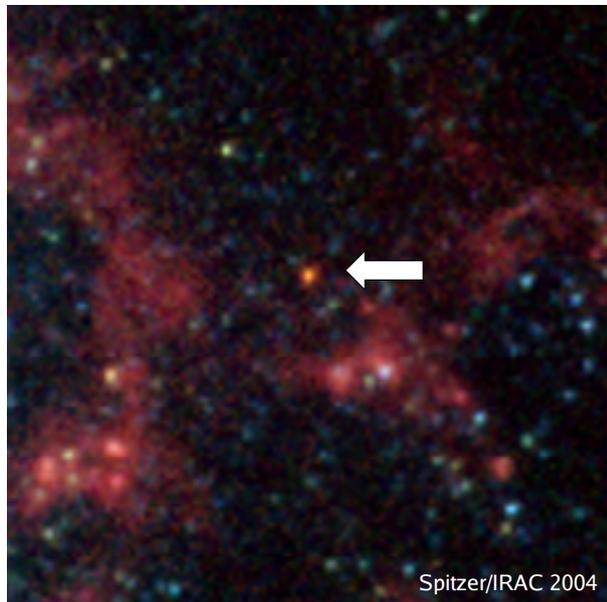
Discovered May 14, 2008 by Berto Monard in NGC 300 (2 Mpc)

$M_V = -13$  mag, spectrum similar to V838 Mon and “SN impostors”

# Progenitor of NGC 300-OT

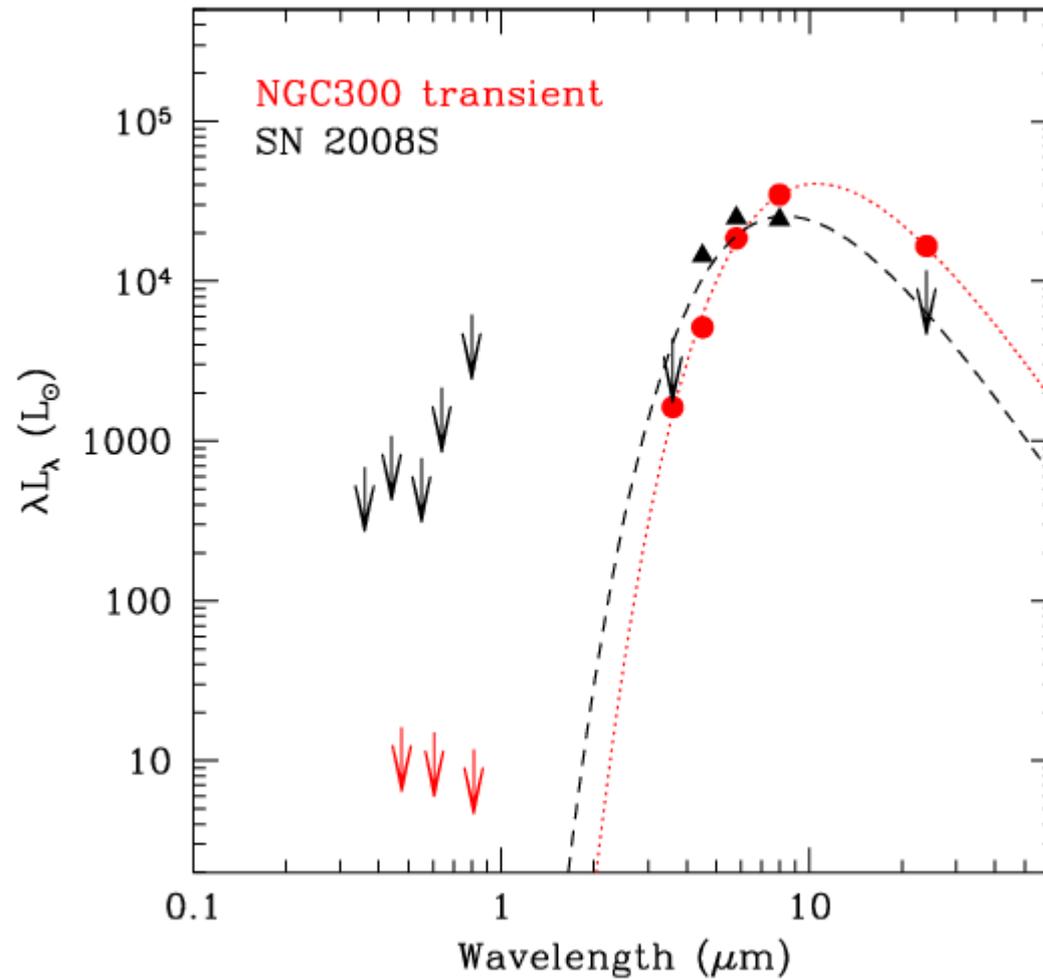


(Bond et al. 2009)



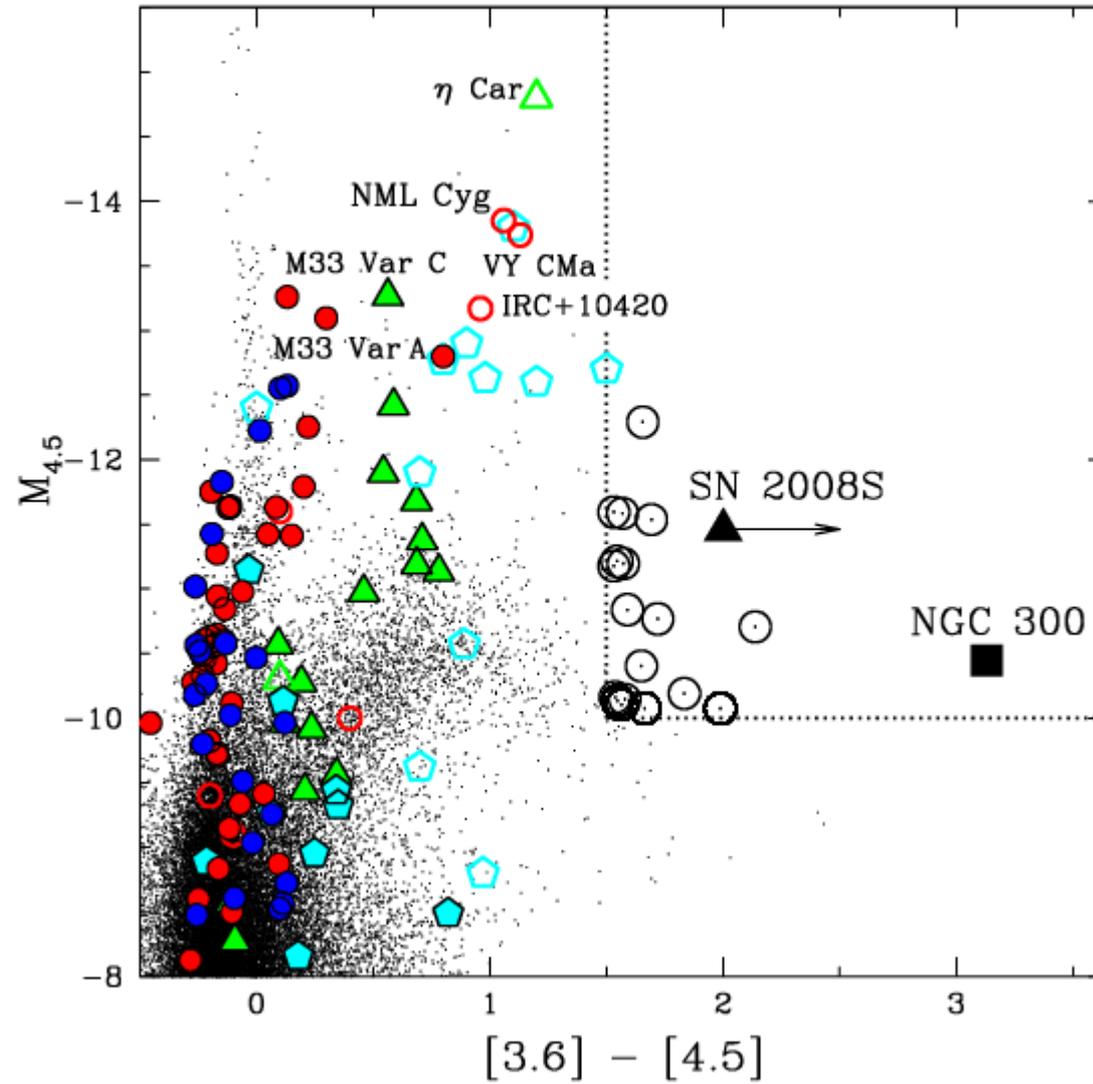
- Progenitor undetected in very deep optical images from HST (Berger & Soderberg 2008; Bond et al. 2009)
- Luminous source in pre-explosion Spitzer images (Prieto 2008; Thompson et al. 2008)

# Progenitor: Dust Enshrouded Star



- Luminosities  $4\text{-}6 \times 10^4$   $L_{\text{sun}}$
- $T = 300\text{-}450$  K
- $R = 150\text{-}300$  AU

# Progenitors as a Class



(Thompson et al. 2008)

# Progenitors and Transients

NGC 300-OT and SN 2008S

- Low peak magnitude  $M_V = -13$  to  $-15$  mag and int. energy
- Spectra of “optically thick wind” with H, He, Ca II in emission (100 to  $>1000$  km/s)
- Massive dust-enshrouded progenitors ( $M_{\text{prog}} = 6-15 M_{\text{sun}}$ )
- Transients are dusty with asymmetric distribution

Prieto et al. (2008, 2009); Thompson et al. (2008); Smith et al. (2009); Bond et al. (2009); Berger et al. (2009); Botticella et al. (2009); Gogarten et al. (2009); Wesson et al. (2009); Patat et al. (2009)

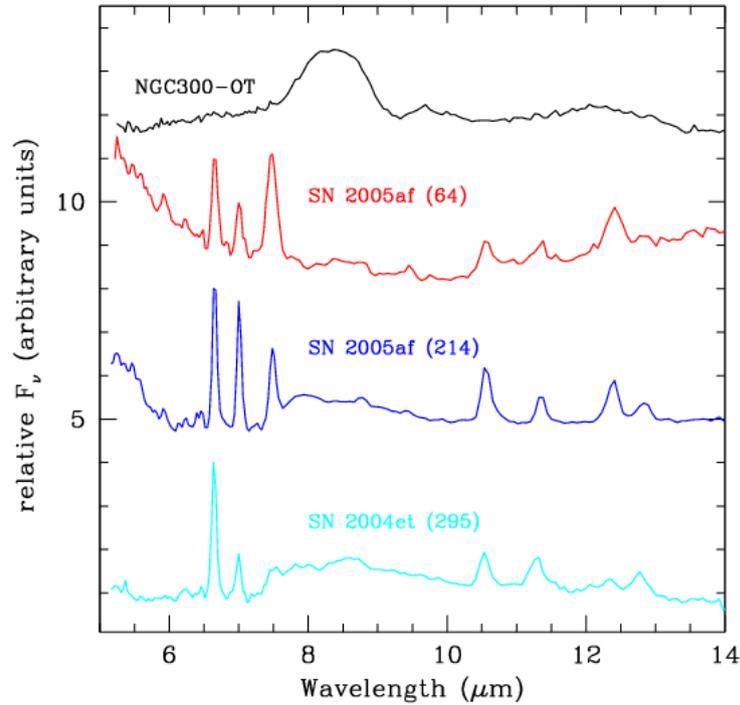
# What are NGC 300-OT and SN 2008S ?

- Electron-capture SNe ( $M_{\text{prog}} = 8-10 \text{ Msun}$ )
- Low-luminosity CC SNe ( $M_{\text{prog}} = 10-12 \text{ Msun}$ )
- Massive Star Eruption ( $M_{\text{prog}} = 10-15 \text{ Msun}$ )
- Massive White Dwarf Birth ( $M_{\text{prog}} = 6-8 \text{ Msun}$ )

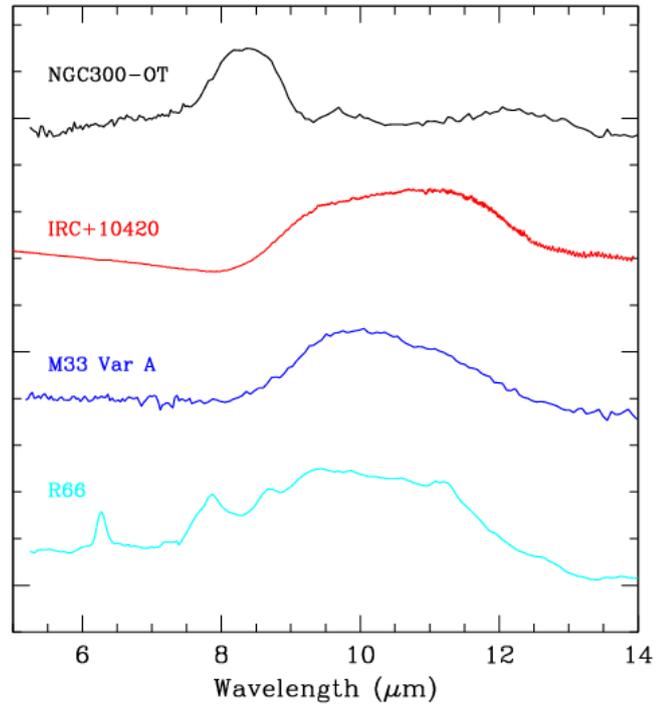
Thompson et al. (2008)

# Spitzer Spectrum of NGC 300-OT

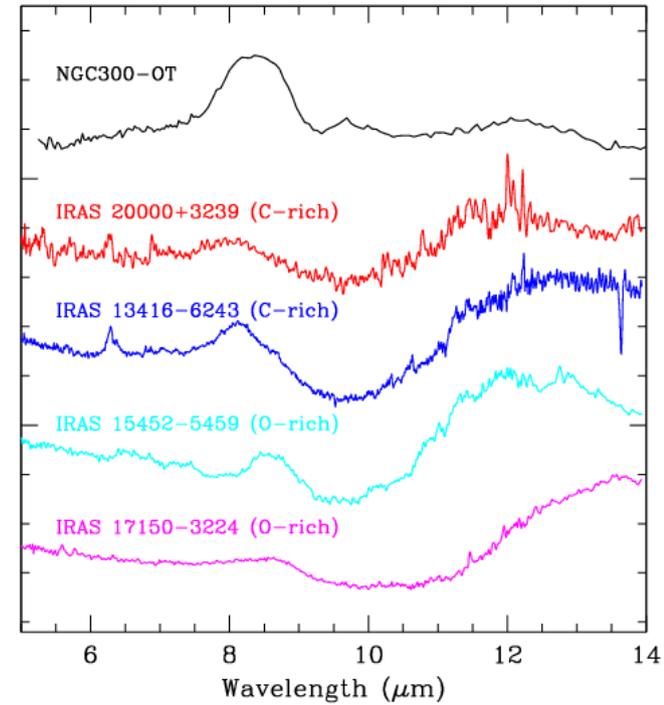
vs SN II-P



vs massive stars



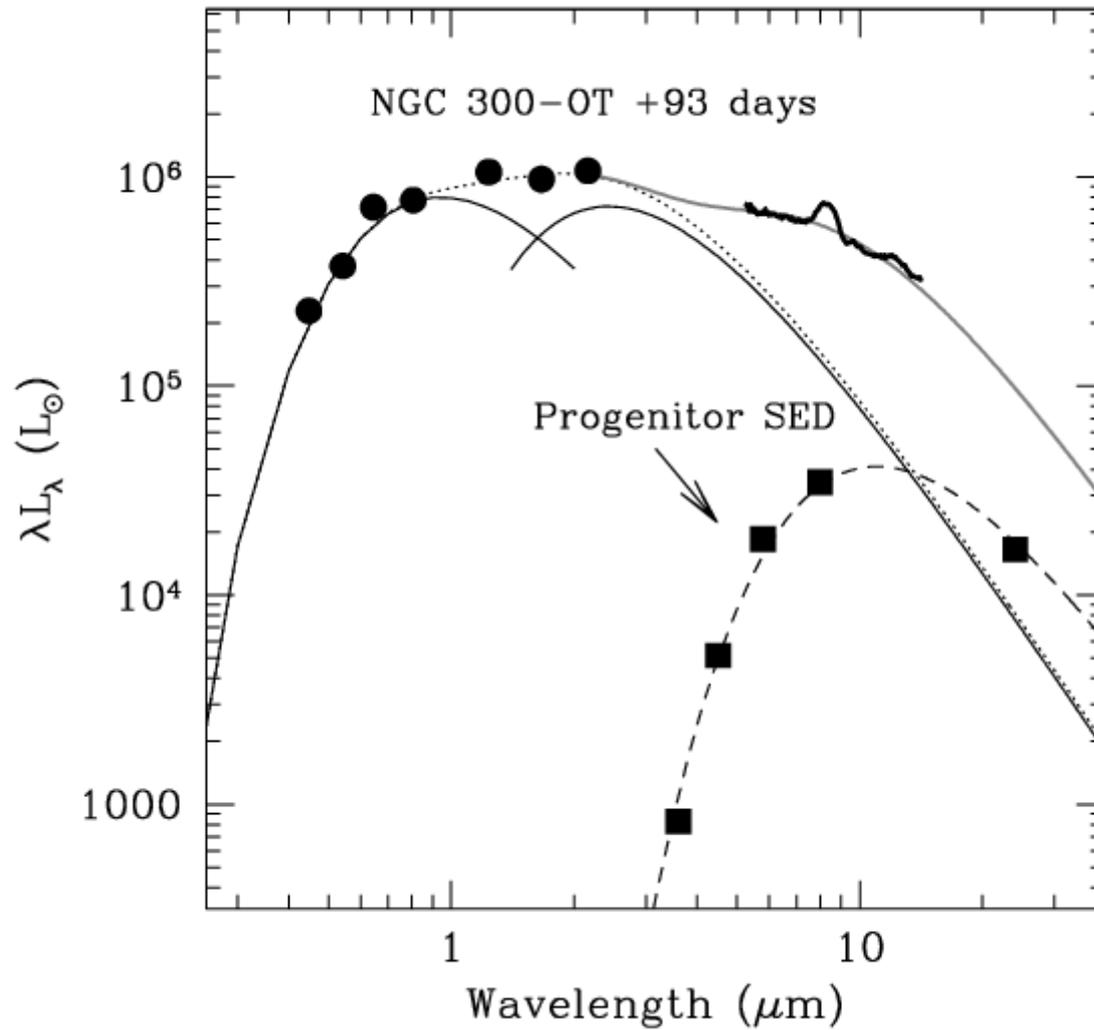
vs proto-PNe



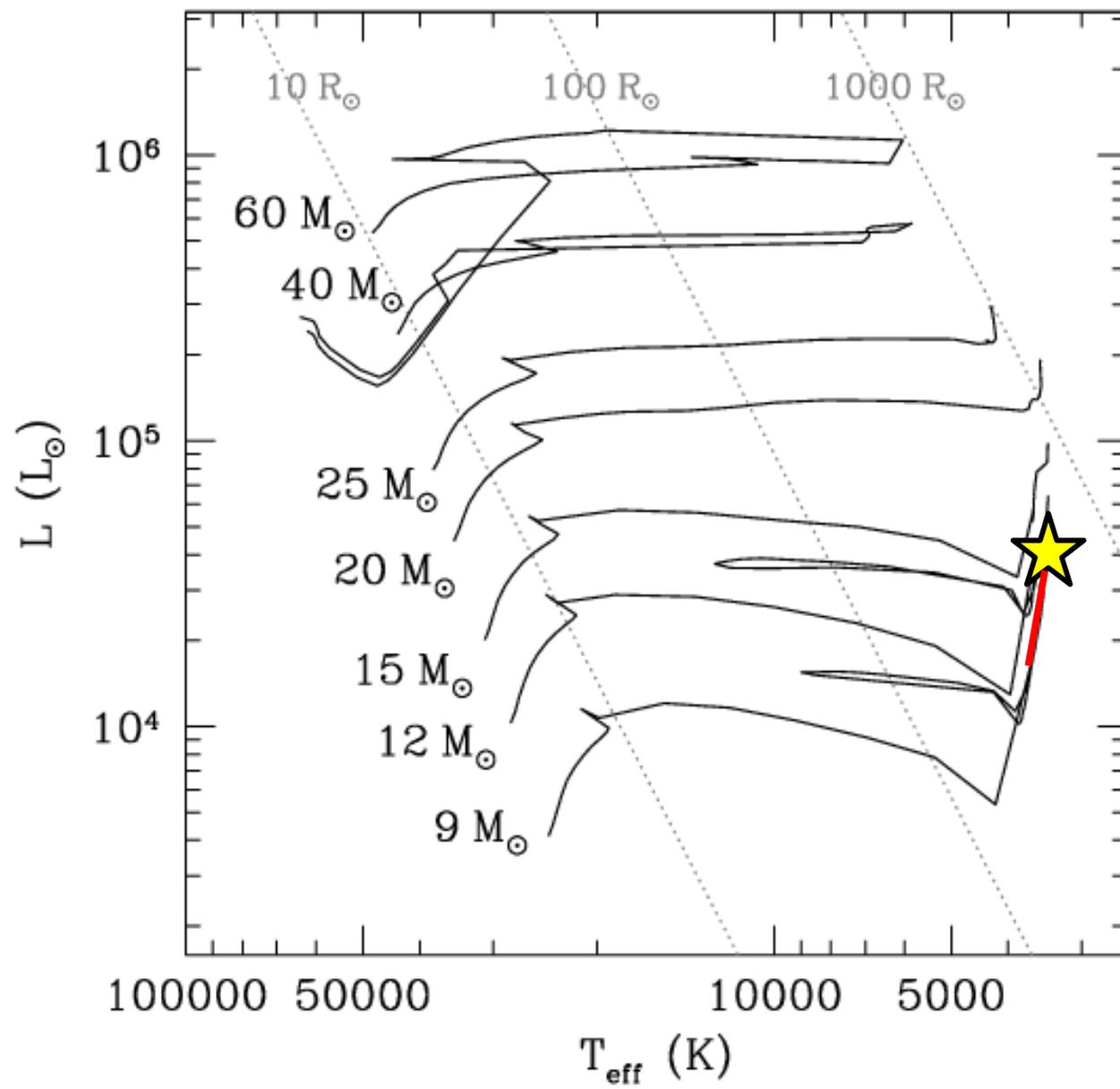
→ Mid-IR features similar to C-rich proto-PNe

(see poster and arXiv:0907.0230)

# Spectral Energy Distribution



→ Progenitor dust survived explosion



Thanks !