

# White Dwarf Mergers

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# Outline

- Inspiral
- Dynamical Stability
- Merger
- Unresolved Questions
- Outcomes

# Inspiral

## Gravitational Radiation

$$\begin{aligned}\tau_{\text{inspiral}} &= \frac{1}{8} \tau_{GR} \equiv |J / \dot{J}| \\ &\approx 98 \text{ yr } (M_{\text{chirp}} / M_{Ch})^{-5/3} (P / \text{min})^{8/3}\end{aligned}$$

where

$$M_{\text{chirp}} \equiv \left( \frac{M_1^3 M_2^3}{M_1 + M_2} \right)^{1/5}$$

## Orbital period at contact

$$P \approx 43 \text{ sec } (M / M_{Ch})^{-1} \left[ 1 - (M / M_{Ch})^{4/3} \right]^{3/4}$$

# Preheating?

Rotation energy if synchronous:

$$E_{rot} \sim 0.01 GM^2 / R \sim 10^{48.5} \text{ erg}$$

- High dissipation rate, concentrated in outer envelope
- Insufficient energy to lift bulk degeneracy of core.

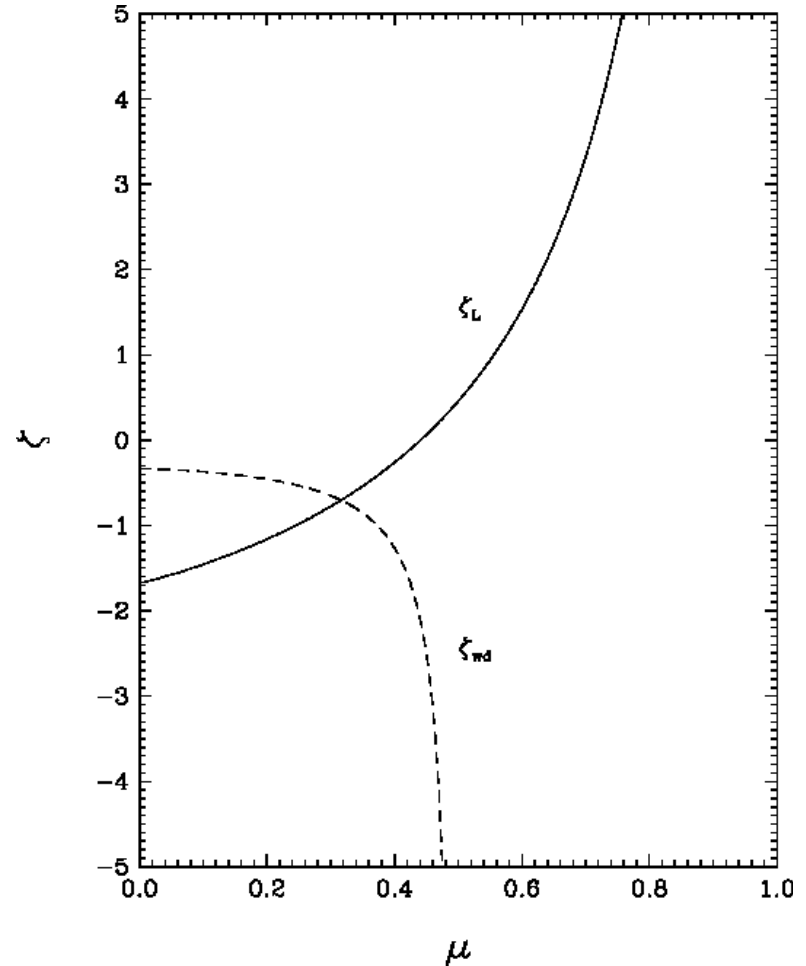
# Stability

$$\zeta_{wd} \geq \zeta_L$$

=> dynamically stable

$$\zeta_L > \zeta_{wd}$$

=> dynamically unstable



# Mass Transfer Rates

Fully degenerate donor star (1) (envelope  $\mu_e'$ , core  $\mu_e$ ):

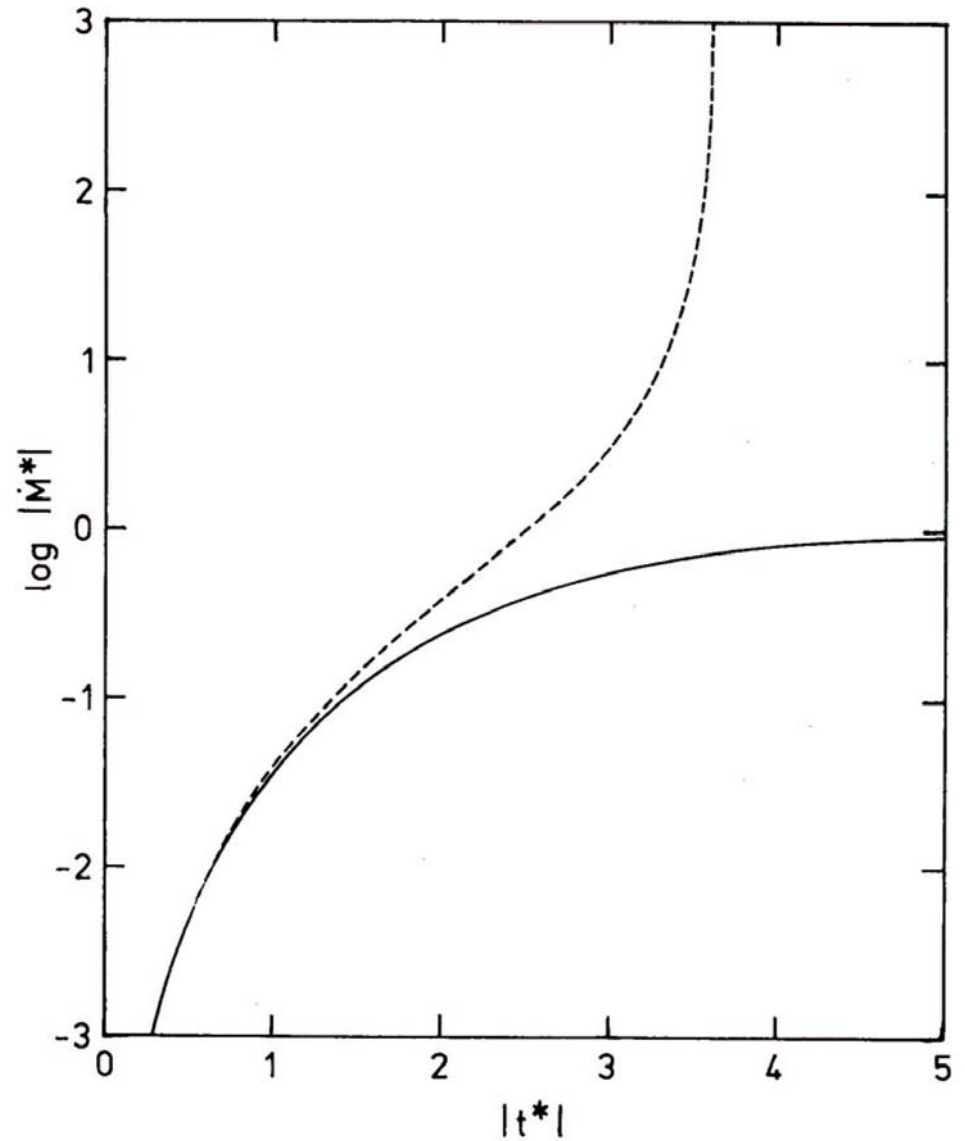
$$\begin{aligned}\dot{M}_1 &= -F(M_1/M_2)E(M_1/M_{Ch})\left(\frac{\mu_e'}{\mu_e}\right)^{5/2}\left(\frac{M_1}{P}\right)\left(\frac{\Delta R_1}{R_1}\right)^3 \\ &\approx 0.15 M_\odot \text{sec}^{-1}\left(\frac{\mu_e'}{\mu_e}\right)^{5/2}\left(\frac{M_1}{M_{Ch}}\right)^2\left(\frac{\Delta R_1}{R_1}\right)^3\end{aligned}$$

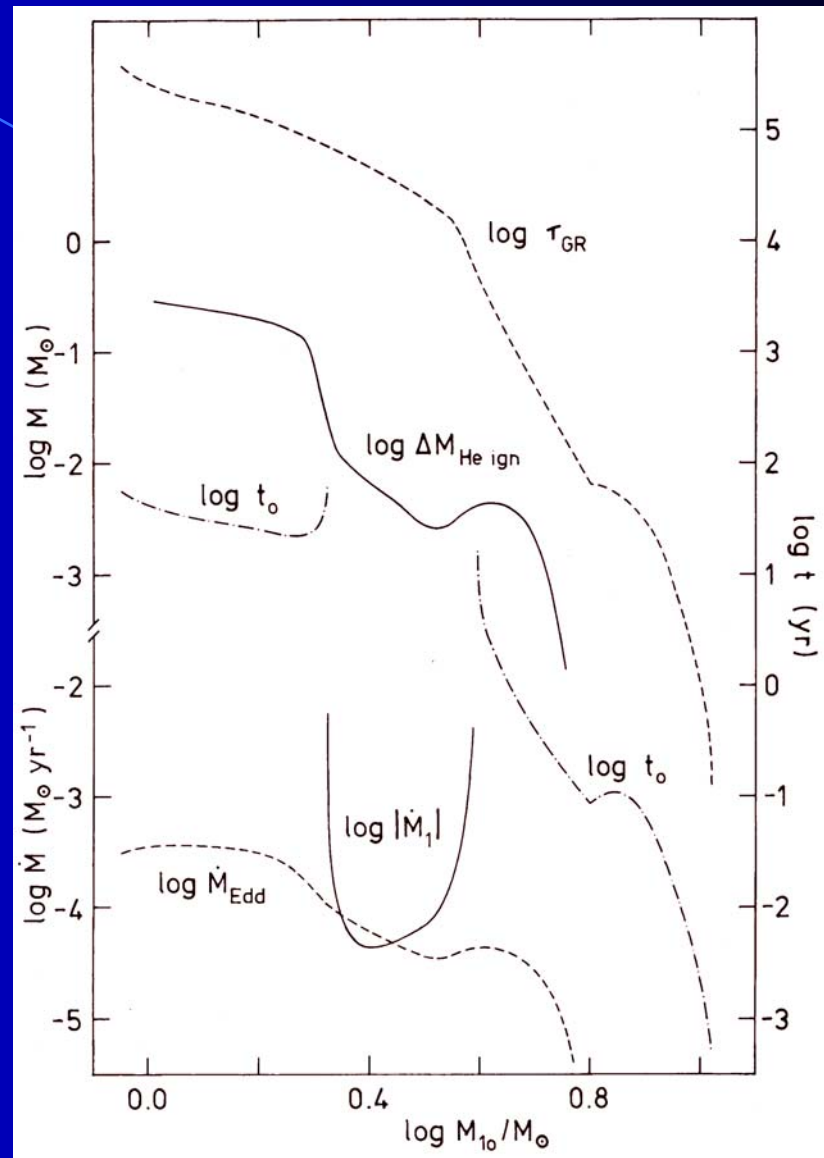
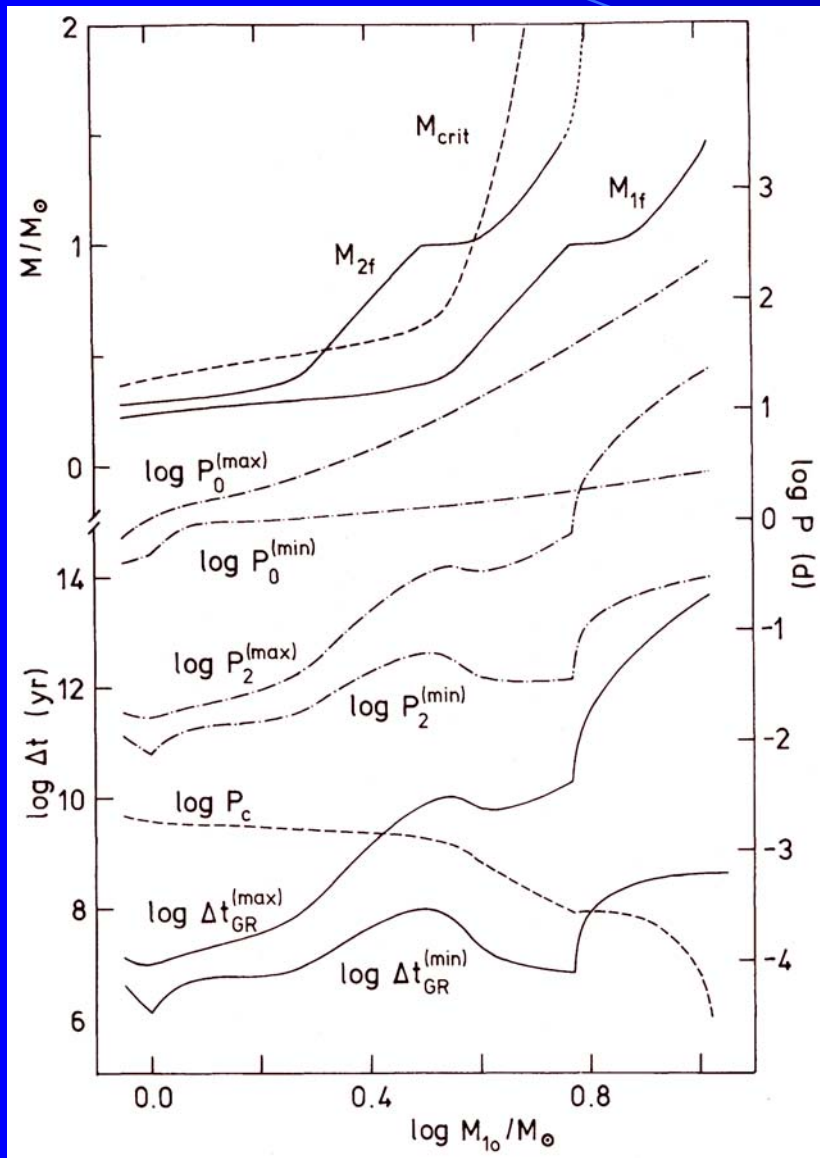
$\dot{M}^*$  in units of

$$\frac{2}{\tau_{GR}} M_1 |\zeta_1 - \zeta_L|^{-1}$$

$t^*$  in units of

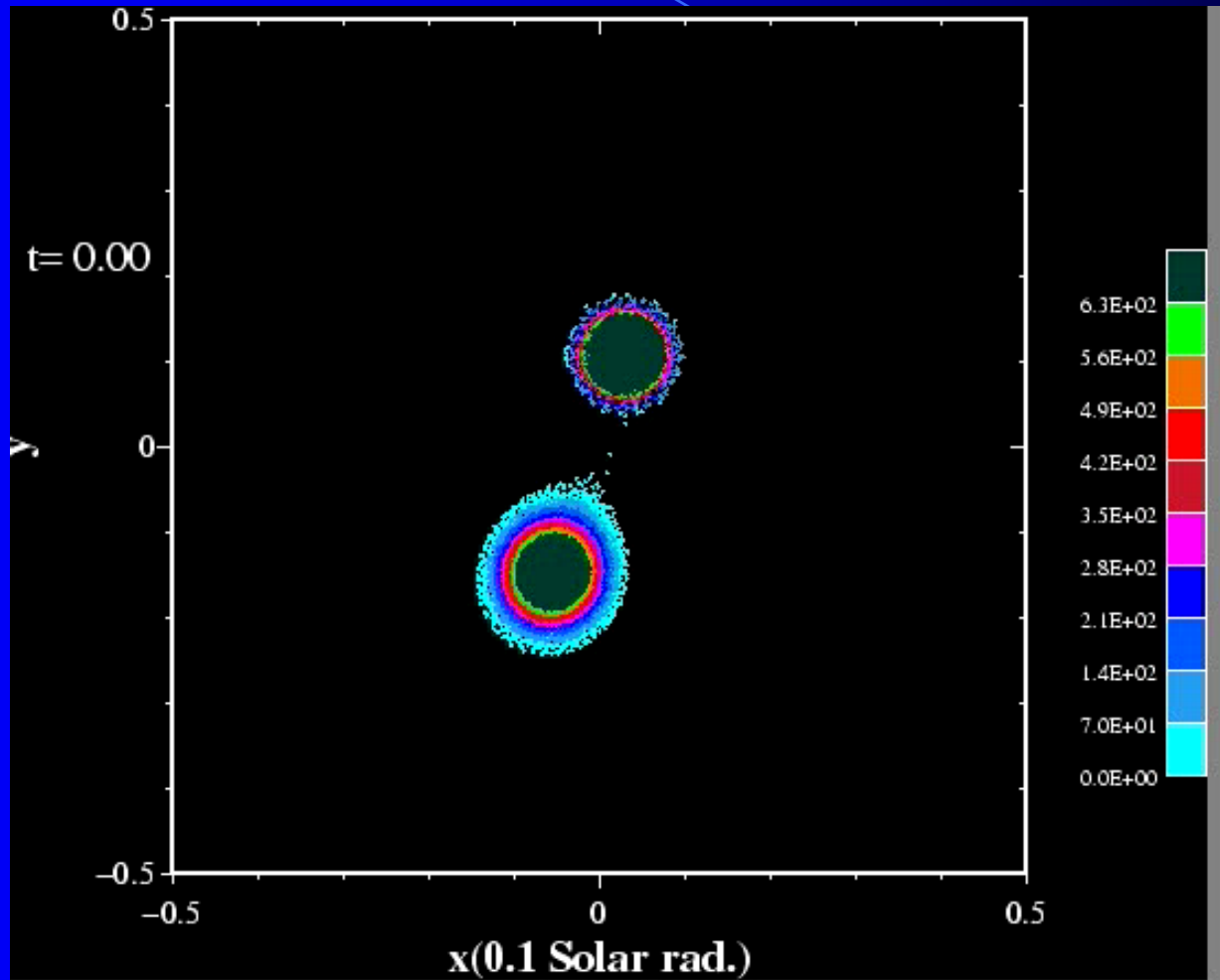
$$\frac{1}{3} \left( \frac{\tau_{GR}^2 P}{4FE |\zeta_1 - \zeta_L|} \right)^{1/3}$$







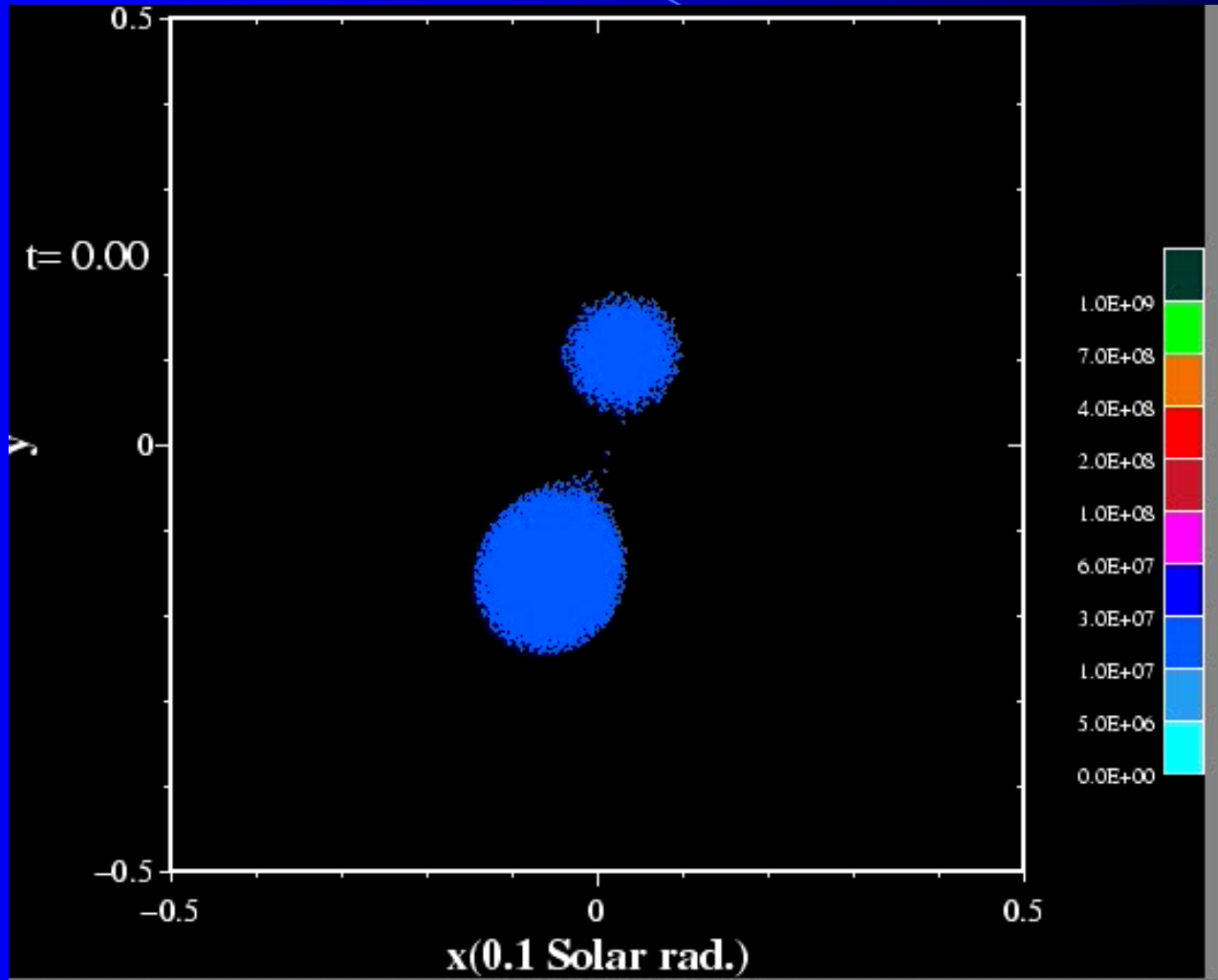
# Merger!



$0.8+1.0 M_{\odot}$

(Guerrero, Garcia-Berro & Isern (2004 A&A 413, 257))

# Heating/Ignition



$0.8+1.0 M_{\odot}$

(Guerrero, Garcia-Berro & Isern (2004 A&A 413, 257))

# Unresolved Questions

- What fraction of the orbital angular momentum of the binary is lost in the dynamical merger? (Dynamical mass loss is minimal - 0.5-2%)
- Over what fraction of the donor is electron degeneracy lifted?
- What physical process removes angular momentum from the heavy disk remnant of the donor? (And on what time scale?)
- Can observable counterparts to heavy disk systems be identified?