

DARK MATTER AND HIGGS BOSONS IN THE MSSM

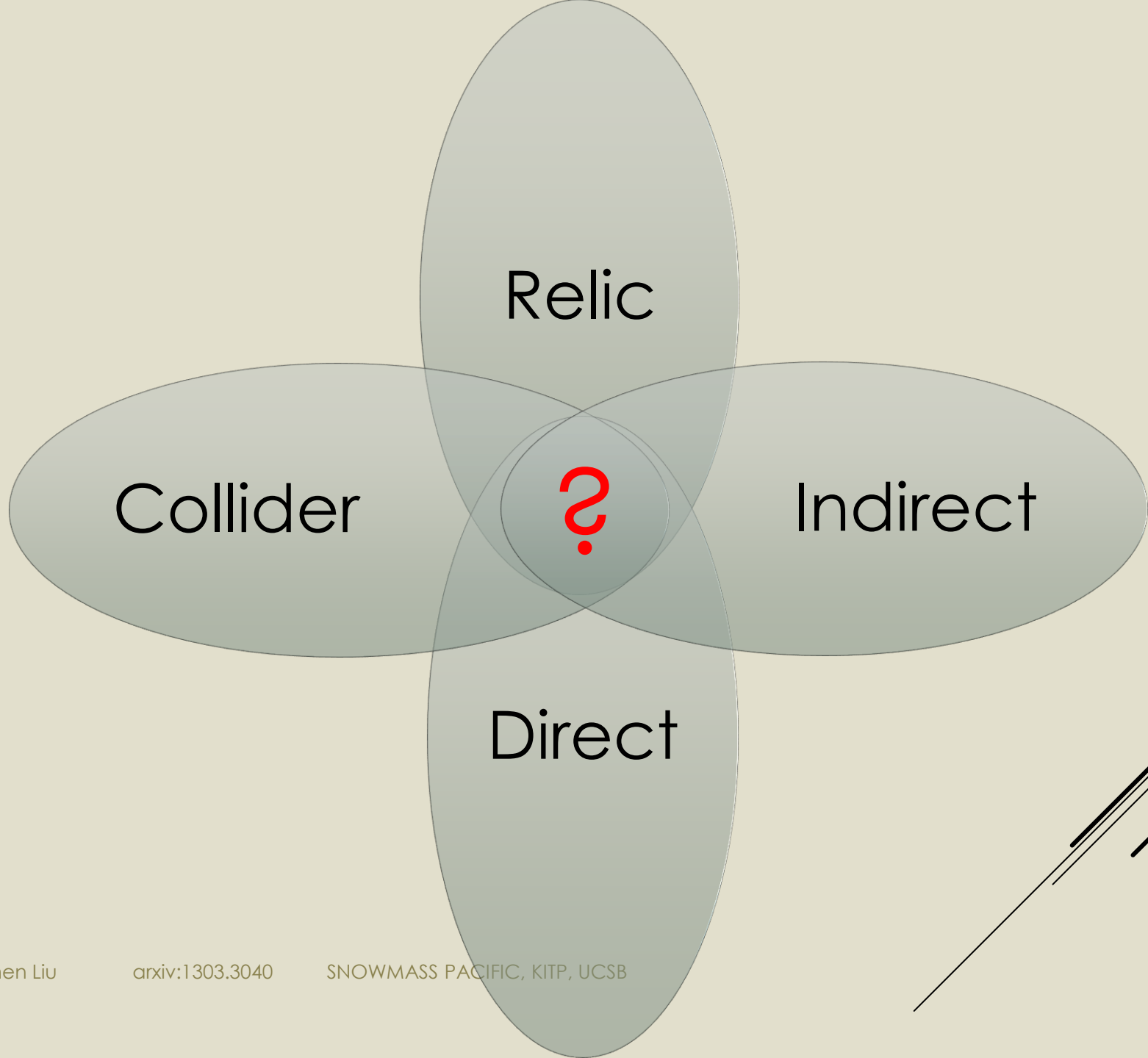
Zhen Liu

Based upon work with

Tao Han and Aravind Natarajan, arxiv:1303.3040

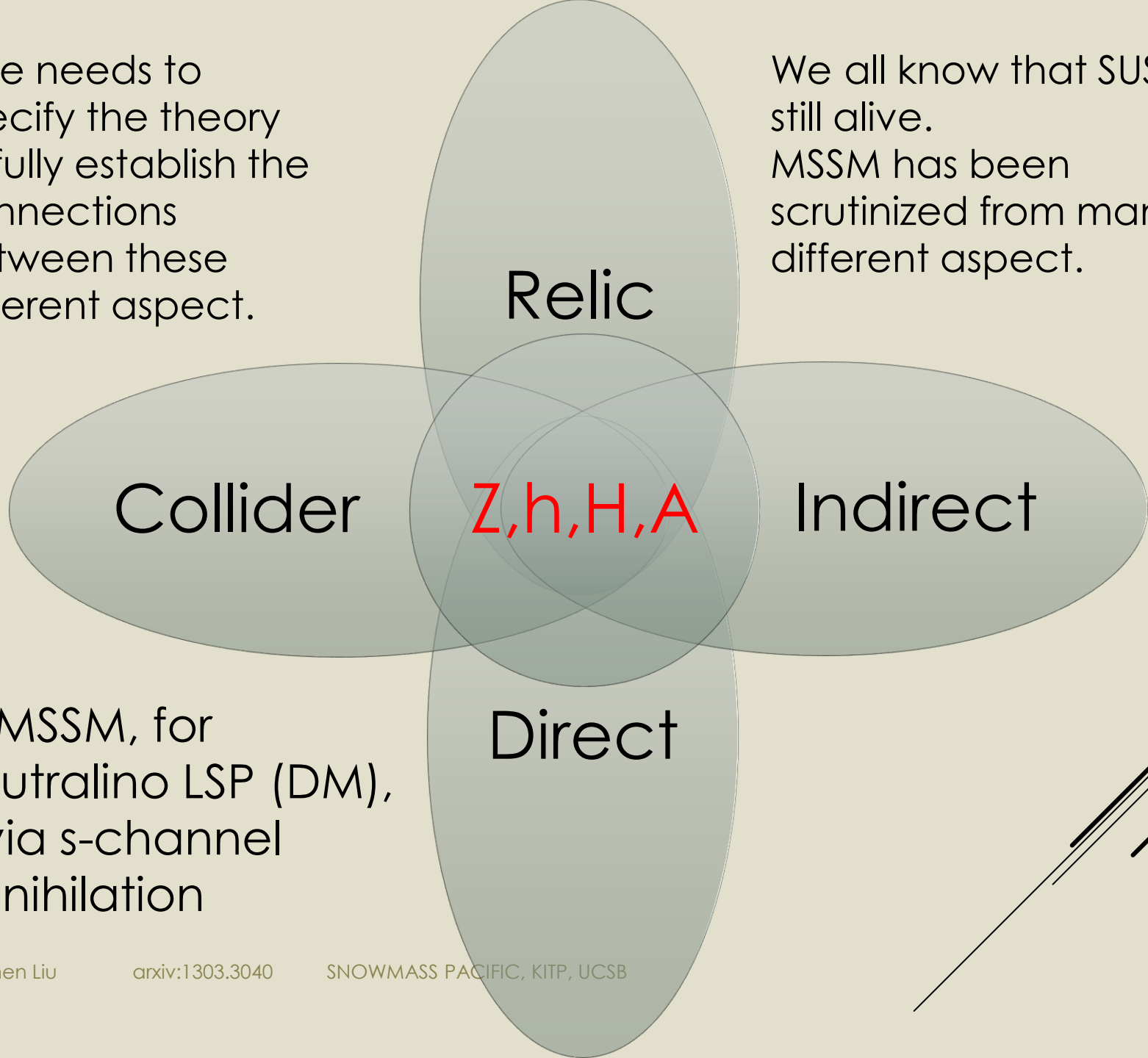


SNOWMASS
PACIFIC
KITP,UCSB

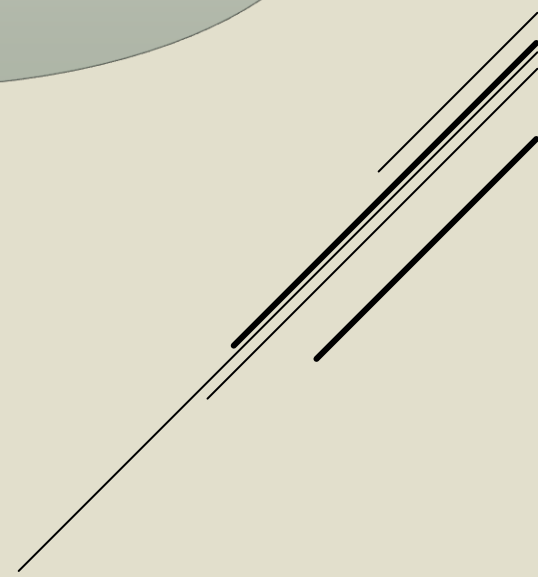


One needs to specify the theory to fully establish the connections between these different aspect.

We all know that SUSY is still alive. MSSM has been scrutinized from many different aspect.



In MSSM, for neutralino LSP (DM),
If via s-channel annihilation



Neutralino Lightest Supersymmetric Particle (LSP)

$$\chi_1^0 = N_{11}\tilde{B} + N_{12}\tilde{W}_3 + N_{13}\tilde{H}_d + N_{14}\tilde{H}_u$$

$$\begin{bmatrix} M_1 & 0 & -m_Z \cos \beta \sin \theta_w & m_Z \sin \beta \sin \theta_w \\ 0 & M_2 & m_Z \cos \beta \cos \theta_w & -m_Z \sin \beta \cos \theta_w \\ -m_Z \cos \beta \sin \theta_w & m_Z \cos \beta \cos \theta_w & 0 & -\mu \\ m_Z \sin \beta \sin \theta_w & -m_Z \sin \beta \cos \theta_w & -\mu & 0 \end{bmatrix}$$

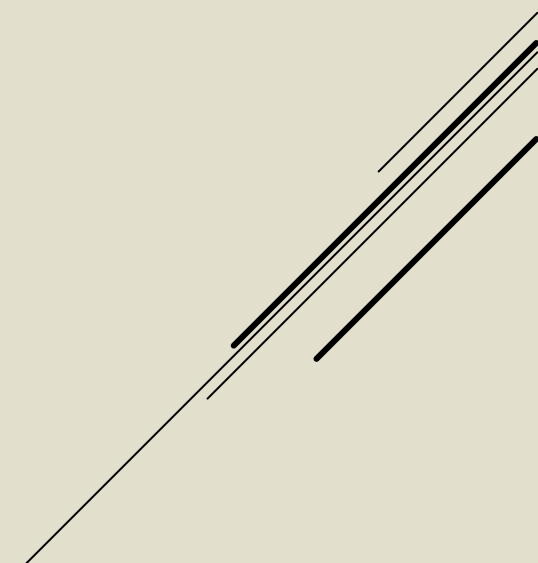
DM CANDIDATE

- ▶ Higgs discovery and other collider searches;
Including LEP, Tevatron and LHC.

Diphoton rate $> 80\%$ of SM value.

- ▶ B physics;
- ▶ The thermal relic density;
- ▶ DM direct detection.

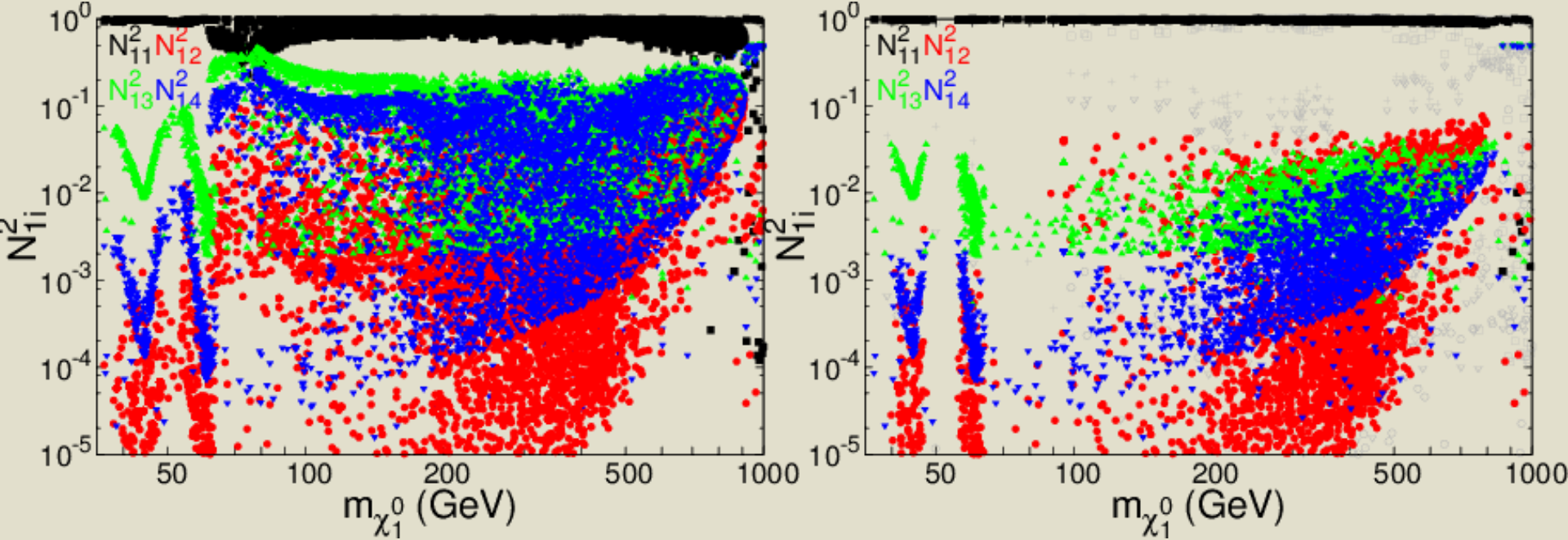
CONSTRAINTS



Parameters ranges:

$$\begin{aligned} 5 \text{ GeV} < |M_1| < 2000 \text{ GeV}, & \quad 100 \text{ GeV} < |M_2, \mu| < 2000 \text{ GeV}, \\ 3 < \tan \beta < 55, & \quad 80 \text{ GeV} < M_A < 1000 \text{ GeV}, \\ -4000 \text{ GeV} < A_t < 4000 \text{ GeV}, & \quad 100 \text{ GeV} < M_{Q3}, M_{U3} < 3000 \text{ GeV}, \\ -4000 \text{ GeV} < A_b < 4000 \text{ GeV}, & \quad 100 \text{ GeV} < M_{D3} < 3000 \text{ GeV}, \\ -4000 \text{ GeV} < A_\tau < 4000 \text{ GeV}, & \quad 100 \text{ GeV} < M_{L3}, M_{E3} < 3000 \text{ GeV}. \end{aligned}$$

SCAN SCHEME



Left: before XENON-100

Right: after XENON-100

Black, **Bino** fraction

Red, **Wino** fraction

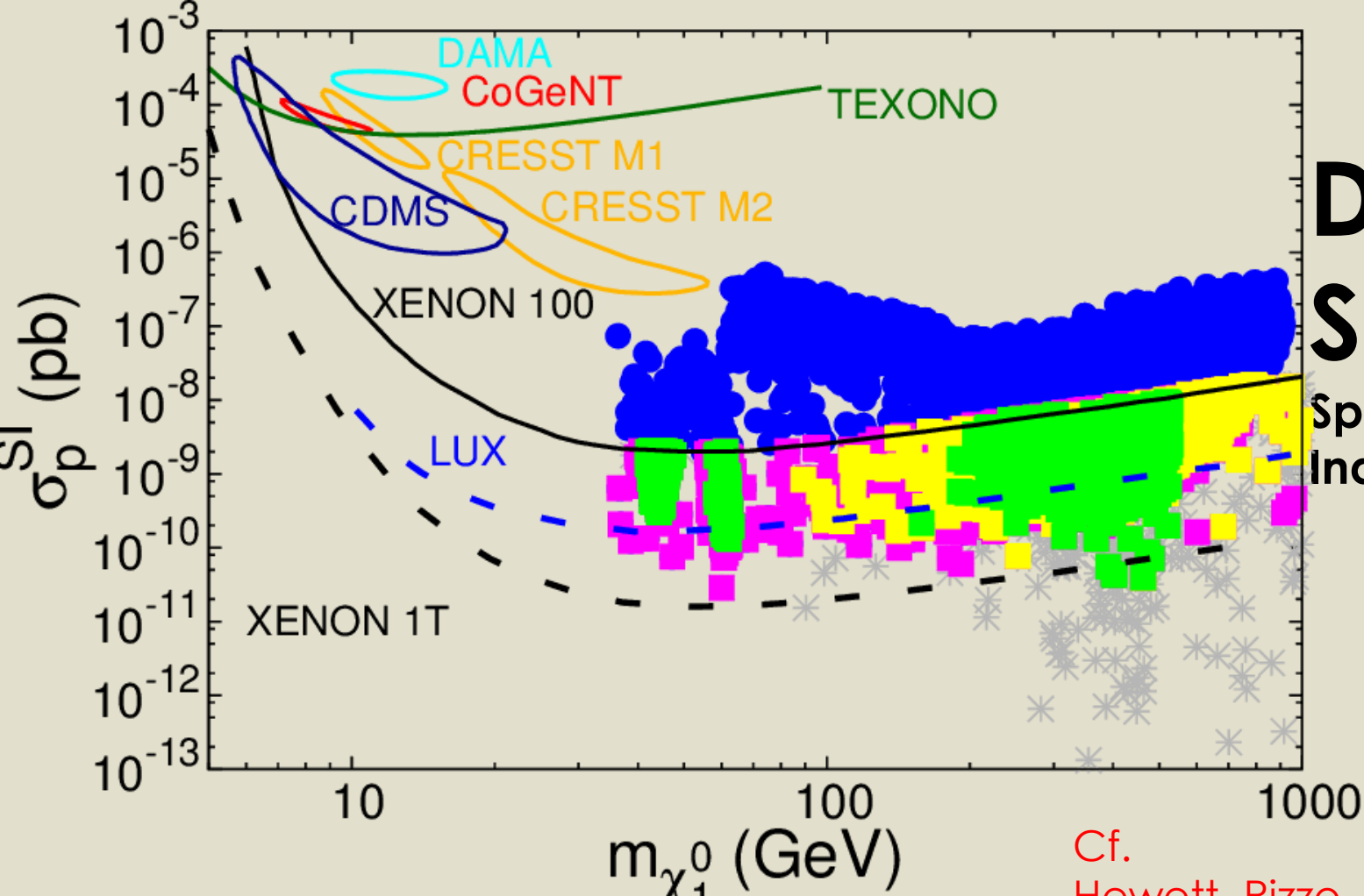
Green, Higgsino 1 (\tilde{H}_d) fraction

Blue, Higgsino 2 (\tilde{H}_u) fraction

Gaugino and Higgsino Fractions

Direct Search

Spin-Independent



Blue, excluded by XENON-100

Green, **Z-funnel, h-funnel, H/A-funnel**

Yellow, **neutralino/chargedino** coannihilation

Magenta, stermion assistance

—fermion coannihilation, t-channel exchange

Gray, “blind spots”

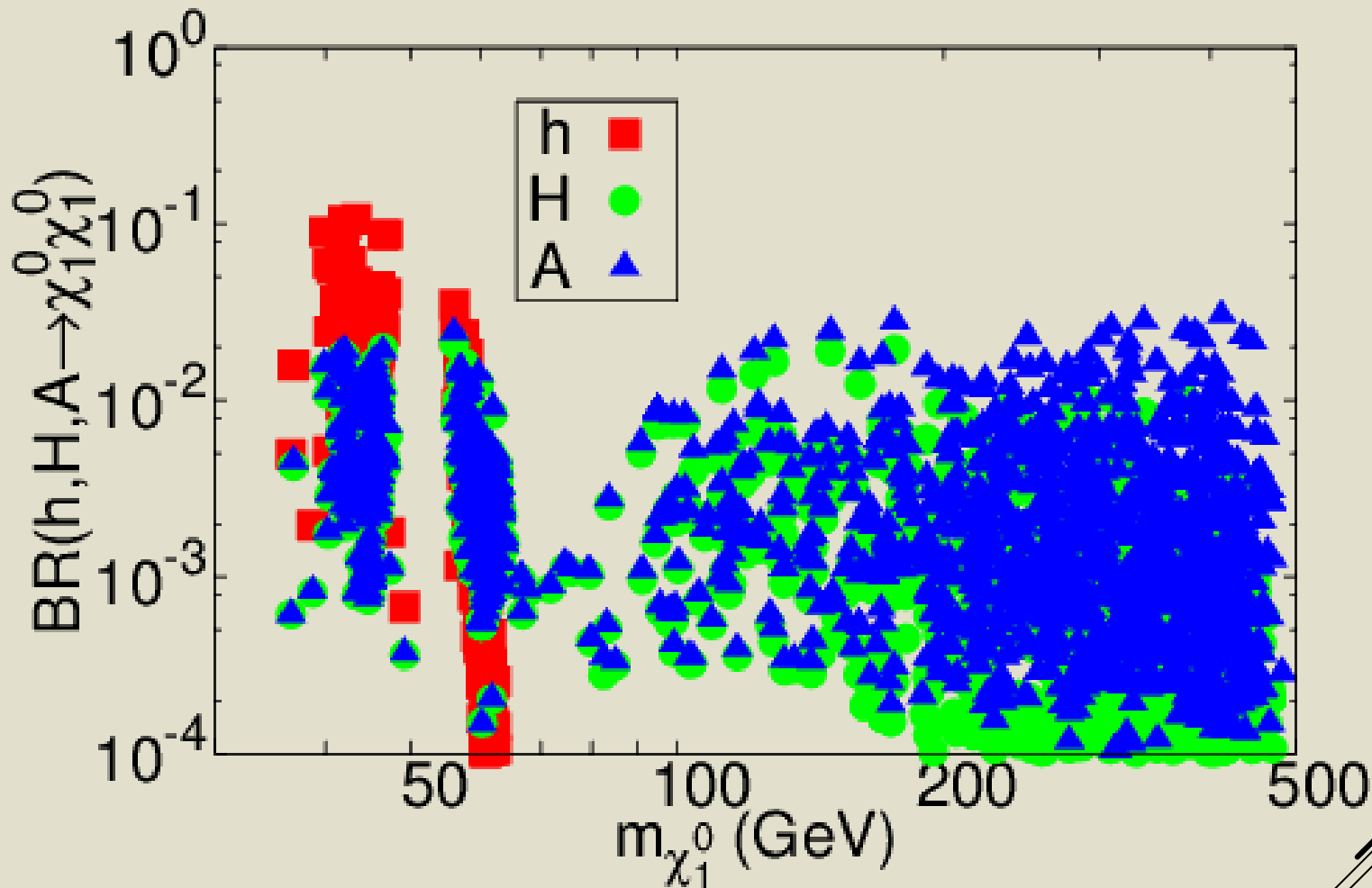
Cf.
 Hewett, Rizzo
 Arbey et al, 1205.2557
 Hooper et al, 1304.2417
 Baer et al, 1304.6732
 Chueng et al, 1211.4873

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arXiv:1308.3040

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Collider



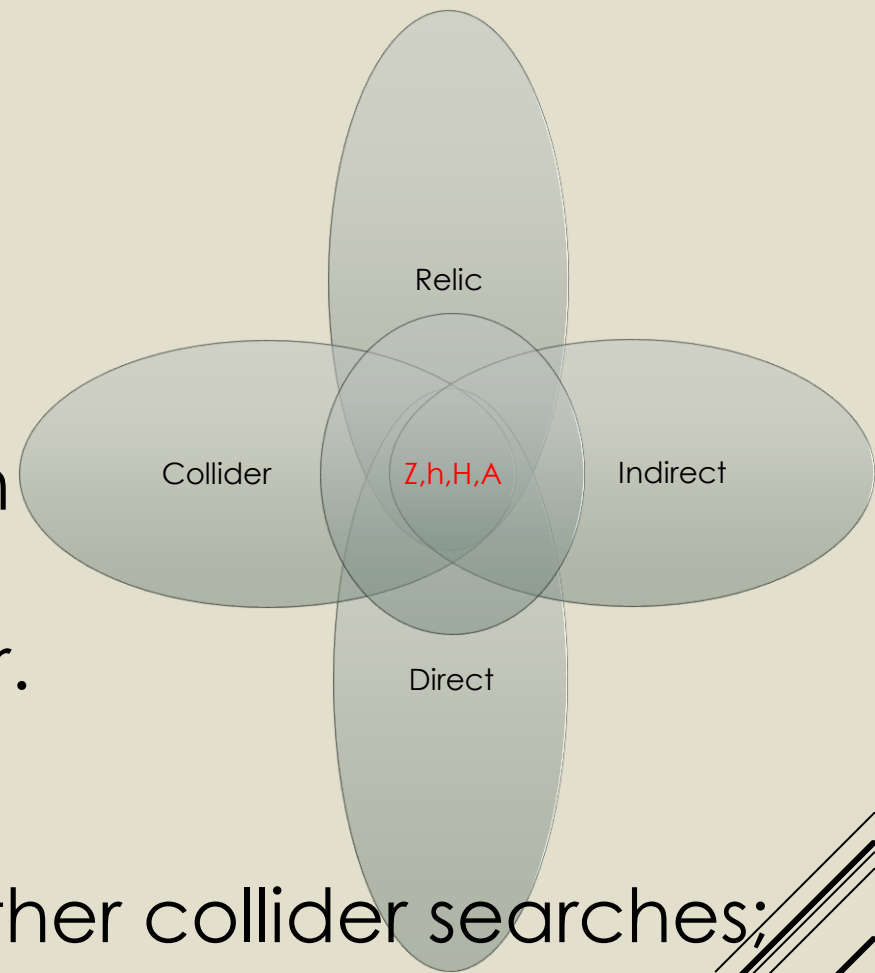
**Needs Precise Higgs Invisible Width
Measurement/Determination**

Type labels	DM mass $m_{\chi_1^0}$	Annihilation channels	Partial waves	$\langle\sigma v\rangle(v \rightarrow 0)$	Collider searches
I-A	$\sim m_Z/2$	$\chi_1^0\chi_1^0 \rightarrow Z$	p	low	$Z, h, H, A \rightarrow \chi_1^0\chi_1^0$
I-B	$\sim m_h/2$	$\chi_1^0\chi_1^0 \rightarrow h$	p	low	$h, H, A \rightarrow \chi_1^0\chi_1^0$
I-C	$\sim m_A/2$	$\chi_1^0\chi_1^0 \rightarrow A$	s	high	$H, A \rightarrow \chi_1^0\chi_1^0$
II-A	$m_{\chi_1^0} \sim m_{\chi_1^\pm}$ $\sim m_{\chi_2^0}$	$\chi_1^0\chi_2^0, \chi_1^0\chi_1^\pm$ $\chi_2^0\chi_2^0, \chi_1^+\chi_1^-$ $\rightarrow SM$	s+p	medium	$H, A \rightarrow \chi_1^0\chi_2^0$ $H, A \rightarrow \chi_2^0\chi_2^0$ $H^\pm \rightarrow \chi_1^0\chi_1^\pm$
II-B	$m_{\chi_1^0} \sim m_{\tilde{\tau}_1}$ $\sim m_{\tilde{\nu}_\tau}$	$\tilde{\tau}_1^+\tilde{\tau}_1^-, \tilde{\nu}_\tau\tilde{\nu}_\tau,$ $\chi_1^0\tilde{\tau}_1^\pm \rightarrow SM$	s+p	medium	$H, A \rightarrow \tilde{\tau}_1^+\tilde{\tau}_1^-$ $H^\pm \rightarrow \tilde{\tau}_1^\pm\tilde{\nu}_\tau$

Simplified Summary Table

Summary

We performed a study on MSSM with a neutralino LSP as the full Dark Matter.



- (a) Higgs discovery and other collider searches;
- (b) B physics;
- (c) The thermal relic density;
- (d) DM direct detection.

Conclusion



- (1) DM Largely a Bino-like neutralino
- (2) Low Mass DM disfavored;
- (3) Funnel and Coannihilation;
- (4) Lower bound on σ_{SI} for funnels
- (5) Indirect search will hit A-funnel
- (6) Higgs decays to neutralino, chargino pairs.

Limitations



Thermal DM (no non-Thermal);
Neutralino DM (no gravitino DM);
Single DM (no multiple DM, dynamical DM, etc.);
Standard Cosmology/Astrophysics
(no special modes for the DM density profile—
uncertainties, etc.)

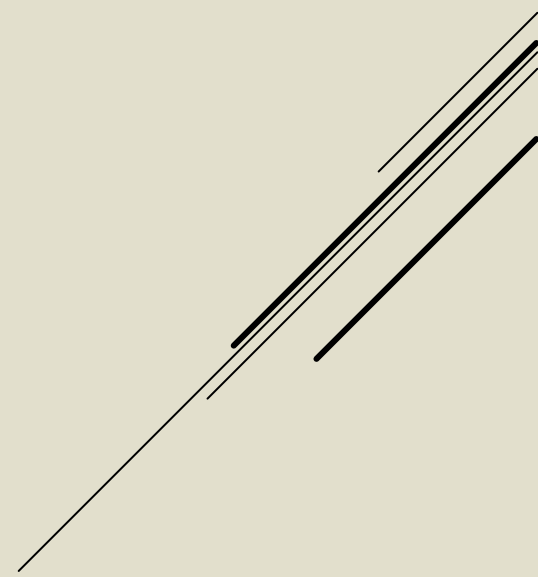
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BACKUP

Zhen Liu

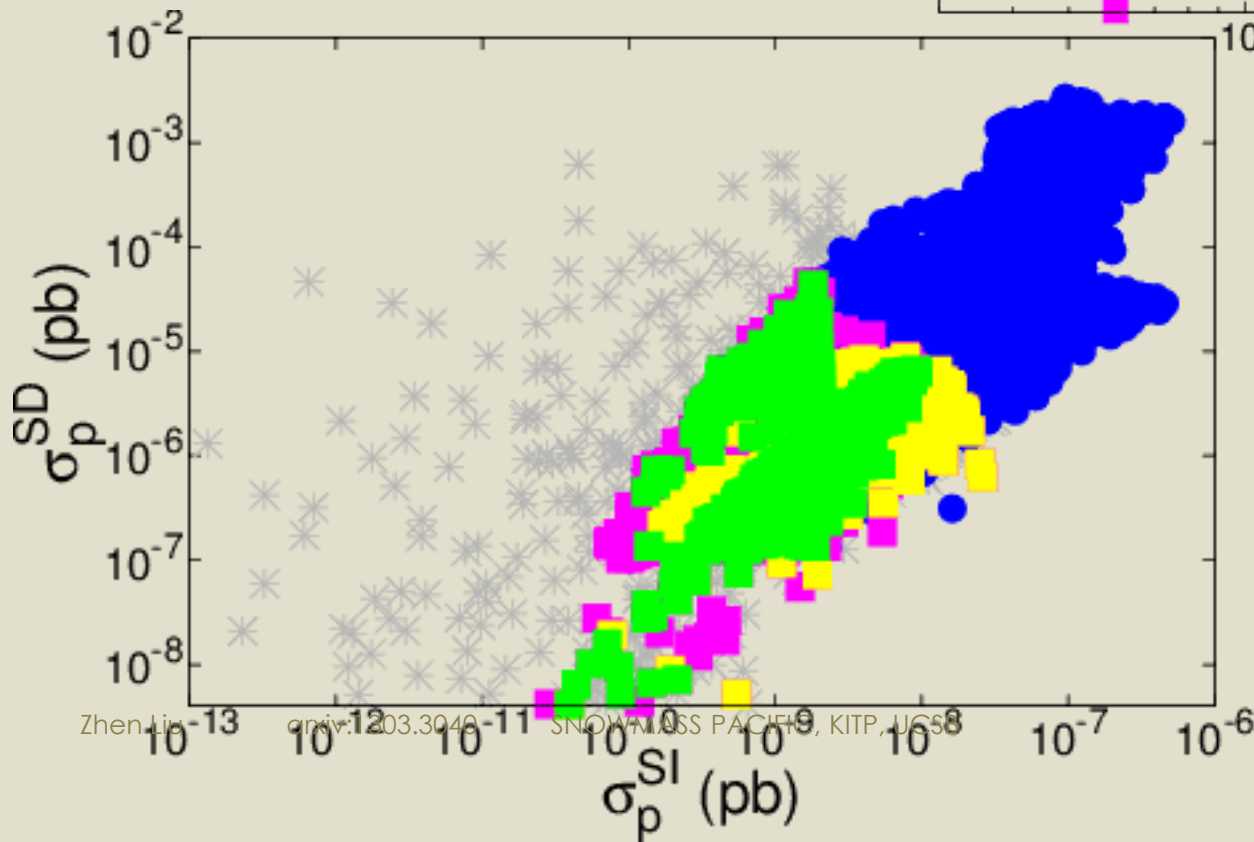
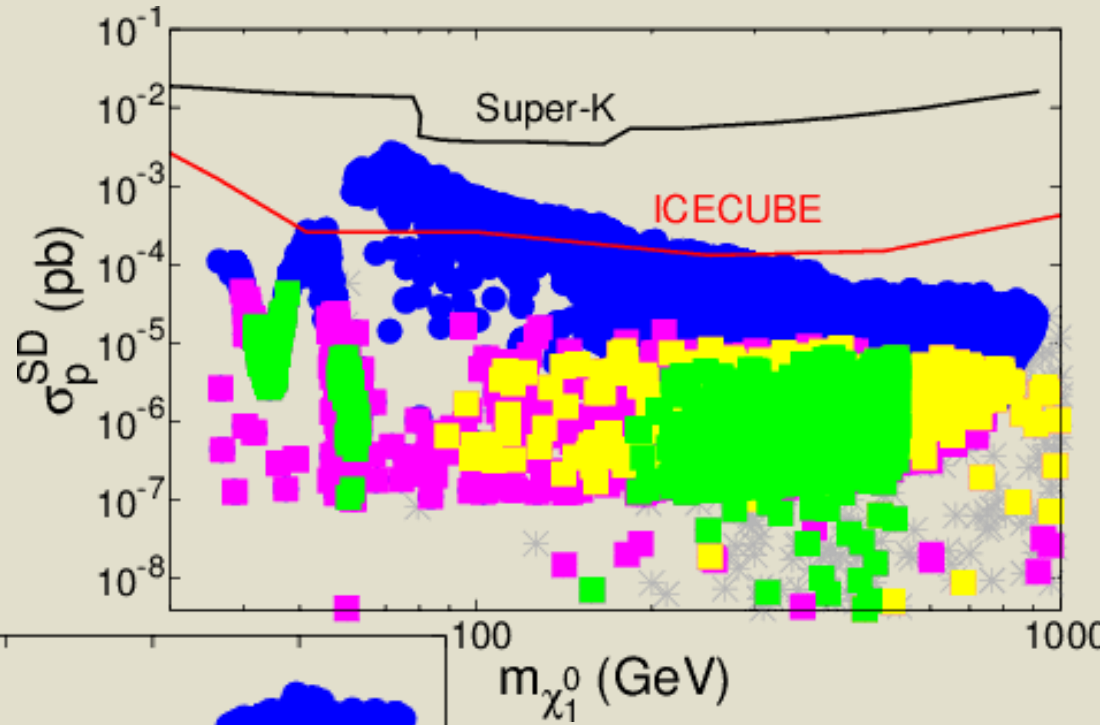
[arxiv:1303.3040](https://arxiv.org/abs/1303.3040)

SNOWMASS PACIFIC, KITP, UCSB

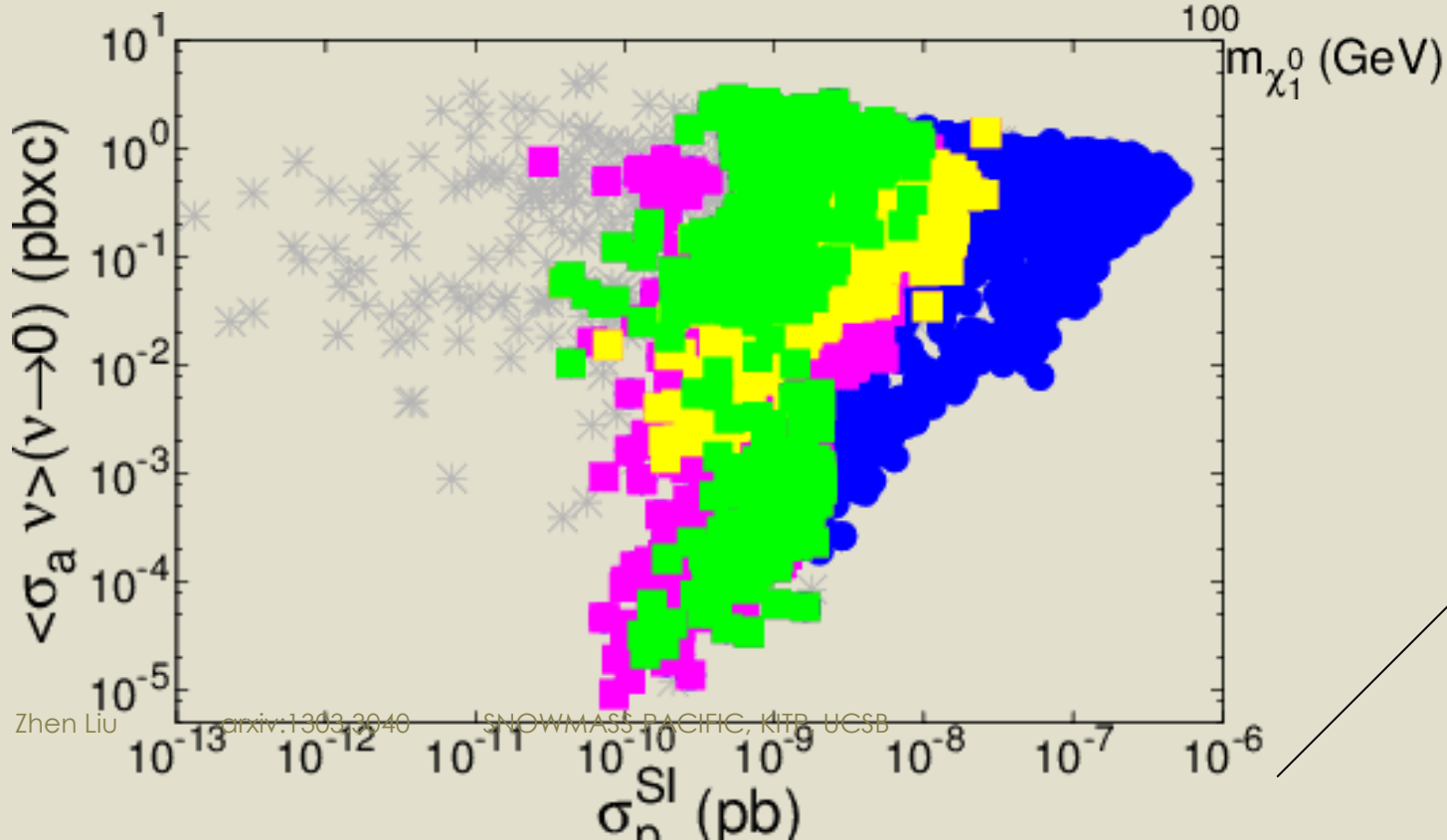
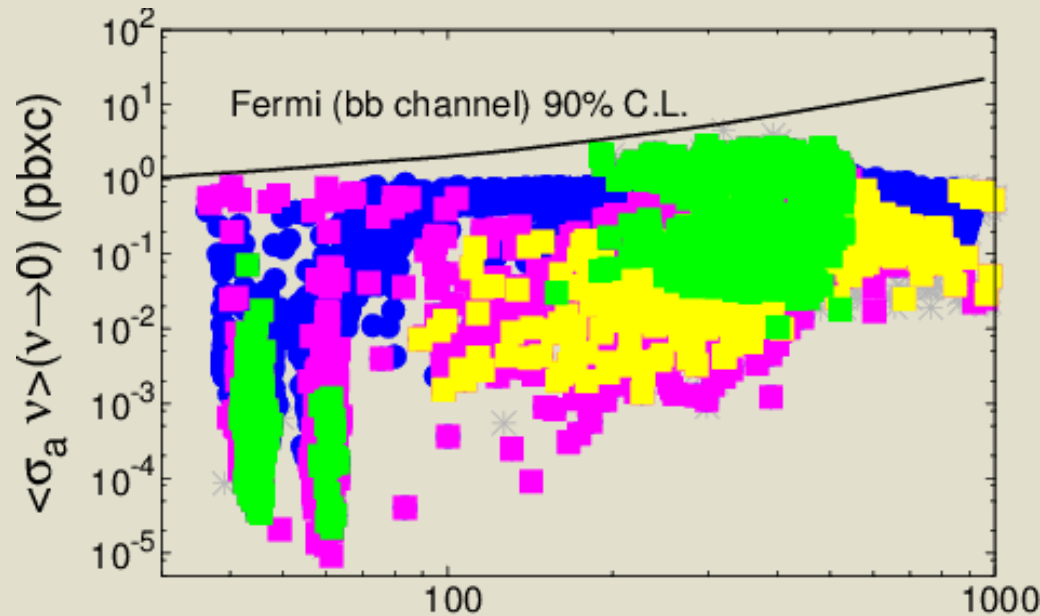


Direct Search

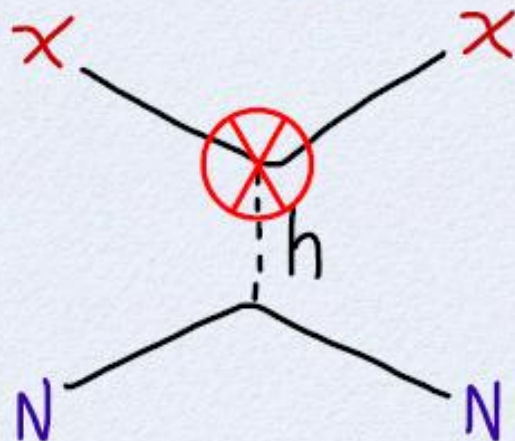
Spin-Dependent



Indirect Search

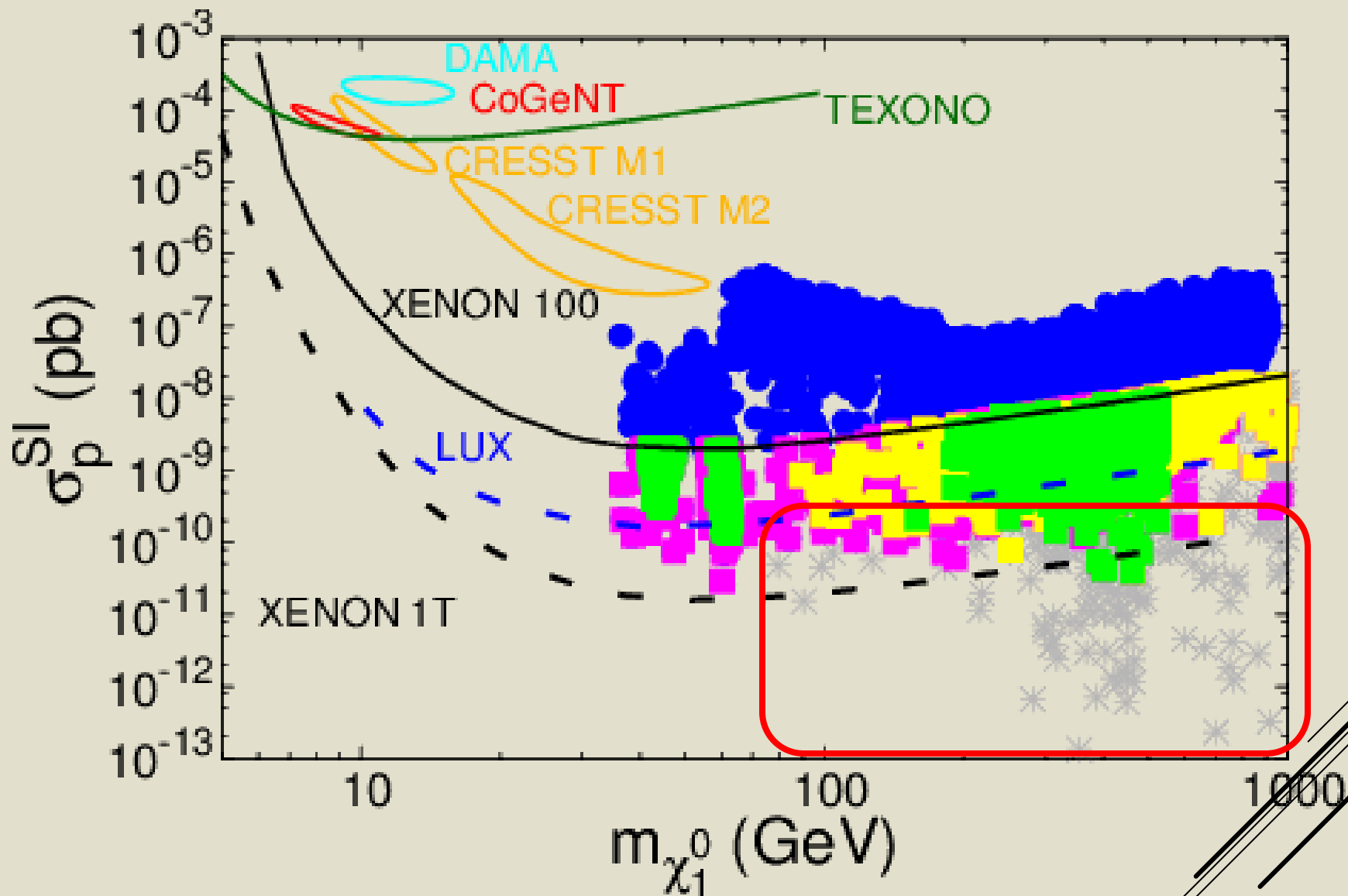


blindspots

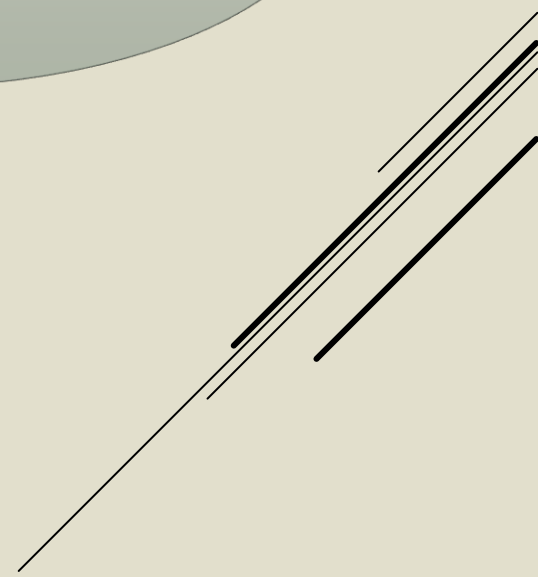
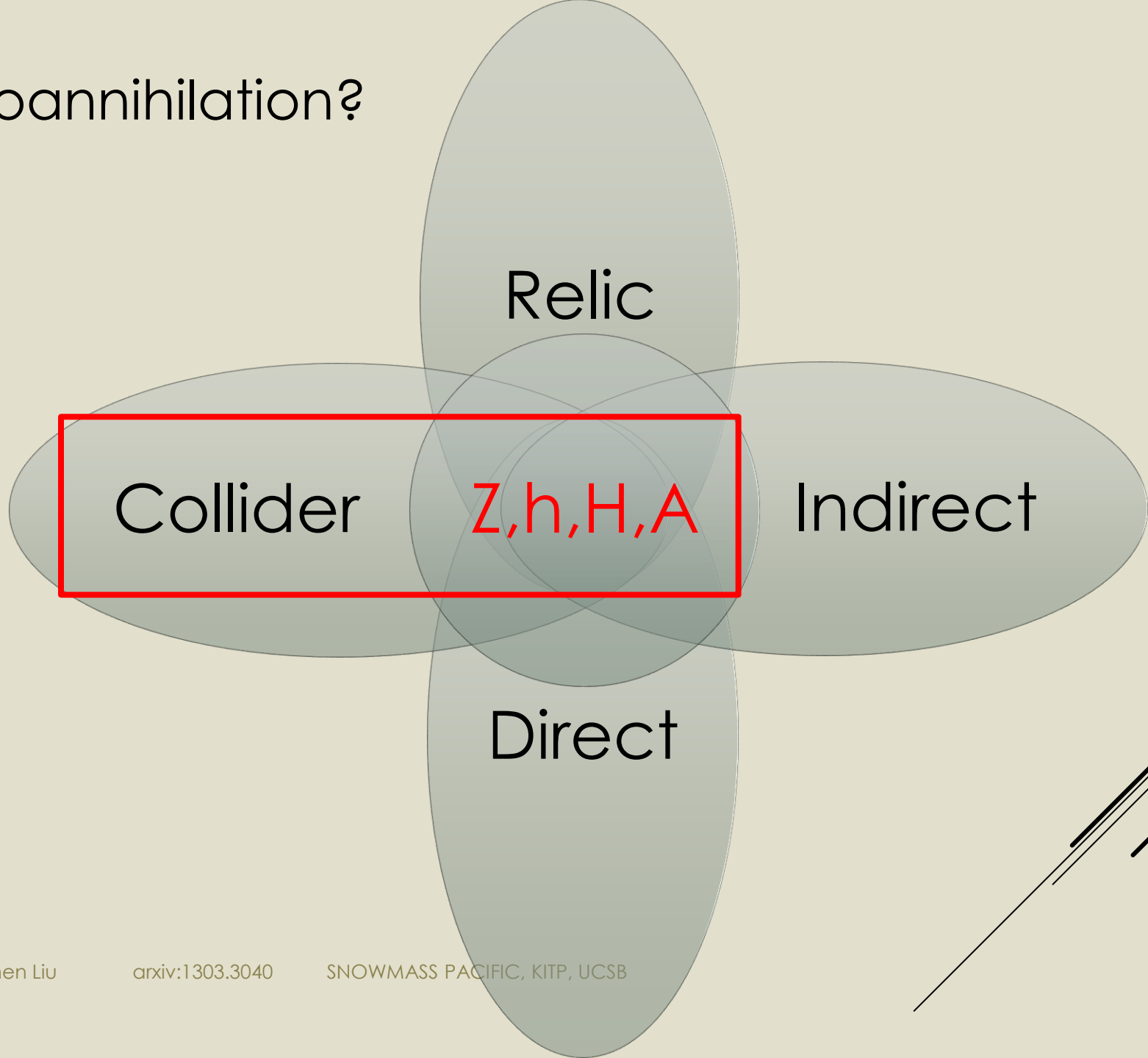


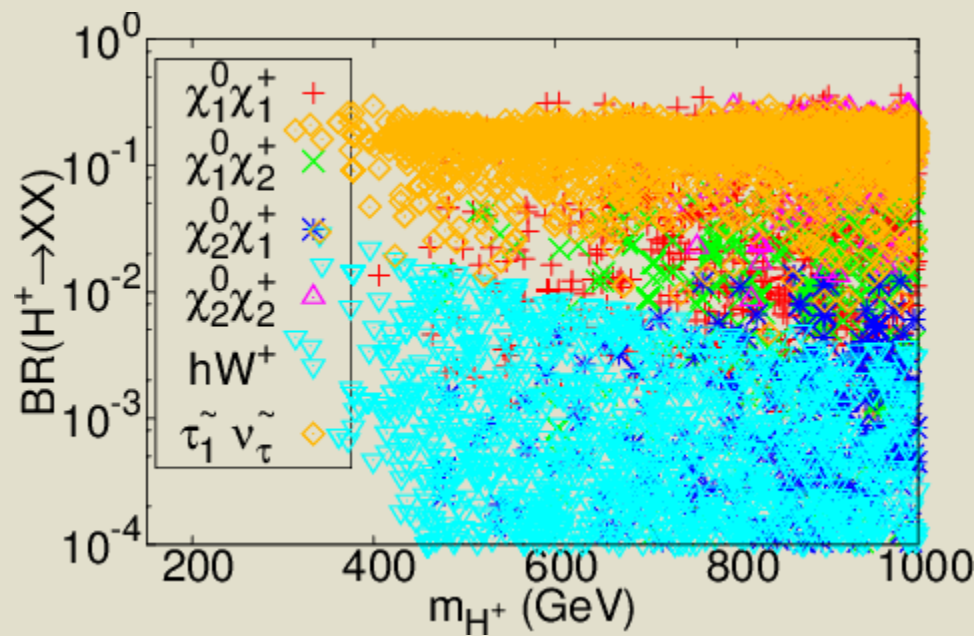
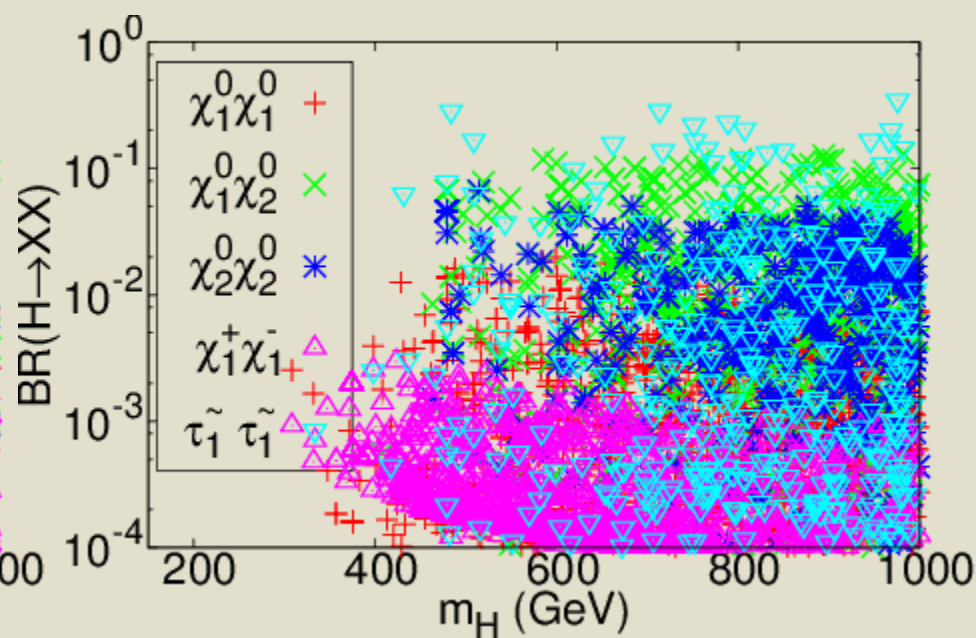
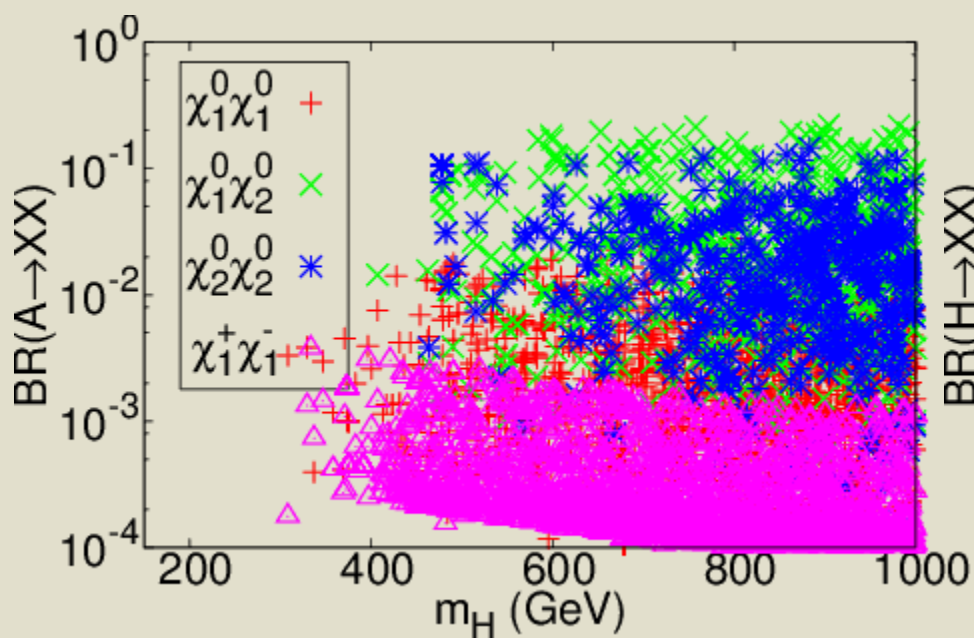
m_χ	condition	signs
bino M_1	$M_1 + \mu \sin 2\beta = 0$	$\text{sign}(M_1/\mu) = -1$
wino M_2	$M_2 + \mu \sin 2\beta = 0$	$\text{sign}(M_2/\mu) = -1$
higgsino $-\mu$	$\tan \beta = 1$	$\text{sign}(M_{1,2}/\mu) = -1$
bino/wino M_2	$M_1 = M_2$	$\text{sign}(M_{1,2}/\mu) = -1$

From Josh Runderman's talk
at Texas A&M DM workshop,
See more "blind spots" on
Arxiv: 1211.4873



Coannihilation?



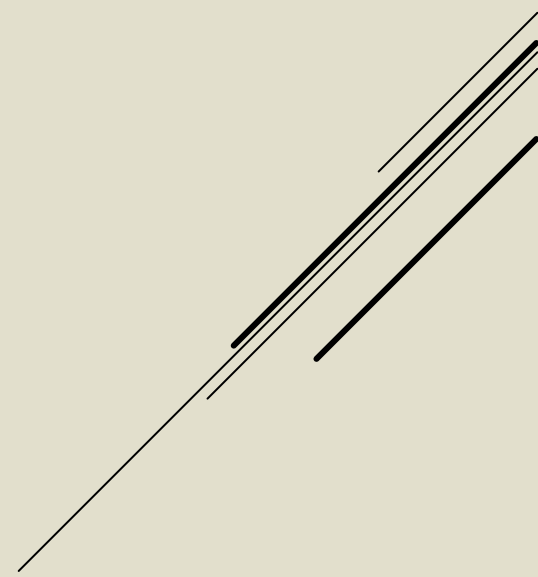


Higgs Decays

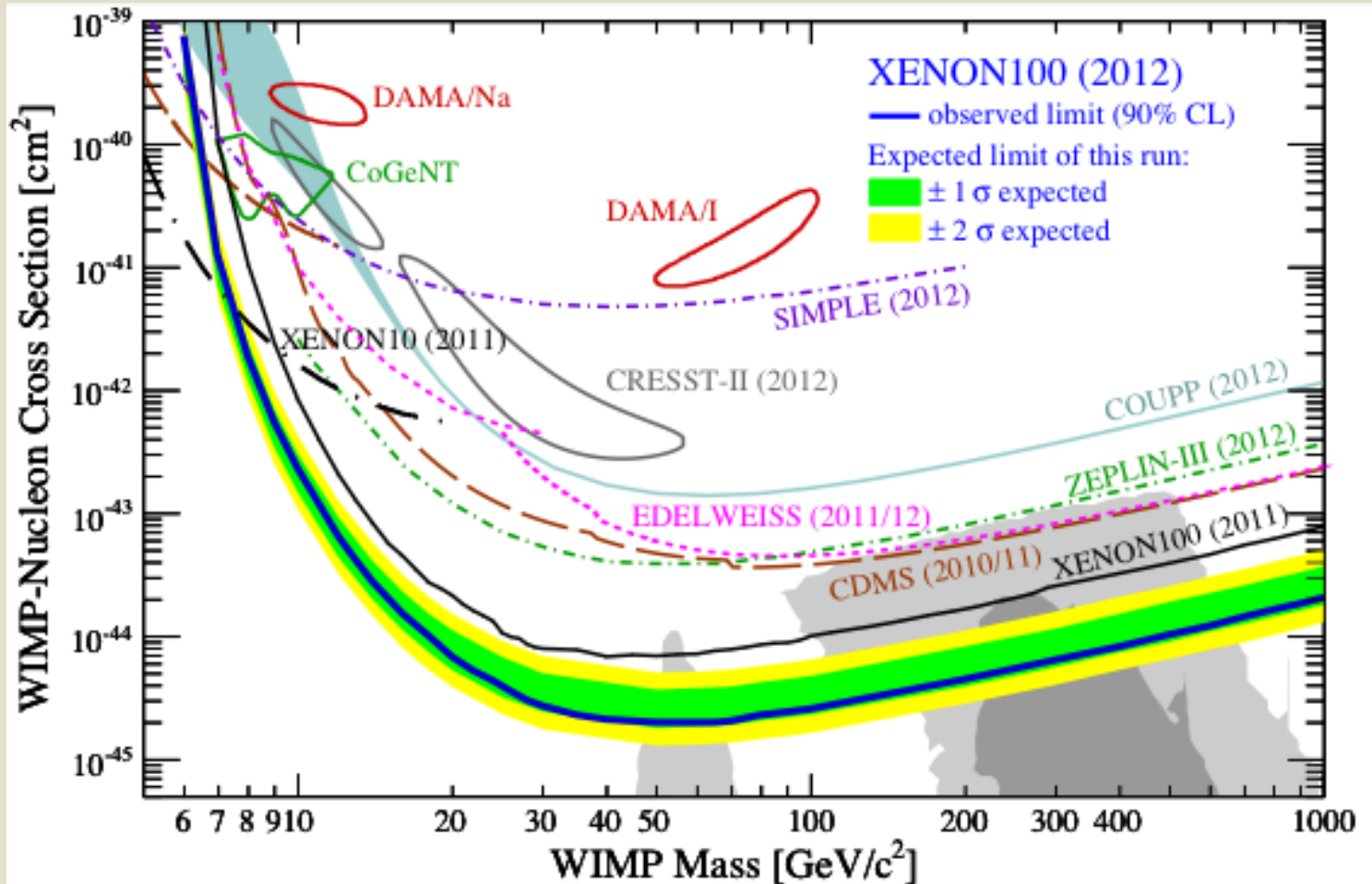
Relic Density

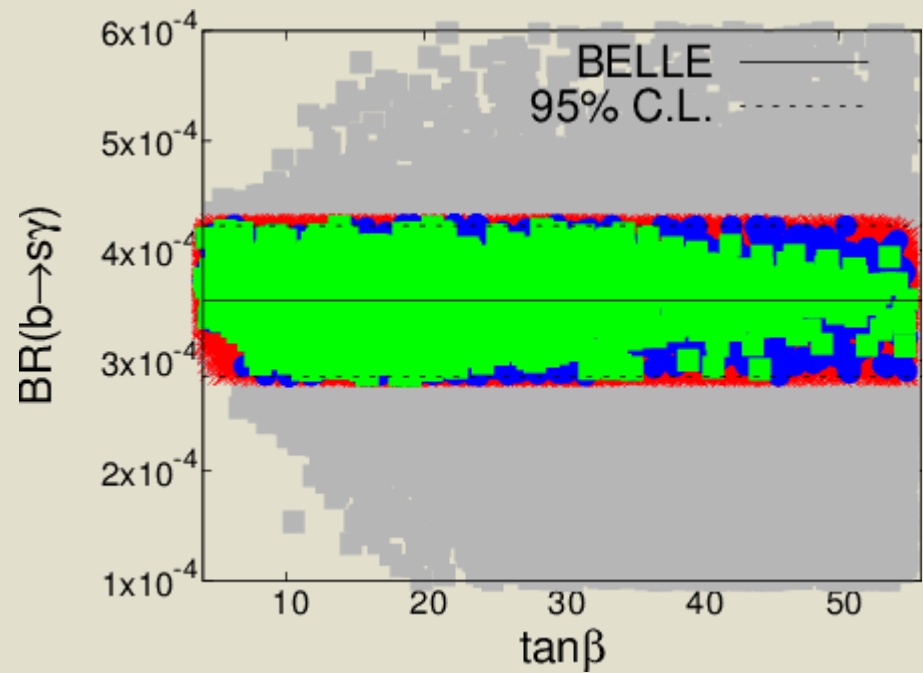
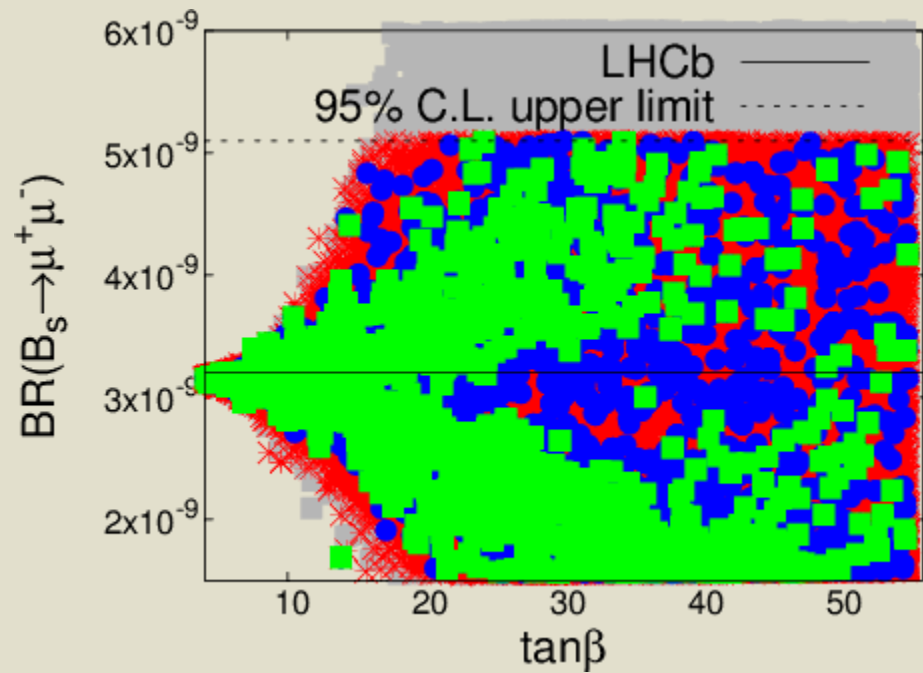
WMAP9/ACT/SPT...+10% theo. Uncertainty:

$$0.0915 < \Omega_{\chi_1^0} h^2 < 0.1381$$

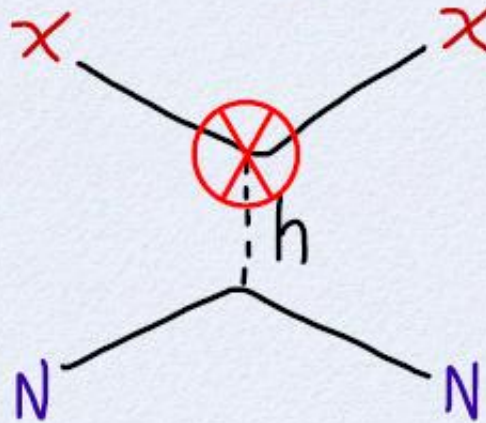


Direct DM Search





blindspots



m_χ	condition	signs
M_1	$M_1 + \mu \sin 2\beta = 0$	$\text{sign}(M_1/\mu) = -1$
M_2	$M_2 + \mu \sin 2\beta = 0$	$\text{sign}(M_2/\mu) = -1$
$-\mu$	$\tan \beta = 1$	$\text{sign}(M_{1,2}/\mu) = -1$
M_2	$M_1 = M_2$	$\text{sign}(M_{1,2}/\mu) = -1$

bino

wino

higgsino

bino/wino

SI cross-section for \tilde{b}/\tilde{h}

