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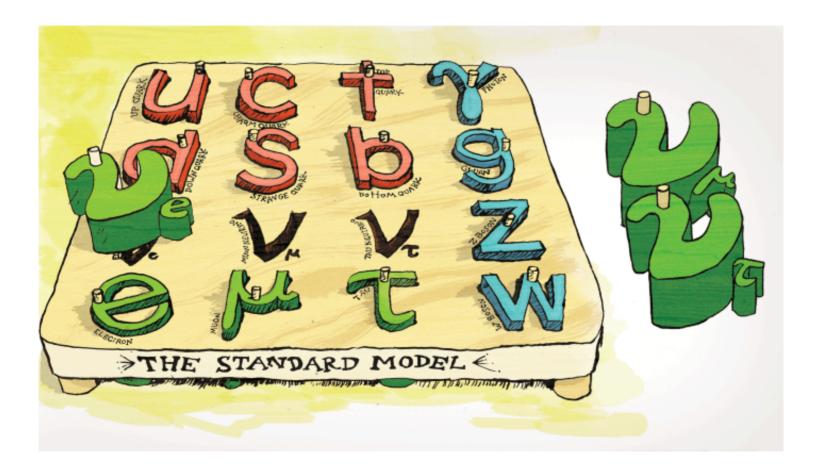
# Town Meeting – A Perspective Talk

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Snowmass In the Pacific UCSB - May 29-31, 2013

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#### I Will Concentrate on One Theme:



Where do Neutrino Masses Come From?

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André de Gouvêa Northwestern TeV mass (eV) 10 <sup>11</sup> b 10<sup>9</sup> GeV NEUTRINOS 10<sup>8</sup> 10 d 10 <sup>6</sup> HAVE MASS MeV е 10<sup>5</sup> 10 <sup>4</sup> [albeit very tiny ones...] 10<sup>3</sup> keV 10<sup>2</sup> 10 eV So What? 10 10  $v_2$ meV 10  $\bar{\nu}_1$ 10 10 2 3 fermion May 30, 2013 - $\nu$  Perspective

#### Neutrino Masses, EWSB, and a New Mass Scale of Nature

The LHC has revealed that the minimum SM prescription for electroweak symmetry breaking — the one Higgs double model — is at least approximately correct. What does that have to do with neutrinos?

The tiny neutrino masses point to three different possibilities.

- 1. Neutrinos talk to the Higgs boson very, very **weakly** (Dirac neutrinos);
- 2. Neutrinos talk to a **different Higgs** boson there is a new source of electroweak symmetry breaking! (Majorana neutrinos);
- 3. Neutrino masses are small because there is **another source of mass** out there a new energy scale indirectly responsible for the tiny neutrino masses, a la the seesaw mechanism (Majorana neutrinos).

Searches for  $0\nu\beta\beta$  help tell (1) from (2) and (3), the LHC, charged-lepton flavor violation, et al may provide more information.

## Accommodating Small Neutrino Masses – Seesaw Mechanism

If  $\mu = \lambda v \ll M$  (Dirac mass), below the mass scale M (right-handed neutrino Majorana mass),

$$\mathcal{L}_5 = \frac{LHLH}{\Lambda}.$$

Neutrino masses are small if  $\Lambda \gg \langle H \rangle$ . Data require  $\Lambda \sim 10^{14}$  GeV.

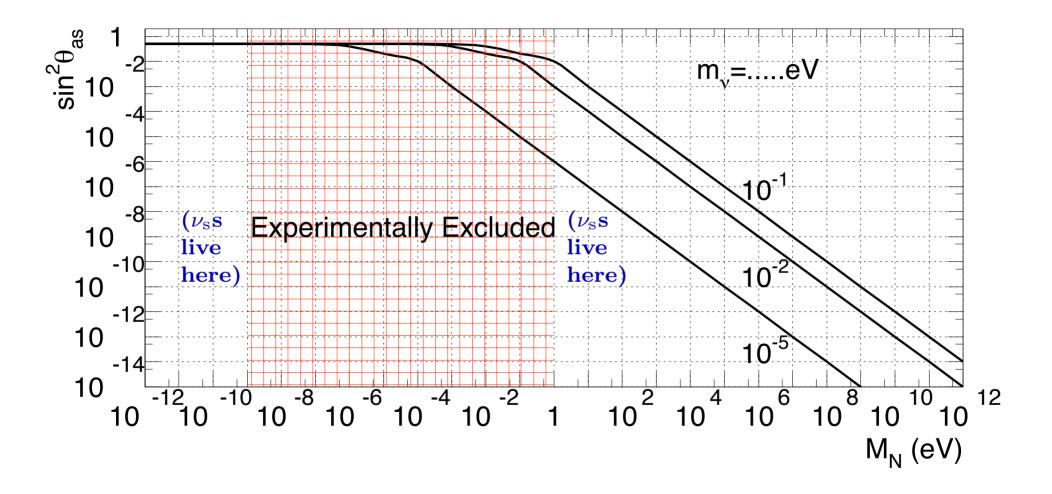
In the case of the seesaw,

$$\Lambda \sim \frac{M}{\lambda^2},$$

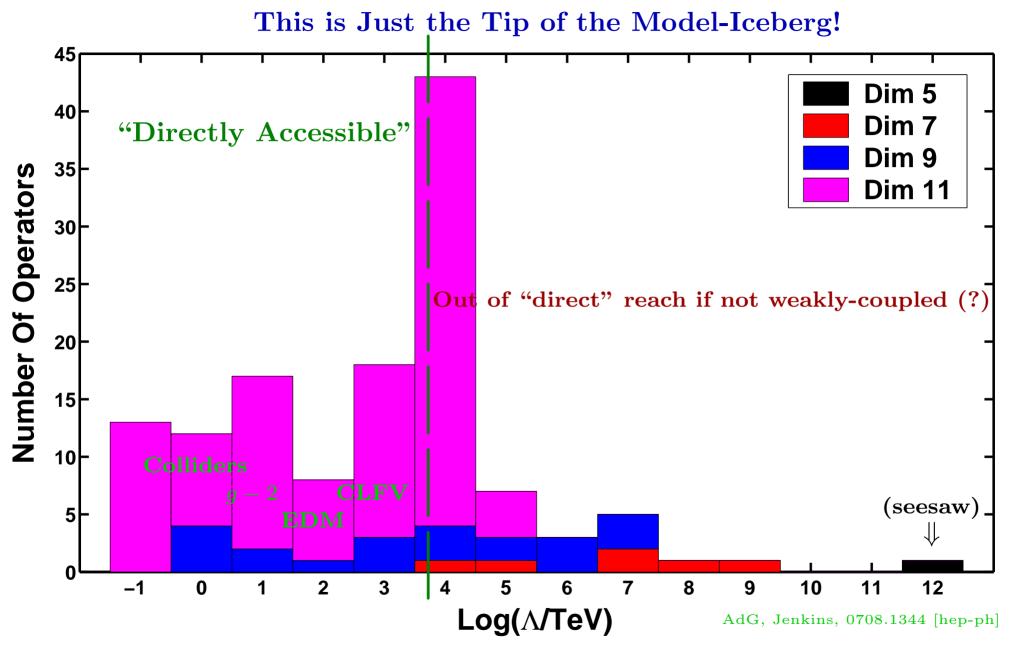
so neutrino masses are small if either

- they are generated by physics at a very high energy scale  $M \gg v$  (high-energy seesaw); or
- they arise out of a very weak coupling between the SM and a new, hidden sector (low-energy seesaw); or
- cancellations among different contributions render neutrino masses accidentally small ("fine-tuning").

#### Constraining the Seesaw Lagrangian



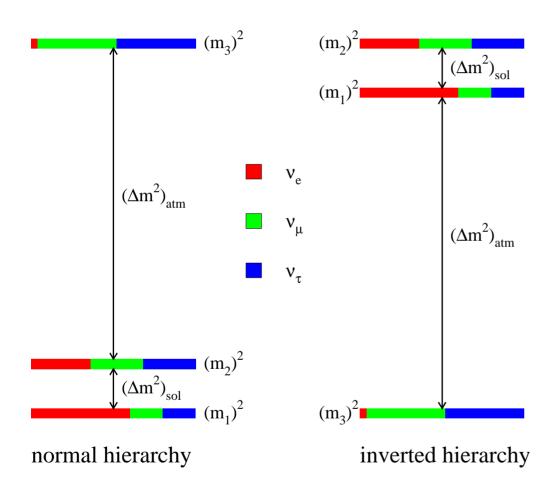
[AdG, Huang, Jenkins, arXiv:0906.1611]



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## What We Know We Don't Know: Missing Oscillation Parameters

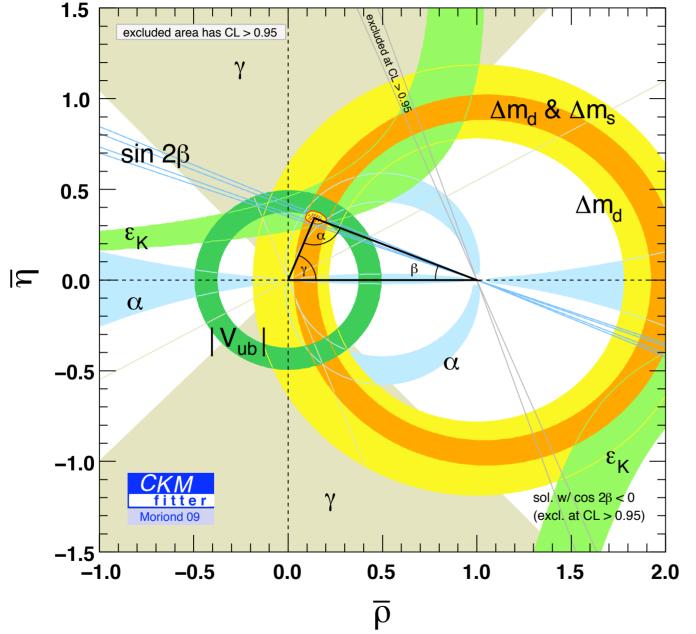


- What is the  $\nu_e$  component of  $\nu_3$ ?  $(\theta_{13} \neq 0!)$
- Is CP-invariance violated in neutrino oscillations?  $(\delta \neq 0, \pi?)$
- Is  $\nu_3$  mostly  $\nu_{\mu}$  or  $\nu_{\tau}$ ?  $(\theta_{23} > \pi/4, \theta_{23} < \pi/4, \text{ or } \theta_{23} = \pi/4?)$
- What is the neutrino mass hierarchy?  $(\Delta m_{13}^2 > 0?)$
- ⇒ All of the above can "only" be addressed with new neutrino oscillation experiments

Ultimate Goal: Not Measure Parameters but Test the Formalism (Over-Constrain Parameter Space)

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## What we ultimately want to achieve:



We need to do <u>this</u> in the lepton sector!

$$\left( egin{array}{c} 
u_e \\

u_{\mu} \\

u_{ au} 
\end{array} 
ight) = \left( egin{array}{ccc} 
U_{e1} & U_{e2} & U_{e3} \\
U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\
U_{ au 1} & U_{ au 2} & U_{ au 3} 
\end{array} 
ight) \left( egin{array}{c} 
u_1 \\

u_2 \\

u_3 
\end{array} 
ight)$$

What we have **really measured** (very roughly):

- Two mass-squared differences, at several percent level many probes;
- $|U_{e2}|^2$  solar data;
- $|U_{\mu 2}|^2 + |U_{\tau 2}|^2 \text{solar data};$
- $|U_{e2}|^2 |U_{e1}|^2 \text{KamLAND};$
- $|U_{\mu 3}|^2 (1 |U_{\mu 3}|^2)$  atmospheric data, K2K, MINOS;
- $|U_{e3}|^2(1-|U_{e3}|^2)$  Double Chooz, Daya Bay, RENO;
- $|U_{e3}|^2 |U_{\mu 3}|^2$  (upper bound  $\rightarrow$  hint) MINOS, T2K.

We still have a ways to go!

## CP-Violation in the Lepton Sector – Why Bother?

The SM with massive Majorana neutrinos accommodates **five** irreducible CP-invariance violating phases.

- One is the phase in the CKM phase. We have measured it, it is large, and we don't understand its value. At all.
- One is  $\theta_{QCD}$  term  $(\theta G\tilde{G})$ . We don't know its value but t is only constrained to be very small. We don't know why (there are some good ideas, however).
- Three are in the neutrino sector. One can be measured via neutrino oscillations. 50% increase on the amount of information.

We don't know much about CP-invariance violation. Is it really fair to presume that CP-invariance is generically violated in the neutrino sector solely based on the fact that it is violated in the quark sector? Why? Cautionary tale: "Mixing angles are small"

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#### Piecing the Neutrino Mass Puzzle

Understanding the origin of neutrino masses and exploring the new physics in the lepton sector will require unique **theoretical** and **experimental** efforts ...

- understanding the fate of lepton-number. Neutrinoless double beta decay!
- A comprehensive long baseline neutrino program. LBNE underground is necessary first step towards the ultimate "superbeam" experiment.
- The next-step is to develop a qualitatively better neutrino beam e.g. muon storage rings (neutrino factories).
- Different baselines and detector technologies a must for both over-constraining the system and looking for new phenomena.
- Probes of neutrino properties, including neutrino scattering experiments.
- Precision measurements of charged-lepton properties (g-2, edm) and searches for rare processes  $(\mu \to e\text{-conversion})$  the best bet at the moment).
- Collider experiments. The LHC and beyond may end up revealing the new physics behind small neutrino masses.
- Neutrino properties affect, in a significant way, the history of the universe (Cosmology). Will we learn about neutrinos from cosmology, or about cosmology from neutrinos?

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### **Synergisms**

I have not discussed the synergy between these types of activities and other aspects of fundamental particle physics and beyond, including

- A large liquid argon detector underground is a diverse long-term facility (think SuperK) to study several phenomena: nucleon decay, atmospheric and solar neutrinos, dark matter searches, neutrinos from supernova explosions . . .
- Neutrinos and their masses may be intimately connected to dark matter. Don't overlook the "neutrino portal"! Also connections to GUTs, dark energy, leptogenesis, and other new physics.
- Neutrinos are, at least, another piece to the flavor puzzle. Unique opportunity to test different ideas, explore potential correlations, including connections to the quark sector.
- Neutrino factories are a natural (necessary?) stepping stone towards a muon collider.

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