

What do we hope to learn this week?

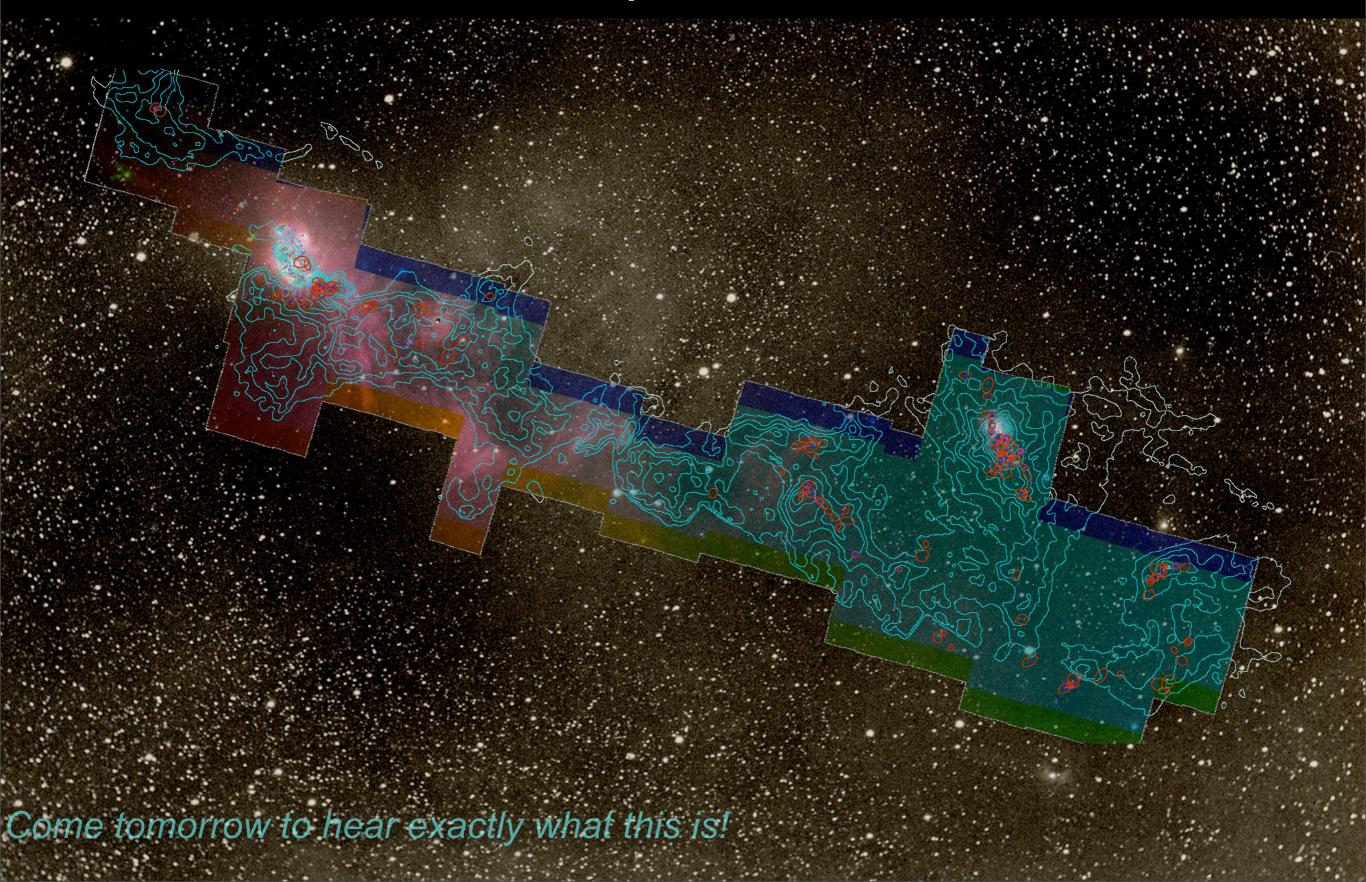
Come tomorrow to hear exactly what this is!

What do we hope to learn this week?



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What do we hope to learn this week?



C PLETE

Alyssa Goodman
Jens Kauffmann
Jaime Pineda
Scott Schnee
Rahul Shetty



Neal Evans
Melissa Enoch
Jens Kauffmann
Phil Myers
Deborah Padgett

FCRAO Taurus Survey

Paul Goldsmith

Spitzer Taurus Survey

Deborah Padgett

Spitzer Cluster Survey

Tom Megeath Phil Myers



Chris McKee, Paolo Padoan, Rahul Shetty & everyone here!



- Have we now found "all" the young stars in regions surveyed?
- What is the "present" mass function in these regions?
 Is that the IMF for those regions?
- What is the full story on multiplicity?
- What is the best way to measure stellar ages?
- What is the prevalence, and what are the properties, of outflows and shells?
- What, if anything, have we learned about timing and or sequence of star-formation?
- Why are protostars faint?
- What can we agree was the biggest stellar "surprise"?
- Stuff AG & CM forgot:



- What is the "true" column density distribution, and how does it vary over space? What does the "CMF" mean?
- What is the temperature distribution?
- What kinematic process(es) best describe gas motions at what scale(s)? (e.g. "turbulence," "shocks," "inflow," "infall," "outflow," etc.)
- What is the prevalence, and what are the properties, of outflows and shells?
- How bound is the gas, to what, and where?
- What's "between" clumps? What's the nature of the "interclump medium"?
- How important is the magnetic field?
- Can we tell how long various structures last?
- What can we agree was the biggest gaseous "surprise"?
- Stuff AG & CM forgot:



- What fraction of gas is converted to stars, when?
- What happens to the rest of the gas?
- How much does "environment" matter? Are star-forming cores always "coherent" and near-thermal?
- How do massive stars form?
- Role of feedback? (Kinematic, thermal, chemical.)
- When are stars "lost" from the "cloud"?
- Can we agree on the "Mouschovias (timing) Diagram's" labeling?
- Stuff AG & CM forgot:

What takes how long?

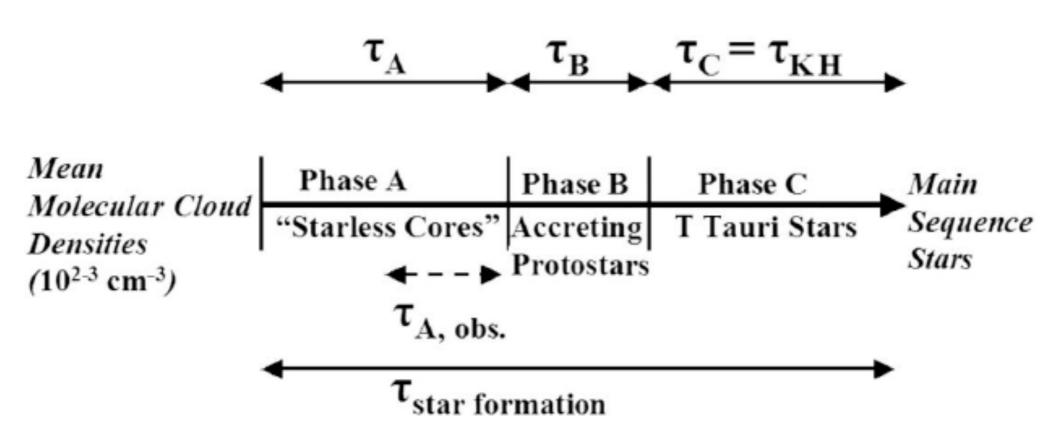


Fig. 1.-Star formation timeline (not to scale).

Q. What's the definition of "starless"?

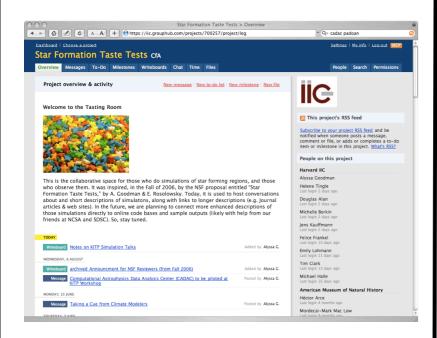
And what about (gas &) dust?

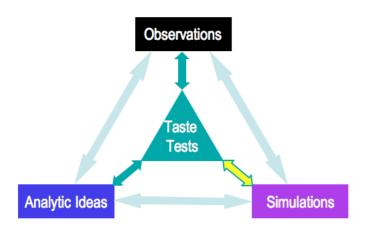
- Are gas & dust temperature coupled? Where?
- Are gas & dust well-mixed? Does dust "settle"?
- How does extinction vary with wavelength, and how much does that vary?
- How does opacity vary, overall, and with wavelength?
- Is there evidence for grain growth? Where?
- Stuff AG forgot:

Collaborative Tools

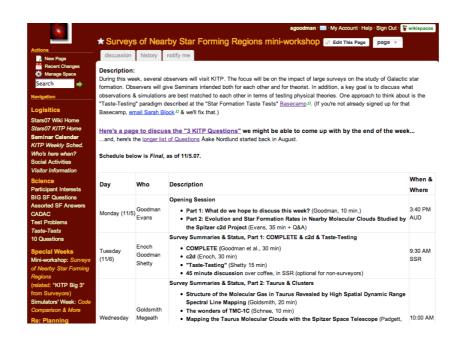
Note: the page icons are hyperlinked in the online version of these slides.

Taste-Testing "Basecamp" Goodman/Shetty



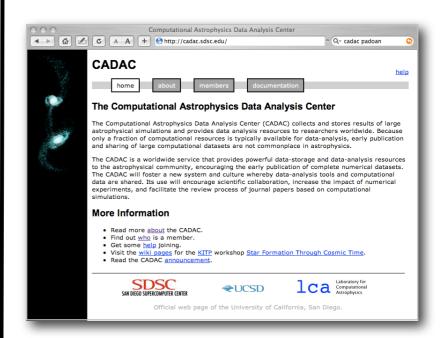


KITP Wiki Offner



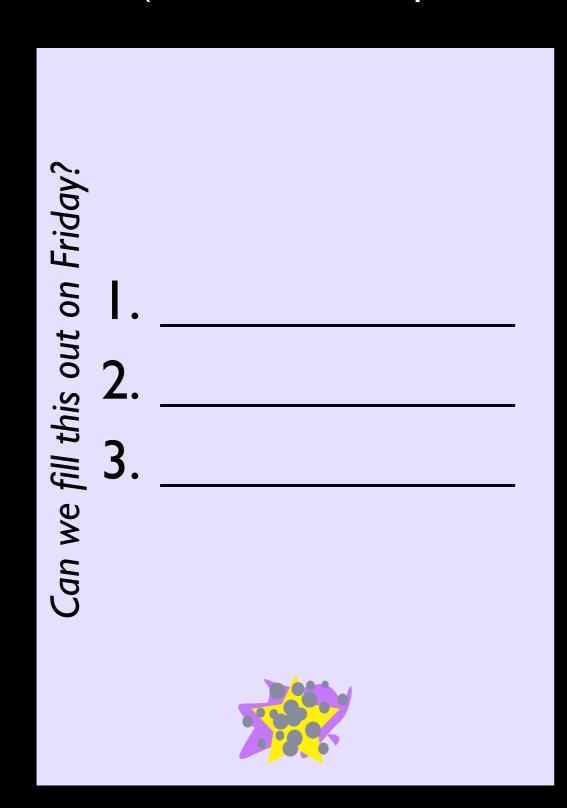
CADAC

Padoan



The "3 KITP Questions" about Star-Formation?

(ones we think present Surveys can best address)



On the KITP Wiki Now (from Åake Nordlund)

0. What is the nature of supersonic MHD turbulence in weakly ionized plasmas? What maintains the turbulence in molecular clouds? 2. What is the role of magnetic fields in star formation? 3. What sets the rate of star formation? What determines the IMF? How do massive stars form? How do star clusters form? How do cores collapse and disks accrete at various stages? OTHER QUESTIONS A. How do stars form in the Galactic Center? B. How do binaries form?