

# A Comparison of AMR and SPH codes for Galaxy Formation Simulations

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Goal: To study the chemical evolution of  
the universe from  $z=30 - 5$

Codes:

Enzo: Eulerian hydrodynamical adaptive Mesh Refinement  
(AMR)/N-body code (Norman & Bryan 1998)

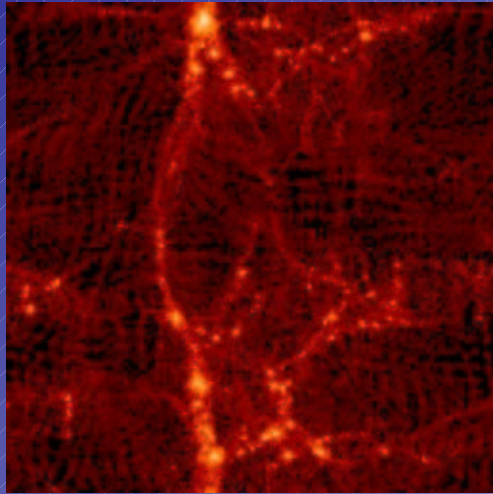
GADGET: Lagrangian Smooth Particle Hydrodynamics (SPH)/  
N-body code (Springel et al. 2001)

Code comparison:

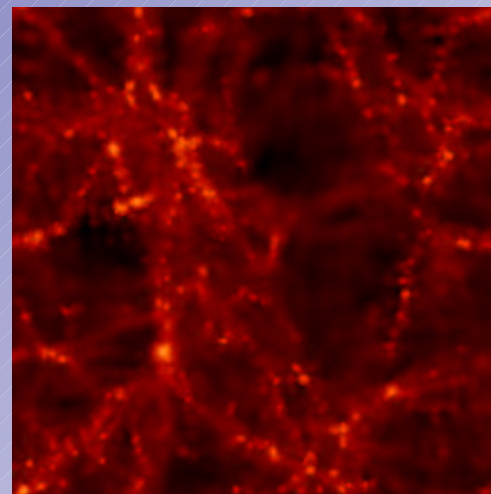
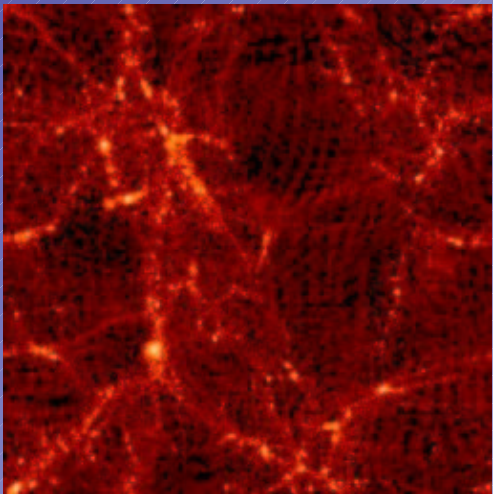
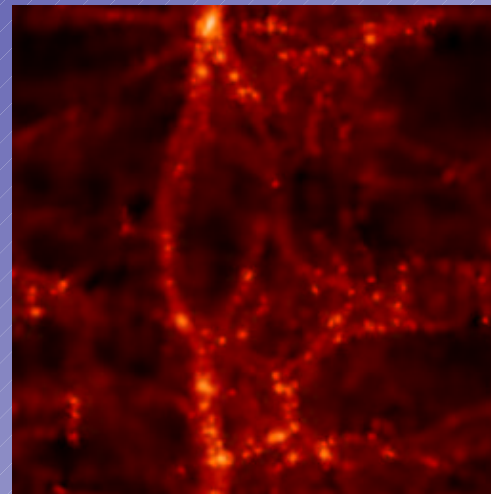
- Dark matter / adiabatic hydro comparison
- Radiative cooling
- Star formation and feedback

# Dark Matter–Only Comparison

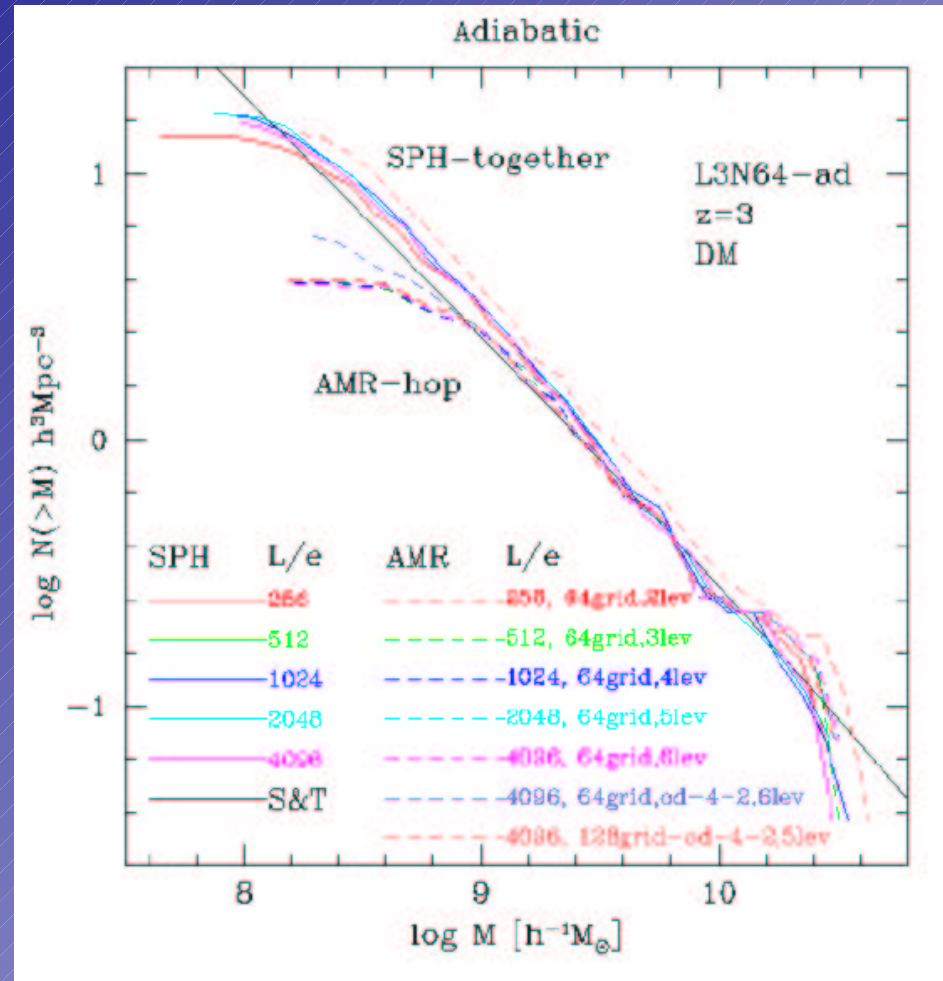
AMR



SPH

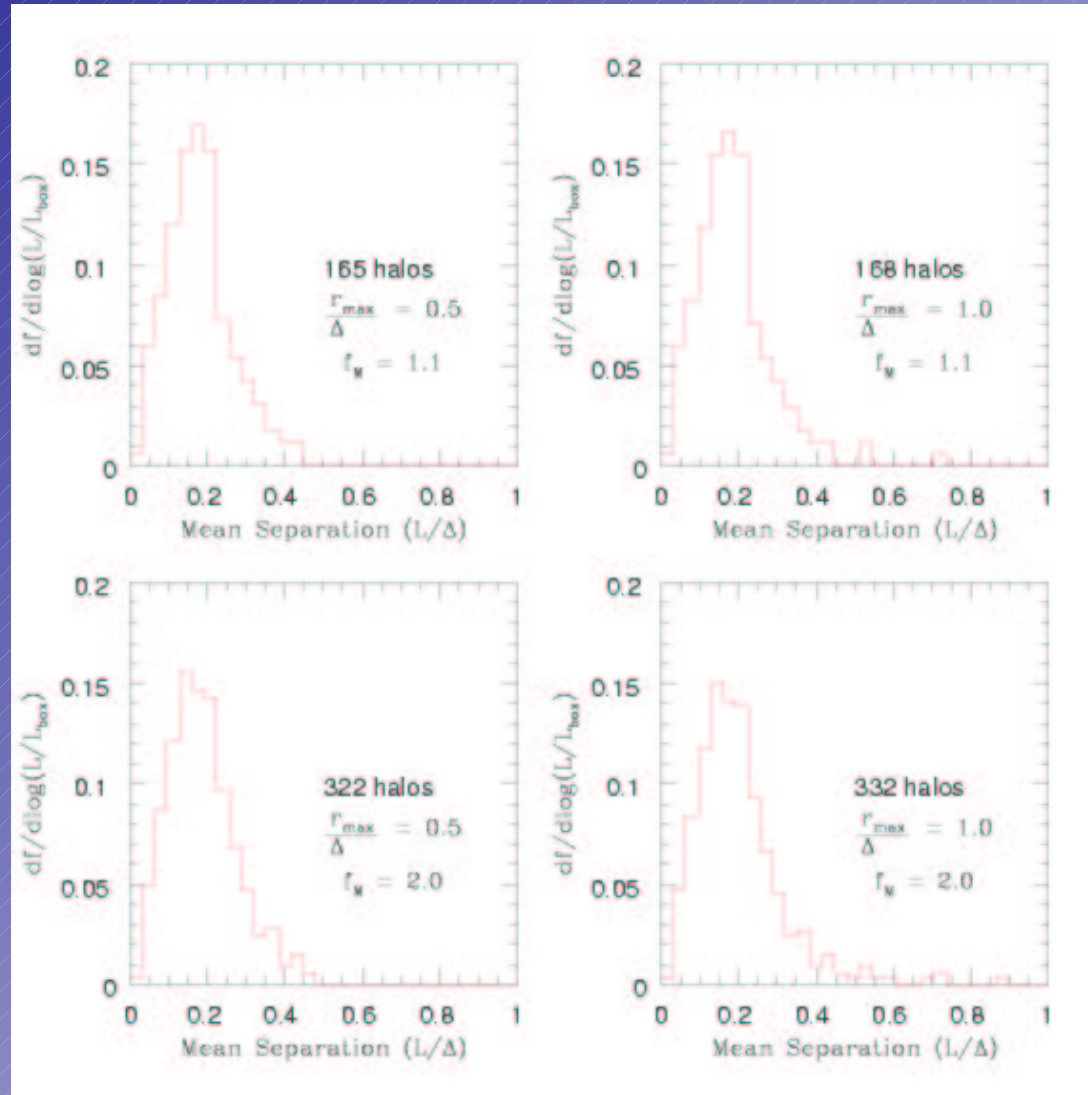


# Dark Matter Mass Function

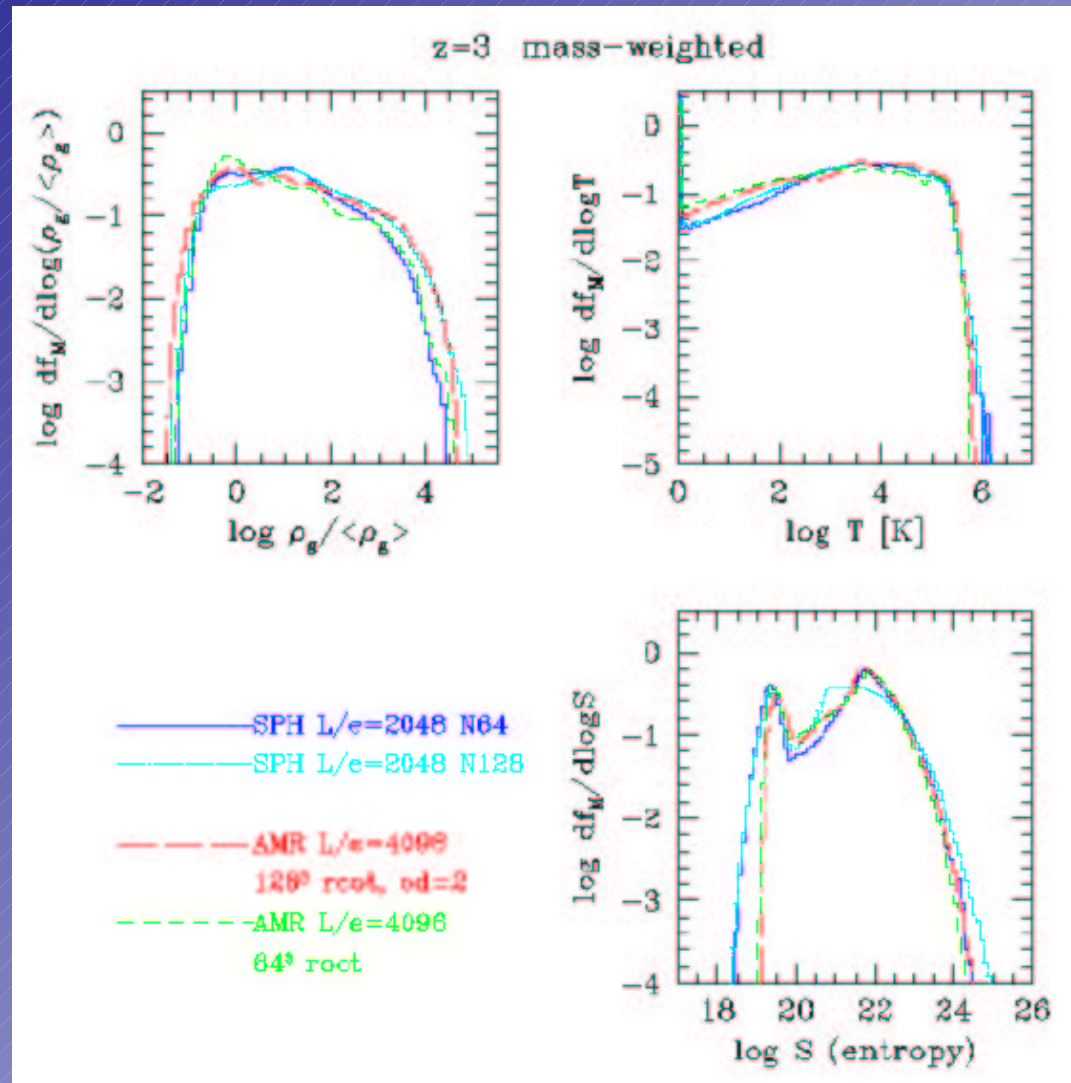


Best results found for  $64^3$  dm particles/ $128^3$  grid cells (AMR) for comparable resolution (due to PM algorithm)

# Mean Separation of DM Halo Peak Densities



# Baryon Distribution Functions



# Gas Mass Fraction

# Conclusions

- Initial results are better than we had expected
- To obtain comparable results the AMR mesh size must be twice the number of particles
- Quite a bit of work remains to be done – detailed comparisons
- Continued agreement with more physics will lend confidence to predictions made with either code