

Spiral wave chaos: Tiling, local symmetries, and asymptotic freedom

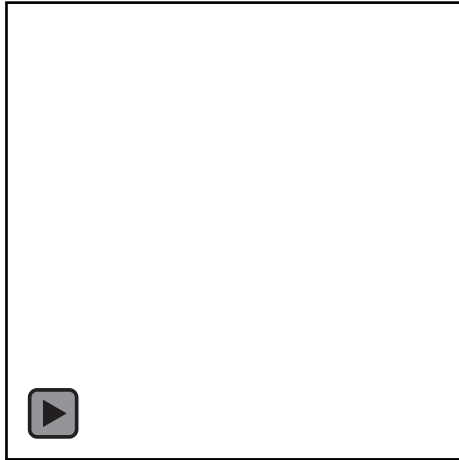
Roman Grigoriev, Chris Marcotte
Center for Nonlinear Science, School of Physics



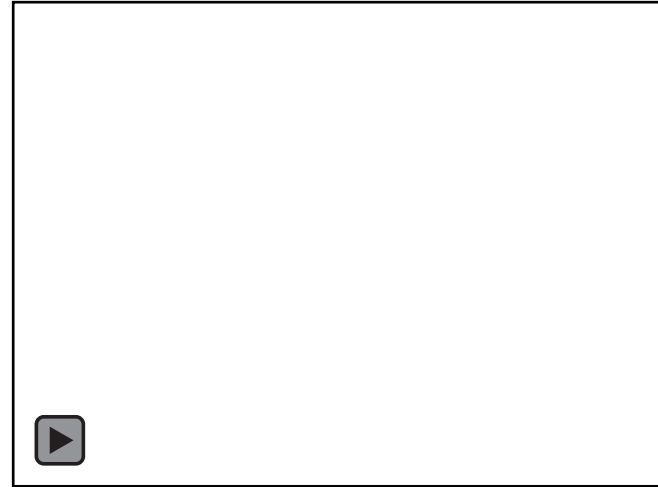
**Georgia
Tech**



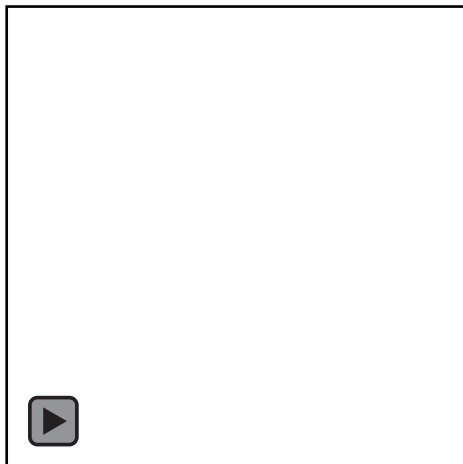
Spiral waves in excitable systems



BZ reaction



Slime mold



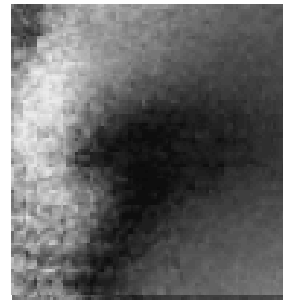
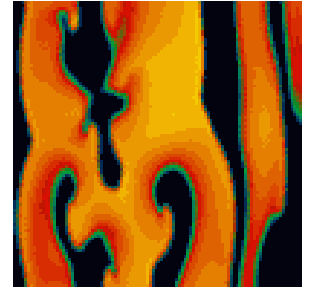
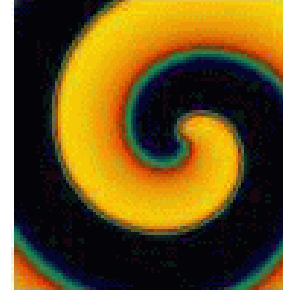
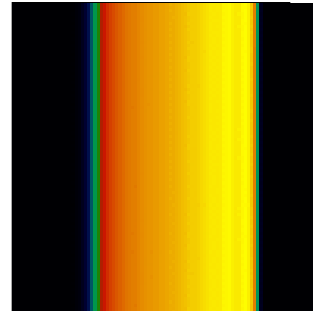
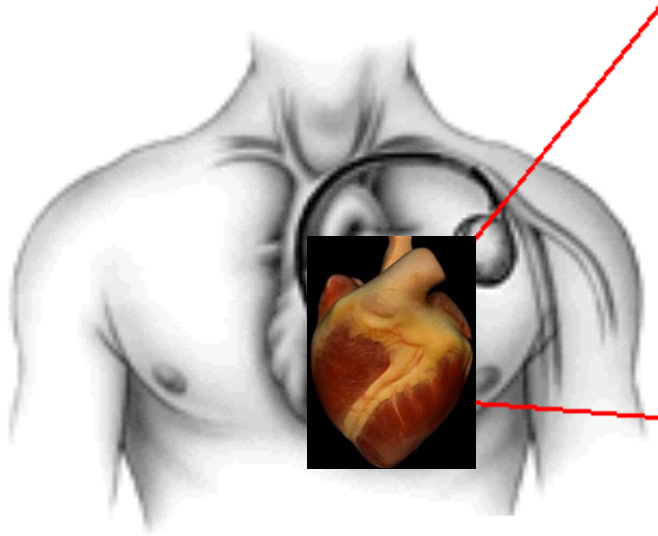
HL-1 cell culture



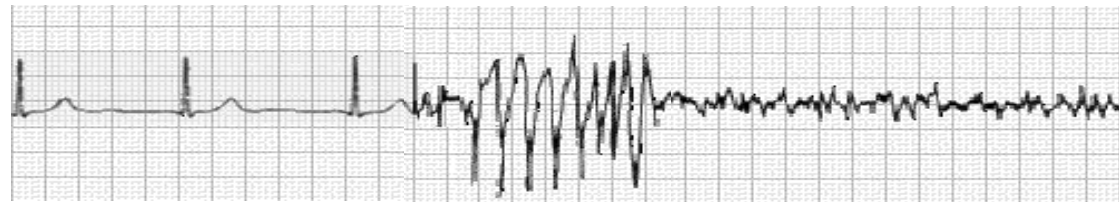
Human crowd

Fenton +
GT students

Cardiac rhythms and arrhythmias



Fenton, Cherry
thevirtualheart.org



Normal
Rhythm

Ventricular
Tachycardia

Ventricular
Fibrillation

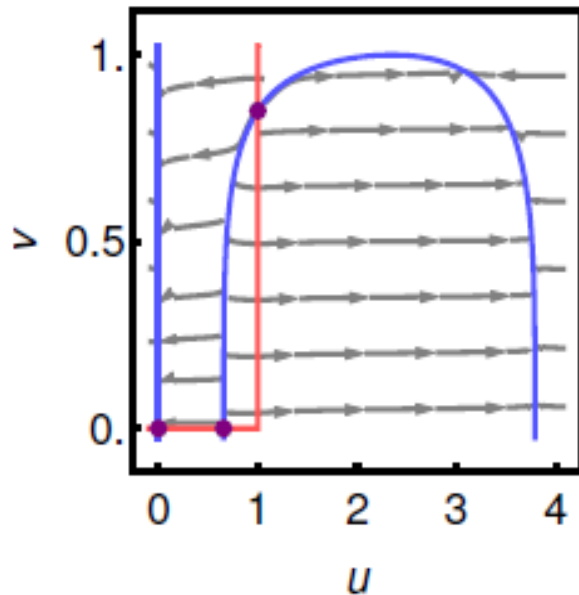
Spiral chaos in a simple model

Reaction-diffusion system:

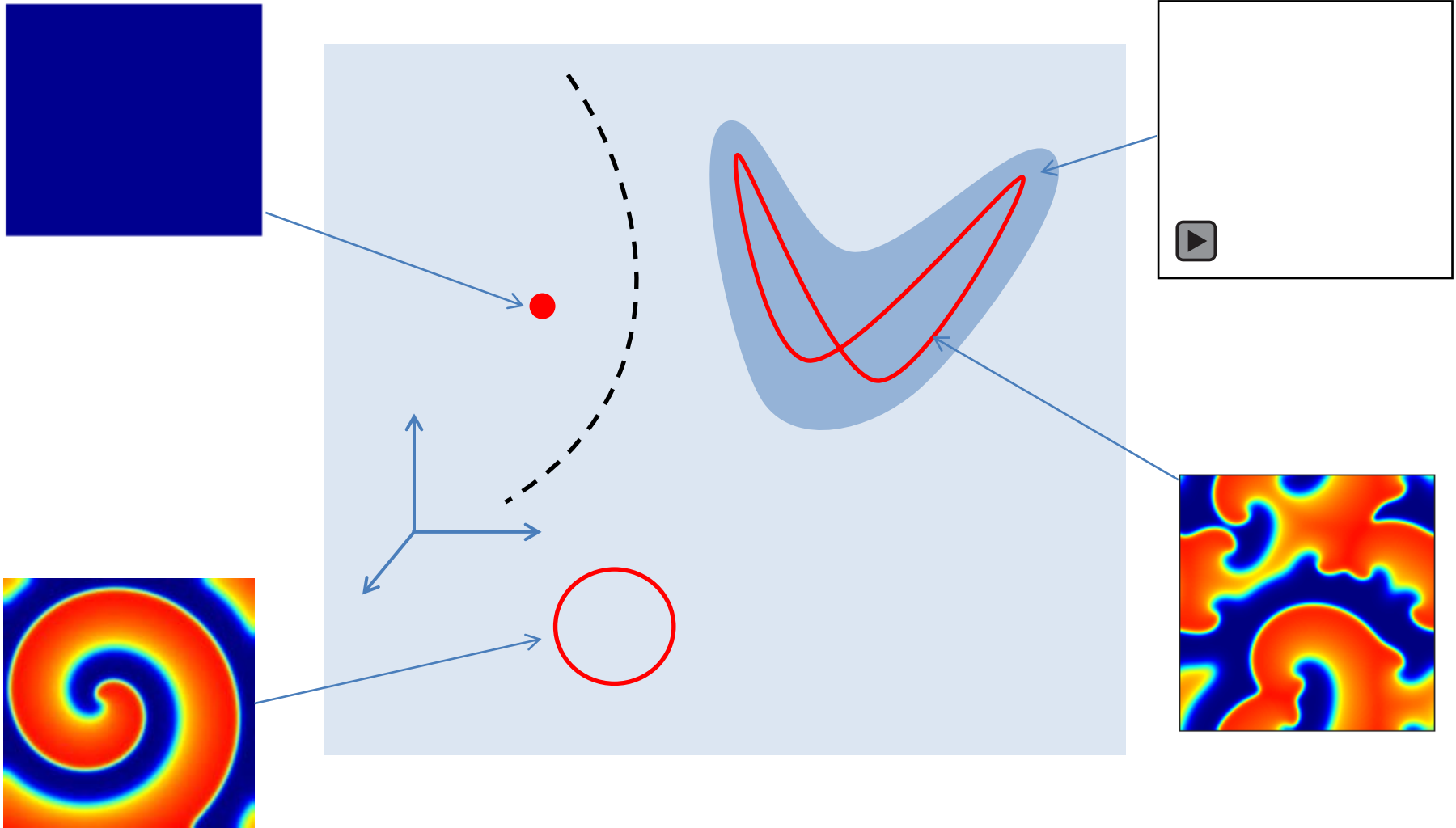
$$\partial_t u = D\nabla^2 u + f(u, v)$$

$$\partial_t v = g(u, v)$$

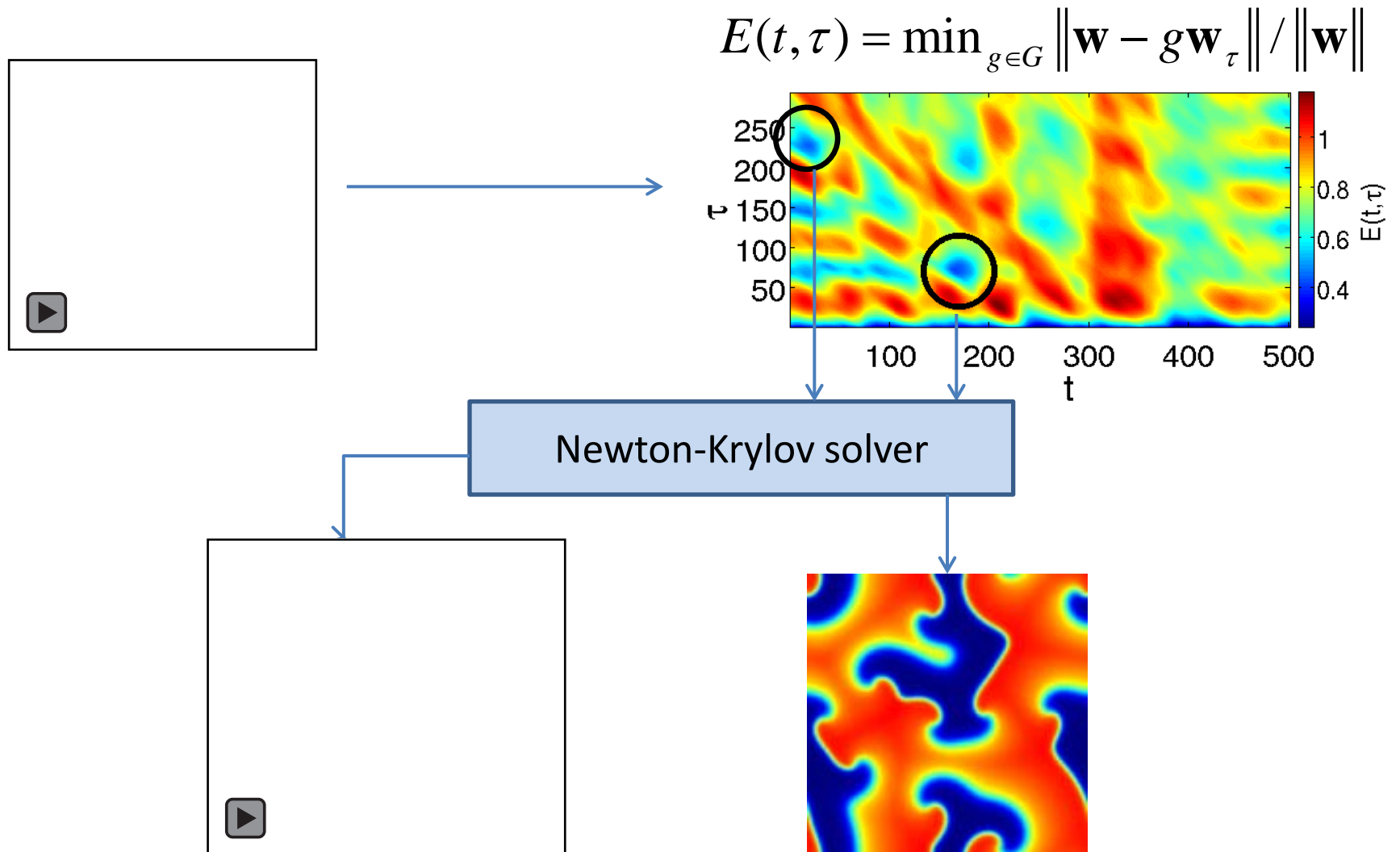
Karma (1994)



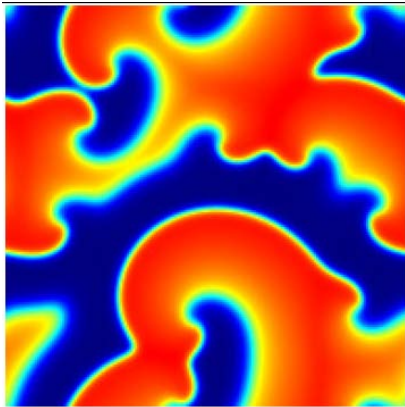
Nonchaotic solutions and ECS



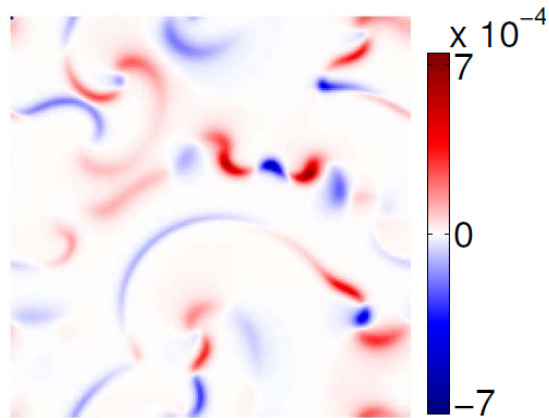
Computing ECS



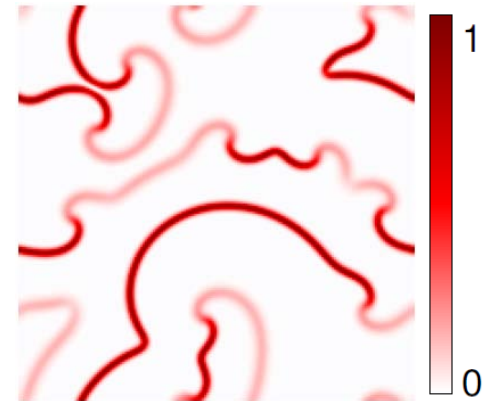
Unstable periodic solutions?



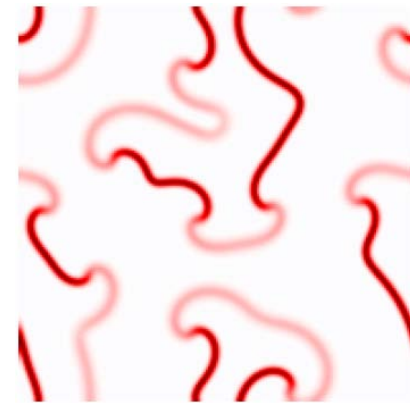
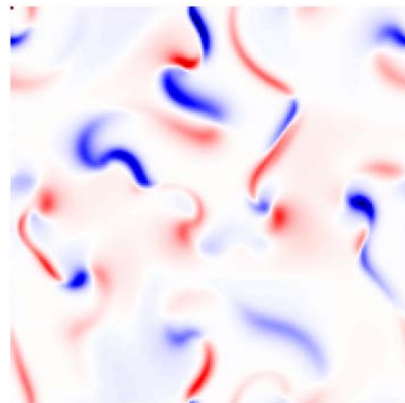
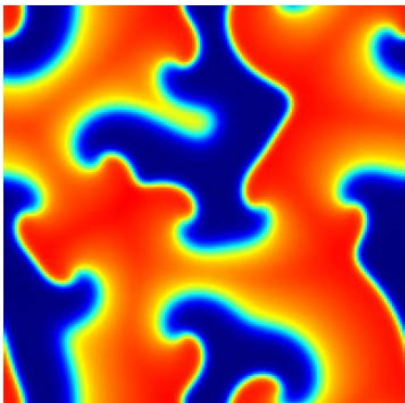
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$$\frac{u - u_T}{\|u\|_\infty}$$



$$\frac{|\nabla u|}{\|\nabla u\|_\infty}$$



Weakly nonlinear waves

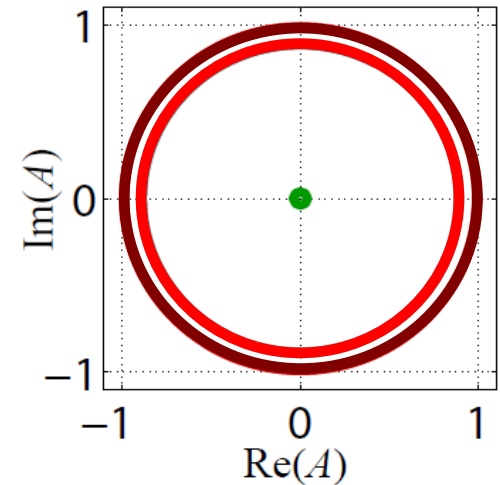
Complex Ginzburg-Landau Equation:

$$\partial_t A = A + (1 + i\alpha)\nabla^2 A - (1 + i\beta)|A|^2 A$$

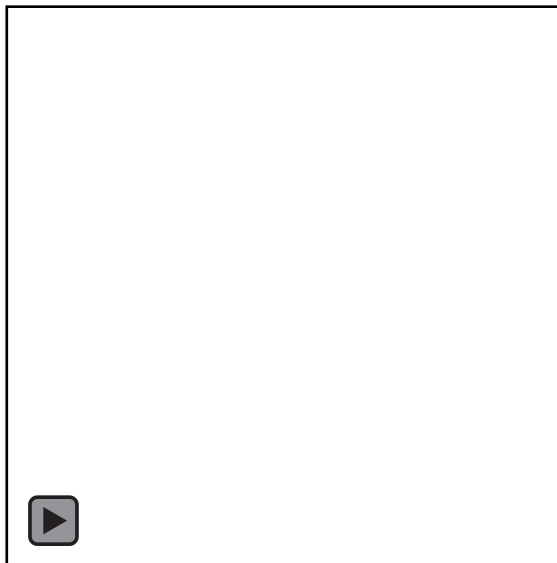
Tile the domain using amplitude ridges:

$$A = \rho e^{i\varphi}$$

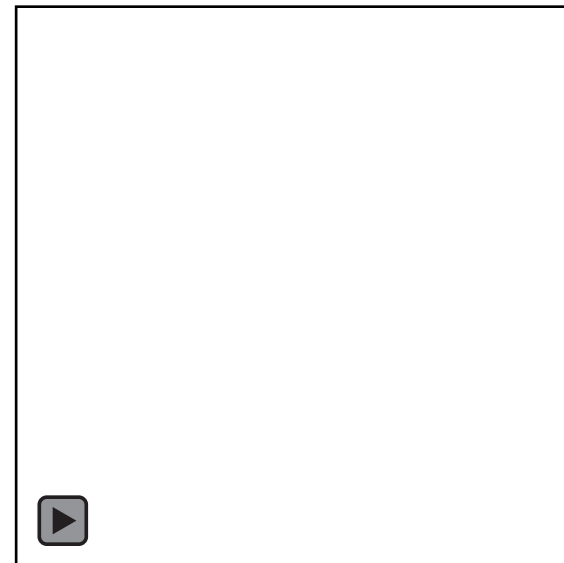
Bohr, Huber, Ott (1996)



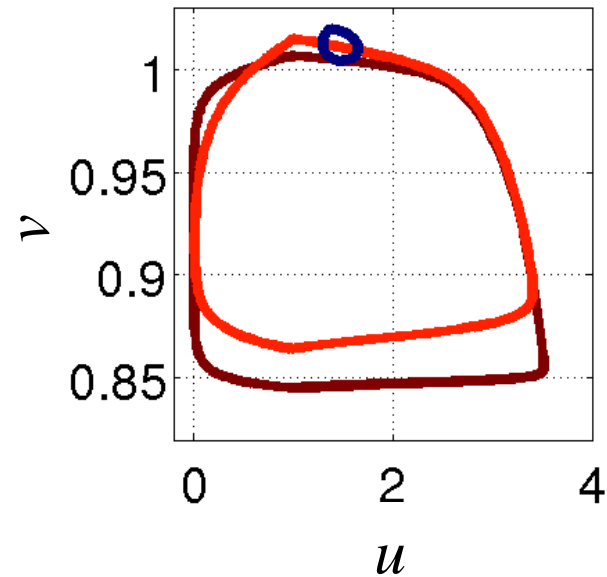
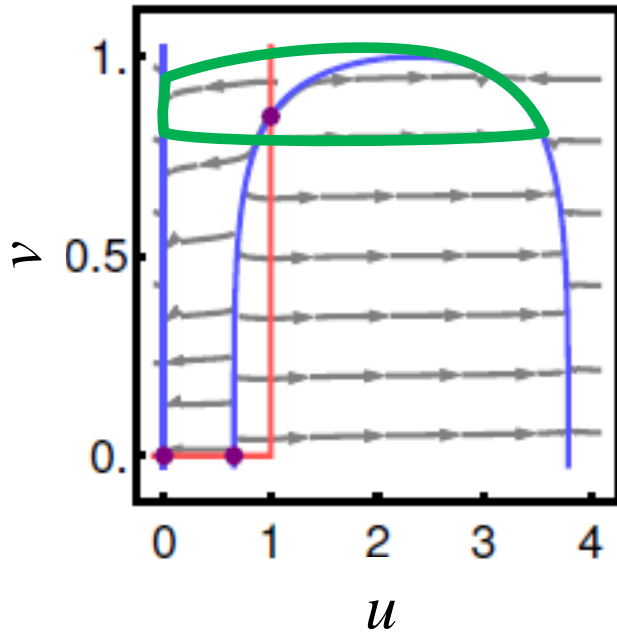
$\text{Re } A$



ρ



Strongly nonlinear waves



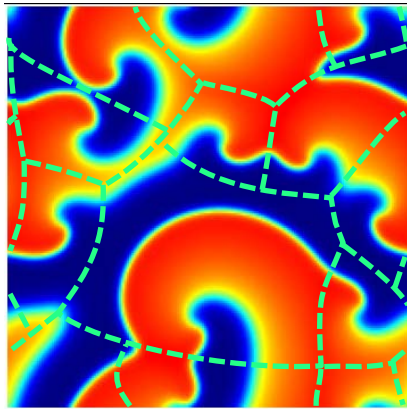
Use cycle area instead of amplitude:

and elapsed time from crossing a Poincare section instead of phase:

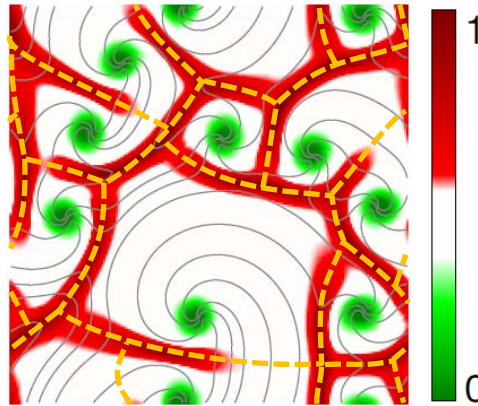
$$I = \oint v du = \int_0^T v \dot{u} dt$$

$$\theta = \int \omega dt, \quad \omega = \frac{2\pi}{T}$$

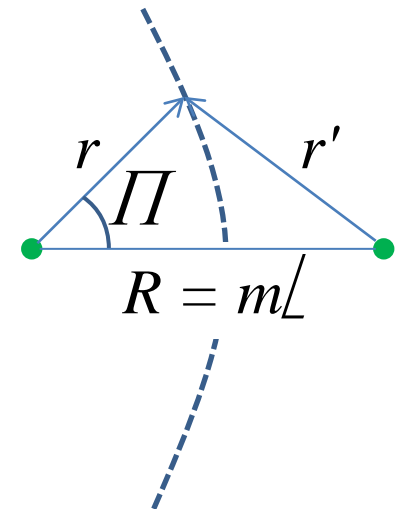
Tiling multispiral states



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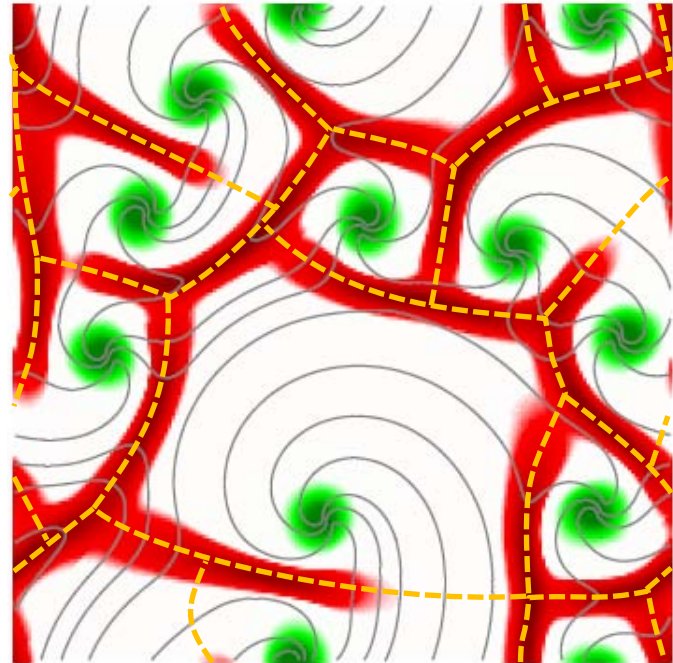
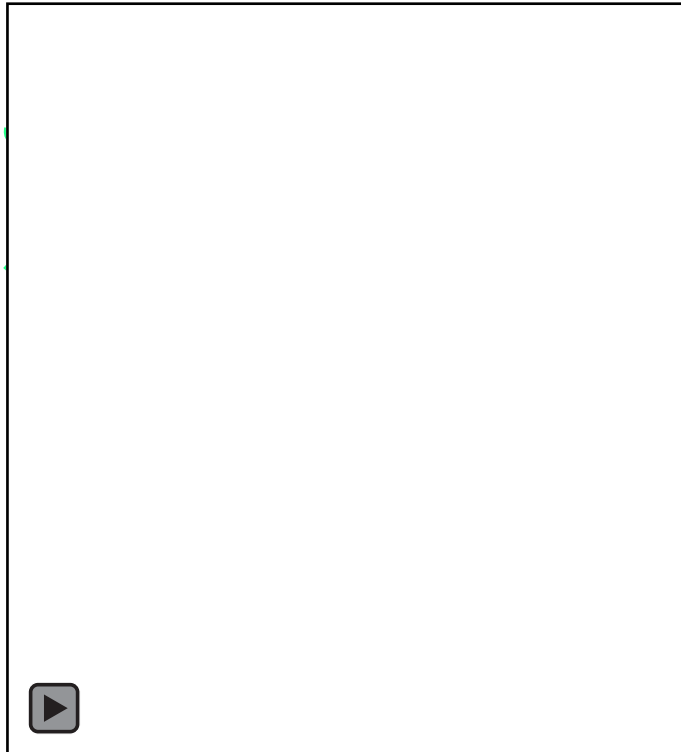


Can also describe tile boundaries analytically:

$$\frac{dr}{d\varphi} = \frac{-\sigma r'^2 - \sigma' r(R \cos \varphi - r) + 2\pi m r' r \sin \varphi}{\sigma' R^2 \sin \varphi + m(r'^2 - r'(r - R \cos \varphi))}$$

Luo, Zhang, Zhan (2009)

Boundary conditions



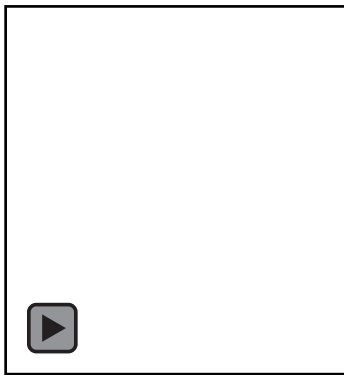
- Tiles are noncircular
- Neumann boundary conditions

Break-up

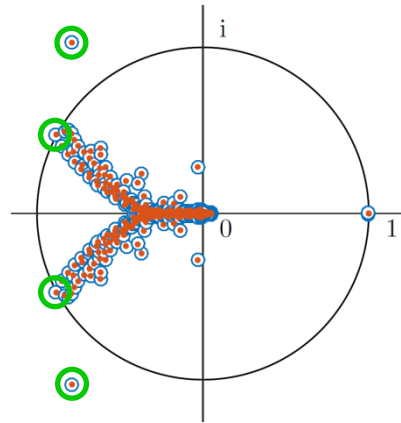
Drift

Collapse

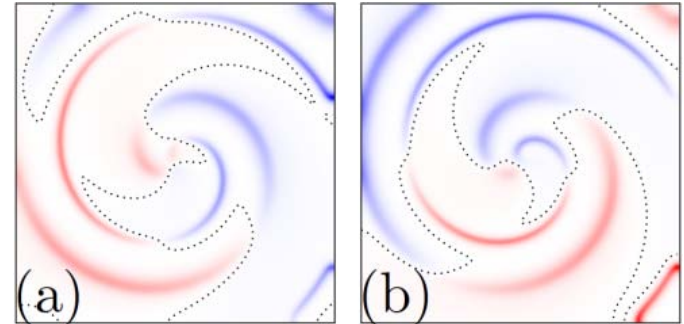
Stability of spiral waves



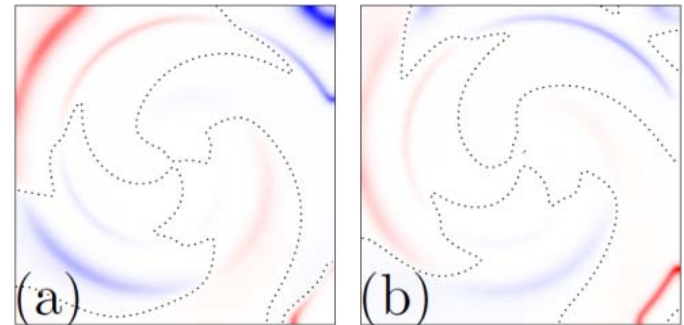
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Floquet multipliers

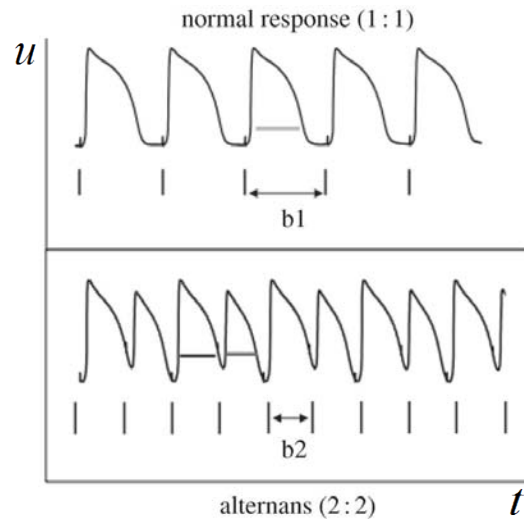


Absolute instability

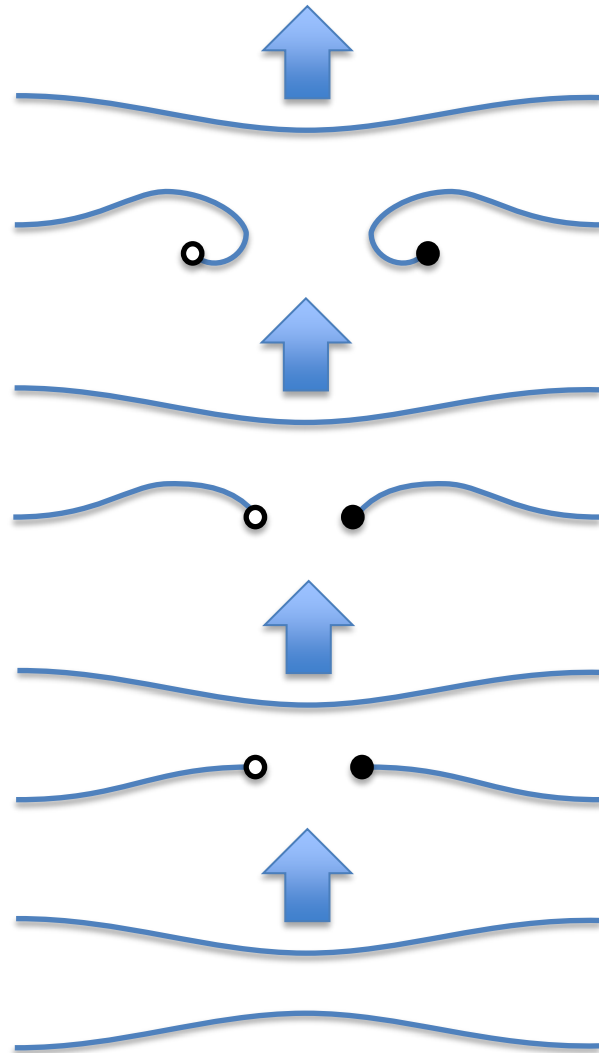


Convective instability

Alternans instability



Petrie, Zhao (2012)



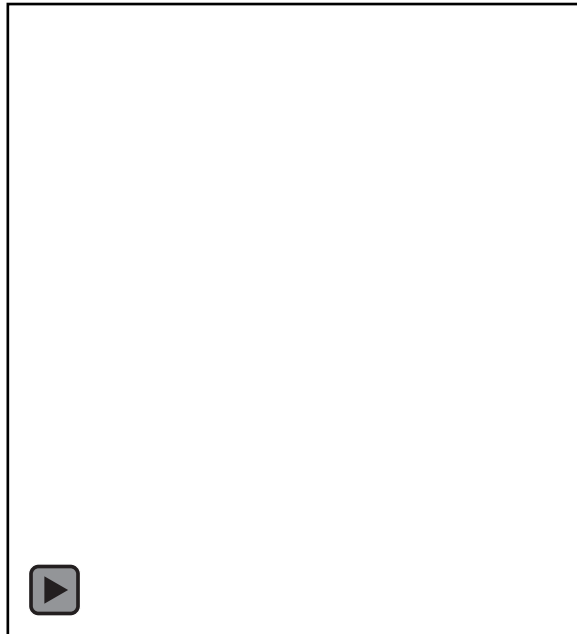
Break-up

Drift

Collapse

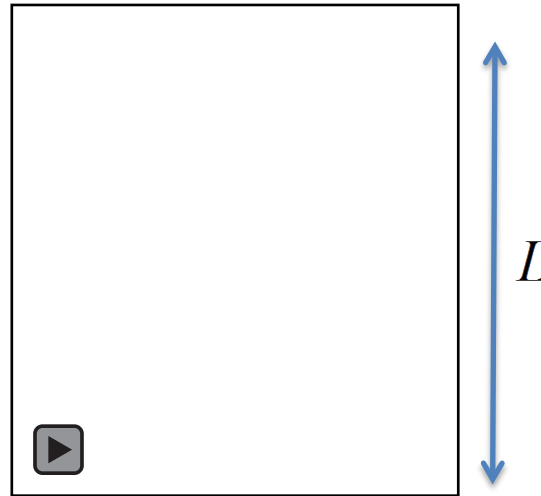
For some initial conditions...

Stroboscopic map:

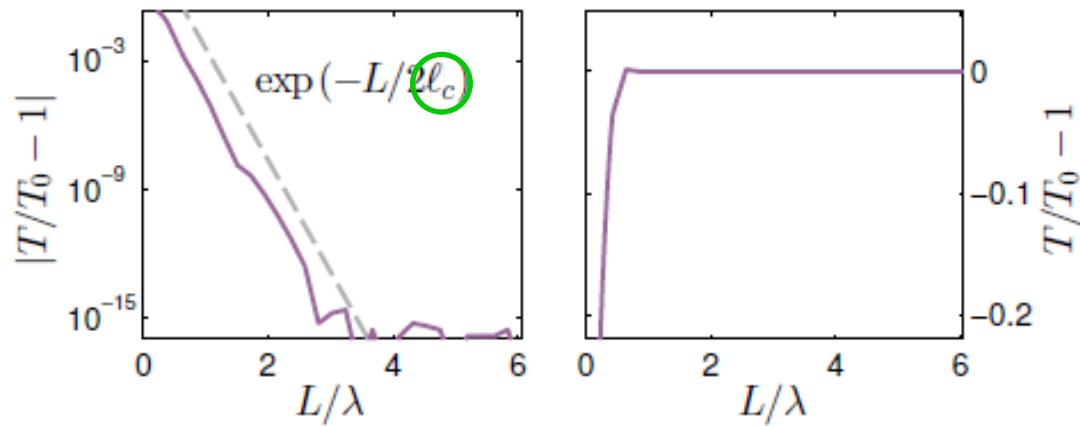


- Cores are drifting \rightarrow tiles have to deform
- Can we understand this drift and deformation?

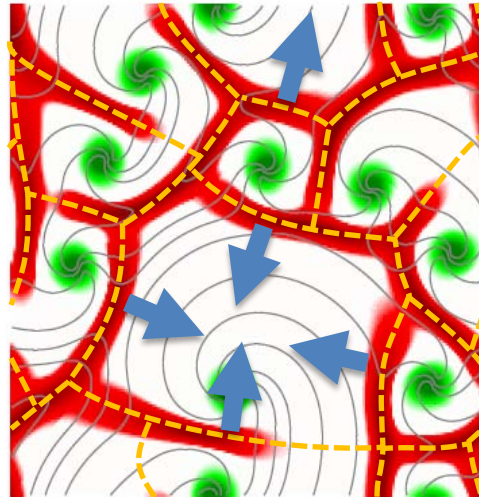
Dynamics of spirals on tiles



Period as a function of tile size



Dynamics of tiles



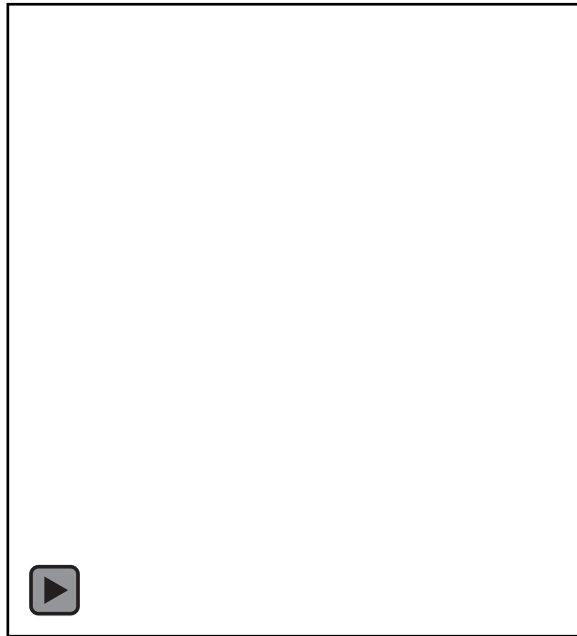
Motion of the boundary:

$$\mathbf{c} = (\omega_1 - \omega_2) \frac{\mathbf{k}_1 - \mathbf{k}_2}{|\mathbf{k}_1 - \mathbf{k}_2|^2}$$

Howard, Kopell (1977)

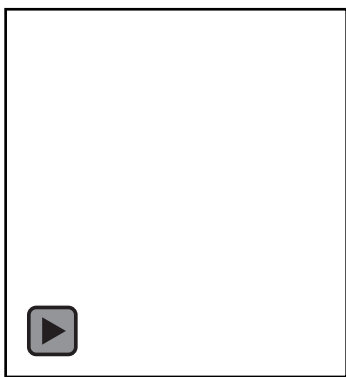
Dynamics of tiles (continued...)

Stroboscopic map:

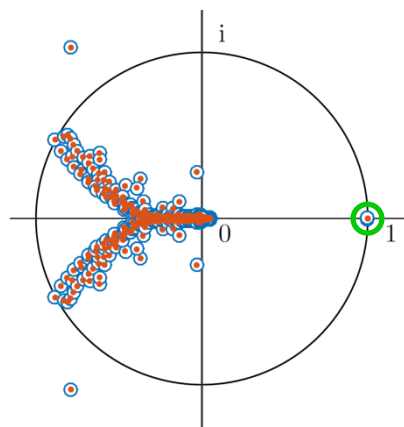


- Why are some cores moving (and others are not)?
- Why is their motion so slow?
- What sets the distance between cores?

Local Euclidean symmetry

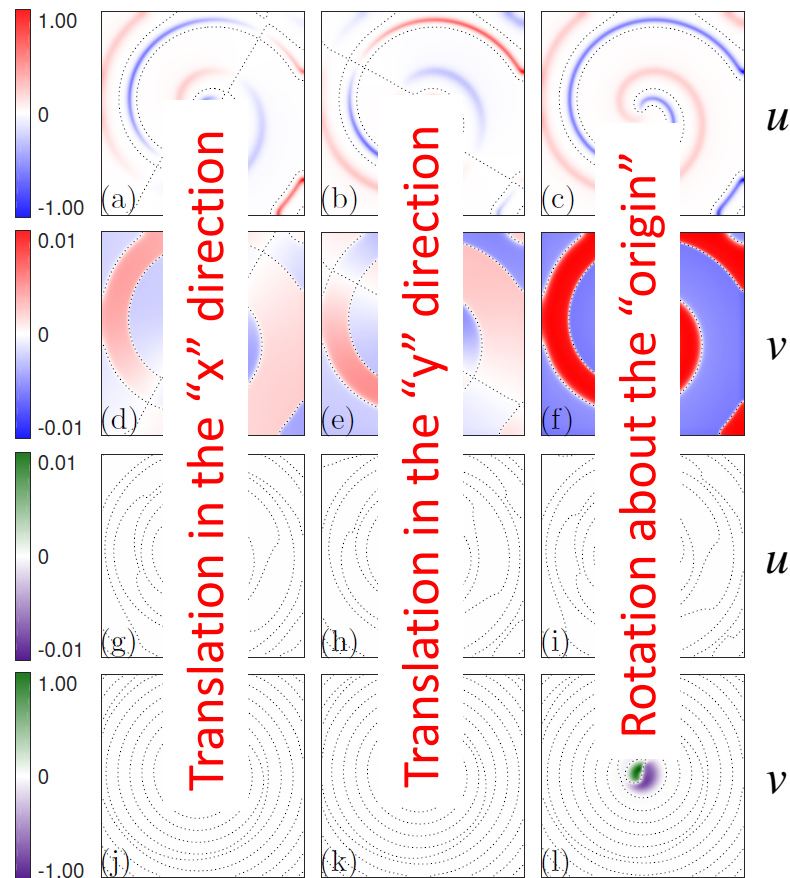


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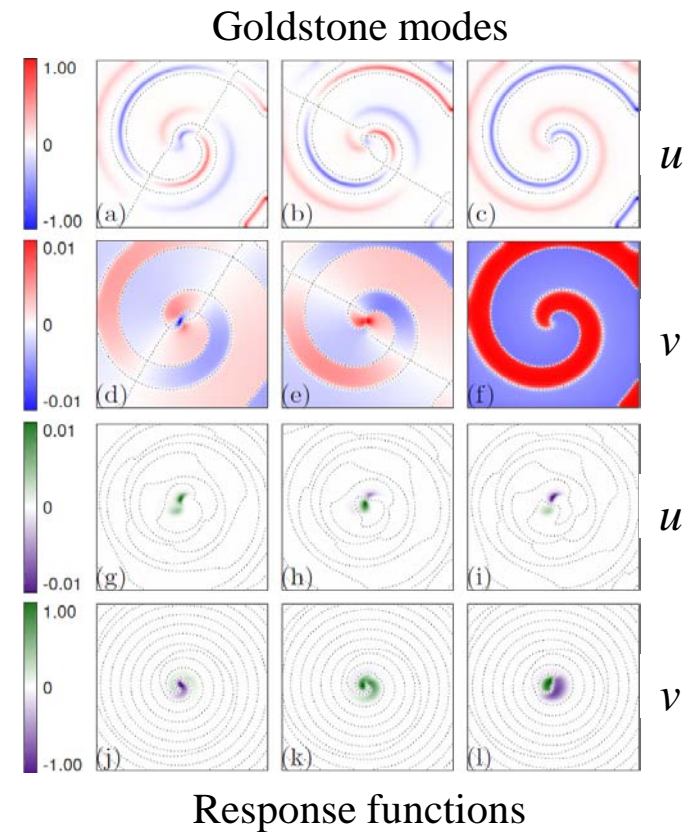
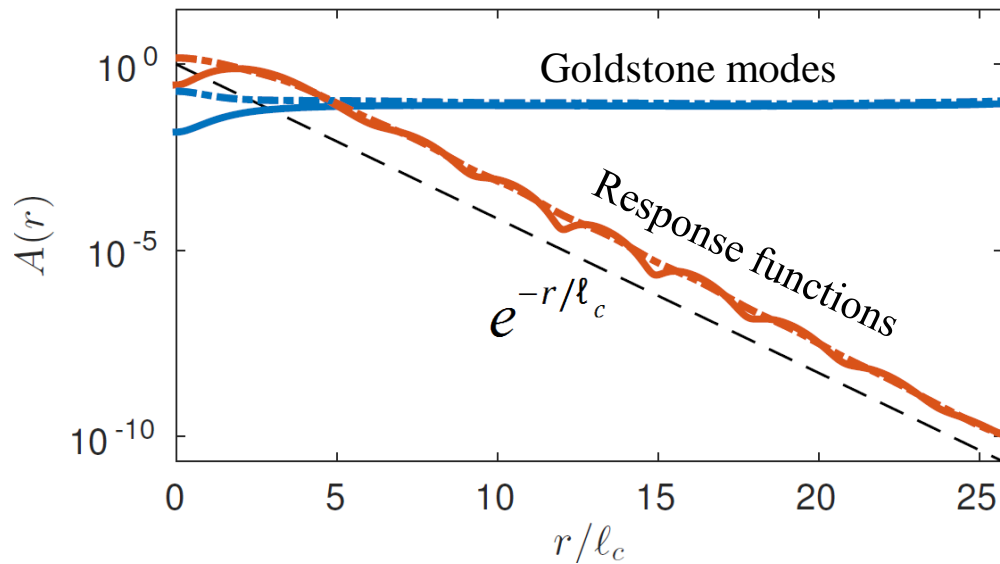
Floquet multipliers

Goldstone modes/right eigenfunctions

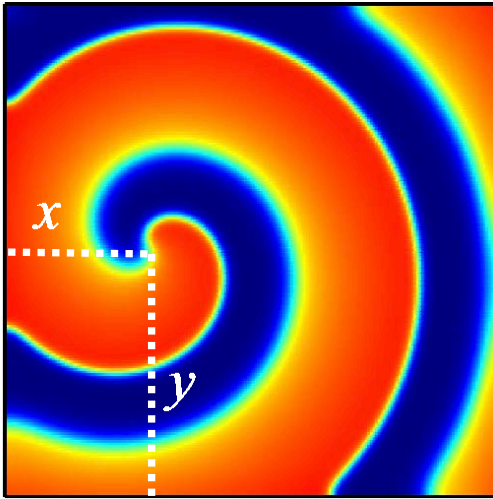


Response functions/left eigenfunctions

Asymptotic freedom

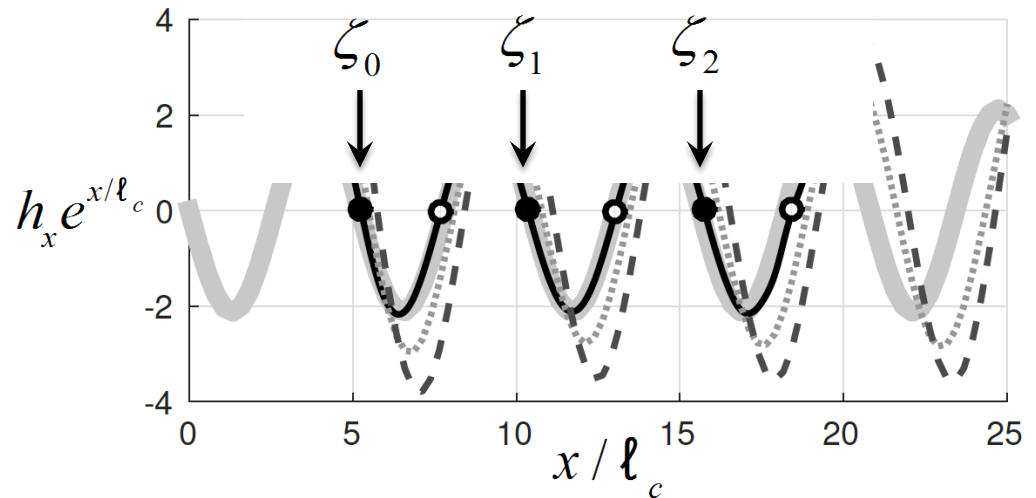
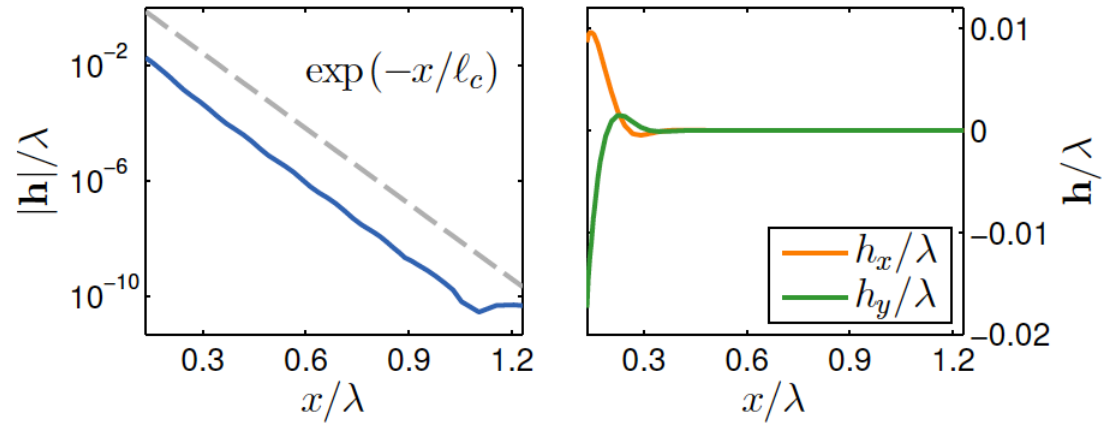


Interaction of cores with boundaries

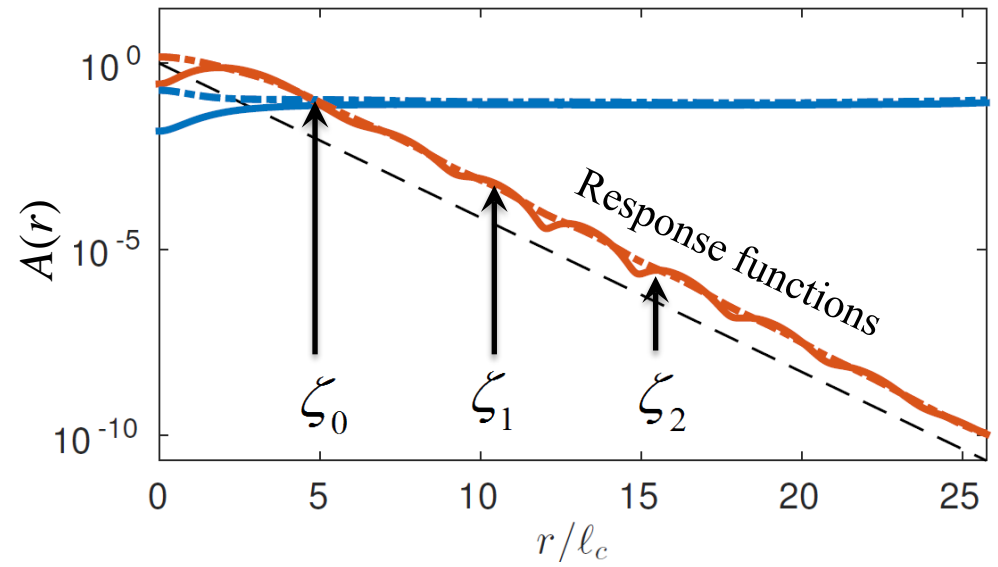
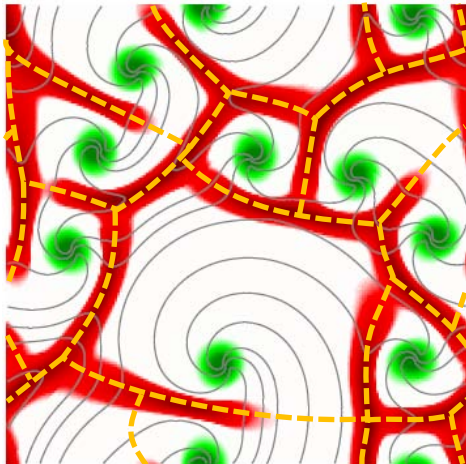
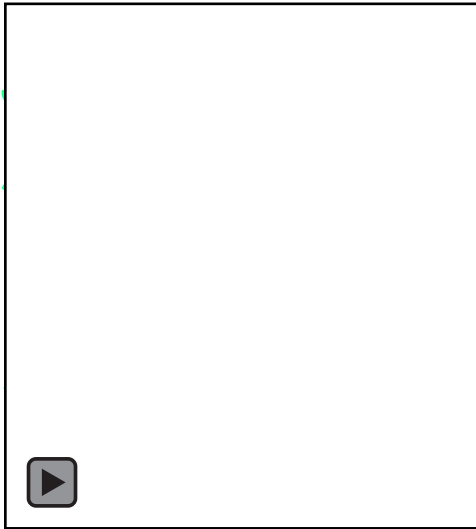


$$\mathbf{x}^{n+1} = \mathbf{x}^n + \mathbf{h}(\mathbf{x}^n),$$

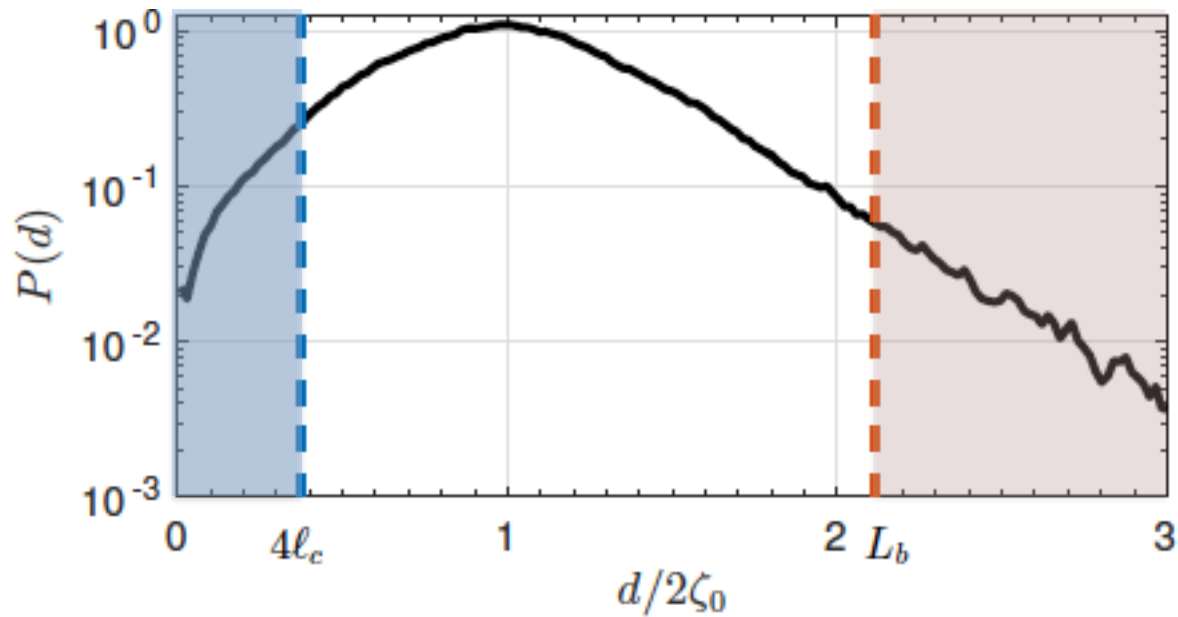
$$\mathbf{x}^n = \mathbf{x}(nT)$$



Core-core interaction



Core-core separation (& tile size)



No time-
periodic
solutions

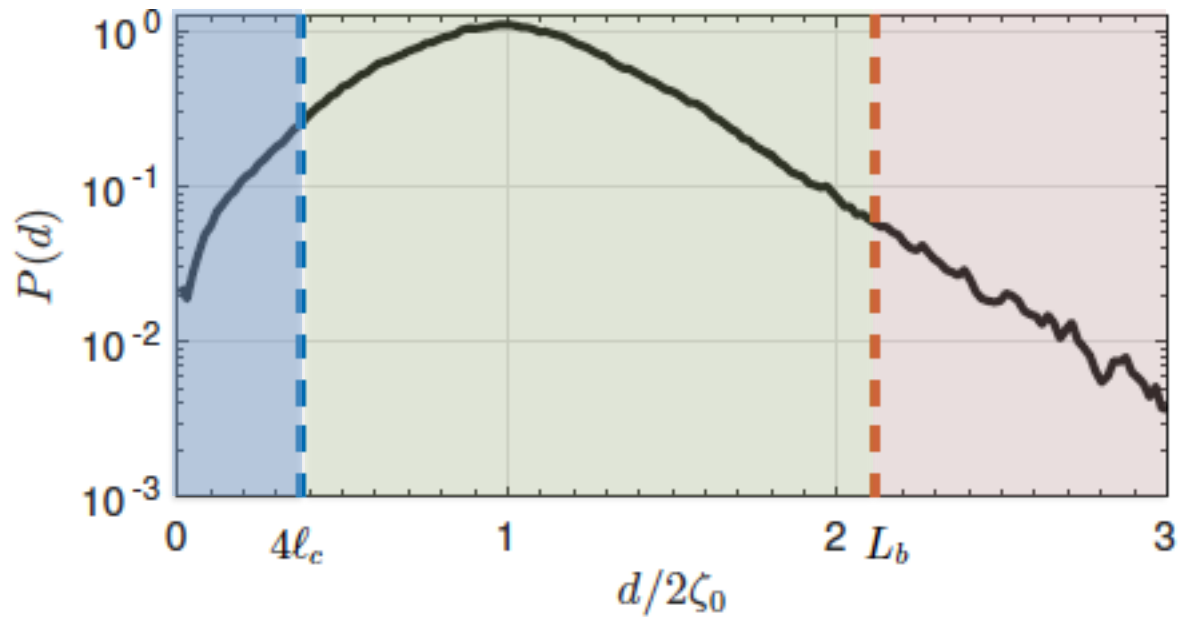
Solutions
unstable
(alternans)

Break-up

Drift

Collapse

Core-core separation (& tile size)

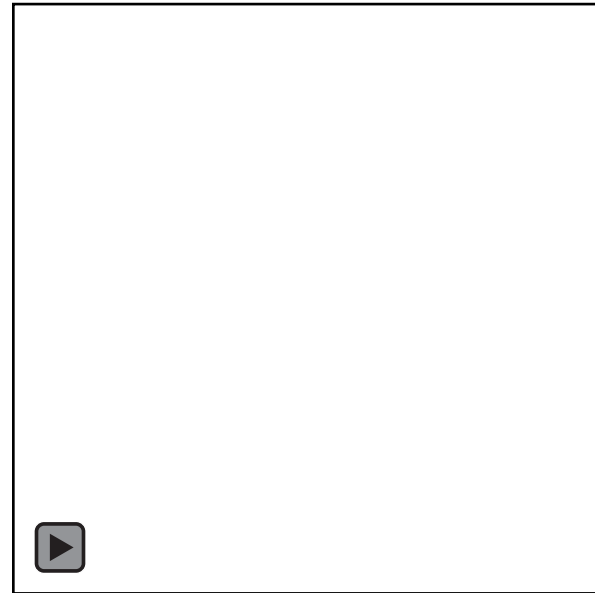
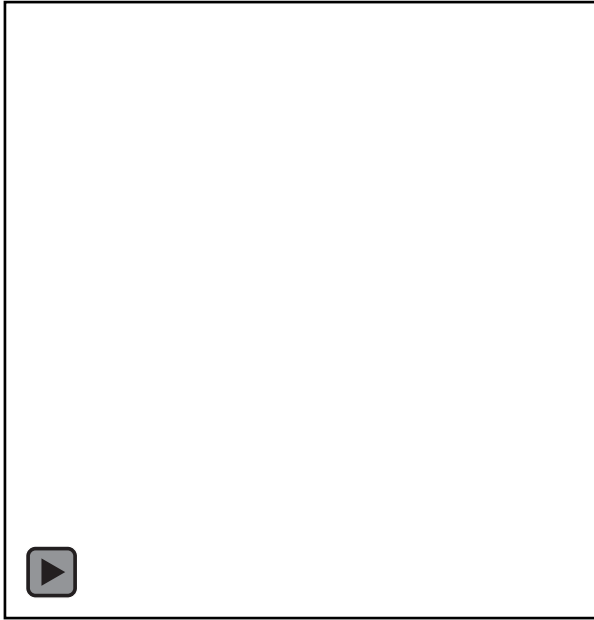


Virtual
pairs

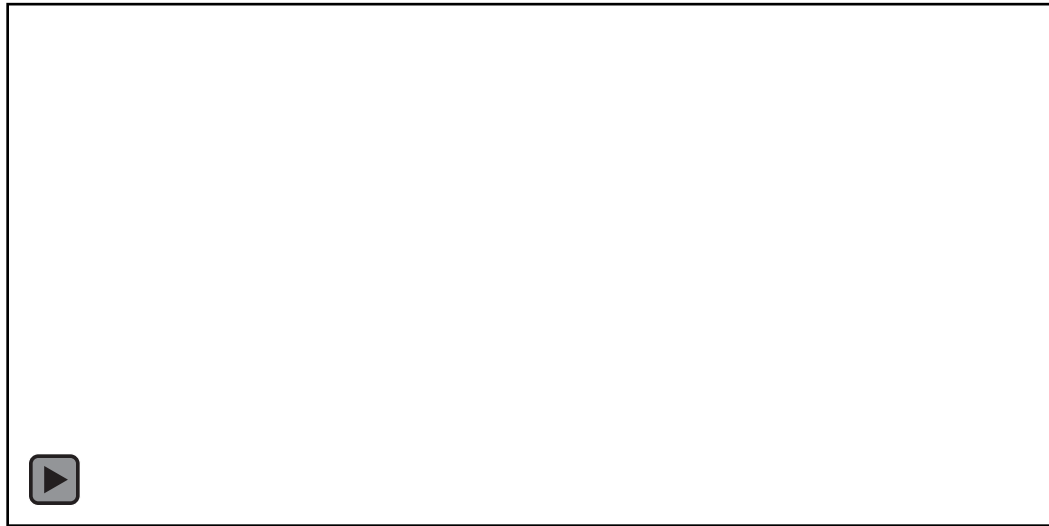
Core drift/
meander/
collapse

Break-up

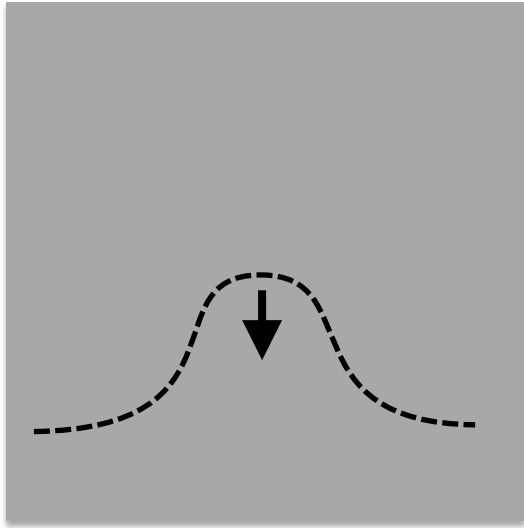
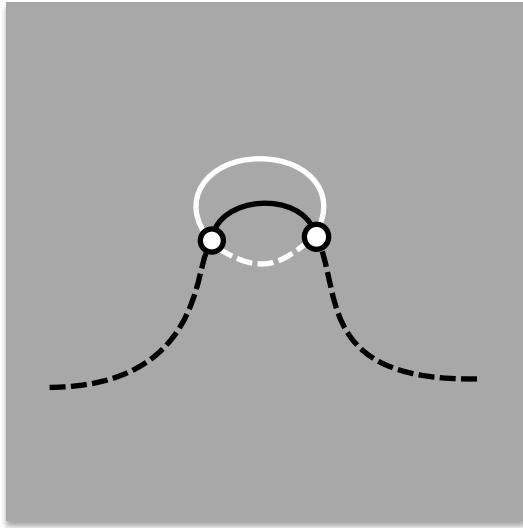
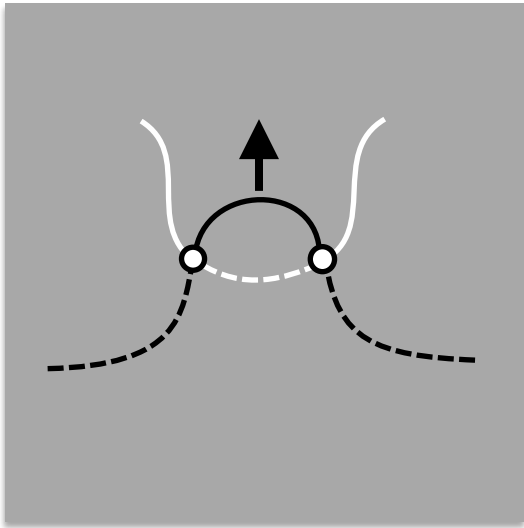
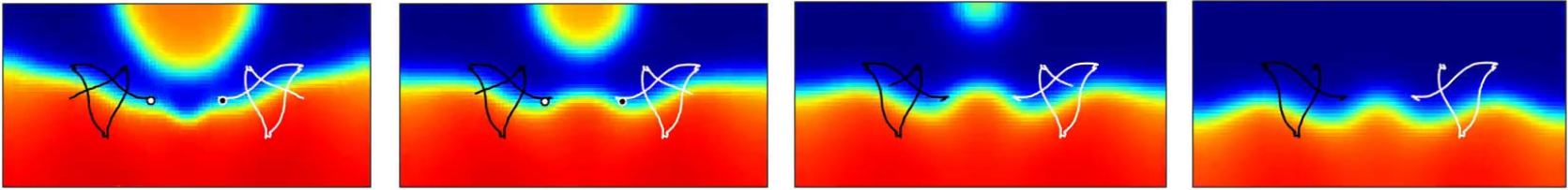
Core meander



Wave collapse



Wave collapse



The mechanism of spiral chaos

- Slow dynamics (of tiles)
 - ✓ The tiles are generally of different size
 - ✓ The frequencies of spirals differ
 - ✓ The tiles boundaries drift (slowly)
 - ✓ Small (fast) spirals grow at the expense of big (slow) ones
- Fast dynamics (of spirals)
 - ✓ Large spirals ($L > L_b$) break up due to alternans instability
 - ✓ Small spirals ($L < 4l_c$) survive for less than one period and collapse (with a neighbor)
 - ✓ Medium size spirals ($4l_c < L < L_b$) interact with each other in nontrivial ways

Implications for fluid turbulence

- ❑ No *global* ECS on domains much larger than the relevant coherence length, no matter what the physics is
- ❑ Need to look for localized solutions that respect *local* Euclidean symmetries and their interactions
- ❑ Coherence length can be defined with the help of *adjoint* eigenfunctions (to-do for fluid dynamicists)
- ❑ Spatial *correlations* may decay exponentially even when solutions do not
- ❑ Does exponential decay of *velocity/energy* imply short spatial correlations? What about *pressure*?