

Multiple Extra Dimensions and Cosmic Sources of KK Particles

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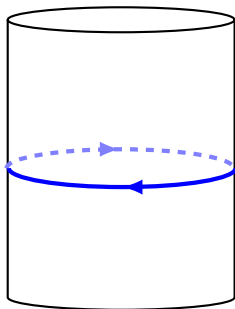


Plan

- 1 General introduction to extra dimensions
- 2 Warped extra dimensions
- 3 Multiple extra dimensions
- 4 Involving students in high energy theory
- 5 Cosmic sources (if time)



S^1 : Wave function



Continuity requires:

$$\Psi(y) = \Psi(y + 2\pi R)$$

$$\Psi = \frac{1}{\sqrt{2\pi R}} \sum_{n=-\infty}^{+\infty} \psi^{(n)} e^{iny/R}$$

$$E^2 = (pc)^2 + (Mc^2)^2$$



S^1 : Mass

$$\begin{aligned}
 E^2 &= (pc)^2 + (Mc^2)^2 = (p_{3D}c)^2 + (p_y c)^2 + (Mc^2)^2 \\
 &= (p_{3D}c)^2 + \underbrace{\left(\frac{n}{R}\hbar c\right)^2 + (Mc^2)^2}_{(m_n c^2)^2}
 \end{aligned}$$

in particle physics, $\hbar = c = 1$:

$$m_n^2 = \frac{n^2}{R^2} + M^2 \quad \text{Kaluza-Klein (KK) particles}$$



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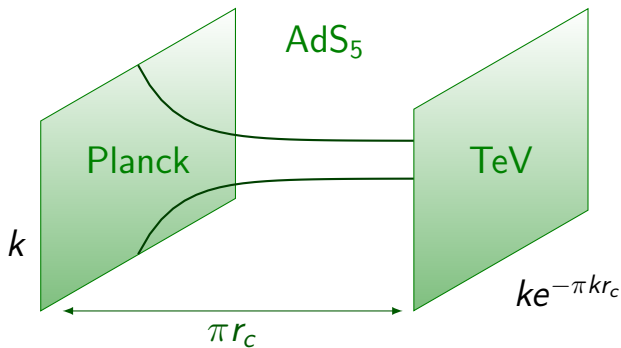
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Randall-Sundrum Model

$$ds^2 = \underbrace{e^{-2kr_c|\phi|}}_{\text{warp factor}} \eta_{\mu\nu} dx^\mu dx^\nu - \underbrace{r_c^2 d\phi^2}_{\text{extra dimension}}$$





Multiple Extra Dimensions

Fermions in Randall-Sundrum models with two additional unwarped extra dimensions

Jeremy Perrin (2013 Apker Finalist)
and Erin De Pree
arXiv:1310.1928 [hep-ph]



6D Models

$$ds^2 = e^{-2kr_c|\phi|} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 d\phi^2 - \underbrace{R^2 d\theta^2}_{\text{new extra dim}}$$

Problem

No bulk mass terms that are also 6D-chiral and have four-components



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7D Warped Extra Dimensions

Instead we add two additional extra dimensions.

Choices

- Compactify on a torus (T_2) or a sphere (S_2)?
- Apply the warp factor to just the 4D dimensions or all other dimensions?

 $\text{AdS}_5 \times T_2$

$$ds^2 = e^{-2kr_c|\phi|} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 d\phi^2 - R^2 (d\theta_1^2 + d\theta_2^2)$$

Set of coupled differential equations for the KK wavefunctions.



AdS₅ × S₂

$$ds^2 = e^{-2kr_c|\phi|} \eta_{\mu\nu} dx^\mu dx^\nu - r_c^2 d\phi^2 \\ - R^2 (d\theta^2 + \sin^2 \theta d\omega^2)$$

Problem

No zero-mode solutions without additional structure

General result for all manifolds of positive curvature.



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Research Group

- 1 Start early
- 2 Work in small groups
- 3 Not every project results in a publication





Questions and Discussion

Thank you

ekdepree@smcm.edu

arXiv:1505.00024 [astro-ph.HE]

arXiv:1310.1928 [hep-ph]



Possible sources of KK particles

Colliders

- Control the source (luminosity, E_{CM} , etc)
- Expensive

Cosmic sources

- Free
- No control over the source



Possible sources of KK particles

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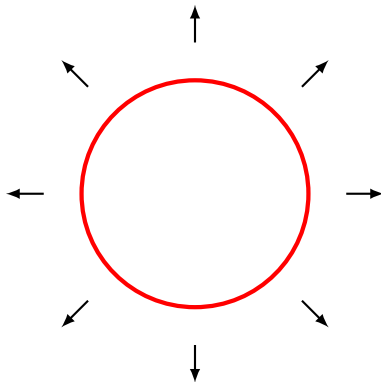
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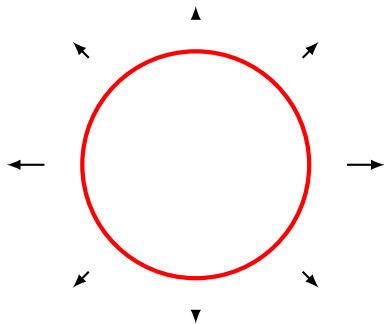
Gamma Ray Bursts (GRBs)



Ian Morgan, Ted Tao, De Pree, Kevin Tennyson
arXiv:1505.00024 [astro-ph.HE]



Observer's frame

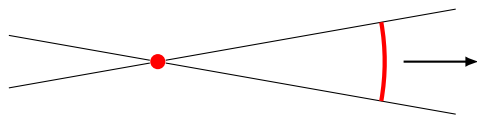


Earth





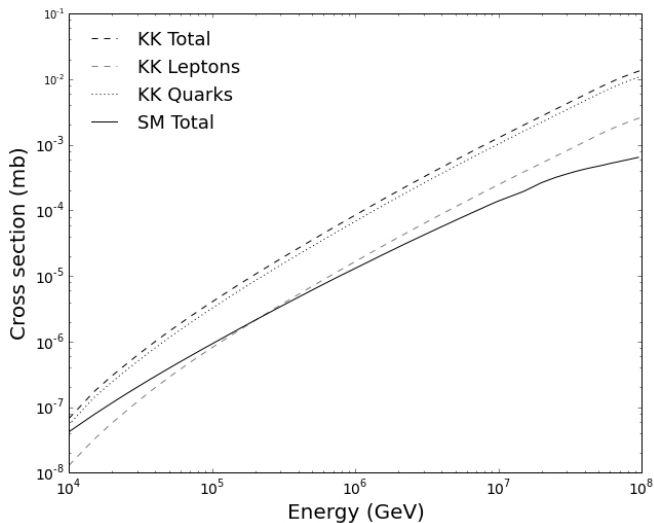
Beaming



Earth




Results





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S^1 : Momentum

$$p^2 = \mathbf{p} \cdot \mathbf{p} = p_{x_1}^2 + p_{x_2}^2 + p_{x_3}^2 + p_y^2 = p_{3D}^2 + p_y^2$$

$$\hat{p} = -i\hbar\nabla, \quad \hat{p}_y = -i\hbar\frac{\partial}{\partial y}$$

$$\hat{p}_y \Psi = \frac{-i\hbar}{\sqrt{2\pi R}} \psi^{(n)} e^{iny/R} \left(\frac{in}{R} \right) = \frac{n\hbar}{R} \Psi$$

Eigenvalue for \hat{p}_y : $n\hbar/R$



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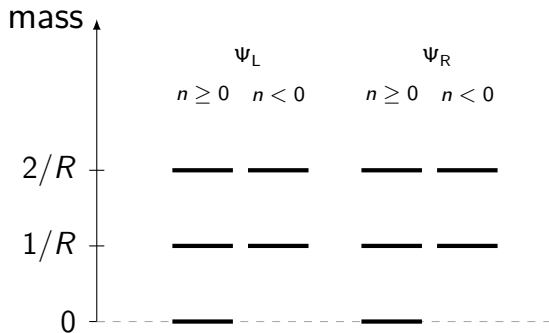
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S^1 : Fermion KK Spectrum



For $M = 0$



AdS₇ Model

$$ds^2 = e^{-2kr_c|\phi|} [\eta_{\mu\nu} dx^\mu dx^\nu - R^2 (d\theta_1^2 + d\theta_2^2)] - r_c^2 d\phi^2$$