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Theoretical Physics

White Dwarfs from
Physics to Astrophysics

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Merger remnants,
double-faced stars and
other curiosities

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Burke Fellow



Caltech

- A (maybe?) moon sized white dwarf
~2000 km in radius
-> ~1.35 Msun
 - Rapidly rotating – 7 minute period
 - Highly magnetized
- > Poster child for a double white dwarf merger remnant



Article

A highly magnetized and rapidly rotating white dwarf as small as the Moon

<https://doi.org/10.1038/s41586-021-03615-y>

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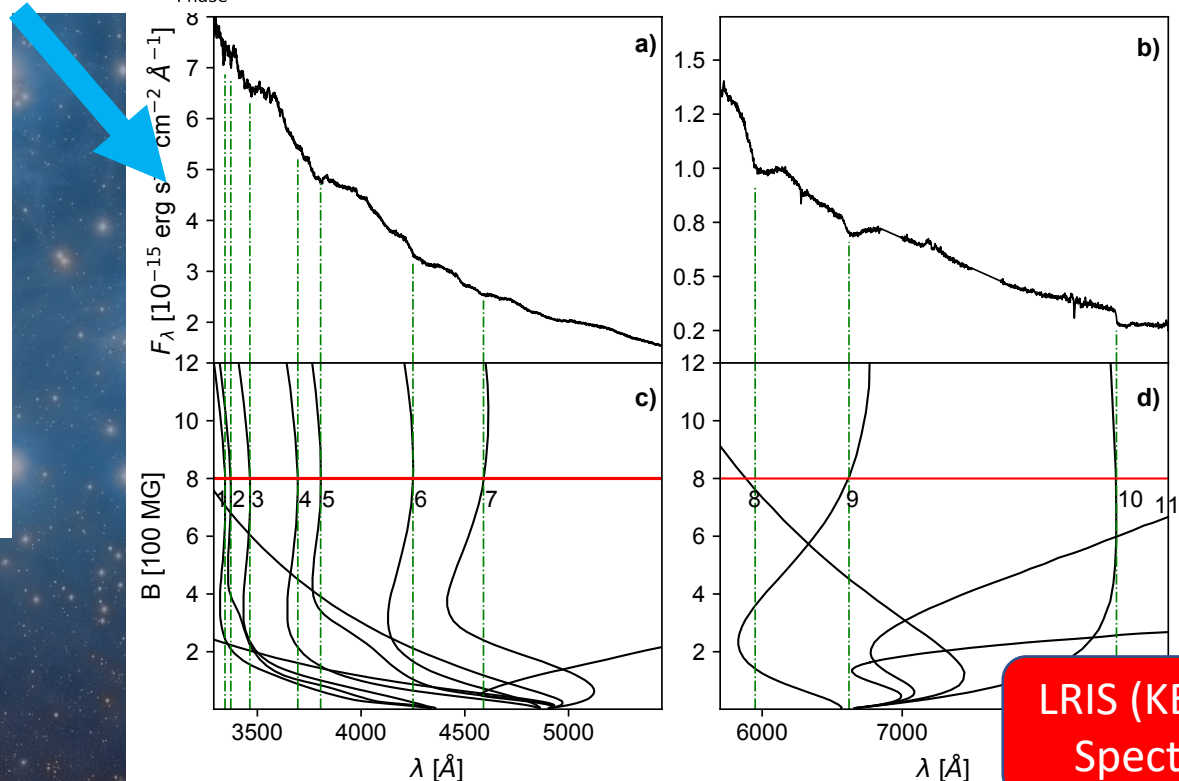
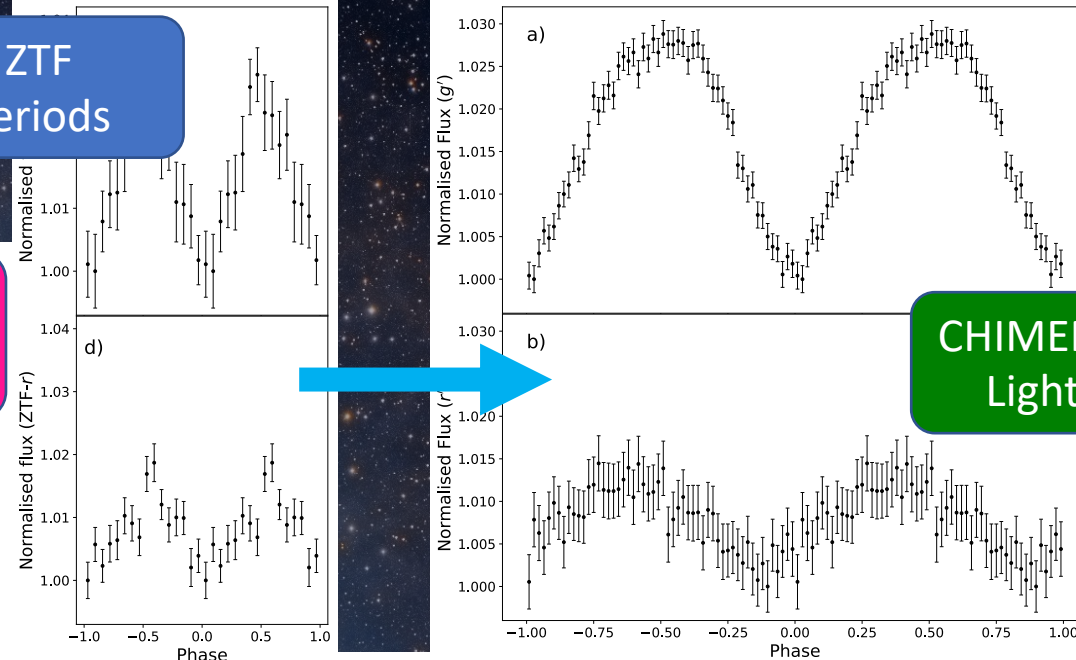
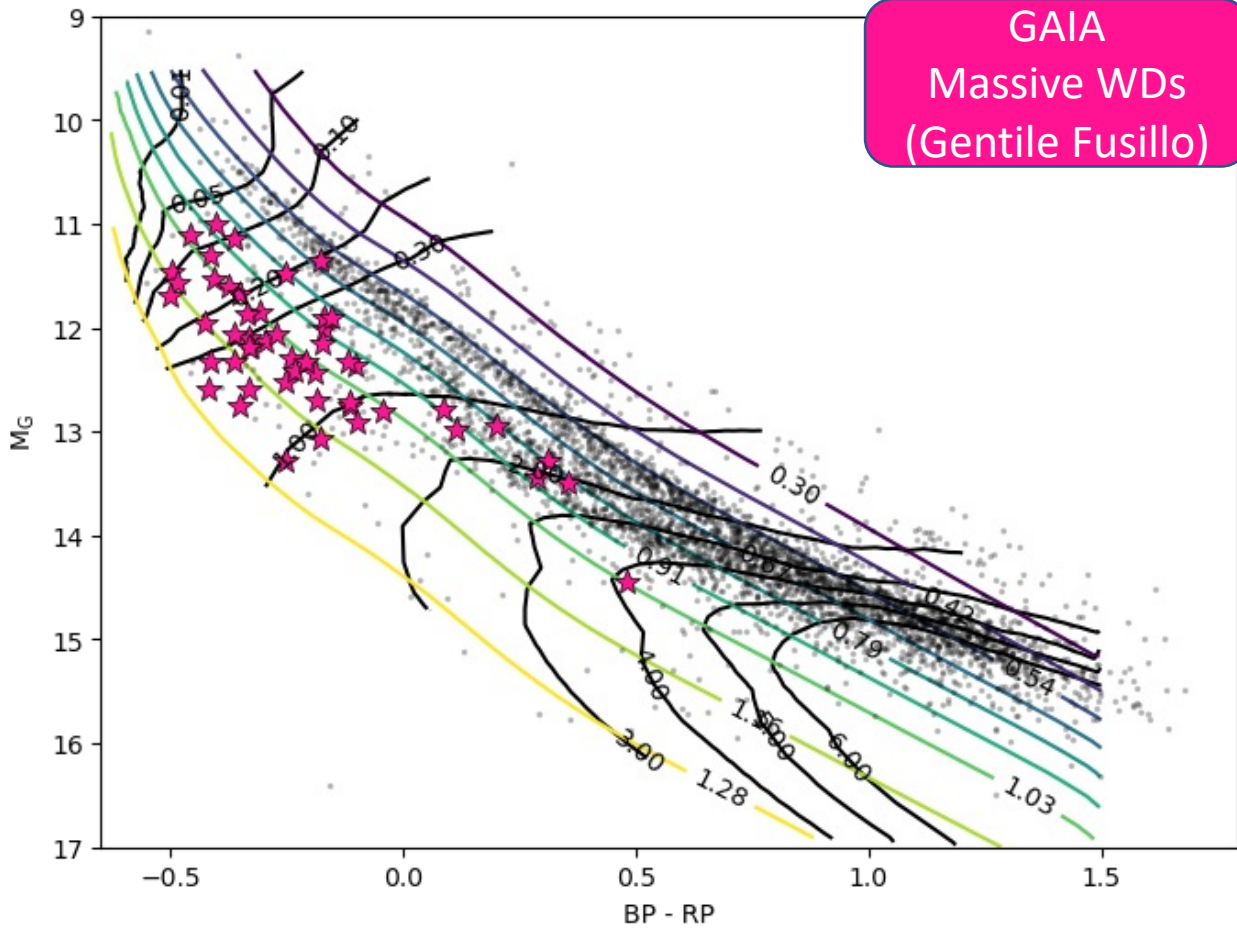
How we found her

ZTF
Periods



GAIA
Massive WDs
(Gentile Fusillo)

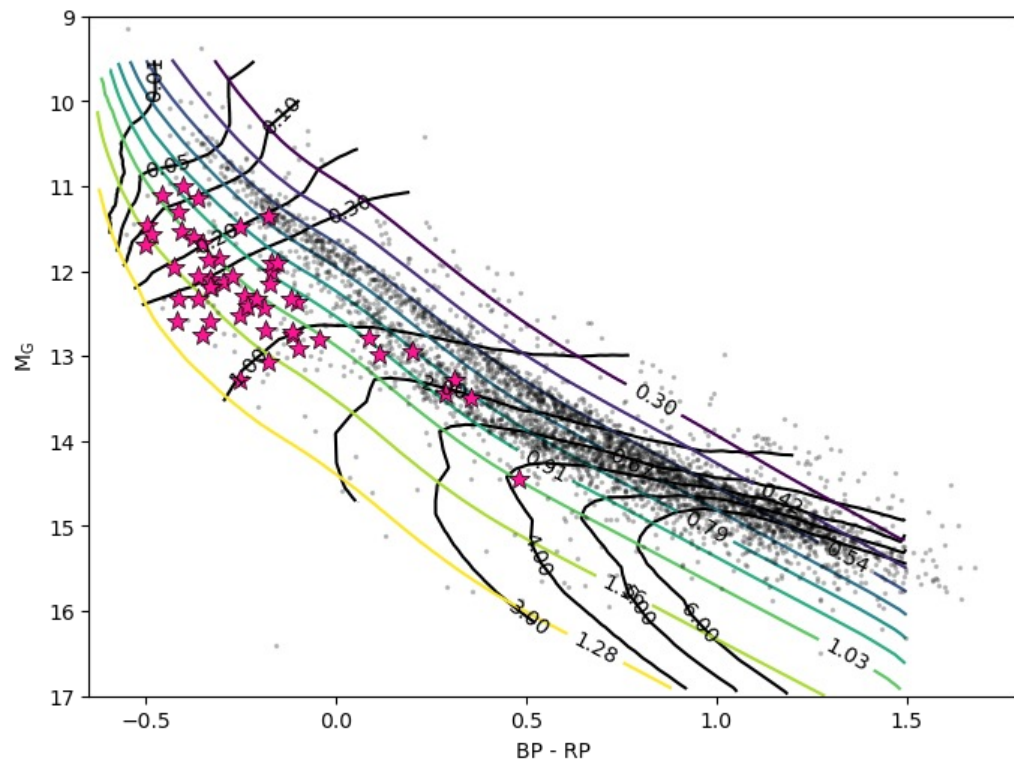
CHIMERA (HALE)
Light curves



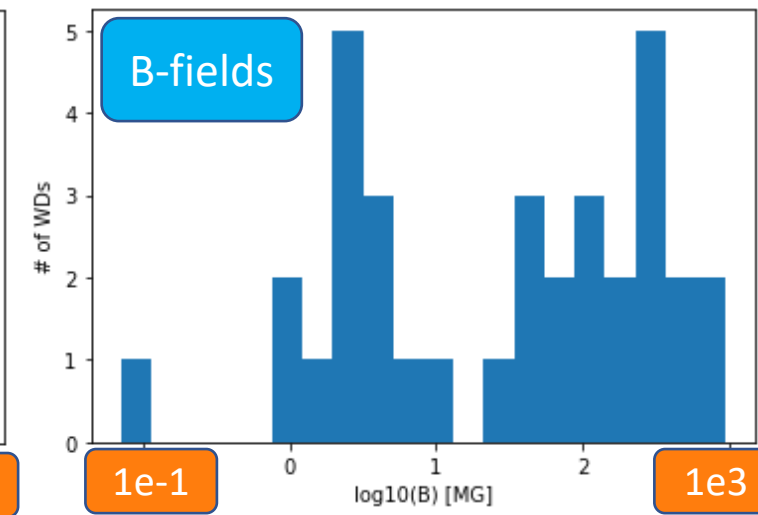
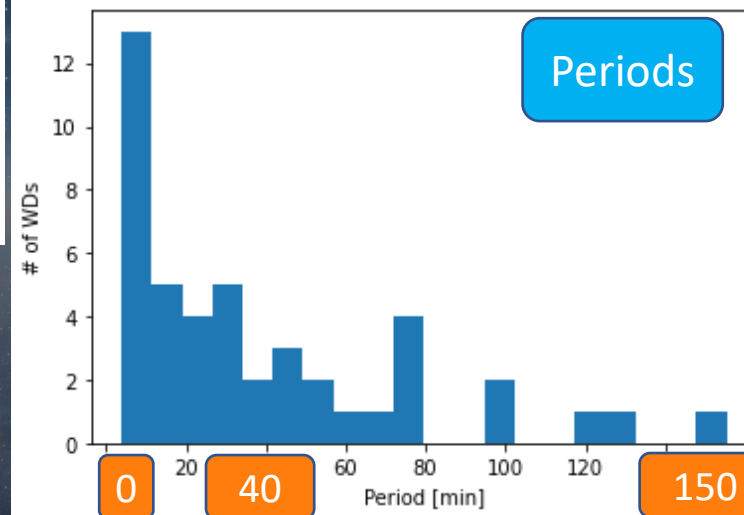
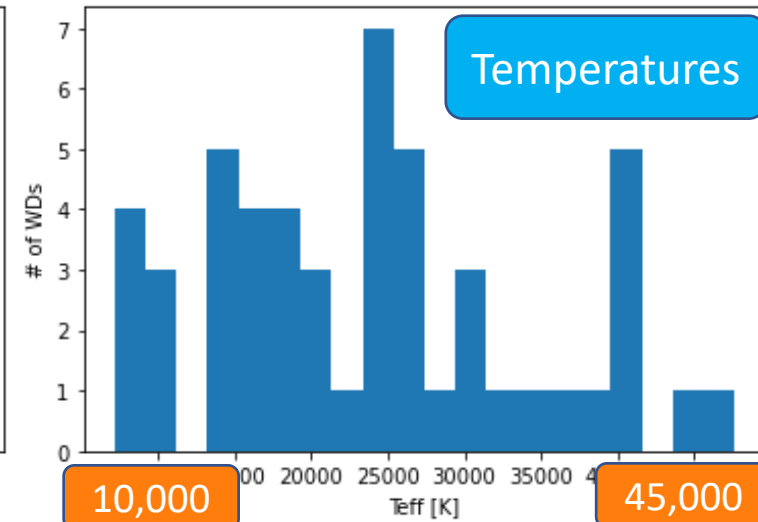
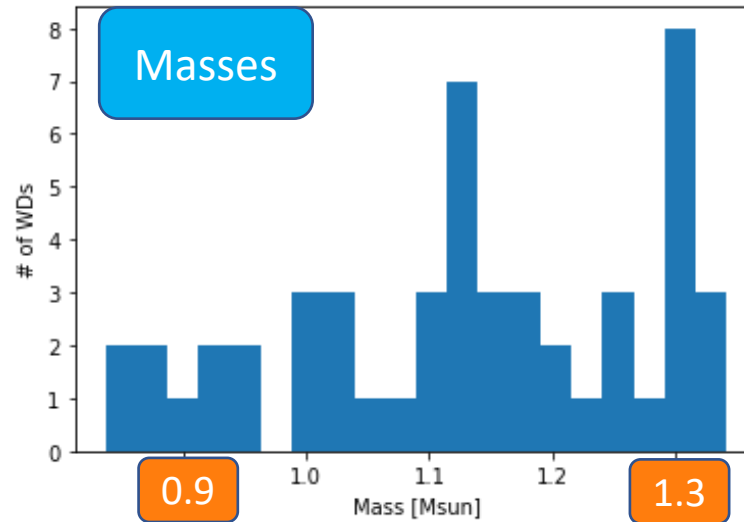
And we are finding many others!

LRIS (KECK)
Spectra

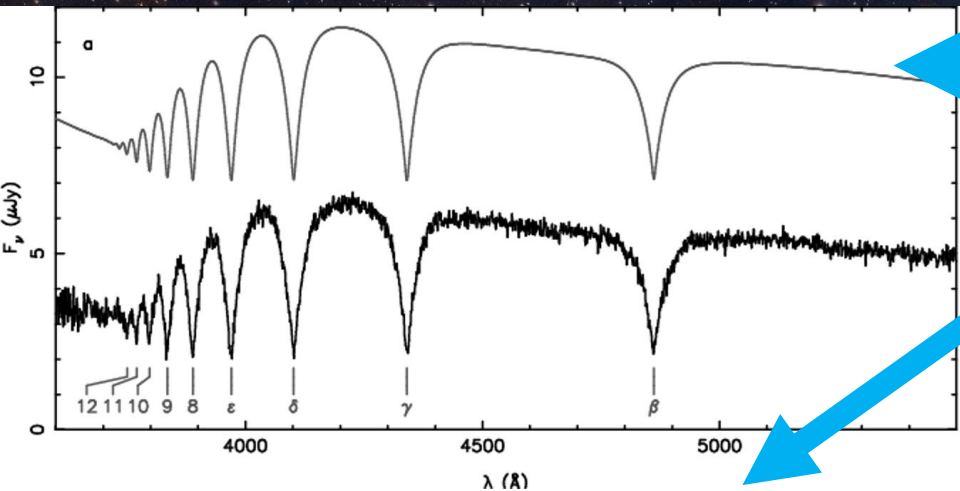
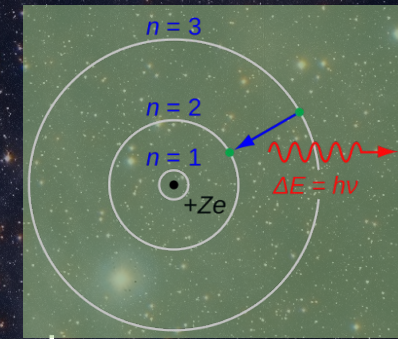
A large sample (~60) of candidate merger remnants!



- Massive
- Rapidly Rotating
- Warm-Hot
- Many obviously magnetic
- Many featureless or strange



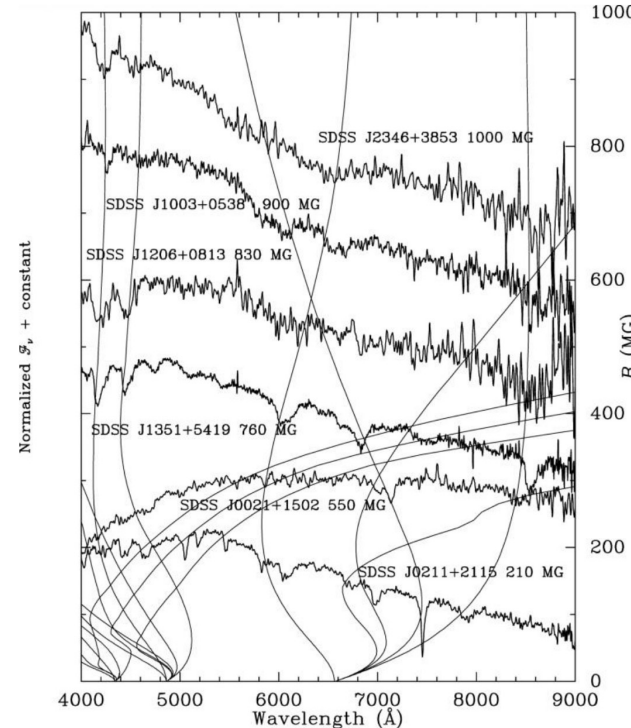
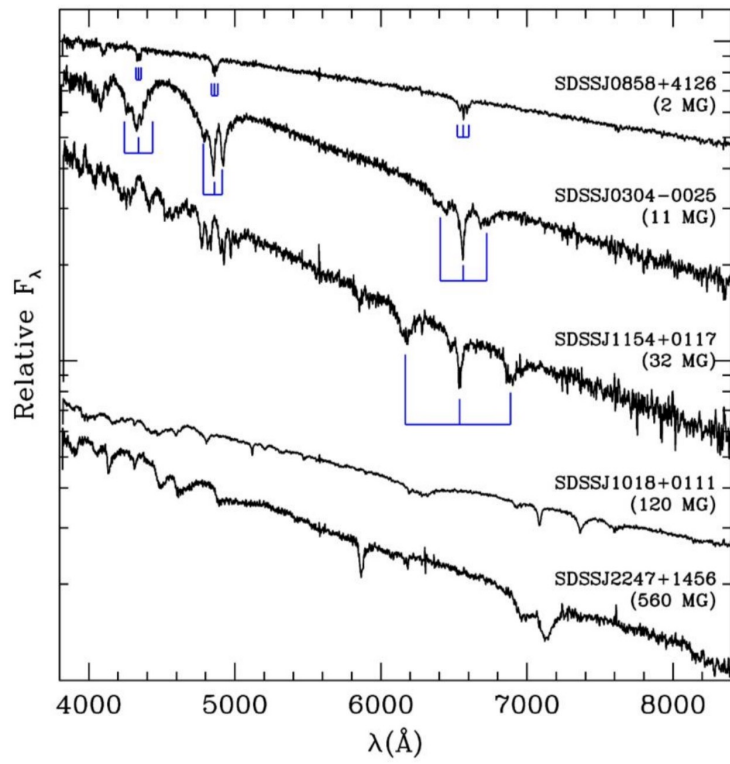
Are they magnetic?



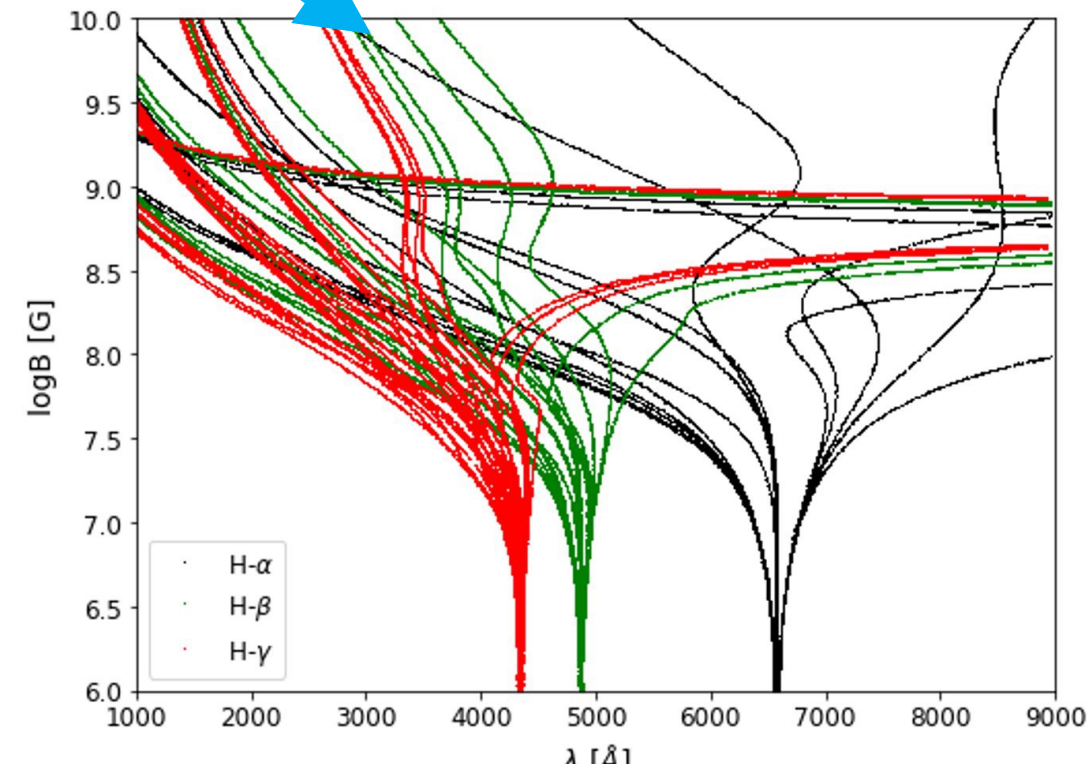
Non-magnetic spectrum

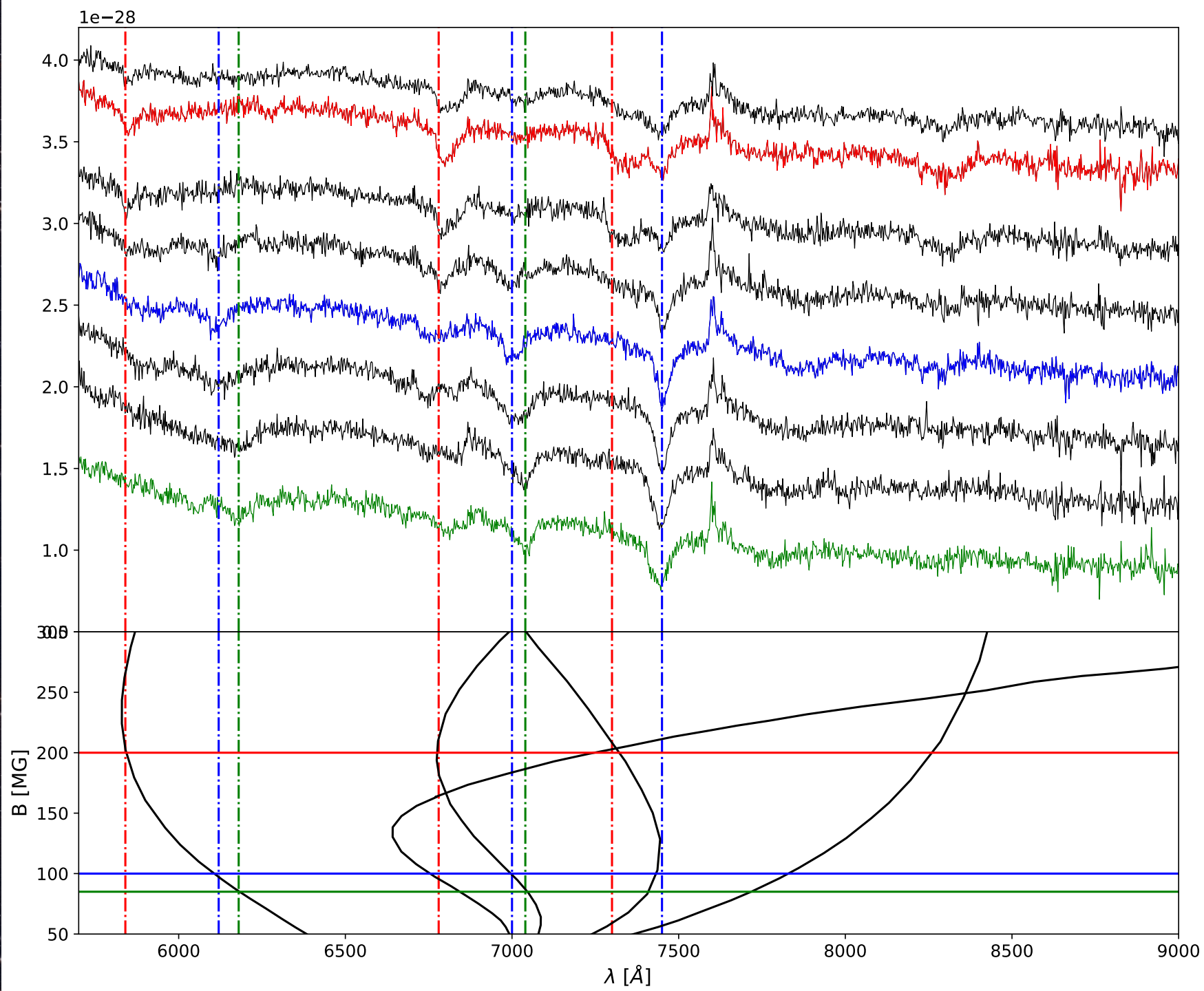
In magnetic WDs the degeneracy of the hydrogen energy levels is lifted \rightarrow Zeeman splitting

At high fields, the energy of each transition gets shifted by a large amount



(Vanlandingham et al. 2005)





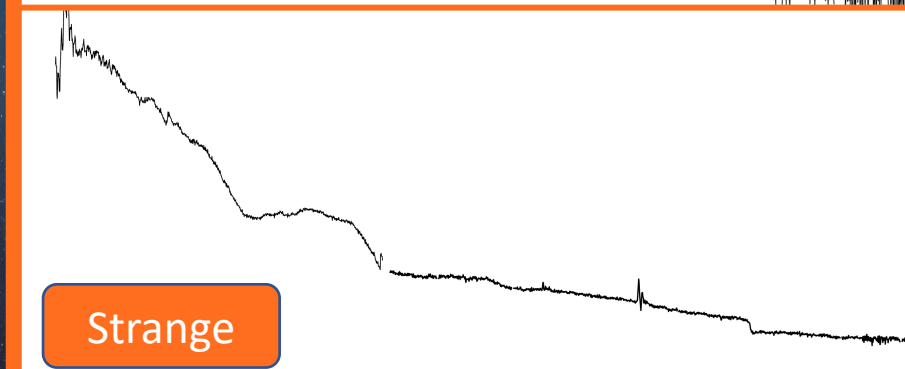
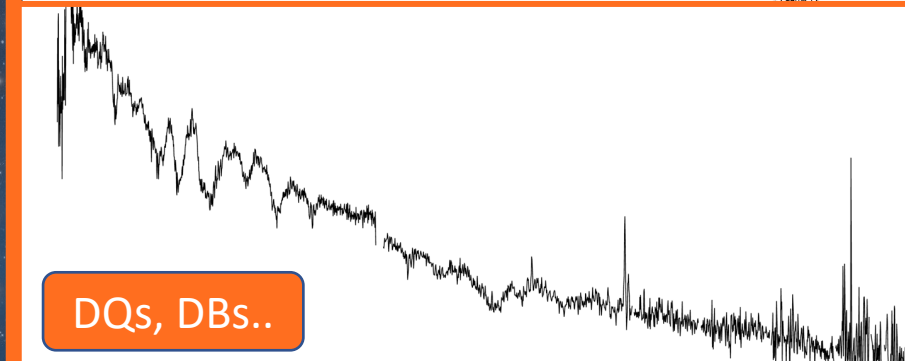
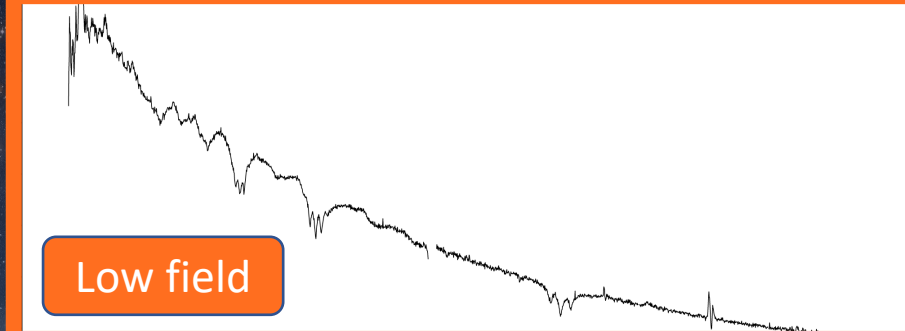
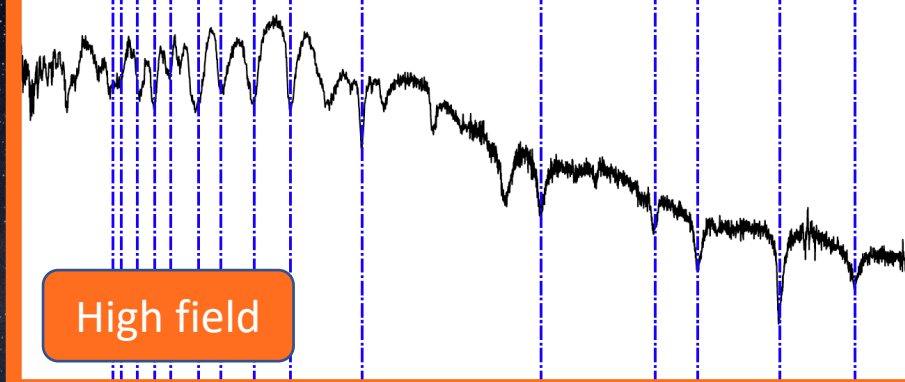


Why should we care?

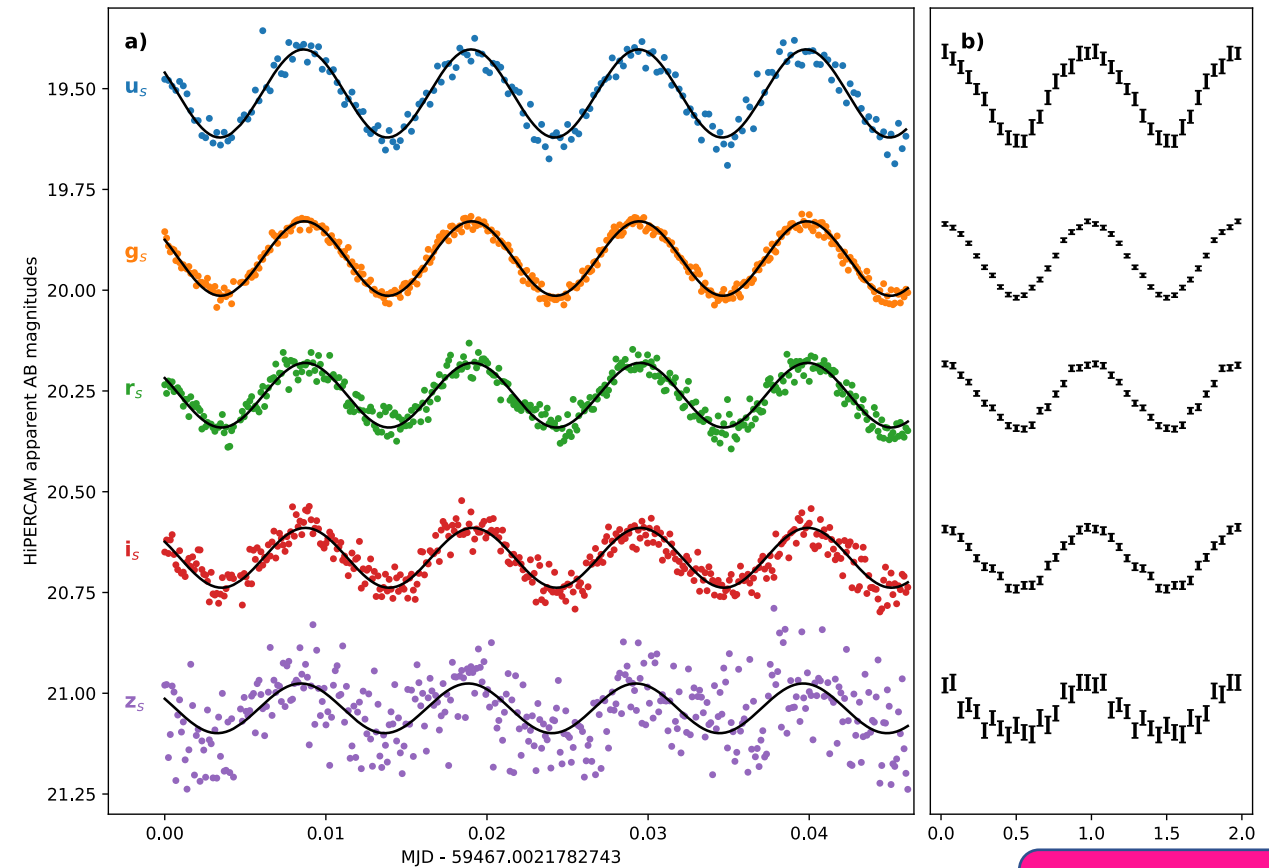
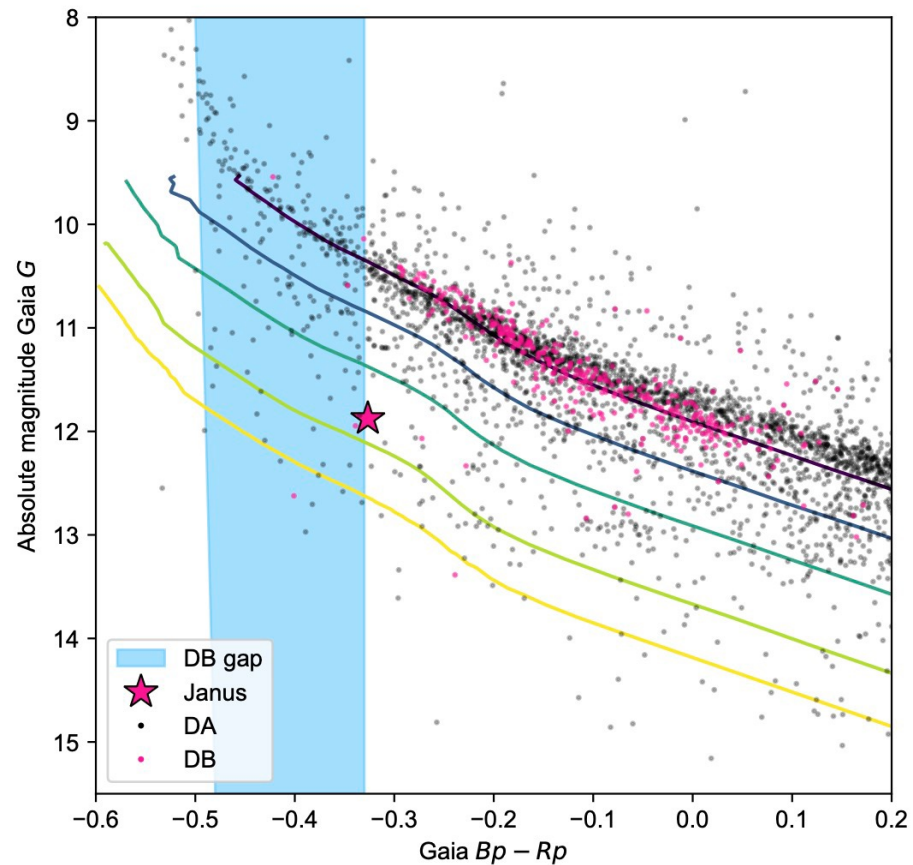
- Type Ia? These are the ones that barely exploded -> we could get a rate and a constraint on the contribution of double-degenerate to type Ia SNe
- LISA? From the rate of mergers we can constrain the population of short-period binaries
- Binary evolution? From their kinematics we should be able to constrain the delay time between the formation of the binary and the merger
- Magnetic white dwarfs? This is a nice sample of hot and warm magnetic white dwarfs at different field strengths

Cool! When should we expect the catalog?

- Collecting data
- Understanding selection effects and completeness
- Very variegate sample! Some objects are quite strange
- Coming soon!

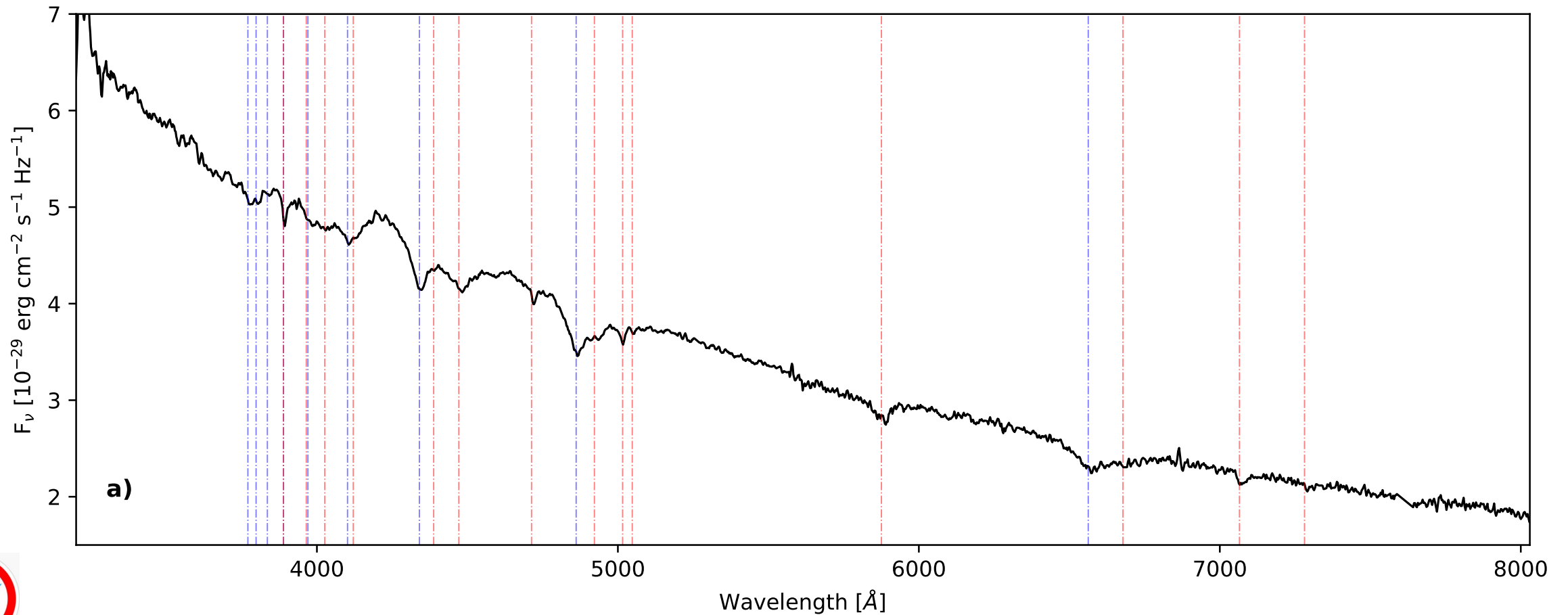


Janus, a new class of variables WDs



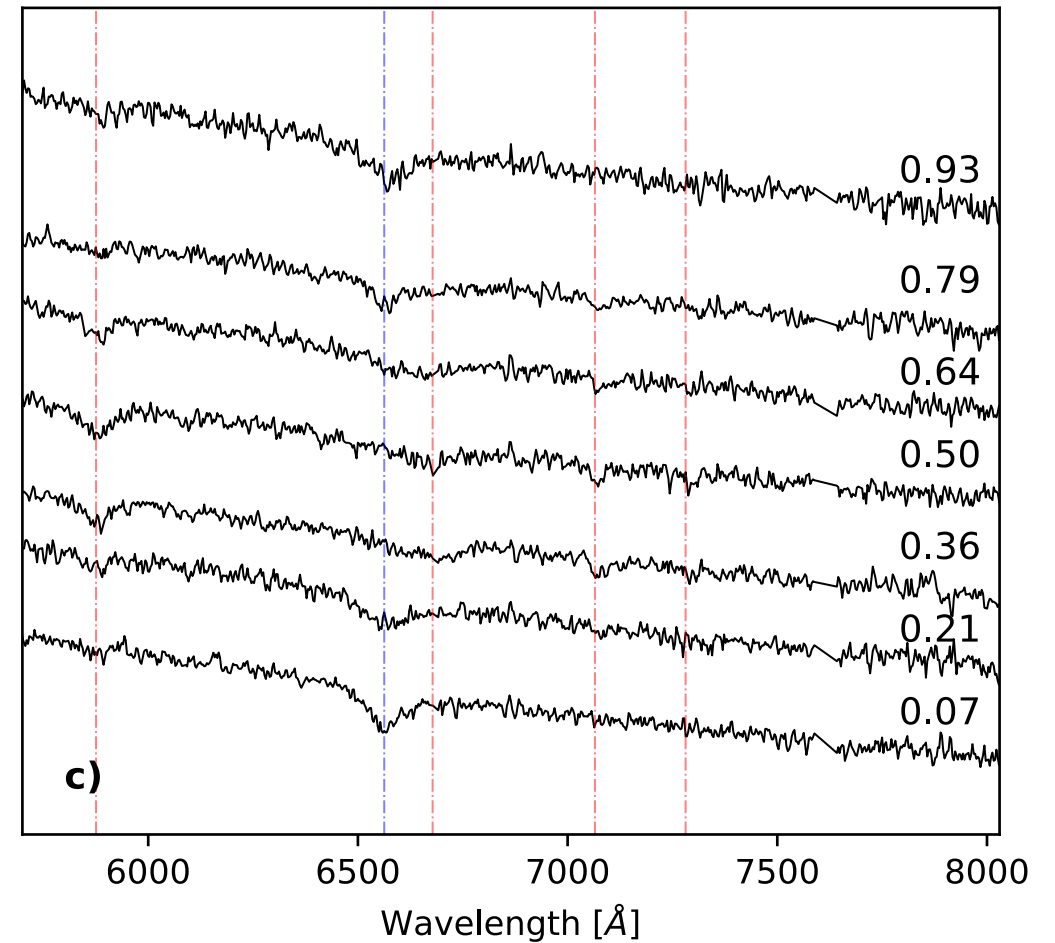
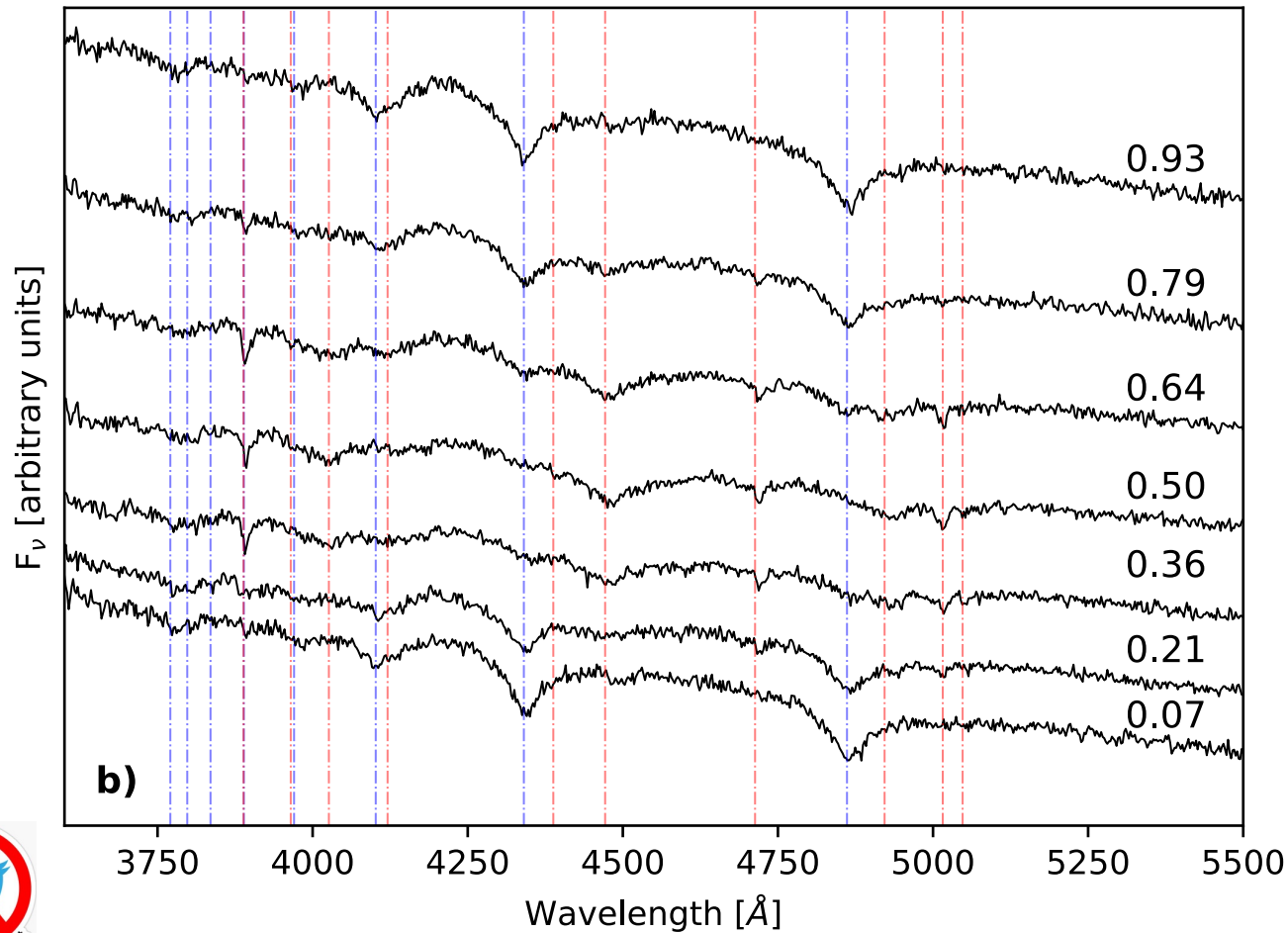
Phase-averaged spectrum

— Hydrogen
— Helium



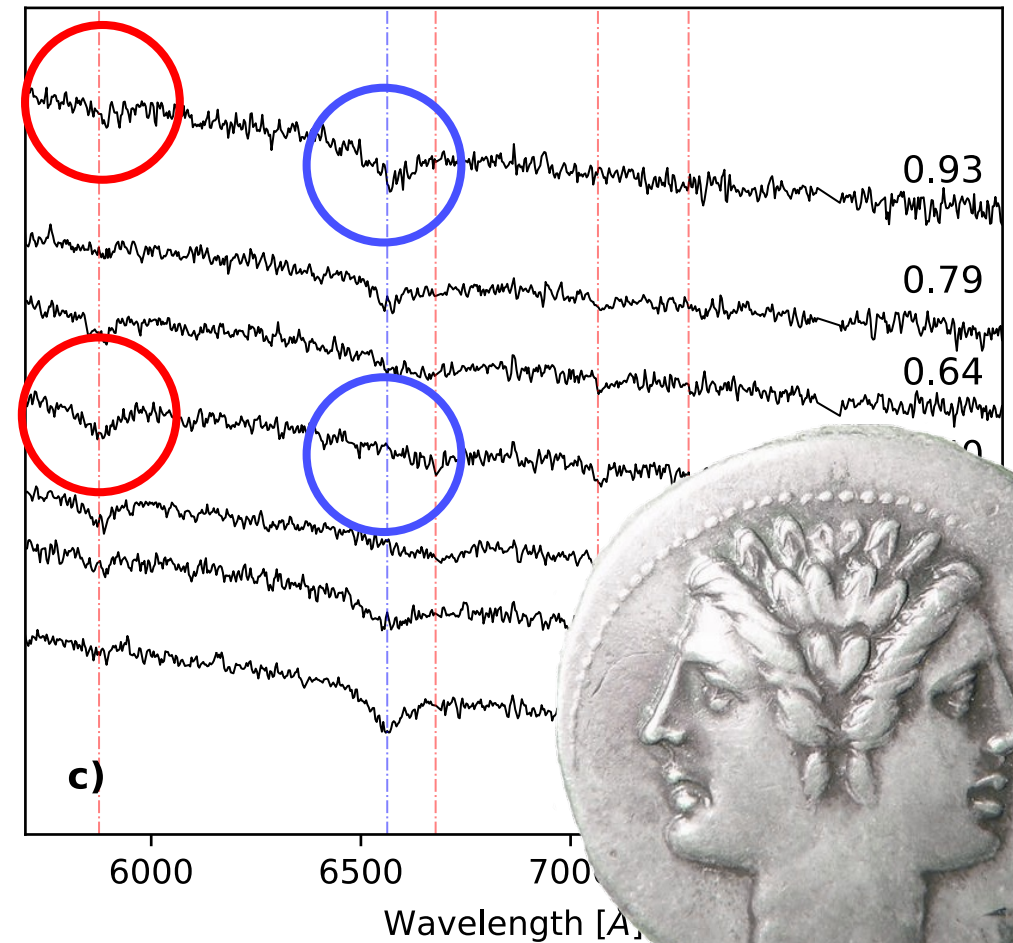
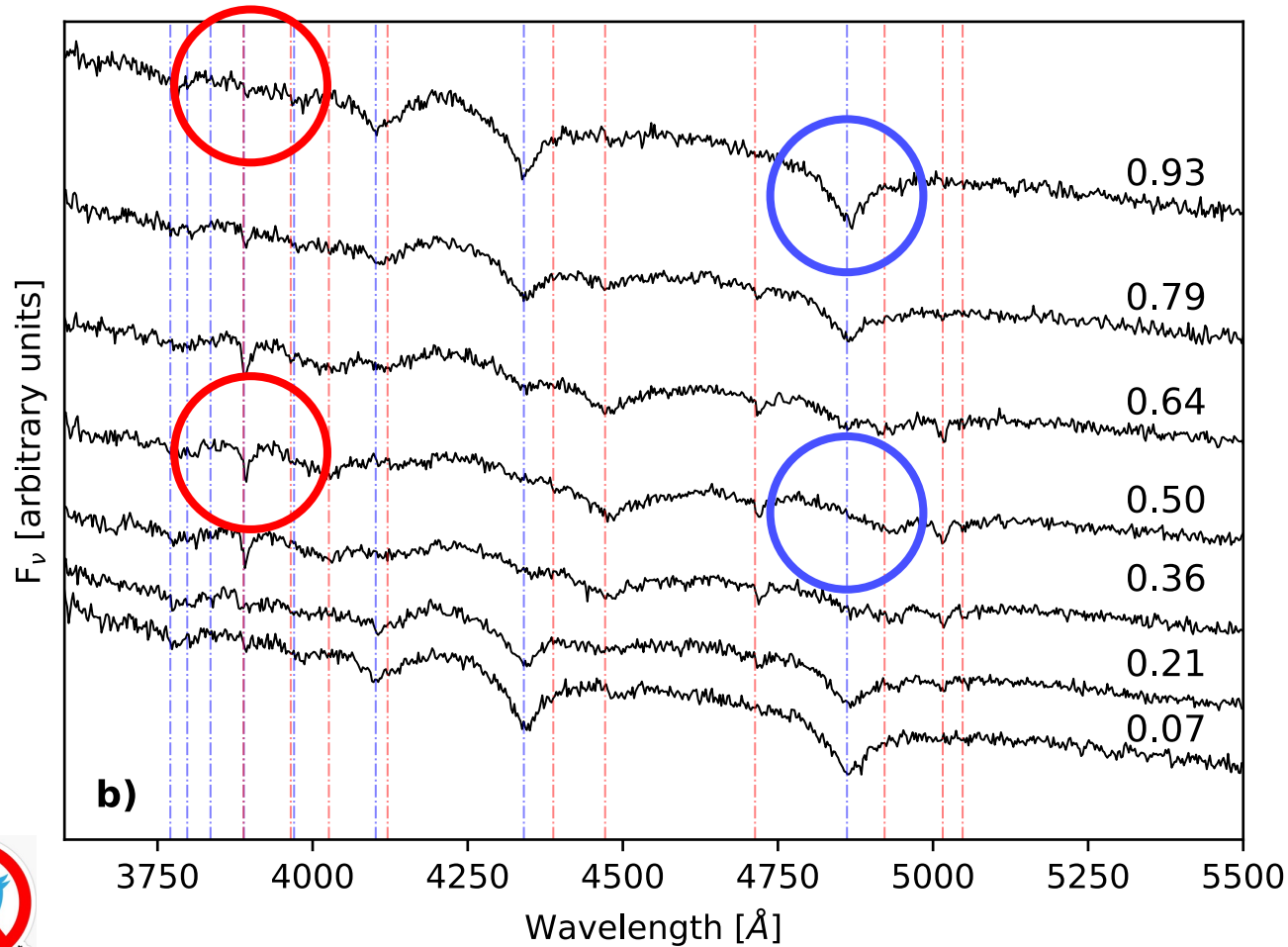
Phase-resolved, a double-faced WD!

— Hydrogen
— Helium



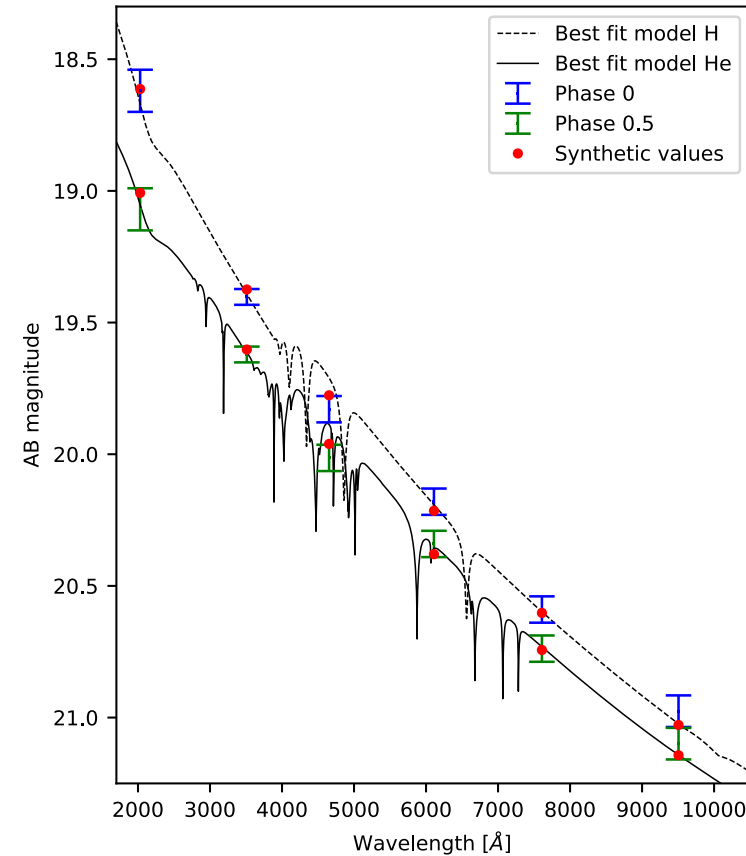
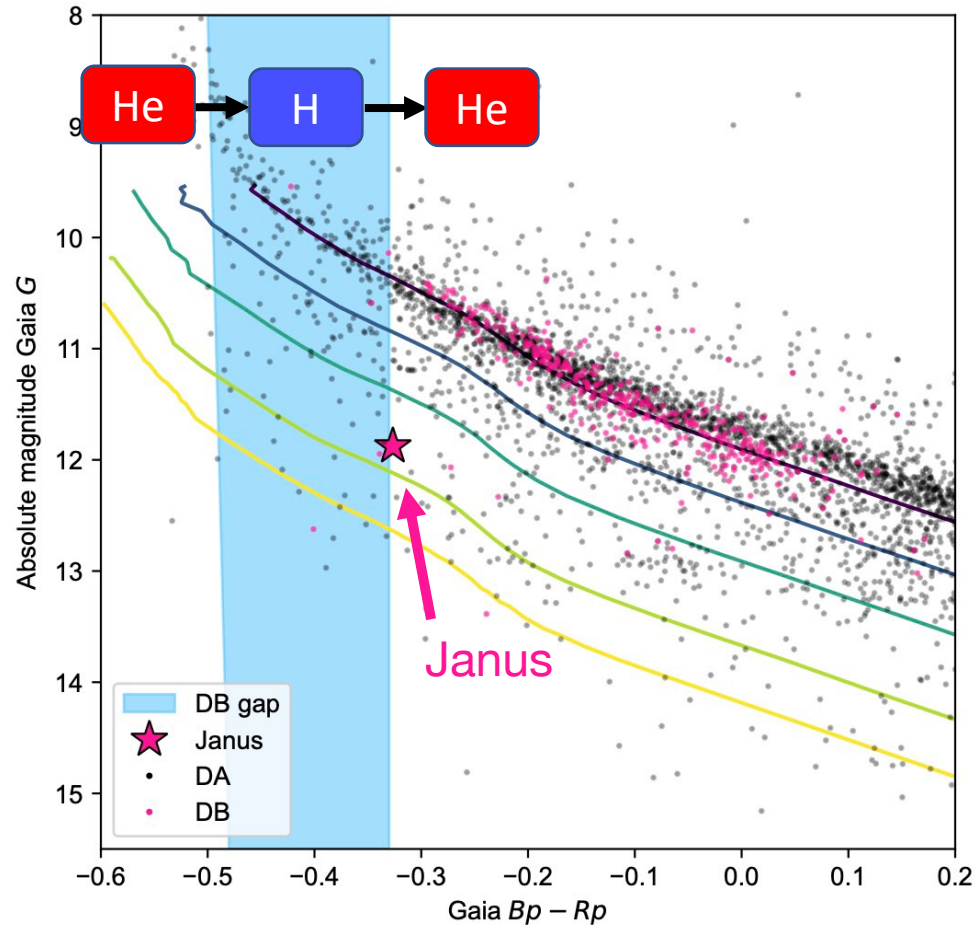
Phase-resolved, a double-faced WD!

— Hydrogen
— Helium



Janus, a transitioning WD?

- At the the low end of the DB gap, strong mixing in the helium layer dilutes the hydrogen: DBs and DBAs appear
- If there is a magnetic field strong enough to inhibit convection on part of the surface, we can still see hydrogen



Janus, a hydrogen ocean?

